## Clear sky 2018 october 12th camera 1 sw-vers. 3

```
In [1]: import matplotlib
  import numpy as np
  import matplotlib.pyplot as plt
  %matplotlib inline
  import csv
  import bisect
  import datetime
  from scipy import interpolate
  from matplotlib.dates import DateFormatter

# User defined functions
  from load_data_from_csv import *
```

## **Import Data**

```
In [2]: day = '2018-10-12'  # select day of observation ('2018-10-12' : camera 2, sw-ve
    rs. 3)
    s_time = ' 07:00:00'  # beginning of observation
    e_time = ' 20:00:00'  # end of observation
    start = day + s_time
    end = day + e_time
```

#### Weather station at Luzern Switzerland

```
In [3]: luz_csv = r'../weather_data/irradiation_luz_2017_2018.csv'

df = process_LUZ(luz_csv)
    df.set_index(df.datetime, inplace=True)
    df_lu = df['gre000z0']  # 10 min mean in W/m²
    lu_rad = df_lu.loc[start:end]  # set distinct observation day

#lu_rad.head(n=2)
```

# Clear Sky solar irradiance from SODA using McClear model

Source: <a href="http://www.soda-pro.com/web-services/radiation/cams-mcclear">http://www.soda-pro.com/web-services/radiation/cams-mcclear</a> (http://www.soda-pro.com/web-services/radiation/cams-mcclear)

```
In [4]: soda_csv = r'../weather_data/irradiation_soda_2017_2018_1min.csv'

df = process_SODA(soda_csv)
    df_soda = df['Clear sky GHI']*60  # given in units of Wh/m²
    mc_rad = df_soda.loc[start:end]  # set distinct observation day

#mc_rad.head(n=2)
```

#### Load relative luminance from images

```
In [5]: calc_csv = r'../lumi_data/20181012_cam1_luminance.csv'

df_calc = process_LUMI(calc_csv)
    lum_hdr = df_calc['lum_hdr']  # mean luminance from HDR image
    lum_hdr_m = df_calc['lum_hdr_m']  # mean luminance from masked HDR image fro
    m raw data
    lum_jpg_m = df_calc['lum_jpg_m']  # mean luminance from masked HDR image bui
    ld from three jpg exposures

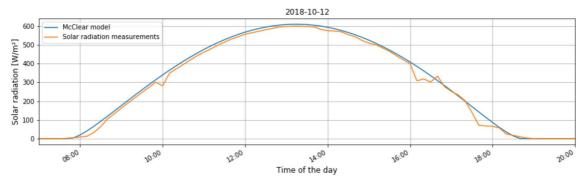
#lum_hdr.head(n=2)
```

### Pyranometer measurements vs. McClear model

```
In [6]: matplotlib.rcParams['timezone'] = 'Europe/Zurich'

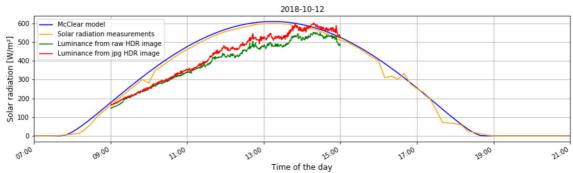
fig = plt.figure(1, figsize=(15,4))
mc_rad.plot(label='McClear model')
lu_rad.plot(label='Solar radiation measurements')

plt.xlabel('Time of the day', fontsize=12)
plt.ylabel('Solar radiation [W/m²]', fontsize=12)
plt.legend(loc='upper left')
plt.grid(b=None, which='major', axis='both')
plt.title(day)
fig.autofmt_xdate()
formatter = DateFormatter('%H:%M')
#formatter.set_tzinfo(timezone('Europe/Zurich'))
plt.gcf().axes[0].xaxis.set_major_formatter(formatter)
```



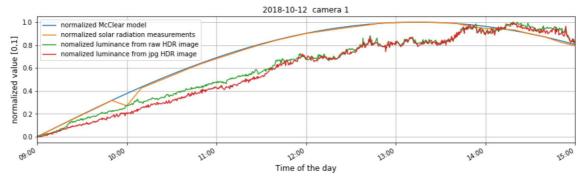
Plot data for 2018-10-12 from 9:00 to 15:00

```
In [7]: matplotlib.rcParams['timezone'] = 'Europe/Zurich'
        day = '2018-10-12'  # day of observation ('2018-10-12' : camera 2, sw-vers. 3)
        s time = ' 07:00:00' # beginning of observation
        e time = ' 21:00:00' # end of observation
        start = day + s time
        end = day + e_time
        lu rad = df lu.loc[start:end]
                                        # weather station
        mc_rad = df_soda.loc[start:end] # McClear model
        lum hdr = lum hdr.loc[start:end] # McClear model
        lum jpg = lum jpg m.loc[start:end] # McClear model
         lum hdr = lum hdr/12
        lum jpg = lum jpg/60
        fig = plt.figure(2, figsize=(15,4))
        mc rad.plot(label ='McClear model', color = 'blue')
        lu rad.plot(label ='Solar radiation measurements', color = 'orange')
        _lum_hdr.plot(label='Luminance from raw HDR image',color = 'green')
        _lum_jpg.plot(label='Luminance from jpg HDR image',color = 'red')
        plt.xlabel('Time of the day', fontsize=12)
        plt.ylabel('Solar radiation [W/m²]', fontsize=12)
        plt.legend(loc='upper left')
        plt.grid(b=None, which='major', axis='both')
        plt.title(day)
        fig.autofmt_xdate()
        formatter = DateFormatter('%H:%M')
        #formatter.set_tzinfo(timezone('Europe/Zurich'))
        plt.gcf().axes[0].xaxis.set_major_formatter(formatter)
```



Plot with normalized data for 2018-10-12 from 9:00 to 15:00

```
In [8]: # Set appropriate timezone
        matplotlib.rcParams['timezone'] = 'Europe/Zurich'
        day = '2018-10-12'
                           # day of observation ('2018-10-12' : camera 2, sw-vers. 3)
        s time = ' 09:00:00' # beginning of observation
        e time = ' 15:00:00' # end of observation
        start = day + s\_time
        end = day + e time
        lu rad = df lu.loc[start:end]
                                         # weather station
        mc rad = df soda.loc[start:end] # McClear model
        #print(lu rad.head(n=5))
        #print(lum hdr.head(n=5))
        lu rad n = ((lu rad-lu rad.min())/(lu rad.max()-lu rad.min()))
        mc rad n = ((mc rad-mc rad.min())/(mc rad.max()-mc rad.min()))
        lum hdr n = ((lum hdr-lum hdr.min())/(lum hdr.max()-lum hdr.min()))
        lum jpg n = ((lum jpg m-lum jpg m.min())/(lum jpg m.max()-lum jpg m.min()))
        fig = plt.figure(2, figsize=(15,4))
        mc rad n.plot(label ='normalized McClear model')
        lu rad n.plot(label ='normalized solar radiation measurements')
        lum hdr n.plot(label='nomalized luminance from raw HDR image')
        lum jpg n.plot(label='nomalized luminance from jpg HDR image')
        plt.xlabel('Time of the day', fontsize=12)
        plt.ylabel('normalized value [0,1]', fontsize=12)
        plt.legend(loc='upper left')
        plt.grid(b=None, which='major', axis='both')
        plt.title(day + ' camera 1')
        fig.autofmt_xdate()
        formatter = DateFormatter('%H:%M')
        plt.gcf().axes[0].xaxis.set_major_formatter(formatter)
```



Scatterplot mean luminance from HDR images vs. pyranometer measurements

```
In [41]: fig = plt.figure(3, figsize=(5,5))
         day = '2018-10-12'  # day of observation ('2018-10-12' : camera 2, sw-vers. 3)
         s time = ' 09:00:00' # beginning of observation
         e time = ' 14:15:00' # end of observation
         start = day + s time
         end = day + e time
         _lu_rad_n = lu_rad_n.loc[start:end]
                                                # weather station Allmend
         lum hdr n = lum hdr n.loc[start:end] # McClear model
         lu rad f = lu rad n.asfreq(freq='40S')
         lu rad ip = lu rad f.interpolate(method = 'linear')
         print('lu rad ip:{} lum hdr:{}'.format(lu rad ip[:-6].size, lum hdr n.size))
         plt.xlabel('Normalized luminance', fontsize=12)
         plt.ylabel('Normalized solar radiation', fontsize=12)
         plt.scatter(x=lu rad ip[:-6], y= lum hdr n, marker='o', color='blue')
         r = np.corrcoef( lum hdr n, lu rad ip[:-6])
         print ('Correlation = ', r[0][1])
```

lu\_rad\_ip:460 lum\_hdr:460
Correlation = 0.9667222324028583

