

DISTINGUISHING BETWEEN SPIN- ALIGNED AND ISOTROPIC BINARY BLACK HOLE POPULATIONS USING GRAVITATIONAL WAVE OBSERVATIONS

Will M. Farr, Simon Stevenson, *M. Coleman Miller*, *Ilya Mandel*,
Alberto Vecchio

Rates & Populations Telecon
23 March 2017

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Not LVC



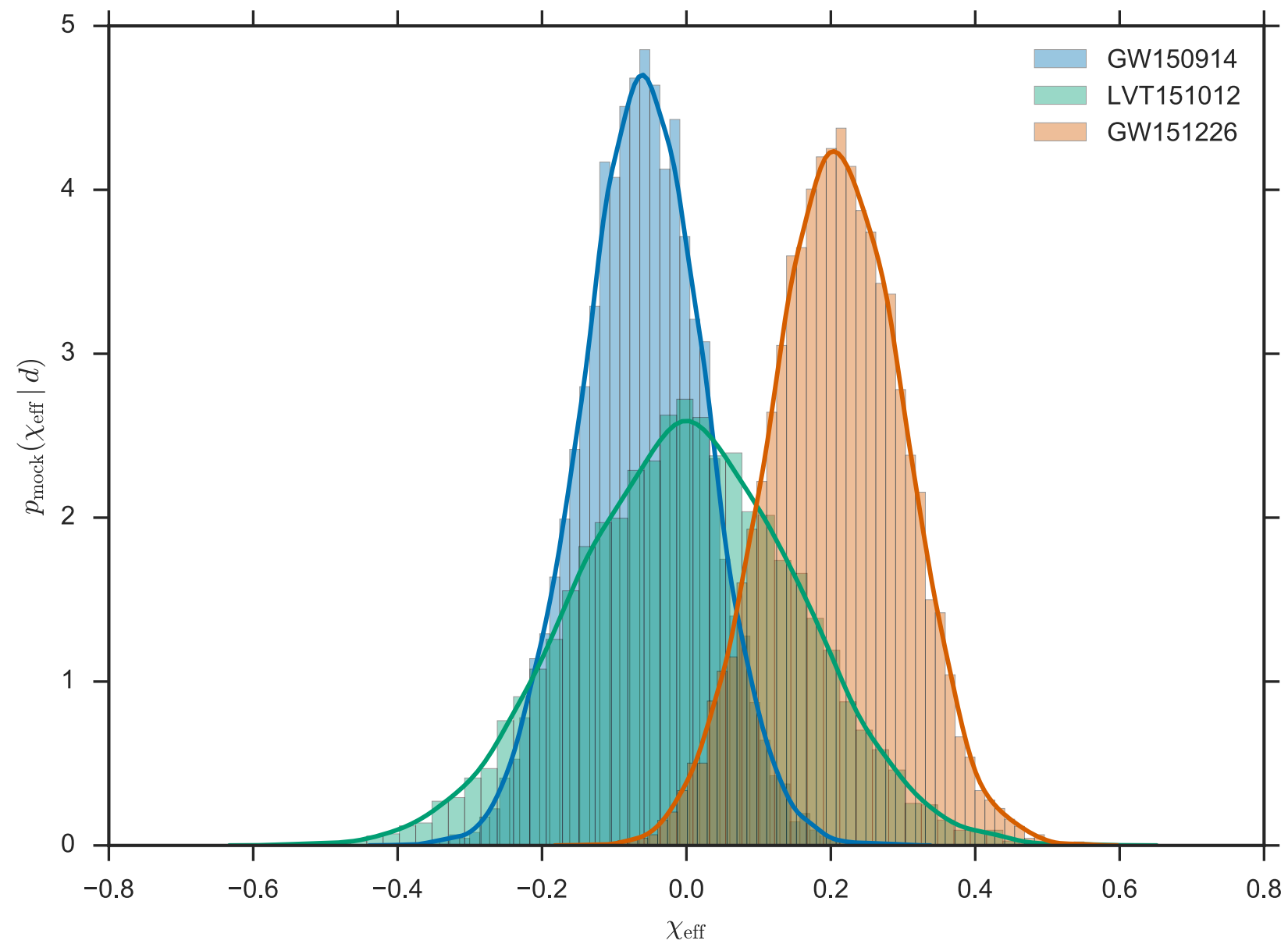
Basic Idea

- Interesting Tension:
 - EM Spin Measurements tend to high spins
 - LIGO Measurements have small χ_{eff} .
- Possible Explanations:
 - LIGO: large spins, but isotropic angles \Rightarrow small χ_{eff}
 - LIGO: different population, so small spins.
- How much can we learn from existing and near-future observations?

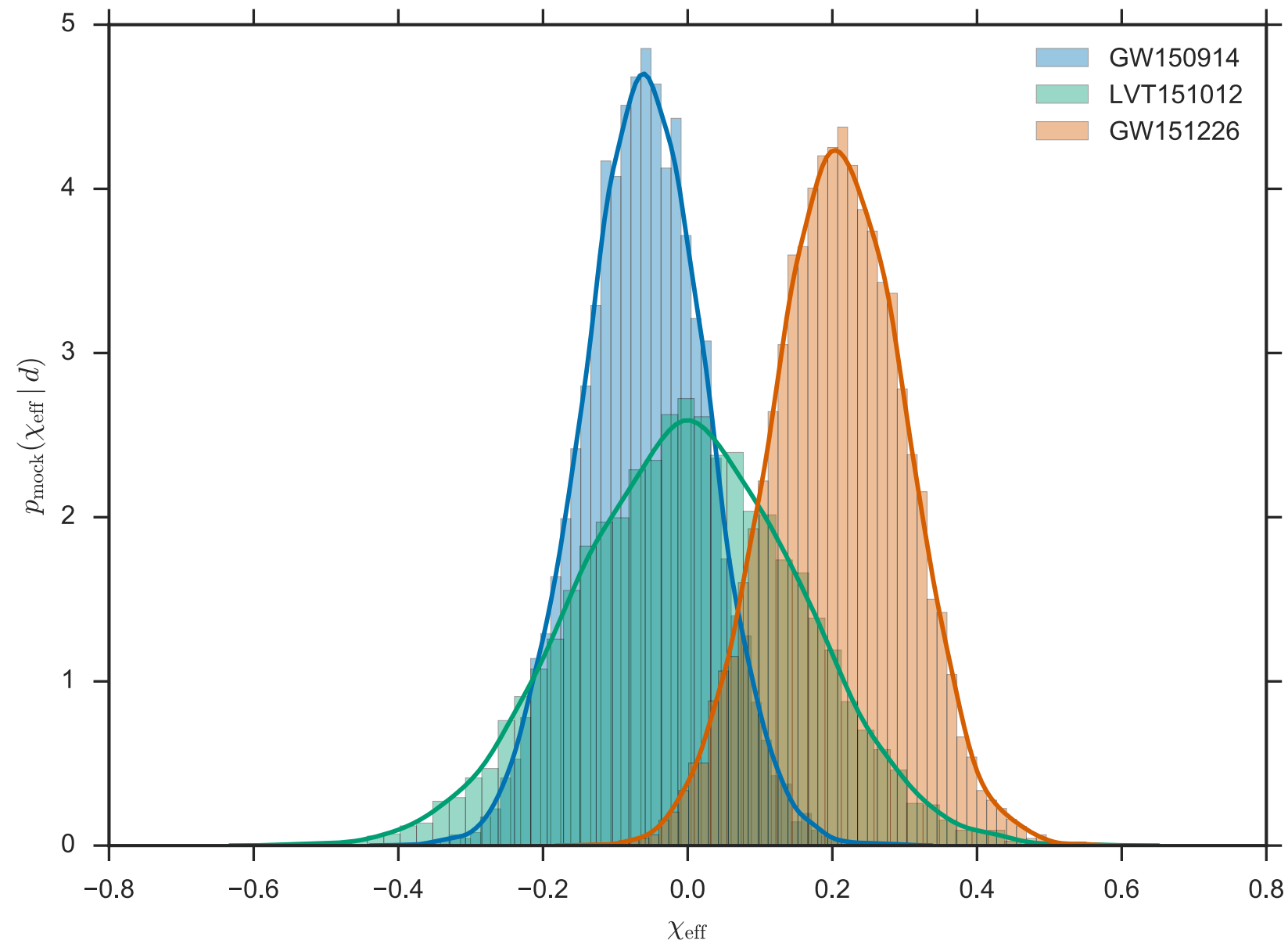
B.L.U.F.

- O1 observations have 1.7σ preference for isotropic over aligned distributions based on χ_{eff} .
- With 10 more observations, get at least 3σ , more probably 5σ angular distribution
- Isotropic **probably** implies dynamical origin => maybe we are seeing a dynamical-formation population?

The 2.9 Events



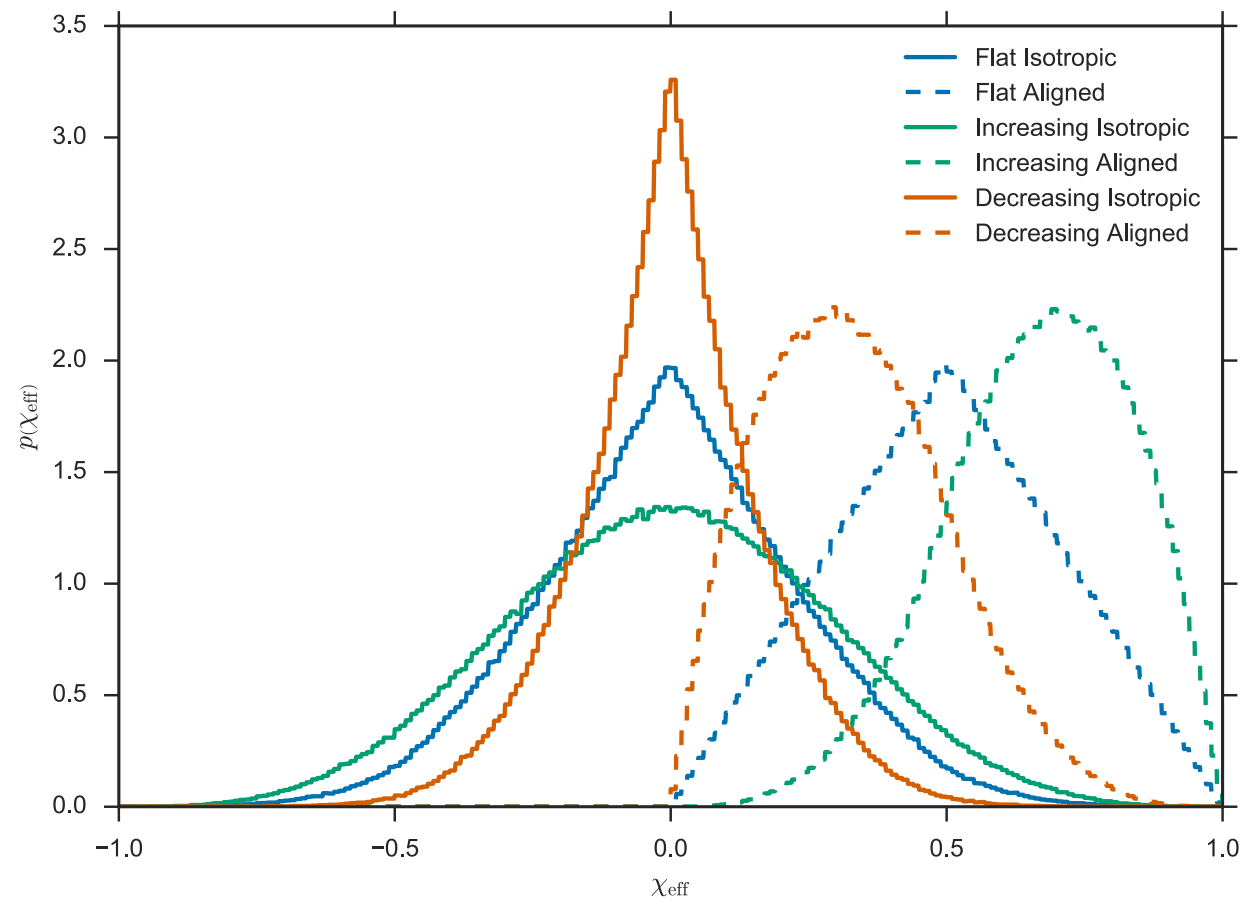
The 2.9 Events



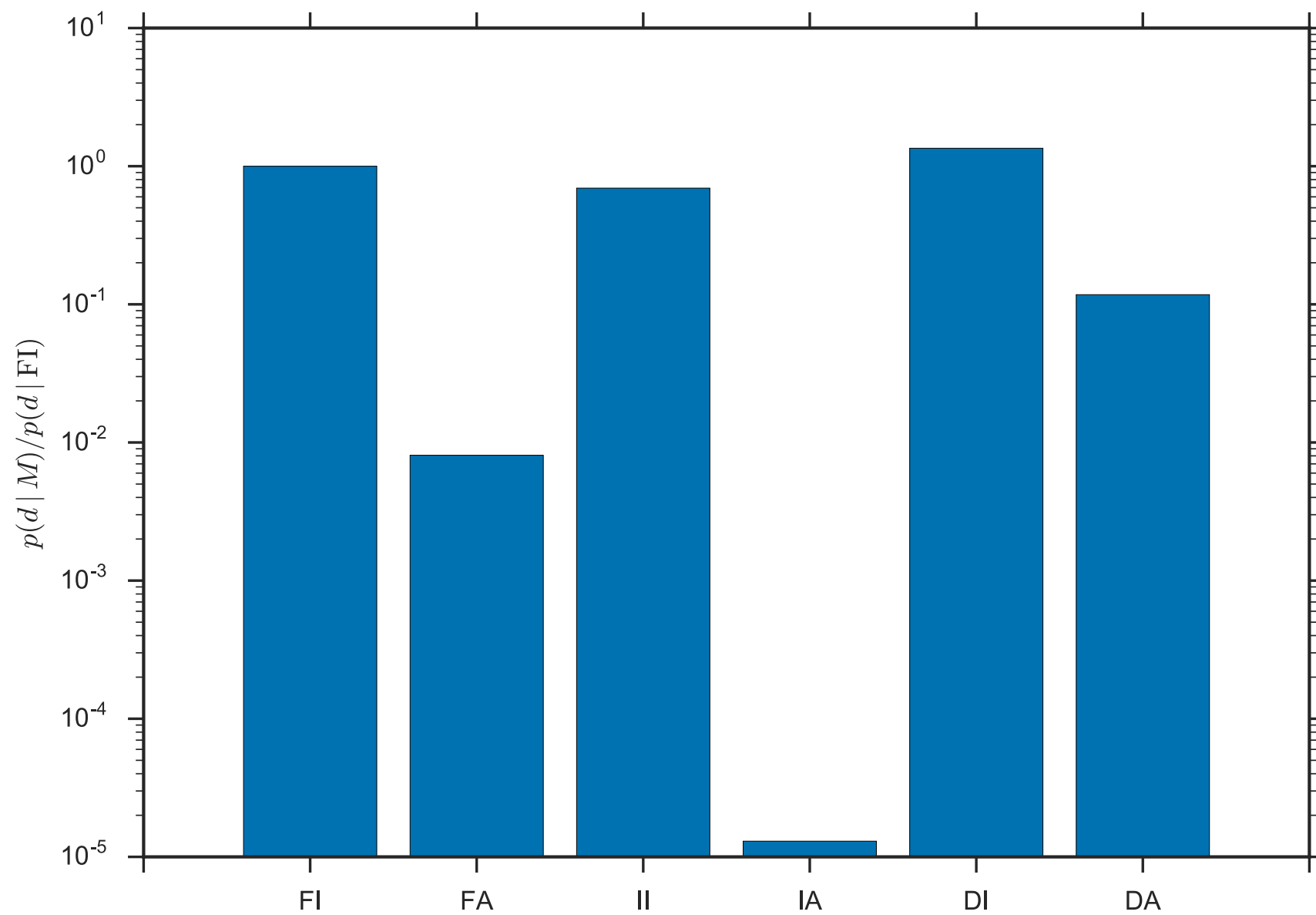
Mock Observations: Same mean and 90% as LIGO Quotes

Some Simple Models

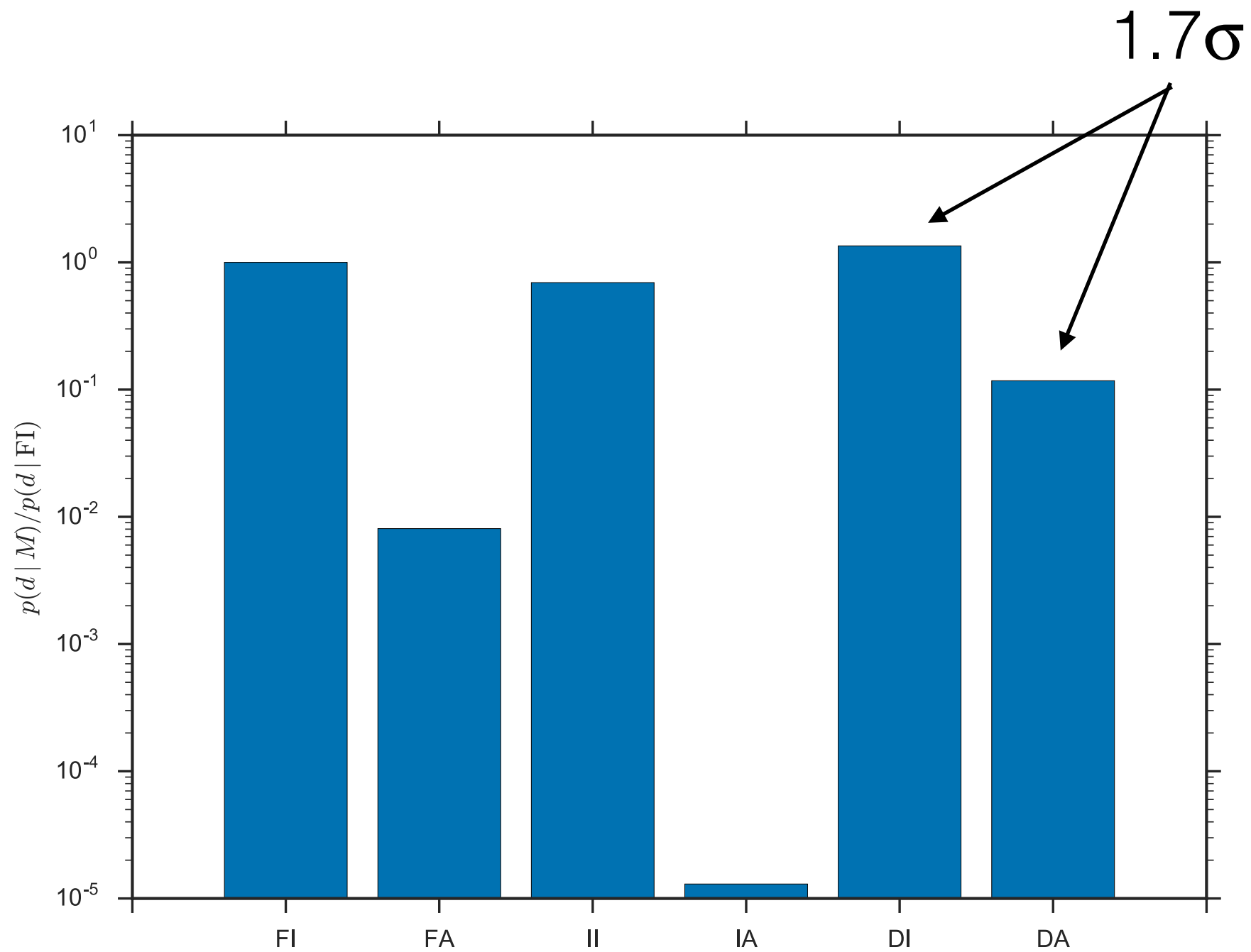
- $p(\chi_{\text{eff}})$ comes from $p(a) \cdot p(\text{angles})$
- $q = 1$ (marginally consistent with existing observations)
- Aligned \Rightarrow tilt angles ≈ 0
- $p(a)$ is:
 - Flat
 - Linearly increasing
 - Linearly decreasing



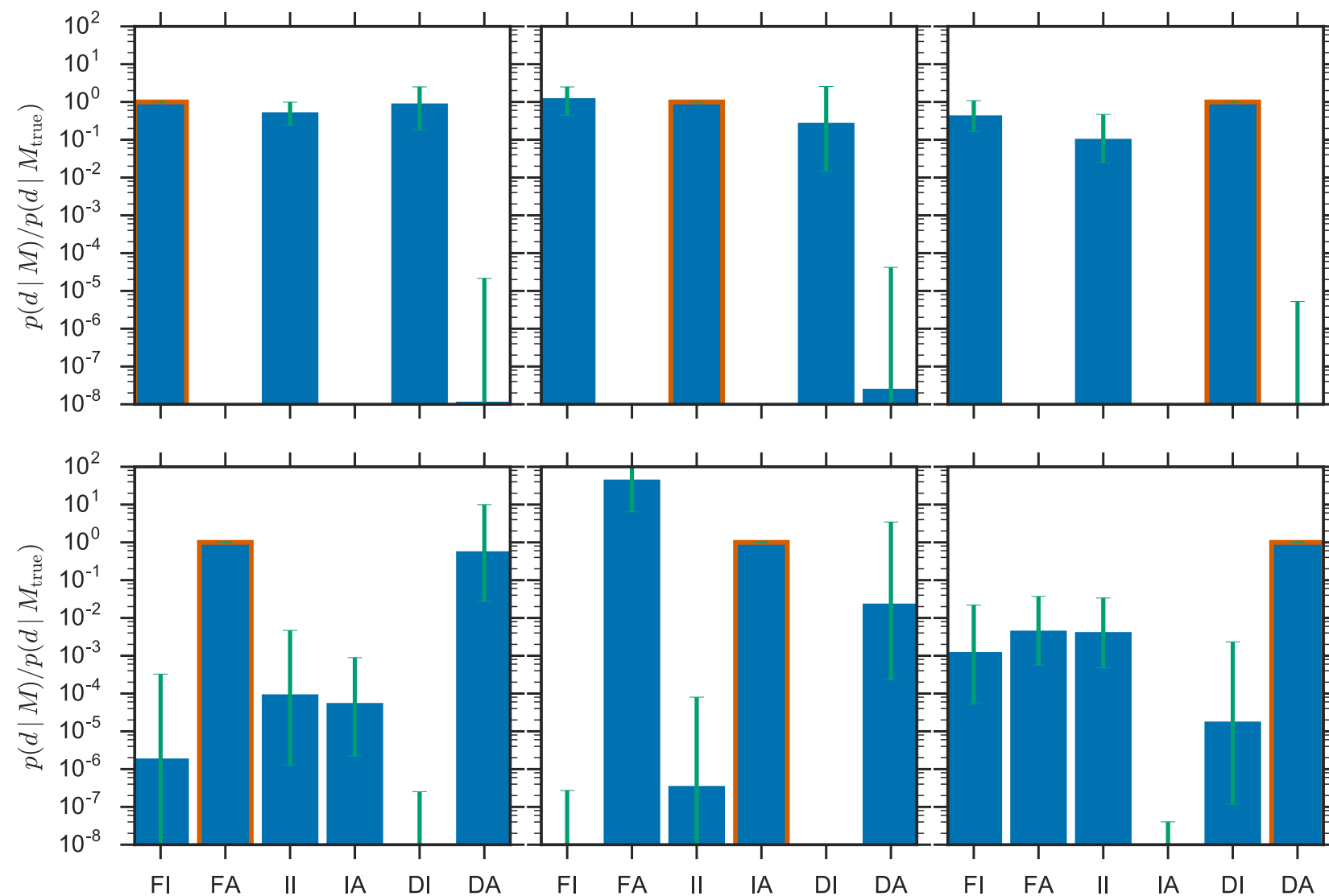
Current Evidence



Current Evidence

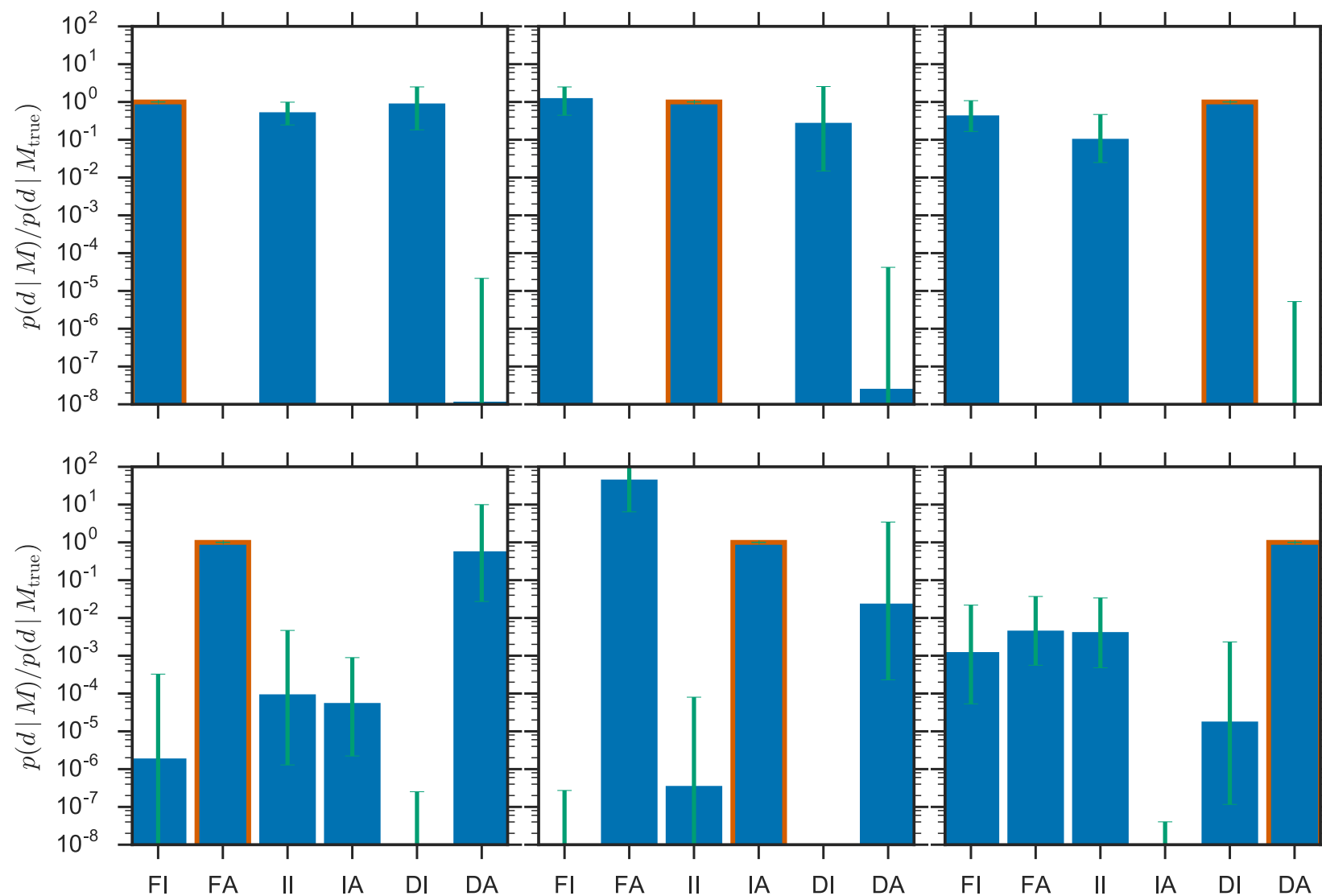


With 10 More Detections



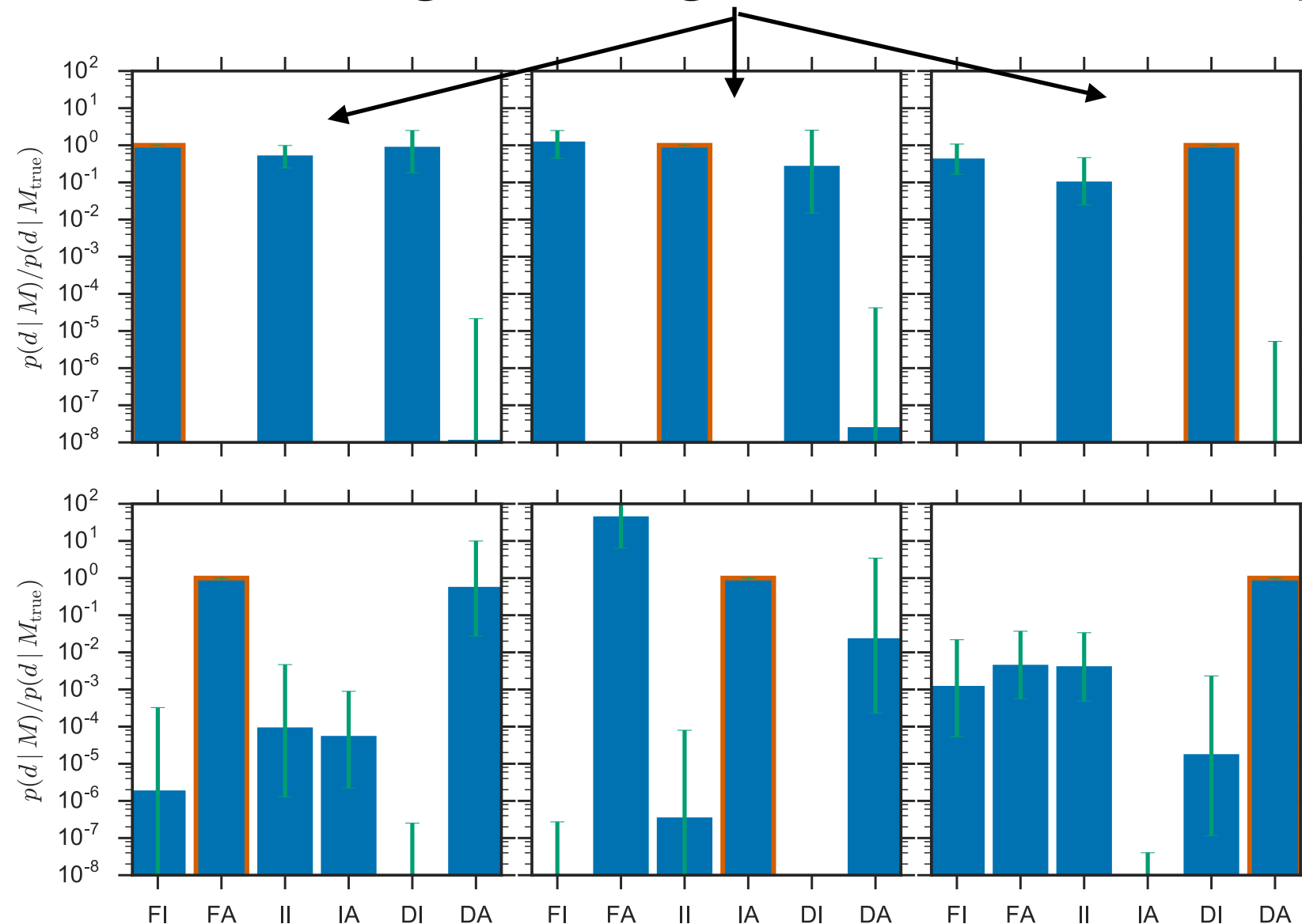
With 10 More Detections

Mostly $> 5\sigma$, all $> 3\sigma$



With 10 More Detections

We know angles long before we know $p(a)$:



Extra: Even Mixture Models Prefer All-Isotropic After O1

