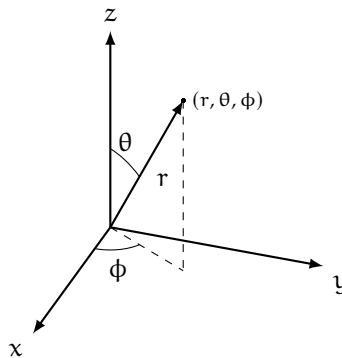
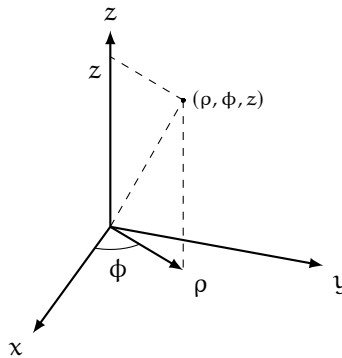
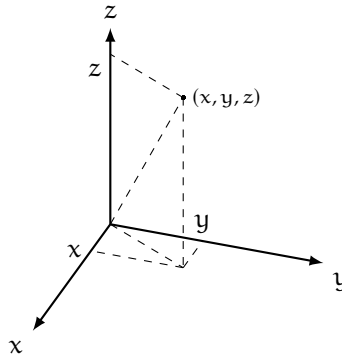


# Things You Just Gotta Know

## Coordinate Systems

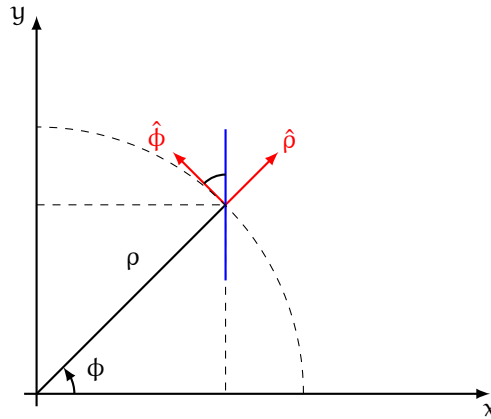


We want to focus on vector-valued functions of coordinates.

$$\vec{V}(\mathbf{r}) = V_x(x, y)\hat{i} + V_y(x, y)\hat{j}.$$

Notice that a vector function uses the coordinate system twice. Once for the function's inputs, once for the vectors themselves.

## Polar Coordinates



We can also express the inputs to  $\vec{V}$  in polar coordinates,  $(\rho, \phi)$ .

$$\vec{V}(\mathbf{r}) = V_\rho(\rho, \phi) \hat{i} + V_\phi(\rho, \phi) \hat{j}.$$

To extract the input functions, we take

$$V_x = \hat{i} \cdot \vec{V}$$

$$V_y = \hat{j} \cdot \vec{V}.$$

Alternatively, we can project  $\vec{V}$  onto the  $\hat{\rho}, \hat{\phi}$  axis:

$$\vec{V}(\mathbf{r}) = V_\rho(\rho, \phi) \hat{\rho} + V_\phi(\rho, \phi) \hat{\phi},$$

and we extract

$$V_\rho = \hat{\rho} \cdot \vec{V}$$

$$V_\phi = \hat{\phi} \cdot \vec{V}.$$

Notice that  $\mathbf{r}$  is an abstract vector; we need to project it onto a basis.

For instance, we can take the position vector and project it onto the cartesian and polar axes:

$$\begin{aligned} \mathbf{s} &= x\hat{i} + y\hat{j} \\ &= \rho \cos \phi \hat{i} + \rho \sin \phi \hat{j} \\ &= \rho \hat{\rho} \\ &= \sqrt{x^2 + y^2} \hat{\rho} \end{aligned}$$

The main reason we avoided using the  $\hat{\rho}, \hat{\phi}$  axis up until this point is that  $\rho$  and  $\phi$  are *position-dependent*, while the  $\hat{i}, \hat{j}$  axis is position-independent.

Now, we must figure out the position-dependence of  $\hat{\rho}$  and  $\hat{\phi}$ :

$$d\mathbf{r} = \frac{\partial \mathbf{r}}{\partial \rho} d\rho + \frac{\partial \mathbf{r}}{\partial \phi} d\phi.$$

If we hold  $\phi$  constant, it must be the case that any change in  $\rho$  is in the  $\hat{\rho}$  direction. Therefore,

$$\hat{\rho} = \frac{\frac{\partial \mathbf{r}}{\partial \rho}}{\left\| \frac{\partial \mathbf{r}}{\partial \rho} \right\|}$$

$$\begin{aligned}
&= \frac{\cos \phi \hat{i} + \sin \phi \hat{j}}{|\cos \phi \hat{i} + \sin \phi \hat{j}|} \\
&= \cos \phi \hat{i} + \sin \phi \hat{j}.
\end{aligned}$$

Similarly,

$$\begin{aligned}
\hat{\phi} &= \frac{\frac{\partial \mathbf{r}}{\partial \phi}}{\left\| \frac{\partial \mathbf{r}}{\partial \phi} \right\|} \\
&= \frac{-\rho \sin \phi \hat{i} + \rho \cos \phi \hat{j}}{\left\| -\rho \sin \phi \hat{i} + \rho \cos \phi \hat{j} \right\|} \\
&= -\sin \phi \hat{i} + \cos \phi \hat{j}.
\end{aligned}$$

Thus, we can see that the  $\hat{\rho}, \hat{\phi}$  axis is orthogonal.

$$\begin{aligned}
\frac{\partial \hat{\rho}}{\partial \phi} &= -\sin \phi \hat{i} + \cos \phi \hat{j} \\
&= \hat{\phi}, \\
\frac{\partial \hat{\phi}}{\partial \phi} &= -\hat{\rho}, \\
\frac{\partial \hat{\phi}}{\partial \rho} &= 0,
\end{aligned}$$

and

$$\frac{\partial \hat{\rho}}{\partial \rho} = 0$$

**Example (Velocity).**

$$\begin{aligned}
\mathbf{v} &= \frac{d\mathbf{s}}{dt} \\
&= \frac{d}{dt} (x\hat{i}) + \frac{d}{dt} (y\hat{j}).
\end{aligned}$$

In the case of cartesian coordinates,  $\hat{i}$  and  $\hat{j}$  are constants.

$$= v_x \hat{i} + v_y \hat{j}$$

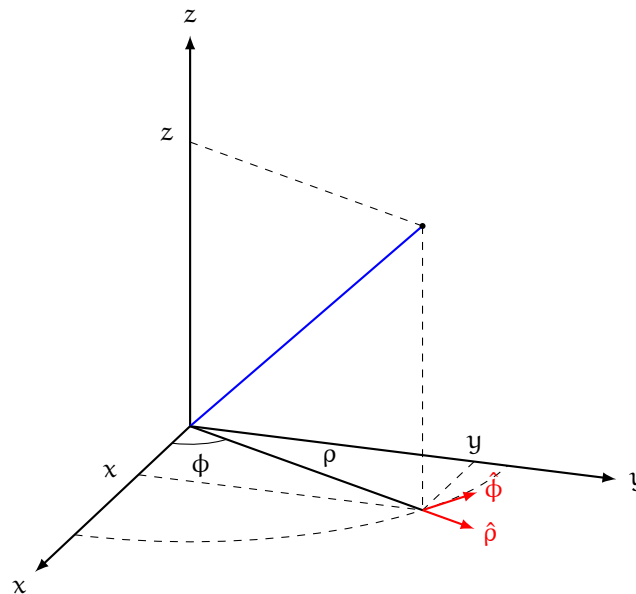
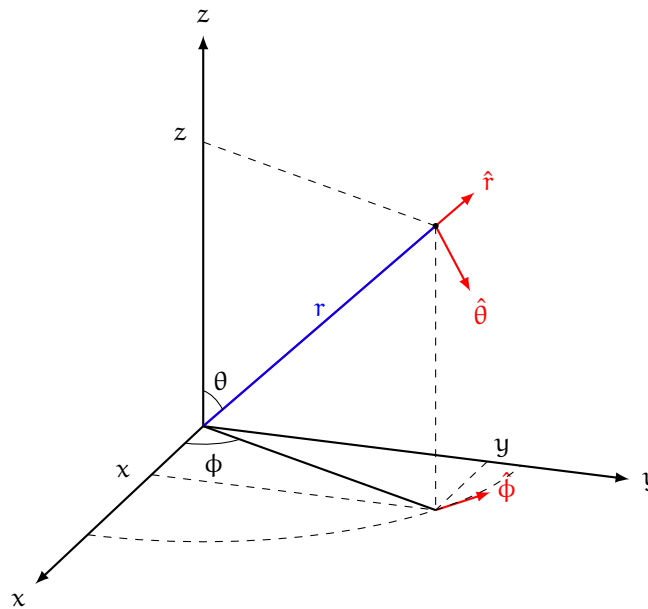
When we examine polar coordinates, since  $\hat{\rho}$  and  $\hat{\phi}$  are position-dependent, we must use the chain rule.<sup>1</sup>

$$\begin{aligned}
\mathbf{v} &= \frac{d\mathbf{s}}{dt} \\
&= \frac{d\rho}{dt} \hat{\rho} + \rho \frac{d\hat{\rho}}{dt} \\
&= \frac{d\rho}{dt} \hat{\rho} + \rho \left( \overset{0}{\cancel{\frac{\partial \hat{\rho}}{\partial \rho} \frac{d\rho}{dt}}} + \underbrace{\frac{\partial \hat{\rho}}{\partial \phi} \frac{d\phi}{dt}}_{=\hat{\phi}} \right) \\
&= \frac{d\rho}{dt} \hat{\rho} + \rho \frac{d\phi}{dt} \hat{\phi} \\
&= \dot{\rho} \hat{\rho} + \rho \dot{\phi} \hat{\phi}.
\end{aligned}$$

Notice that  $\dot{\rho}$  is the radial velocity and  $\dot{\phi} = \omega$  is the angular velocity.

<sup>1</sup>Note that  $\hat{\rho} = \hat{\rho}(\rho, \phi)$  and  $\hat{\phi} = \hat{\phi}(\rho, \phi)$ .

## Spherical and Cylindrical Coordinates



Polar	Cylindrical	Spherical
$\mathbf{s} = s(\rho, \phi)$	$\mathbf{s} = s(\rho, \phi, z)$	$\mathbf{s} = s(r, \phi, \theta)$
$\mathbf{s} = \begin{pmatrix} \rho \cos \phi \\ \rho \sin \phi \end{pmatrix}$	$\mathbf{s} = \begin{pmatrix} \rho \cos \phi \\ \rho \sin \phi \\ z \end{pmatrix}$	$\mathbf{s} = \begin{pmatrix} r \cos \phi \sin \theta \\ r \sin \phi \sin \theta \\ r \cos \theta \end{pmatrix}$

Here,<sup>II</sup>  $\phi$  denotes the polar angle and  $\theta$  denotes the azimuthal angle. Notice that  $\phi \in [0, 2\pi)$  and  $\theta \in [0, \pi]$ .

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<sup>II</sup>Physicists amirite?

We can see that  $\hat{\rho}$ ,  $\hat{\phi}$ , and  $\hat{\theta}$  in spherical coordinates are also position-dependent.

$$\begin{aligned}\hat{r} &= \frac{\frac{\partial \mathbf{s}}{\partial r}}{\left\| \frac{\partial \mathbf{s}}{\partial r} \right\|} \\ &= \sin \theta \cos \phi \hat{i} + \sin \theta \sin \phi \hat{j} + \cos \theta \hat{k} \\ \hat{\phi} &= \frac{\frac{\partial \mathbf{s}}{\partial \phi}}{\left\| \frac{\partial \mathbf{s}}{\partial \phi} \right\|} \\ &= -\sin \phi \hat{i} + \cos \phi \hat{j} \\ \hat{\theta} &= \frac{\frac{\partial \mathbf{s}}{\partial \theta}}{\left\| \frac{\partial \mathbf{s}}{\partial \theta} \right\|} \\ &= \cos \phi \cos \theta \hat{i} + \cos \theta \sin \phi \hat{j} - \sin \theta \hat{k}\end{aligned}$$

### Scale Factors and Jacobians

Coordinate System	Line Element	Area Element	Volume Element
Polar	$d\mathbf{s} = \hat{\rho}d\rho + \rho\hat{\phi}d\phi$	$d\mathbf{a} = r dr d\phi$	—
Cylindrical	$d\mathbf{s} = \hat{\rho}d\rho + \rho\hat{\phi}d\phi + \hat{k}dz$	—	$d\mathbf{v} = r dr d\phi dz$
Spherical	$d\mathbf{s} = \hat{r}dr + r \sin \theta \hat{\phi}d\phi + r\hat{\theta}d\theta$	$d\mathbf{a} = r^2 \sin \theta d\phi d\theta$	$d\mathbf{v} = r^2 \sin \theta dr d\phi d\theta$

In cylindrical coordinates, we can use the chain rule to find the value of  $d\mathbf{r}$ :

$$d\mathbf{r} = \hat{\rho}d\rho + \rho\hat{\phi}d\phi + \hat{k}dz.$$

The extra factor of  $\rho$  in the expression of  $\rho\hat{\phi}d\phi$  is the *scale factor* on  $\phi$ .

Similarly, in spherical coordinates, we have

$$d\mathbf{r} = \hat{r}dr + r \sin \theta \hat{\phi}d\phi + r\hat{\theta}d\theta,$$

with scale factors of  $r \sin \theta$  on  $\hat{\phi}d\phi$  and  $r$  on  $\hat{\theta}d\theta$ .

When we go from line elements (of the form  $d\mathbf{r}$ ) to area elements (of the form  $d\mathbf{a}$ ), we can see that the area element in polar coordinates is  $d\mathbf{a} = \rho d\rho d\phi$  — we need the extra factor of  $\rho$  to account for the fact that the magnitude of the area element scales with the radius.

Similarly, the volume element in cylindrical coordinates is  $d\mathbf{v} = r dr d\phi dz$  and the volume element in spherical coordinates is  $r^2 \sin \theta dr d\phi d\theta$ .

Recall that the definition of an angle  $\phi$  that subtends an arc length  $s$  is  $\phi = \frac{s}{r}$ , where  $r$  is the radius of a circle. We can imagine a similar concept on a sphere — a *solid angle* measured in steradians is of the form  $\Omega = \frac{A}{r^2}$ , where  $A$  denotes the surface area subtended by the angle  $\Omega$ . In particular, since  $d\Omega = \frac{dA}{r^2}$ , we find that  $d\Omega = \sin \theta d\phi d\theta$ .

When we are dealing with products of scale factors, we need to use the Jacobian to determine the proper scale factor on any given element:

$$\begin{aligned}d\mathbf{a} &= dx dy \\ &= |J| du dv,\end{aligned}$$

where  $|J|$  denotes the determinant of the Jacobian matrix. We write the Jacobian as follows:

$$J = \frac{\partial(x, y)}{\partial(u, v)} = \begin{pmatrix} \frac{\partial x}{\partial u} & \frac{\partial y}{\partial u} \\ \frac{\partial x}{\partial v} & \frac{\partial y}{\partial v} \end{pmatrix}.$$

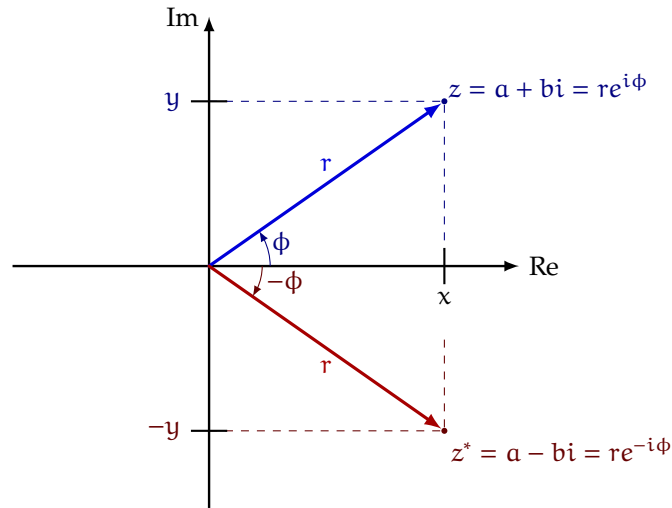
We specifically desire the determinant:

$$|J| = \frac{\partial x}{\partial u} \frac{\partial y}{\partial v} - \frac{\partial y}{\partial u} \frac{\partial x}{\partial v}.$$

## Complex Numbers

Quantity	Expression and/or Criterion
Cartesian form	$z = a + bi$
Polar form	$z = re^{i\phi}$
$r$	$\sqrt{a^2 + b^2}$
$\phi$	$\arg z = \arctan\left(\frac{b}{a}\right)$
Cartesian $z^*$	$z^* = a - bi$
Polar $z^*$	$z = re^{-i\phi}$
$ z $	$\sqrt{zz^*}$
$\operatorname{Re}(z)$	$\operatorname{Re}(z) = \frac{z+z^*}{2}$
$\operatorname{Im}(z)$	$\operatorname{Im}(z) = \frac{z-z^*}{2i}$
$\cos \phi$	$\frac{e^{i\phi} + e^{-i\phi}}{2}$
$\sin \phi$	$\frac{e^{i\phi} - e^{-i\phi}}{2i}$
$e^{i\phi}$	$\cos \phi + i \sin \phi$
$e^{in\phi}$	$\cos(n\phi) + i \sin(n\phi)$

## Introduction



A complex number is denoted

$$z = a + bi$$

where  $i^2 = -1$  and  $a, b \in \mathbb{R}$ . This is known as the cartesian representation. However, we can also imagine  $z$  as the polar representation:

$$z = re^{i\phi},$$

where  $\phi = \arg z$  is known as the argument, and  $r = |z|$  is the modulus. We can see the relation between the cartesian and polar representations through Euler's identity:<sup>III</sup>

$$r(\cos \phi + i \sin \phi) = re^{i\phi}.$$

We denote the conjugate of  $z$  as  $z^*$ <sup>IV</sup>, found by  $z^* = a - bi = re^{-i\phi}$ .

We find  $\text{Re}(z)$  and  $\text{Im}(z)$ , the real and imaginary parts of  $z$ , by

$$\begin{aligned}\text{Re}(z) &= \frac{z + z^*}{2} \\ \text{Im}(z) &= \frac{z - z^*}{2i}.\end{aligned}$$

We say that a complex number of the form  $e^{i\phi}$  is a *pure phase*, as  $|e^{i\phi}| = 1$ .

To find if some complex number  $z$  is purely real or purely imaginary, we can use the following criterion:

$$\begin{aligned}z \in \mathbb{R} &\Leftrightarrow z = z^* \\ z \in i\mathbb{R} &\Leftrightarrow z = -z^*.\end{aligned}$$

**Example** (Real, Imaginary, or Complex?). Consider

$$z_1 = i^i.$$

To find if this is purely real or complex, we take

$$z_1^* = (-i)^{-i}$$

<sup>III</sup>This can be proven relatively easily through substitution into the Taylor series, which is allowed because  $e^z$  is entire.

<sup>IV</sup>Physicists amirite?

$$\begin{aligned}
 &= \left( \frac{1}{-i} \right)^i \\
 &= i^i.
 \end{aligned}$$

Thus,  $z_1 \in \mathbb{R}$ . In order to determine the value of  $i^i$ , we substitute the polar form:

$$\begin{aligned}
 z_1 &= \left( e^{i\frac{\pi}{2}} \right)^i \\
 &= e^{-\frac{\pi}{2}}.
 \end{aligned}$$

### Some Trigonometry with Complex Exponentials

Consider  $z = \cos \phi + i \sin \phi$ . We can see that

$$\begin{aligned}
 \operatorname{Re}(z) &= \cos \phi \\
 &= \frac{(\cos \phi + i \sin \phi) + (\cos \phi - i \sin \phi)}{2} \\
 &= \frac{e^{i\phi} + e^{-i\phi}}{2} \\
 \operatorname{Im}(z) &= \sin \phi \\
 &= \frac{(\cos \phi + i \sin \phi) - (\cos \phi - i \sin \phi)}{2i} \\
 &= \frac{e^{i\phi} - e^{-i\phi}}{2i}.
 \end{aligned}$$

We can actually define  $\sin \phi$  and  $\cos \phi$  with the above derivation.

**Theorem** (De Moivre).

$$\begin{aligned}
 e^{inx} &= \cos(nx) + i \sin(nx) \\
 &= \left( e^{ix} \right)^n \\
 &= (\cos x + i \sin x)^n.
 \end{aligned}$$

**Example** (Finding  $\cos(2x)$  and  $\sin(2x)$ ).

$$\begin{aligned}
 \cos(2x) + i \sin(2x) &= (\cos x + i \sin x)^2 \\
 &= (\cos^2 x - \sin^2 x) + i(2 \sin x \cos x).
 \end{aligned}$$

Since the real parts and imaginary parts have to be equal, this means

$$\begin{aligned}
 \cos 2x &= \cos^2 x - \sin^2 x \\
 \sin^2 x &= 2 \sin x \cos x.
 \end{aligned}$$

In particular, we can see that  $e^{in\pi} = (-1)^n$  and  $e^{in\frac{\pi}{2}} = i^n$ .<sup>v</sup>

Additionally, we can see that for  $z = re^{i\phi}$ ,

$$\begin{aligned}
 z^{1/m} &= \left( re^{i\phi+2\pi n} \right)^{1/m} \\
 &= r^{1/m} e^{i\frac{1}{m}(\phi+2\pi n)},
 \end{aligned}$$

where  $n \in \mathbb{N}$  and  $m$  is fixed. For  $r = 1$ , we call these values the  $m$  roots of unity.

<sup>v</sup>This will be especially useful when we get to Fourier series.



**Example** (Waves and Oscillations). Recall that for a wave with spatial frequency  $k$ , angular frequency  $\omega$ , and amplitude  $A$ , the wave is represented by

$$f(x, t) = A \cos(kx - \omega t).$$

The speed of a wave  $v$  is equal to  $\frac{\omega}{k}$ .

Simple harmonic motion is characterized by the solution to the differential equation  $\ddot{x} = -\omega^2 x$ , where  $x$  denotes position. In simple harmonic motion, there is no spatial motion, meaning our function is only of time:

$$\begin{aligned} f(t) &= A \cos \omega t \\ &= \operatorname{Re} \left( A e^{i\omega t} \right). \end{aligned}$$

As a result of the representation of complex numbers in polar form, we can do math entirely in exponentials, then take the real part of our solution to find  $f(t)$ .

Unfortunately, in the real world, there is friction; as a result, our oscillation is damped by an exponential factor.

**Example** (Hyperbolic Sine and Hyperbolic Cosine). We wish to calculate  $\cos ix$  and  $\sin ix$ .

$$\begin{aligned} \cos ix &= \frac{1}{2} \left( e^{i(ix)} + e^{-i(ix)} \right) \\ &= \frac{e^{-x} + e^x}{2} \end{aligned}$$

We define  $\cosh x = \cos(ix)$ . Additionally,

$$\begin{aligned} -i \sin ix &= -i \frac{1}{2i} \left( e^{i(ix)} - e^{-i(ix)} \right) \\ &= i \frac{e^{ix} - e^{-ix}}{2i} \\ &= \frac{e^x - e^{-x}}{2}. \end{aligned}$$

We define  $\sinh x = -i \sin(ix)$ .

Similar to how  $\cos^2 x + \sin^2 x = 1$ , we can find that  $\cosh^2 x - \sinh^2 x = 1$ .

## Index Algebra

We usually denote vectors by either  $\vec{A}$ ,  $\mathbf{A}$ , or

$$\begin{pmatrix} a_1 \\ a_2 \\ \vdots \\ a_n \end{pmatrix},$$

which is defined by a basis.

If we imagine we are in  $n$ -dimensional space, we can let  $A_i$  where  $i = 1, 2, \dots, n$  denote both

- the  $i$ th component of  $\vec{A}$ ;
- the entire vector  $\vec{A}$  (since  $i$  can be arbitrary).

### Contractions and Dummy Indices

Consider  $C = AB$ , where  $A, B$  are  $n \times m$  and  $m \times p$  matrices respectively.

$$C = \begin{pmatrix} A_{11} & A_{12} & \cdots & A_{1m} \\ \vdots & \vdots & \ddots & \vdots \\ A_{n1} & A_{n2} & \cdots & A_{nm} \end{pmatrix} \begin{pmatrix} B_{11} & B_{12} & \cdots & B_{1p} \\ \vdots & \vdots & \ddots & \vdots \\ B_{m1} & B_{m2} & \cdots & B_{mp} \end{pmatrix}.$$

**Definition** (Matrix Multiplication in Index Notation). For matrices  $A$  and  $B$ , where  $A$  is an  $m \times n$  and  $B$  is a  $n \times p$  matrix, we write

$$C_{ij} = \sum_{k=1}^n A_{ik} B_{kj}$$

We say that  $k$  is a dummy index, since  $k$  takes values from 1 to  $n$ . Note that the value we calculate is  $C_{ij}$ ; in other words, in the sum  $\sum_k A_{ik} B_{kj}$ , the indices of the form  $ij$  are the “net indices” from the multiplication.

Note that if  $C = BA$ , then

$$\begin{aligned} C_{ij} &= \sum_{k=1}^n B_{ik} A_{kj} \\ &= \sum_{k=1}^n A_{kj} B_{ik} \\ &\neq \sum_{k=1}^n A_{ik} B_{kj}. \end{aligned}$$

The corresponding fact is that  $AB \neq BA$  necessarily.

Note that the index that is summed over always appears exactly twice.

**Definition** (Symmetric Matrix). Let  $C$  be a matrix. Then, we say  $C$  is symmetric if

$$C_{ij} = C_{ji}$$

**Definition** (Antisymmetric Matrix). Let  $C$  be a matrix. We say  $C$  is antisymmetric if

$$C_{ij} = -C_{ji}.$$

We can always decompose a random matrix into the sum of a symmetric matrix and an antisymmetric matrix.

### Two Special Tensors

Name	Notation	Definition
Kronecker Delta	$\delta_{ij}$	$\delta_{ij} = \begin{cases} 1 & i = j \\ 0 & i \neq j \end{cases}$
Levi-Civita Symbol	$\epsilon_{ijk}$	$\epsilon_{ijk} = \begin{cases} 1 & (i, j, k) = (1, 2, 3) \text{ cyclically} \\ -1 & (i, j, k) = (2, 1, 3) \text{ cyclically} \\ 0 & \text{else} \end{cases}$

Order of (i, j, k)	Value of $\epsilon_{ijk}$
1, 2, 3	1
3, 1, 2	1
2, 3, 1	1
1, 3, 2	-1
2, 1, 3	-1
3, 2, 1	-1
else	0

Value	Index Notation
$\mathbf{A} \times \mathbf{B}$	$\sum_{i,j,k} \epsilon_{ijk} A_i B_j \hat{e}_k$
$(\mathbf{A} \times \mathbf{B})_\ell$	$\sum_{i,j} \epsilon_{ij\ell} A_i B_j$
$(\hat{e}_i \times \hat{e}_j) \cdot \hat{e}_k$	$\epsilon_{ijk}$
$B_i$	$\sum_{\alpha} B_{\alpha} \delta_{\alpha i}$
$\mathbf{A} \cdot \mathbf{B}$	$\sum_{i,j} A_i B_j \delta_{ij}$
$\sum_{j,k} \epsilon_{mjk} \epsilon_{njk}$	$2\delta_{mn}$
$\sum_{\ell} \epsilon_{mnl} \epsilon_{ijl}$	$\delta_{mi} \delta_{nj} - \delta_{mj} \delta_{ni}$

**Definition** (Kronecker Delta). The Kronecker Delta,  $\delta_{ij}$ , is the tensor that denotes the identity matrix.

$$\delta_{ij} = \begin{cases} 1 & i = j \\ 0 & i \neq j \end{cases}$$

**Example** (Extracting an Index). Consider  $A$  as vector. Then,

$$\sum_i A_i \delta_{ij} = A_j.$$

In other words, the Kronecker Delta collapses the sum to the  $j$ th index.

**Example** (Orthonormal Basis from Kronecker Delta). Let  $\{\hat{e}_i\}_{i=1}^n$  be a basis for some vector space  $V$ . If

$$\hat{e}_i \cdot \hat{e}_j = \delta_{ij}$$

for every  $i, j$ , then  $\{\hat{e}_i\}_{i=1}^n$  is an orthonormal basis for  $V$ .

**Definition** (Levi-Civita Symbol). In two dimensions, as a matrix, we write

$$\epsilon_{ij} = \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix},$$

meaning

$$\epsilon_{ij} = \begin{cases} 1 & i = 1, j = 2 \\ -1 & i = 2, j = 1 \\ 0 & \text{else} \end{cases}.$$

The Levi–Civita Symbol is antisymmetric, just as the Kronecker Delta is symmetric.

In three dimensions, we define

$$\epsilon_{ijk} = \begin{cases} 1 & (i, j, k) = (1, 2, 3) \text{ cyclically} \\ -1 & (i, j, k) = (2, 1, 3) \text{ cyclically} \\ 0 & \text{else} \end{cases}$$

In other words,  $\epsilon_{ijk} = -\epsilon_{jik}$ .

**Exercise** (Relations between  $\delta_{ij}$  and  $\epsilon_{ijk}$ ).

$$\begin{aligned} \sum_{j,k} \epsilon_{mjk} \epsilon_{njk} &= 2\delta_{mn} \\ \sum_{\ell} \epsilon_{mnl} \epsilon_{ijl} &= \delta_{mi} \delta_{nj} - \delta_{mj} \delta_{ni} \end{aligned}$$

**Definition** (Dot Product). Let  $\{\hat{e}_i\}_{i=1}^n$  be an orthonormal basis for  $V$ . Let  $\mathbf{A} = \sum_i A_i \hat{e}_i$  and  $\mathbf{B} = \sum_i B_i \hat{e}_i$ . Then,

$$\begin{aligned} \mathbf{A} \cdot \mathbf{B} &= \sum_{i,j} (A_i \hat{e}_i) \cdot (B_j \hat{e}_j) \\ &= \sum_{i,j} A_i B_j (\hat{e}_i \cdot \hat{e}_j) \\ &= \sum_{i,j} A_i B_j \delta_{ij} \\ &= \sum_i A_i B_i \end{aligned}$$

**Definition** (Cross Product). Let  $\{\hat{e}_i\}_{i=1}^3$  be the standard basis over  $\mathbb{R}^3$ . Let  $\mathbf{A} = \sum_i A_i \hat{e}_i$  and  $\mathbf{B} = \sum_i B_i \hat{e}_i$ . Then,

$$\begin{aligned} \mathbf{A} \times \mathbf{B} &= \sum_{i,j} (A_i \hat{e}_i) \times (B_j \hat{e}_j) \\ &= \sum_{i,j} A_i B_j (\hat{e}_i \times \hat{e}_j) \\ &= \sum_{i,j,k} A_i B_j (\epsilon_{ijk} \hat{e}_k) \end{aligned}$$

Instead of asking about  $\mathbf{A} \times \mathbf{B}$ , we ask about  $(\mathbf{A} \times \mathbf{B})_\ell$ , yielding

$$\begin{aligned} (\mathbf{A} \times \mathbf{B})_\ell &= (\mathbf{A} \times \mathbf{B}) \cdot \hat{e}_\ell \\ &= \left( \sum_{i,j,k} A_i B_j (\epsilon_{ijk} \hat{e}_k) \right) \cdot \hat{e}_\ell \\ &= \sum_{i,j} \epsilon_{ij\ell} A_i B_j \end{aligned}$$

**Remark:** This notation for  $\mathbf{A} \times \mathbf{B}$  automatically shows us that

$$(\mathbf{B} \times \mathbf{A})_\ell = \sum_{i,j} \epsilon_{ij\ell} B_i A_j$$

$$\begin{aligned}
&= - \sum_{i,j} \epsilon_{jil} B_i A_j \\
&= - \sum_{i,j} \epsilon_{jil} A_j B_i \\
&= - \sum_{i,j} \epsilon_{ijl} A_i B_j \quad i, j \text{ are dummy indices} \\
&= - (\mathbf{A} \times \mathbf{B})_\ell.
\end{aligned}$$

**Example** (Central Force and Angular Momentum). A central force is defined by

$$\mathbf{F} = f(r)\hat{\mathbf{r}},$$

where  $\hat{\mathbf{r}}$  is a radial vector.

Angular momentum is defined by

$$\mathbf{L} = \mathbf{r} \times \mathbf{p},$$

where  $\mathbf{r}$  denotes position and  $\mathbf{p}$  denotes momentum. Then,

$$\begin{aligned}
\frac{d\mathbf{L}}{dt} &= \frac{d}{dt} (\mathbf{r} \times \mathbf{p}) \\
&= \left( \frac{d}{dt} \mathbf{r} \times \mathbf{p} \right) + \mathbf{r} \times \left( \frac{d\mathbf{p}}{dt} \right) \\
&= m \left( \frac{d}{dt} \mathbf{r} \times \frac{d}{dt} \mathbf{r} \right) + \mathbf{r} \times (f(r)\hat{\mathbf{r}}) \\
&= f(r) (\mathbf{r} \times \hat{\mathbf{r}}).
\end{aligned}$$

This implies that  $\frac{d\mathbf{L}}{dt} = 0$  under a central force.

**Example** (Determinant). Let  $\mathbf{M} = M_{ij}$  be square. We denote  $\mathbf{M}_i$  to be the vector denoting the  $i$ th-row. Then,

$$\begin{aligned}
m &= |\mathbf{M}| \\
&= \mathbf{M}_1 \cdot (\mathbf{M}_2 \times \mathbf{M}_3) \\
&= \mathbf{M}_3 \cdot (\mathbf{M}_1 \times \mathbf{M}_2) \\
&= \mathbf{M}_2 \cdot (\mathbf{M}_3 \times \mathbf{M}_1).
\end{aligned}$$

**Example** (Trace). Let  $\mathbf{M} = M_{ij}$  be a square matrix. We define  $\text{tr}(\mathbf{M}) = \sum_i M_{ii}$ . Equivalently,

$$\begin{aligned}
\text{tr}(\mathbf{M}) &= \sum_{ij} M_{ij} \delta_{ij} \\
&= \sum_i M_{ii}.
\end{aligned}$$

Note that

$$\begin{aligned}
\text{tr}(\mathbf{I}_n) &= \sum_i \delta_{ii} \\
&= n.
\end{aligned}$$

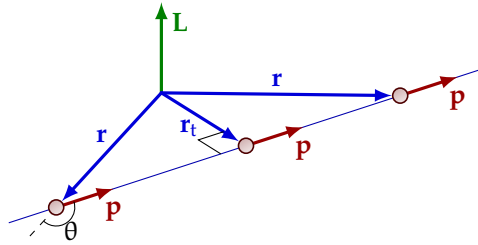
When we upgrade to 3 matrices, we take

$$\text{tr}(ABC) = \sum_{i,j} \left( \sum_{k,\ell} A_{ik} B_{k\ell} C_{\ell j} \right) \delta_{ij}$$

$$\begin{aligned}
&= \sum_{i,k,\ell} A_{ik} B_{k\ell} C_{\ell i} \\
&= \sum_{i,k,\ell} C_{\ell i} A_{ik} B_{k\ell} \\
&= \text{tr}(CAB).
\end{aligned}$$

In other words, the trace is invariant under cyclic permutations.

**Example** (Angular Momentum, Revisited).



Recall that

$$\begin{aligned}
\mathbf{L} &= \mathbf{r} \times \mathbf{p}, \\
&= I\boldsymbol{\omega}.
\end{aligned}$$

where  $\mathbf{p} = m\dot{\mathbf{x}}$ , and  $I$  denotes the moment of inertia. Note that  $I \sim m r^2$ . On a more fundamental level, it is the case that the first equation,  $\mathbf{L} = \mathbf{r} \times \mathbf{p}$ , is the “true” definition of  $\mathbf{L}$ .

Consider a small portion  $m_\alpha$  about some axis at radius  $r_\alpha$  and momentum  $\mathbf{p}_\alpha$ . Then, we have

$$\begin{aligned}
\mathbf{L}_\alpha &= \sum_{\alpha} \mathbf{r}_\alpha \times \mathbf{p}_\alpha \\
&= \sum_{\alpha} m_\alpha (\mathbf{r}_\alpha \times (\boldsymbol{\omega} \times \mathbf{r}_\alpha)).
\end{aligned}$$

In the infinitesimal case (i.e., as  $\alpha \rightarrow 0$ ), we get

$$\mathbf{L} = \int \mathbf{r} \times (\boldsymbol{\omega} \times \mathbf{r}) \rho \, d\tau,$$

where  $\rho$  denotes volume density. Applying the identity  $\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) = \mathbf{B}(\mathbf{A} \cdot \mathbf{C}) - \mathbf{C}(\mathbf{A} \cdot \mathbf{B})$ , we find

$$\mathbf{L} = \int (\boldsymbol{\omega} (\mathbf{r} \cdot \mathbf{r}) - \mathbf{r} (\mathbf{r} \cdot \boldsymbol{\omega})) \rho \, d\tau.$$

Switching to index notation, we have

$$\begin{aligned}
L_i &= \int \left( \omega_i r^2 - r_i \sum_j r_j \omega_j \right) \rho \, d\tau \\
&= \sum_j \int \omega_j \left( \delta_{ij} r^2 - r_i r_j \right) \rho \, d\tau \\
&= \sum_j \omega_j \underbrace{\left( \int \left( \delta_{ij} r^2 - r_i r_j \right) \rho \, d\tau \right)}_{\text{moment of inertia tensor}} \\
&= \sum_j I_{ij} \omega_j.
\end{aligned}$$

## Binomial Theorem

The binomial theorem allows us to calculate the expansion

$$(x + y)^n = \sum_{m=0}^n \binom{n}{m} x^{n-m} y^m.$$

In the case of  $(x + y)^2 = x^2 y^0 + 2x^1 y^1 + x^0 y^2 = x^2 + 2xy + y^2$ . Recall that

$$\binom{n}{m} = \frac{n!}{m! (n - m)!}.$$

Recall that  $0! = 1$ .

## Infinite Series

Let

$$S = \sum_{k=0}^{\infty} a_k$$

be an infinite series. We are often curious as to the convergence of this sum (for a variety of reasons). Formally, we have to invoke partial sums

$$S_N = \sum_{k=0}^N a_k,$$

and see if the sequence of partial sums is convergent. However, we will prefer to use series convergence tests.

**Example** (Geometric Series). Let

$$\begin{aligned} S &= \sum_{k=0}^{\infty} r^k \\ &= 1 + r + r^2 + \dots \end{aligned}$$

Then, we have

$$\begin{aligned} S_N &= \sum_{k=0}^N r^k \\ rS_N &= \sum_{k=0}^N r^{k+1}. \end{aligned}$$

Subtracting, we get

$$\begin{aligned} (1 - r)S_N &= 1 - r^{N+1} \\ S_N &= \frac{1 - r^{N+1}}{1 - r}. \end{aligned}$$

In the limit, we expect that if  $r \rightarrow \infty$ , and  $r < 1$ , then  $r^{N+1} \rightarrow 0$ . In the infinite case, we have

$$\begin{aligned} S &= \sum_{k=0}^{\infty} r^k \\ &= \frac{1}{1 - r}, \end{aligned}$$

if  $r < 1$ .

There are a few prerequisites for series convergence:

- there exists some  $K$  for which for all  $k \geq K$ ,  $a_{k+1} \leq a_k$ ;
- $\lim_{k \rightarrow \infty} a_k < \infty$ ;
- we need the series to reduce “quickly” enough.

**Example (Ratio Test).** A series  $S = \sum_k a_k$  converges if the ratio of consecutive terms is (eventually) less than 1:

$$r = \lim_{k \rightarrow \infty} \frac{a_{k+1}}{a_k} < 1.$$

**Example (Applying the Ratio Test).** Consider  $S = \sum_k \frac{1}{k!}$ . Then,

$$\begin{aligned} r &= \lim_{k \rightarrow \infty} \frac{\frac{1}{(k+1)!}}{\frac{1}{k!}} \\ &= \lim_{k \rightarrow \infty} \frac{1}{k+1} \\ &= 0 < 1 \end{aligned}$$

**Example (Riemann Zeta Function).** We write

$$\zeta(s) = \sum_{k=1}^{\infty} \frac{1}{k^s}.$$

In order to evaluate the convergence of the Riemann zeta function. We have

$$\begin{aligned} r &= \lim_{k \rightarrow \infty} \frac{\frac{1}{(k+1)^s}}{\frac{1}{k^s}} \\ &= \lim_{k \rightarrow \infty} \left( \frac{k}{k+1} \right)^s \\ &= 1. \end{aligned}$$

Unfortunately, this means the ratio test is inconclusive.

For examples of evaluations of the zeta function, we have

$$\begin{aligned} \zeta(1) &= 1 + \frac{1}{2} + \frac{1}{3} + \cdots \\ \zeta(2) &= 1 + \frac{1}{4} + \frac{1}{9} + \cdots \\ &= \frac{\pi^2}{6}. \end{aligned}$$