

Новое *NN*-взаимодействие JISP: описание *NN*-рассеяния и ядер *s*- и *p*-оболочек в подходе *ab initio*

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Ab initio:

- Без модельных предположений (например, без введения инертного кора)
- *Ab initio* подходы:
 - уравнения Фаддеева;
 - метод гиперсферических функций;
 - Green function's Monte Carlo;
 - no-core shell model;
 - coupled-cluster approach

Coupled cluster approach:

$$|\Psi\rangle = \exp(T)|\Phi_0\rangle . \quad (1)$$

Here $|\Phi_0\rangle$ is an uncorrelated reference Slater determinant which might be either the Hartree-Fock (HF) state or a naive filling of the oscillator single-particle basis. Correlations are introduced through the exponential $\exp(T)$ operating on $|\Phi_0\rangle$. The operator T is a sum of n -particle- n -hole excitation operators $T = T_1 + T_2 + \dots$ of the form,

$$T_n = \sum_{a_1 \dots a_n, i_1 \dots i_n} t_{i_1 \dots i_n}^{a_1 \dots a_n} a_{a_1}^\dagger \cdots a_{a_n}^\dagger a_{i_n} \cdots a_{i_1} , \quad (2)$$

where i_1, i_2, \dots are summed over hole states and a_1, a_2, \dots are summed over particle states. One obtains the algebraic equation for the excitation amplitudes $t_{ij\dots}^{ab\dots}$ by left-projecting the similarity-transformed Hamiltonian with an n -particle- n -hole excited Slater determinant giving

$$\langle \Phi_{ij\dots}^{ab\dots} | (H_N \exp(T))_C | \Phi_0 \rangle = 0 , \quad (3)$$

Модель оболочек.

- Две основные современные схемы:
- Monte Carlo shell model
- *m*-scheme + Lanczos (в том числе no-core shell model); осцилляторный базис

m -scheme + Lanczos:

- Идея: неполная диагонализация в большом базисе быстрее и проще, чем построение “правильного” базиса с определенными J , L и т.д.
- Translational invariant SM (координаты Якоби, J , L и т.д.) \Rightarrow
 - \Rightarrow no-core SM (детерминанты Слэтера с фиксированным m , но не J , L и т.д.); J приобретает определенное значение в результате диагонализации.
- Получаем сразу несколько уровней одной четности, но с разными J .

Lanczos iterations:

$$\tilde{a}_1 = H a_0;$$

$$\tilde{a}_1 \implies a_1 : \langle a_1 | a_0 \rangle = 0;$$

$$\tilde{a}_2 = H a_1;$$

$$\tilde{a}_2 \implies a_2 : \langle a_2 | a_0 \rangle = \langle a_2 | a_1 \rangle = 0;$$

$$\tilde{a}_3 = H a_2;$$

.....

$\langle a_i | H | a_j \rangle$ — трехдиагональная матрица.

- Диагонализация сравнительно небольшой трехдиагональной матрицы, дающей хорошее основное и нижайшие возбужденные состояния

Lanczos iterations:

$$a_0 = \alpha_0 x_0 + \alpha_1 x_1 + \alpha_2 x_2 + \dots$$

$$a_1 = E_0 \alpha_0 x_0 + E_1 \alpha_1 x_1 + E_2 \alpha_2 x_2 + \dots$$

$$a_2 = E_0^2 \alpha_0 x_0 + E_1^2 \alpha_1 x_1 + E_2^2 \alpha_2 x_2 + \dots$$

.....

$$a_n = E_0^n \alpha_0 x_0 + E_1^n \alpha_1 x_1 + E_2^n \alpha_2 x_2 + \dots$$

No-core shell model:

$$H_A = \frac{1}{A} \sum_{i < j}^A \frac{(\mathbf{p}_i - \mathbf{p}_j)^2}{2m} + \sum_{i < j}^A V_{NN,ij} \quad (1)$$
$$+ \sum_{i < j < k}^A V_{NNN,ijk},$$

where m is the nucleon mass, $V_{NN,ij}$ is the two-nucleon interaction (including both strong and electromagnetic components), and $V_{NNN,ijk}$ is the three-nucleon interaction, should be arranged as spurious-free linear combinations of basis states.

To achieve this, the auxiliary Hamiltonian

$$H_{\text{NCSM}} = H_A + \beta \tilde{Q}_0 \quad (2)$$

is conventionally diagonalized within the NCSM instead of the Hamiltonian (1). Here,

$$\tilde{Q}_0 \equiv H_{\text{CM}} - \frac{3}{2} \hbar \Omega, \quad (3)$$

$$H_{\text{CM}} = T_{\text{CM}} + U_{\text{CM}} \quad (4)$$

is the harmonic oscillator CM Hamiltonian, T_{CM} is the CM kinetic energy operator, and

$$U_{\text{CM}} = \frac{1}{2} A m \Omega^2 \mathbf{R}^2, \quad (5)$$

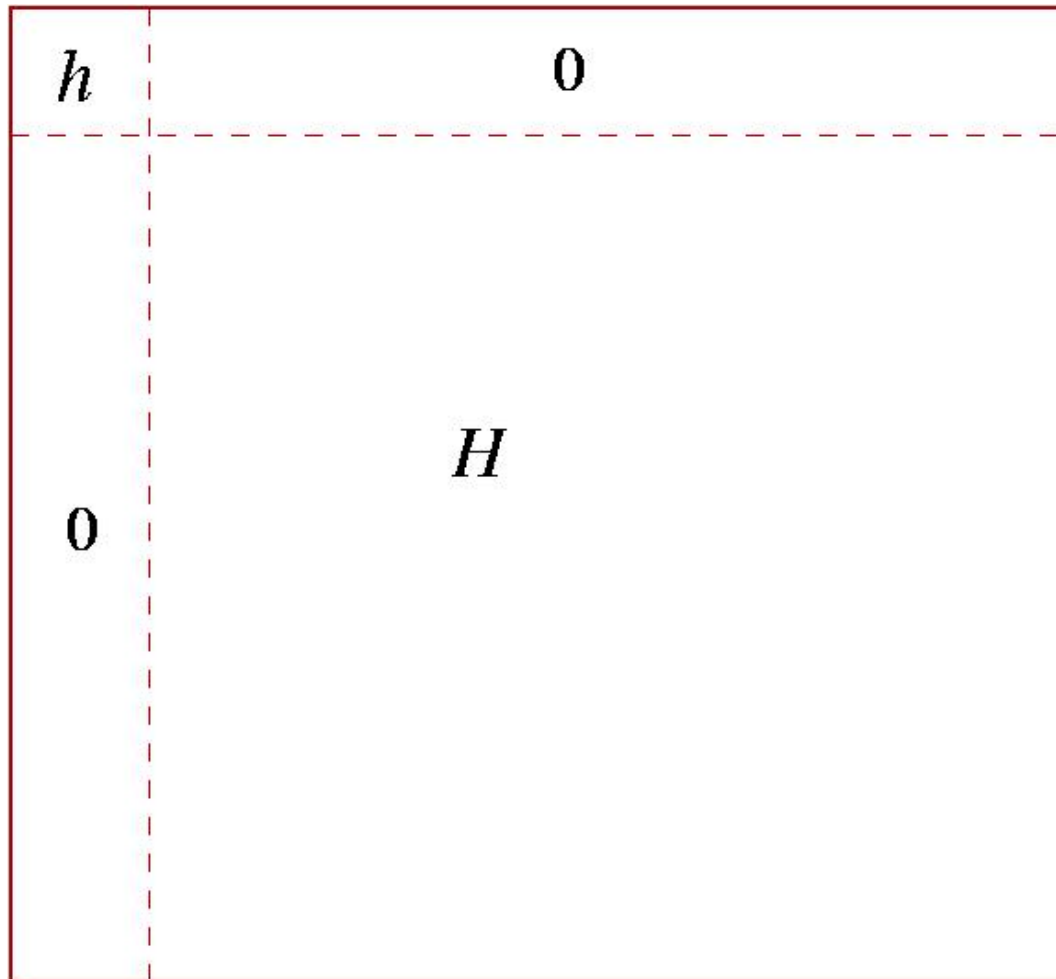
where

$$\mathbf{R} = \frac{1}{A} \sum_{i=1}^A \mathbf{r}_i. \quad (6)$$

Effective interactions:

- “Обычная” (с кором) модель оболочек (тяжелые ядра): G -матрица или просто феноменология.
- No-core SM: Lee–Suzuki transformation, т.е. *ab initio* NN -взаимодействие, полученное из исходного “голого” NN -взаимодействия.

Lee–Suzuki transformation:



“Кластерное” разложение:

$$H \implies h_2 + h_3 + \dots$$

Обычно ограничиваются h_2 или $h_2 + h_3$.

Highlights

JISP = J-matrix inverse scattering potential

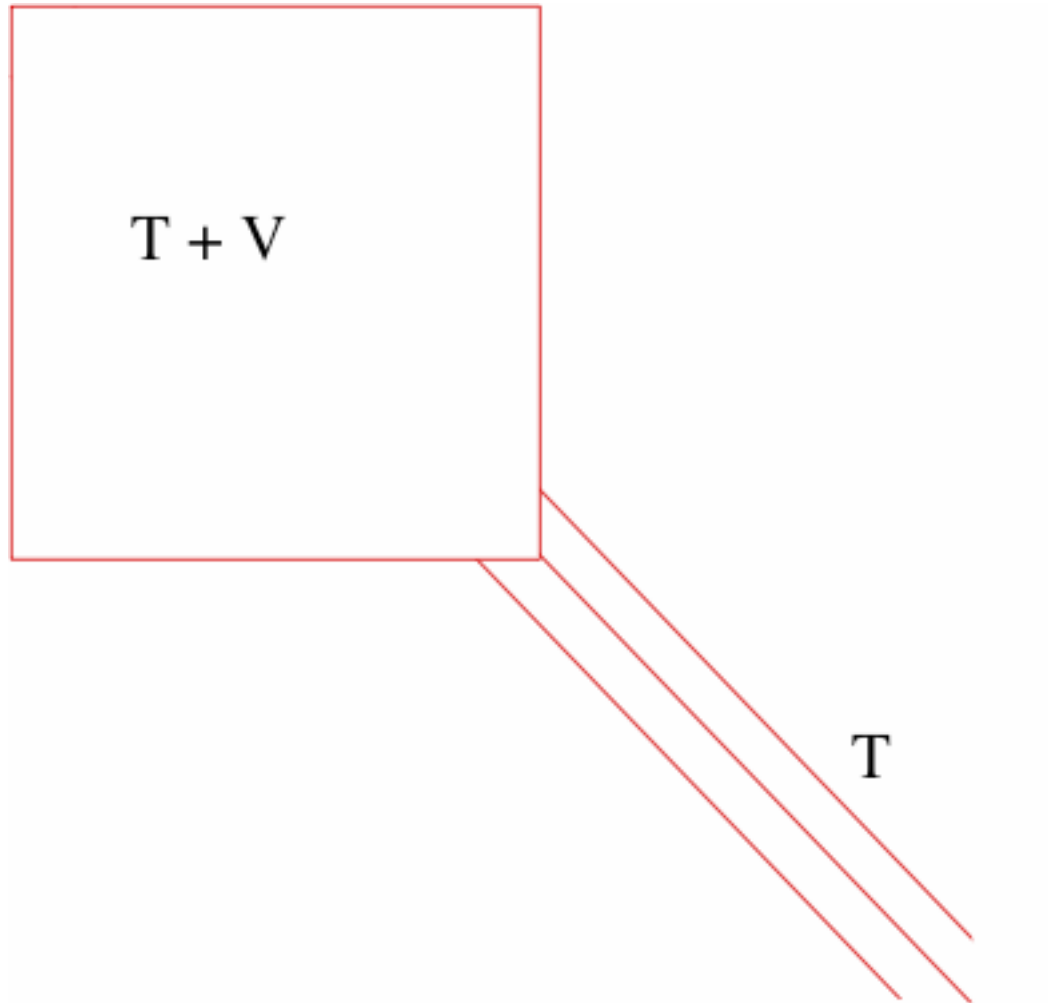
PETs = phase-equivalent transformations

No-core shell model: *ab initio* \Leftrightarrow *ab exitu*
approach

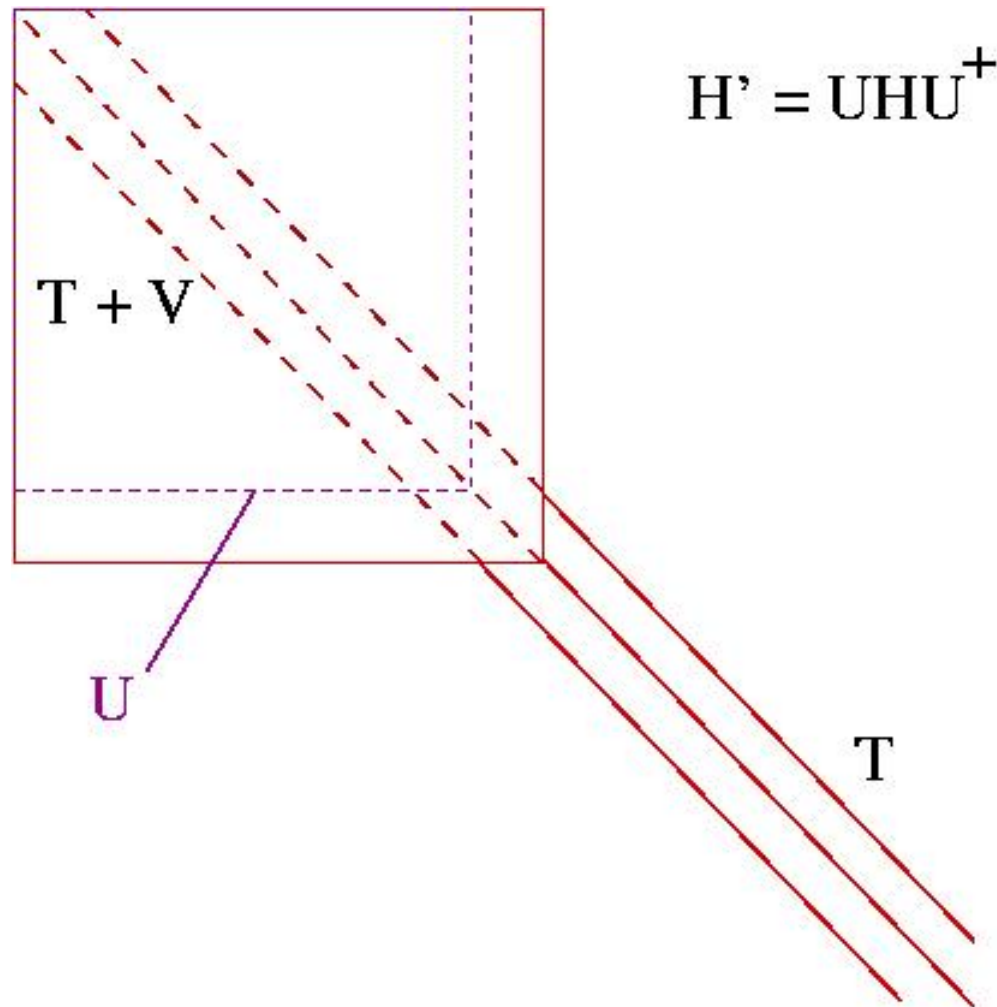
No three-nucleon forces

Results for nuclei with $A \leq 16$

J-matrix formalism: scattering in the oscillator basis



PETs



ab initio \Leftrightarrow *ab exitu*

- JISP16: J-matrix inverse scattering $9h\Omega$ *NN* potential with $h\Omega = 40$ MeV fitted to nuclei up through ^{16}O
- Only simplest PETs generated by 2x2 unitary matrix U are used
- *Ab exitu* approach:
- PETs: *sd* wave - fitting deuteron properties (rms radius and quadrupole moment)
various p and one of d waves - fitting few levels of ^6Li and binding energy of ^{16}O in relatively small model spaces
- All the rest NCSM results (other nuclei, larger model spaces) are *ab initio*

JISP16 properties

- 1992 *np* data base (2514 data): $\chi^2/\text{datum} = 1.03$
- 1999 *np* data base (3058 data): $\chi^2/\text{datum} = 1.05$

Table I: Deuteron properties.

Potential	E_d , MeV	d state probability, %	rms radius, fm	Q , fm ²	As. norm. const. \mathcal{A}_s , fm ^{-1/2}	$\eta = \frac{\mathcal{A}_d}{\mathcal{A}_s}$
JISP16	-2.224575	4.1360	1.9643	0.2886	0.8629	0.0252
Nijmegen-II	-2.224575	5.635	1.968	0.2707	0.8845	0.0252
AV18	-2.224575	5.76	1.967	0.270	0.8850	0.0250
CD-Bonn	-2.224575	4.85	1.966	0.270	0.8846	0.0256
Nature	-2.224575(9)	—	1.971(6)	0.2859(3)	0.8846(9)	0.0256(4)

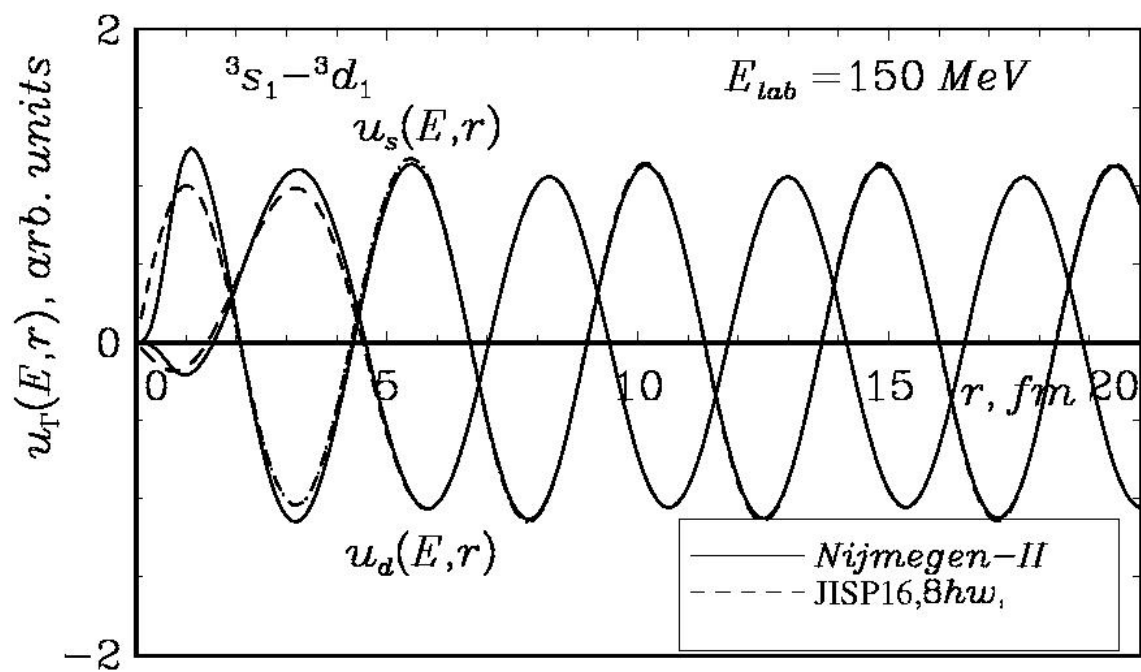
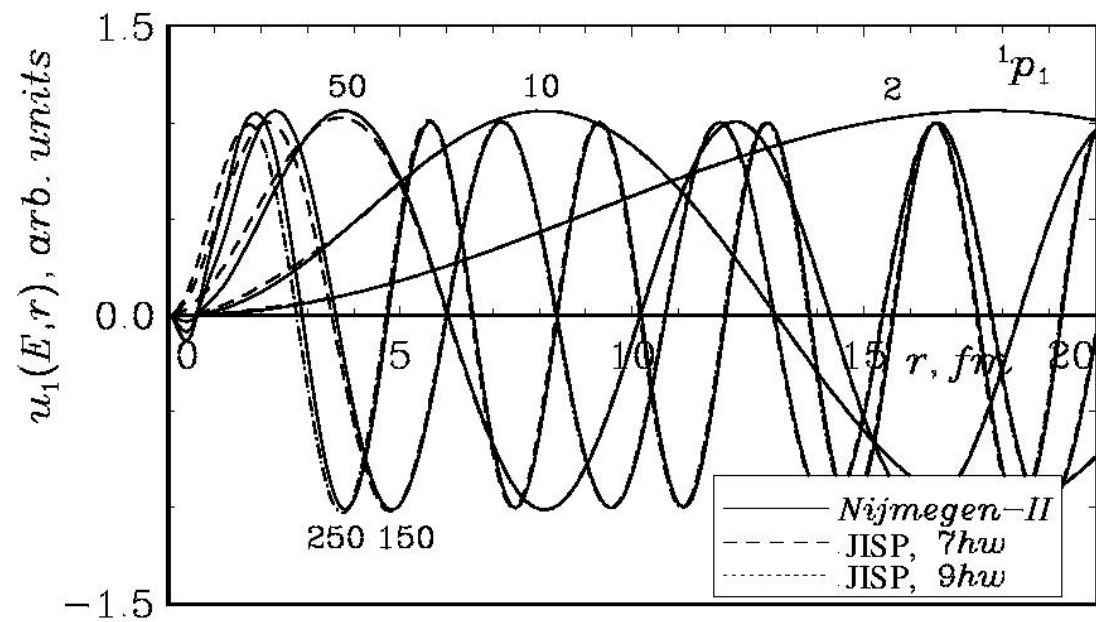


Table II: The binding energies of ^3H , ^3He , ^4He , ^6He and ^6Li nuclei.

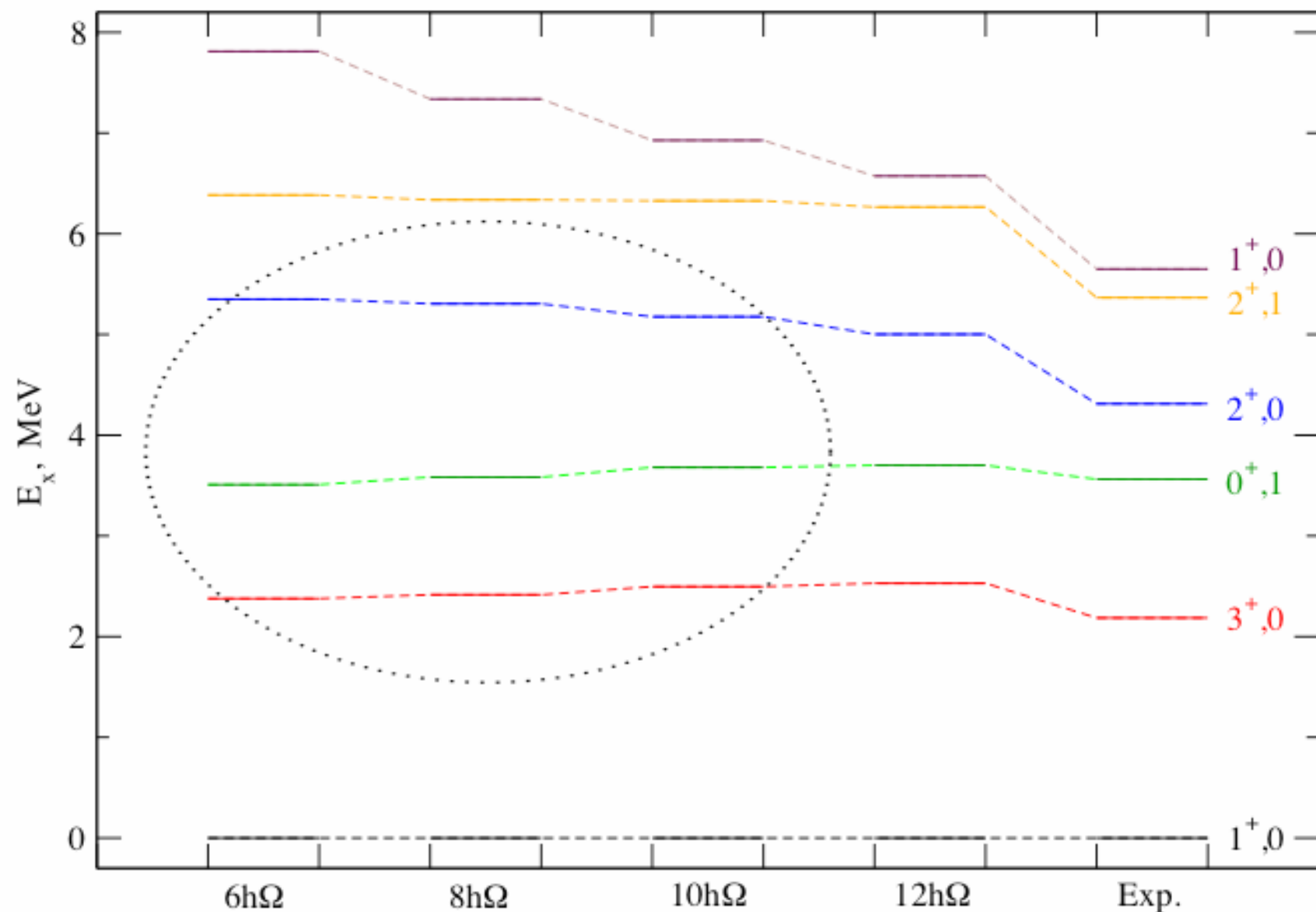
Potential	^3H	^3He	^4He	^6He	^6Li
JISP16, NCSM	8.496(20)	7.797(17)	28.374(57)	28.32(28)	31.00(31)
CD-Bonn+TM, Faddeev [1]	8.480	7.734	29.15		
AV18+TM, Faddeev [1]	8.476	7.756	28.84		
AV18+TM', Faddeev [1]	8.444	7.728	28.36		
NijmI+TM, Faddeev [1]	8.392	7.720	28.60		
NijmII+TM, Faddeev [1]	8.386	7.720	28.54		
AV18+UrbIX, Faddeev [1]	8.478	7.760	28.50		
AV18+UrbIX, GFMC [2]	8.47(1)		28.30(2)	27.64(14)	31.25(11)
AV8'+TM', NCSM [3]				28.189	31.036
Nature	8.48	7.72	28.30	29.269	31.995

[1] A. Nogga *et al*, Phys. Rev. Lett. **85**, 944 (2000).

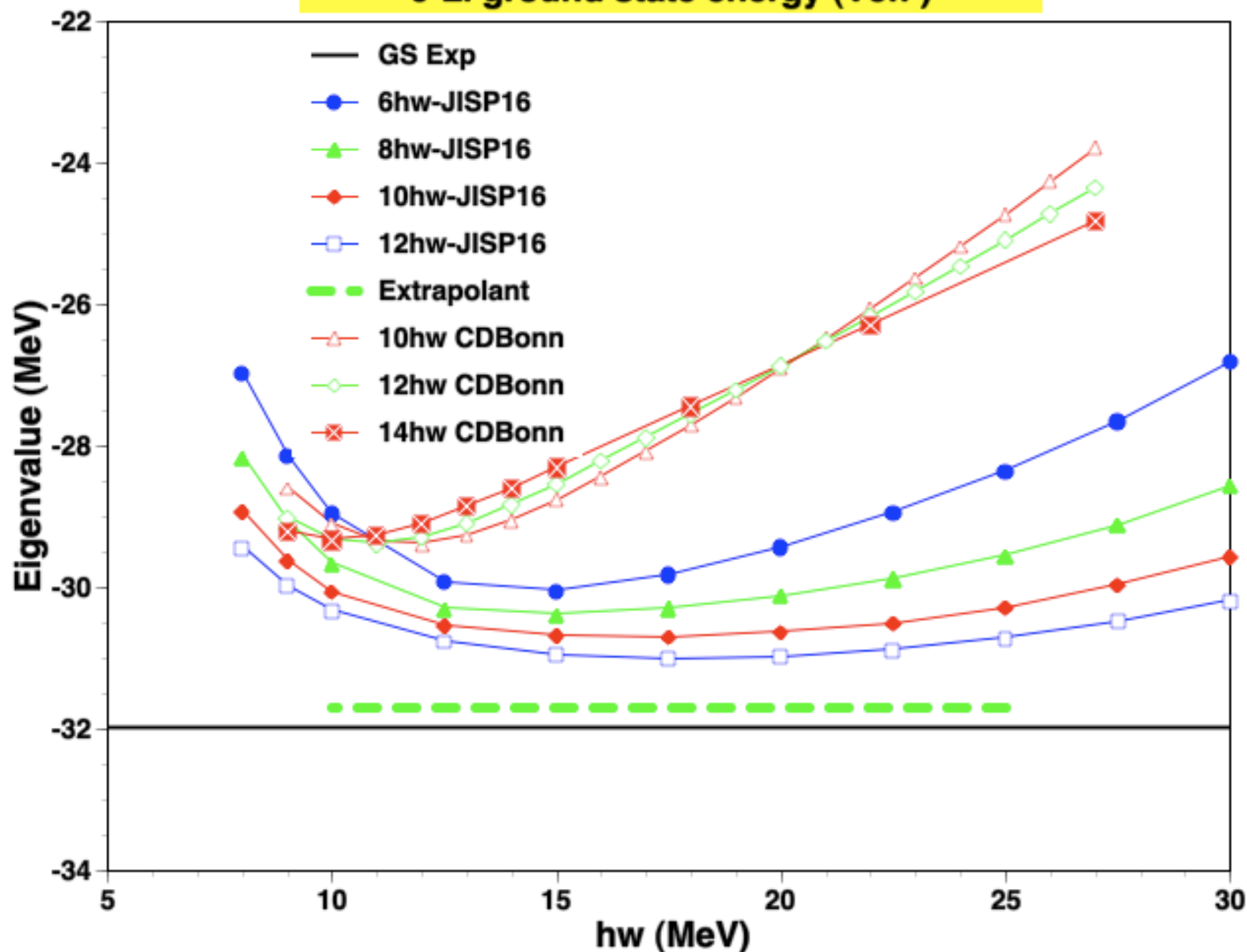
[2] B. S. Pudliner *et al.*, Phys. Rev. **C 56**, 1720 (1997).

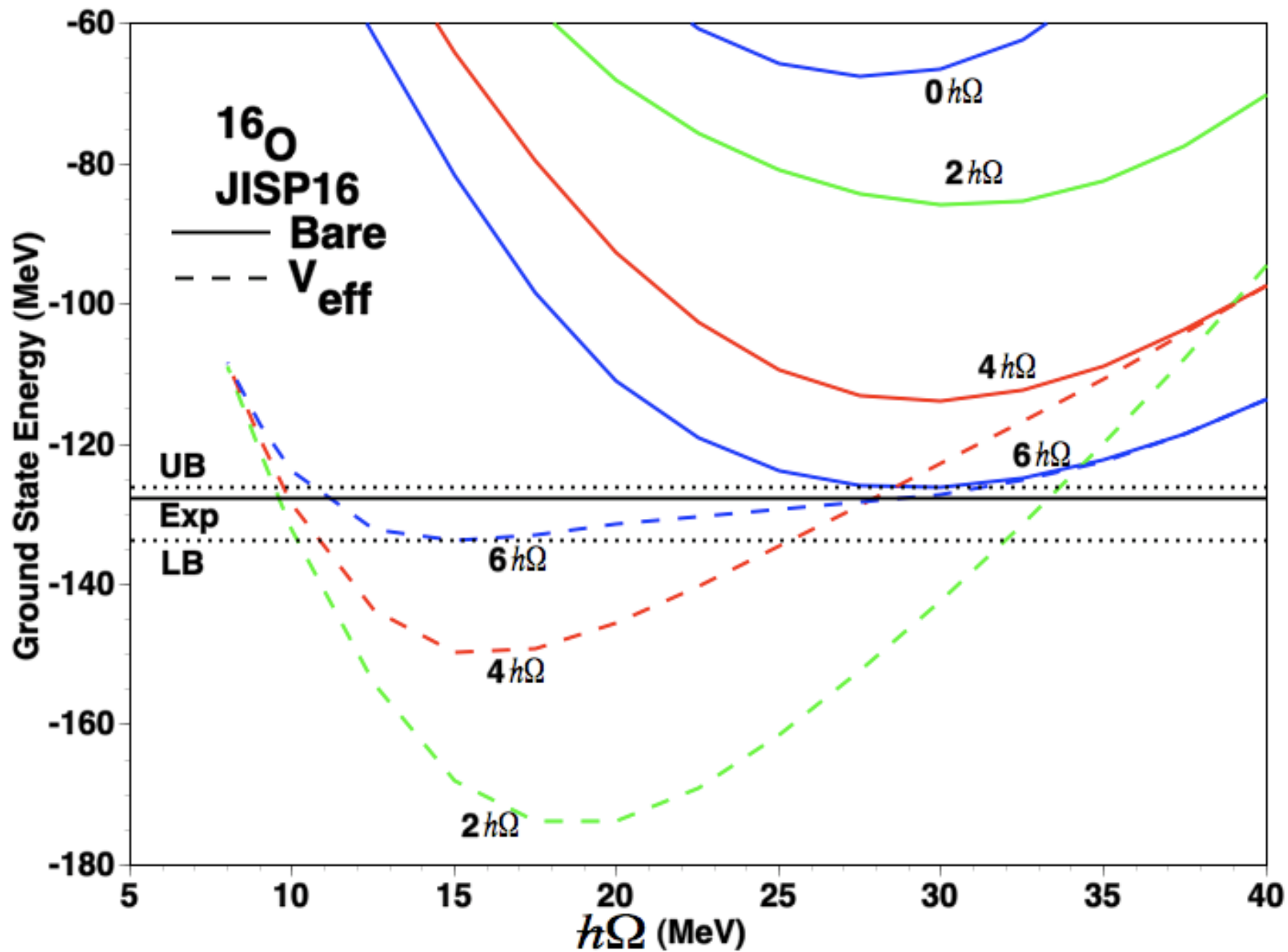
[3] P. Navrátil and E. Ormand, Phys. Rev. **c 68**, 034305 (2003).

${}^6\text{Li}$ spectrum with JISP16 NN interaction, $\hbar\Omega=17.5$ MeV



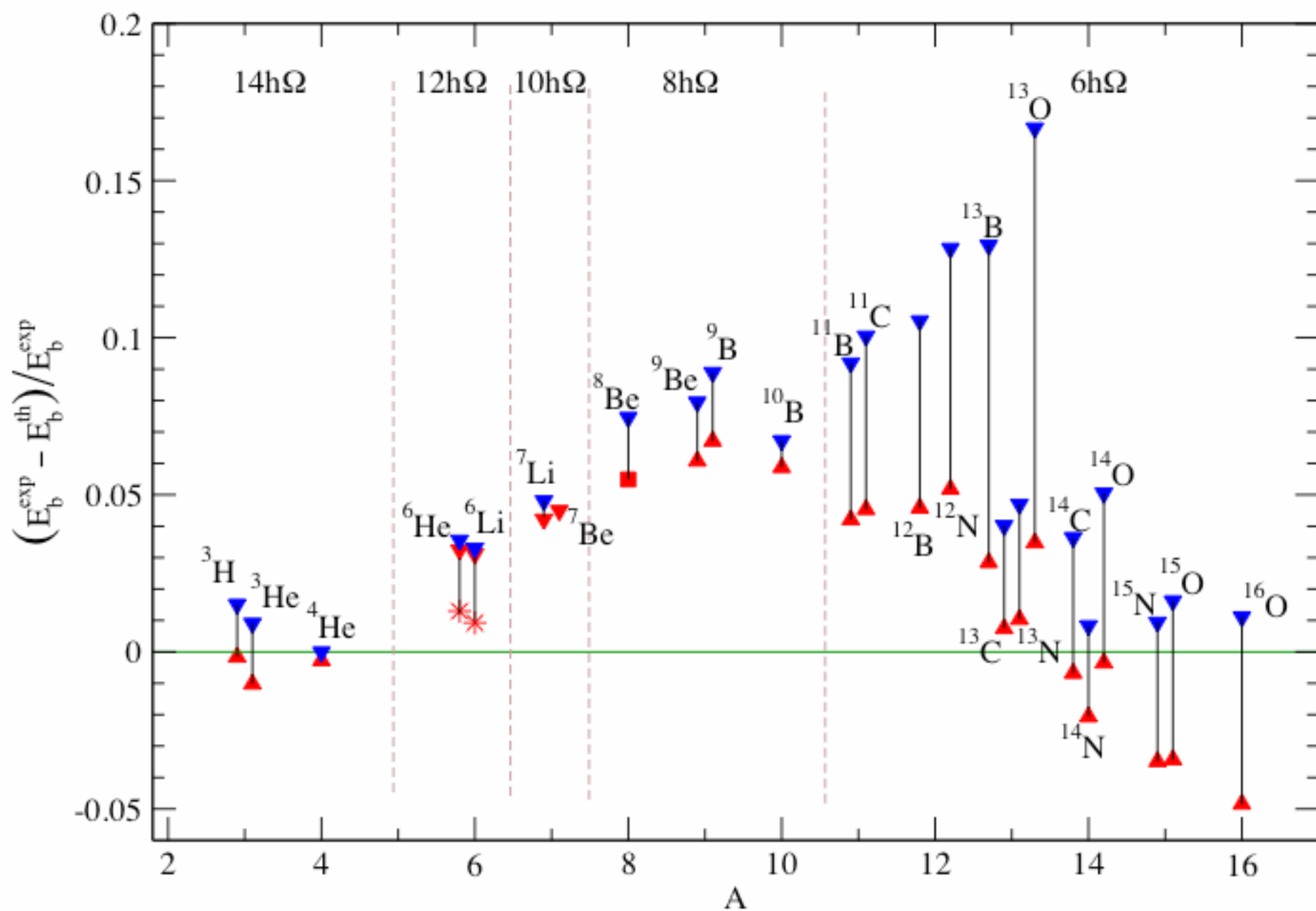
6-Li ground state energy (V_{eff})



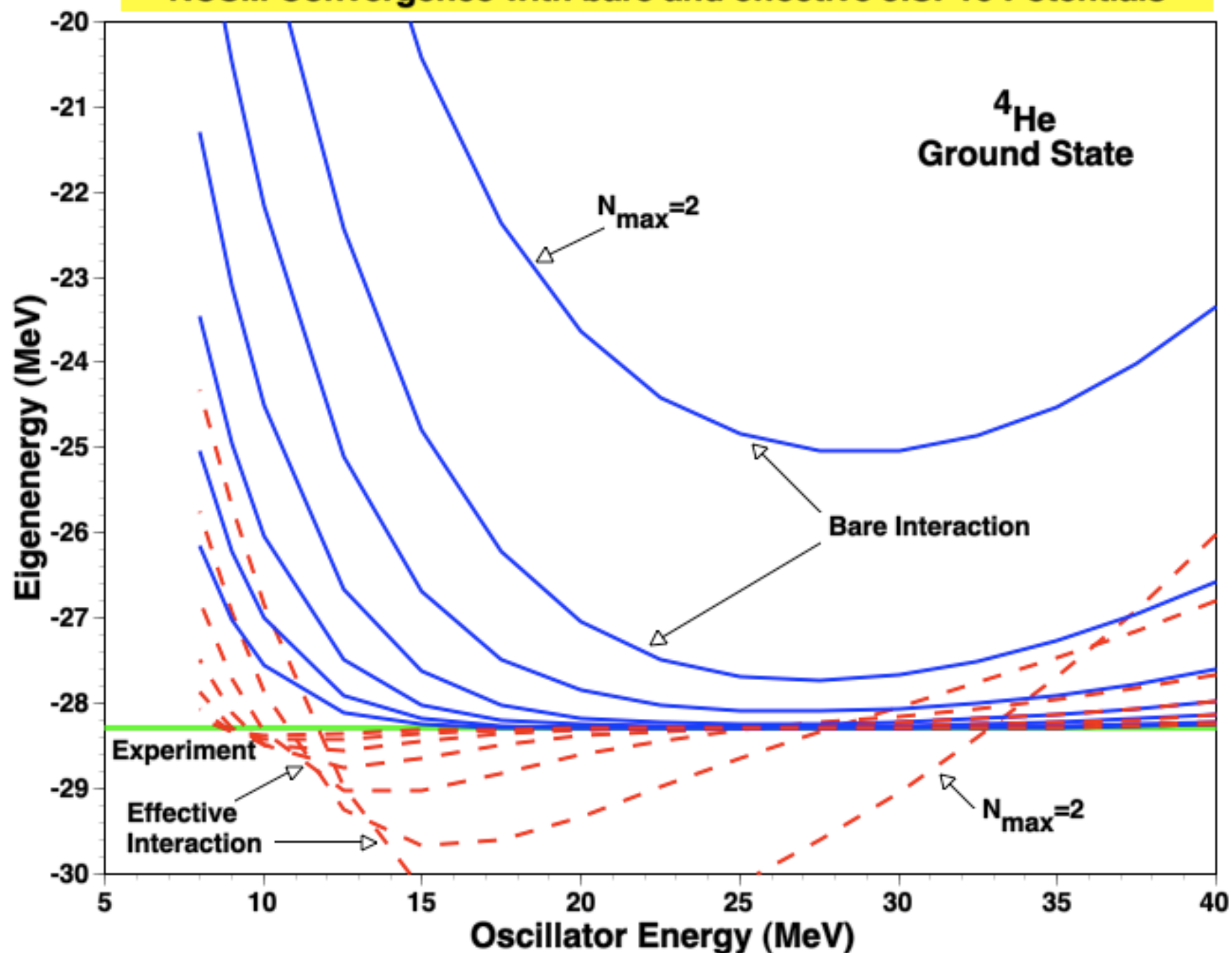


Nucleus	Nature	Bare	Effective	$\hbar\omega$ (MeV)	Model space
^3H	8.482	8.354	8.496(20)	7	$14\hbar\omega$
^3He	7.718	7.648	7.797(17)	7	$14\hbar\omega$
^4He	28.296	28.297	28.374(57)	10	$14\hbar\omega$
^6He	29.269		28.32(28)	17.5	$12\hbar\omega$
^6Li	31.995		31.00(31)	17.5	$12\hbar\omega$
^7Li	39.245		37.59(30)	17.5	$10\hbar\omega$
^7Be	37.600		35.91(29)	17	$10\hbar\omega$
^8Be	56.500		53.40(10)	15	$8\hbar\omega$
^9Be	58.165	53.54	54.63(26)	16	$8\hbar\omega$
^9B	56.314	51.31	52.53(20)	16	$8\hbar\omega$
^{10}Be	64.977	60.55	61.39(20)	19	$8\hbar\omega$
^{10}B	64.751	60.39	60.95(20)	20	$8\hbar\omega$
^{10}C	60.321	55.26	56.36(67)	17	$8\hbar\omega$
^{11}B	76.205	69.2	73.0(31)	17	$6\hbar\omega$
Nucleus	Nature	Bare	Effective	$\hbar\omega$ (MeV)	Model space
^{11}C	73.440	66.1	70.1(32)	17	$6\hbar\omega$
^{12}B	79.575	71.2	75.9(48)	15	$6\hbar\omega$
^{12}C	92.162	87.4	91.0(49)	17.5	$6\hbar\omega$
^{12}N	74.041	64.5	70.2(48)	15	$6\hbar\omega$
^{13}B	84.453	73.5	82.1(67)	15	$6\hbar\omega$
^{13}C	97.108	93.2	96.4(59)	19	$6\hbar\omega$
^{13}N	94.105	89.7	93.1(62)	18	$6\hbar\omega$
^{13}O	75.558	63.0	72.9(62)	14	$6\hbar\omega$
^{14}C	105.285	101.5	106.0(93)	17.5	$6\hbar\omega$
^{14}N	104.659	103.8	106.8(77)	20	$6\hbar\omega$
^{14}O	98.733	93.7	99.1(92)	16	$6\hbar\omega$
^{15}N	115.492	114.4	119.5(126)	16	$6\hbar\omega$
^{15}O	111.956	110.1	115.8(126)	16	$6\hbar\omega$
^{16}O	127.619	126.2	133.8(158)	15	$6\hbar\omega$

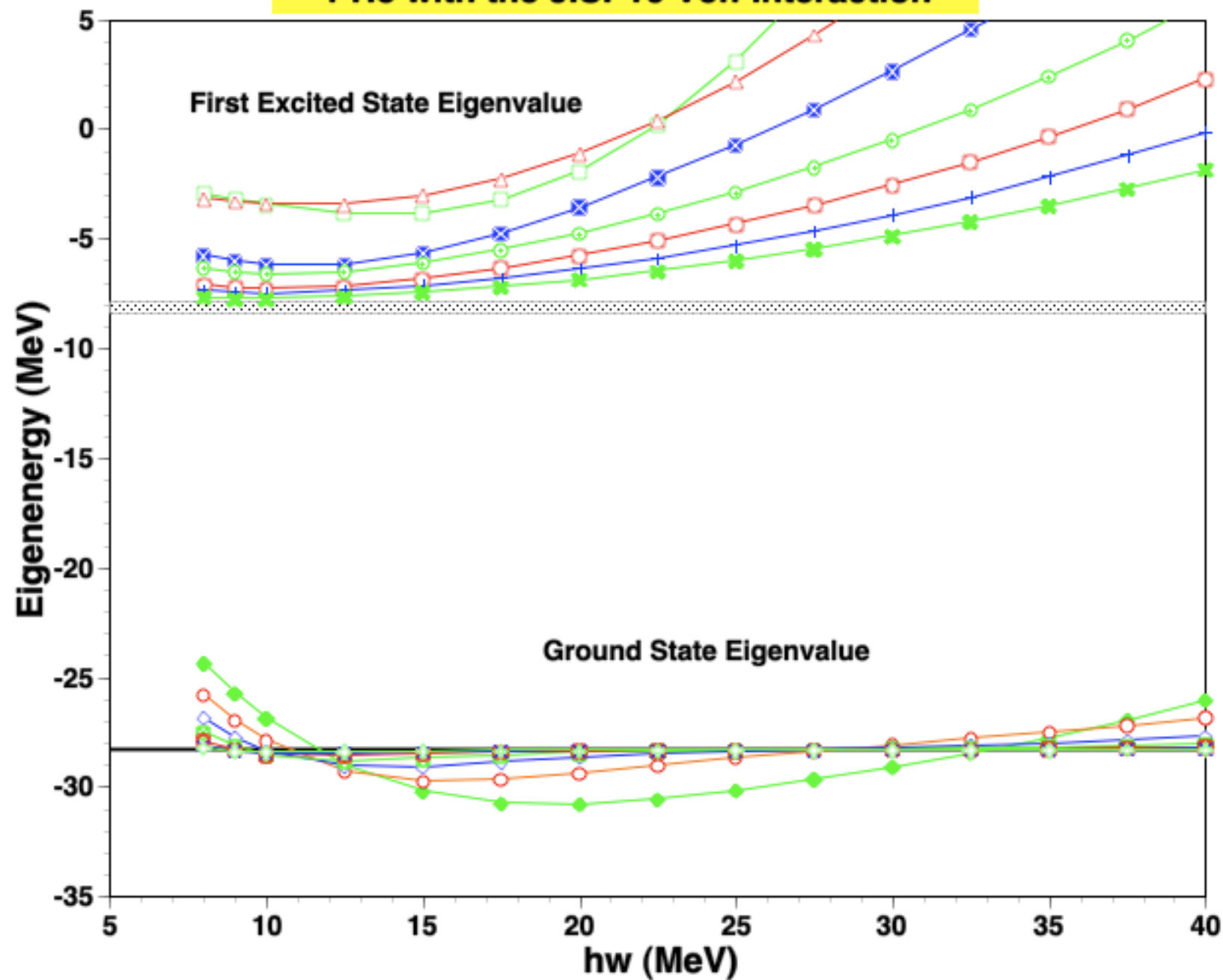
Binding energies



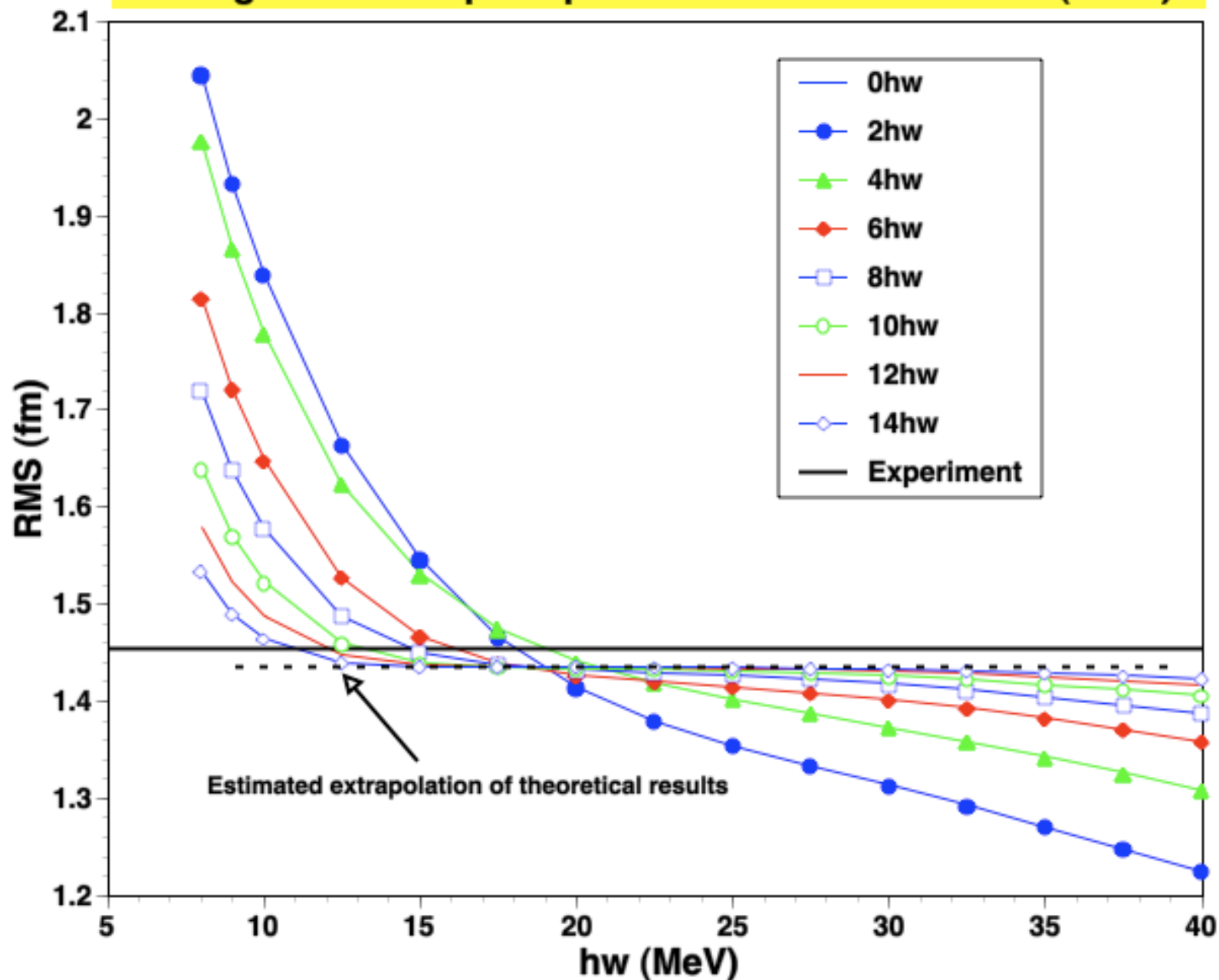
NCSM Convergence with bare and effective JISP16 Potentials



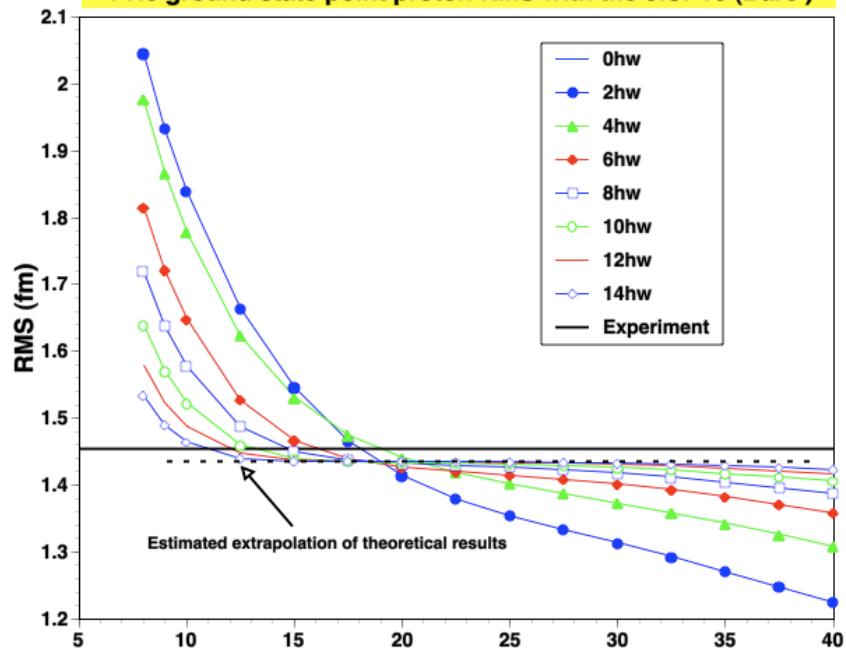
4-He with the JISP16 Veff Interaction



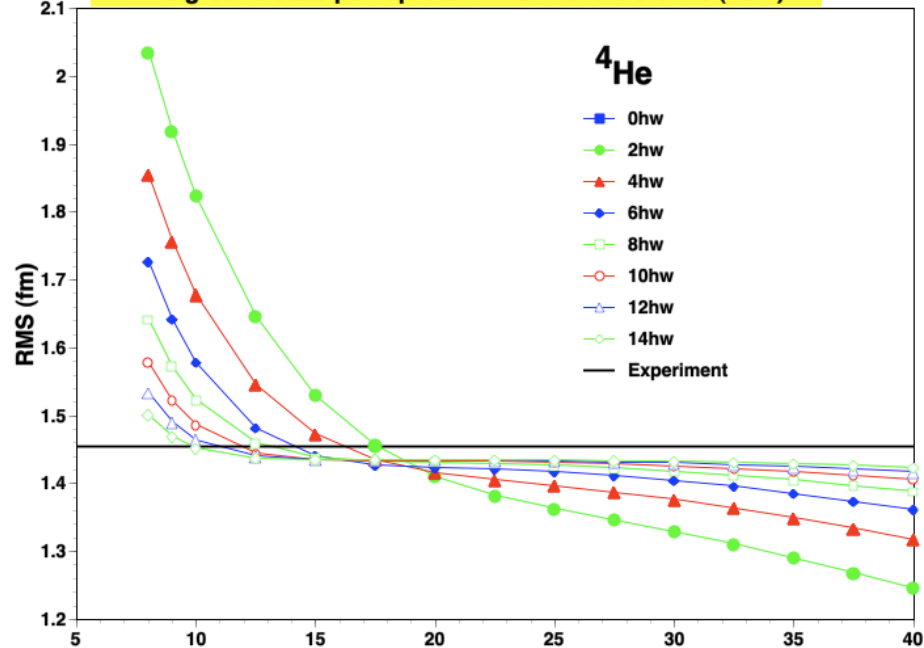
4-He ground state point proton RMS with the JISP16 (Bare)



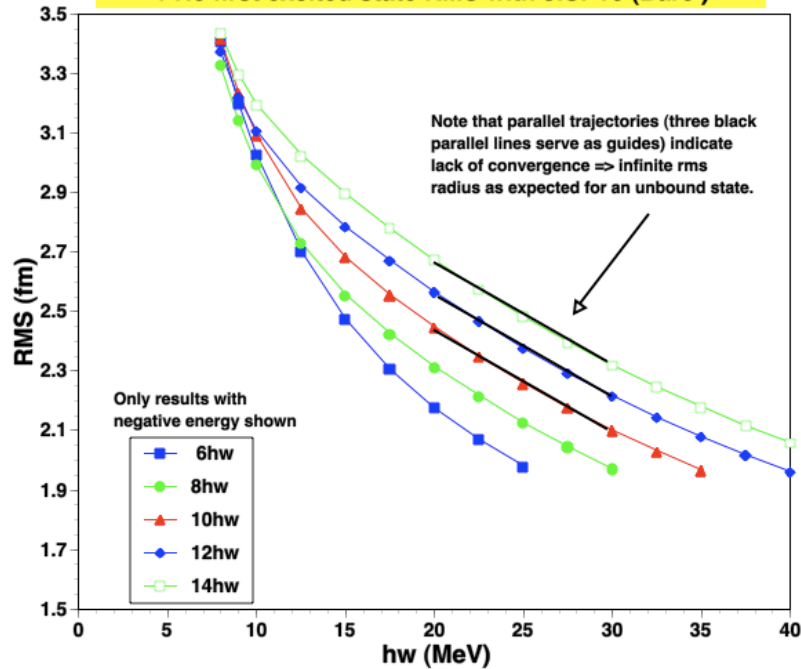
4-He ground state point proton RMS with the JISP16 (Bare)



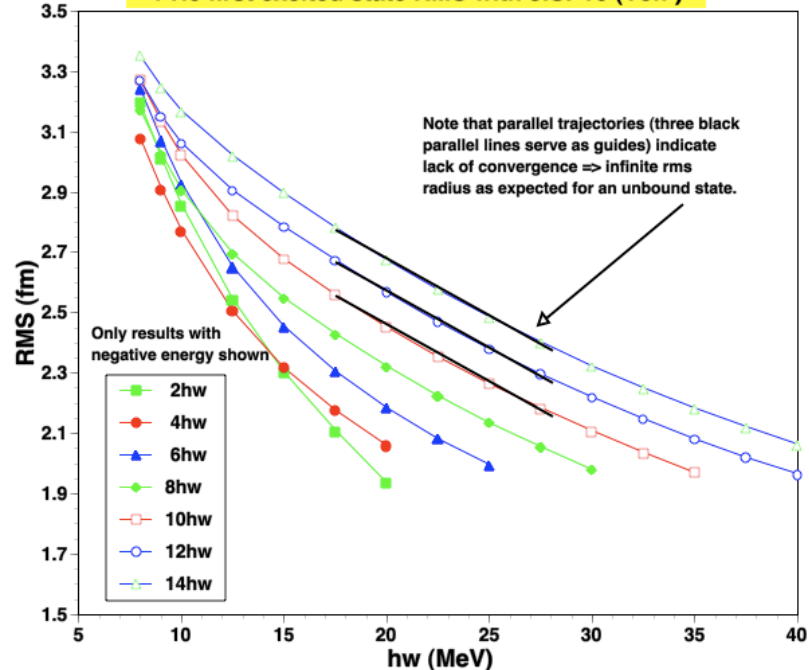
4-He ground state point proton RMS with the JISP16 (Veff)



4-He first excited state RMS with JISP16 (Bare)



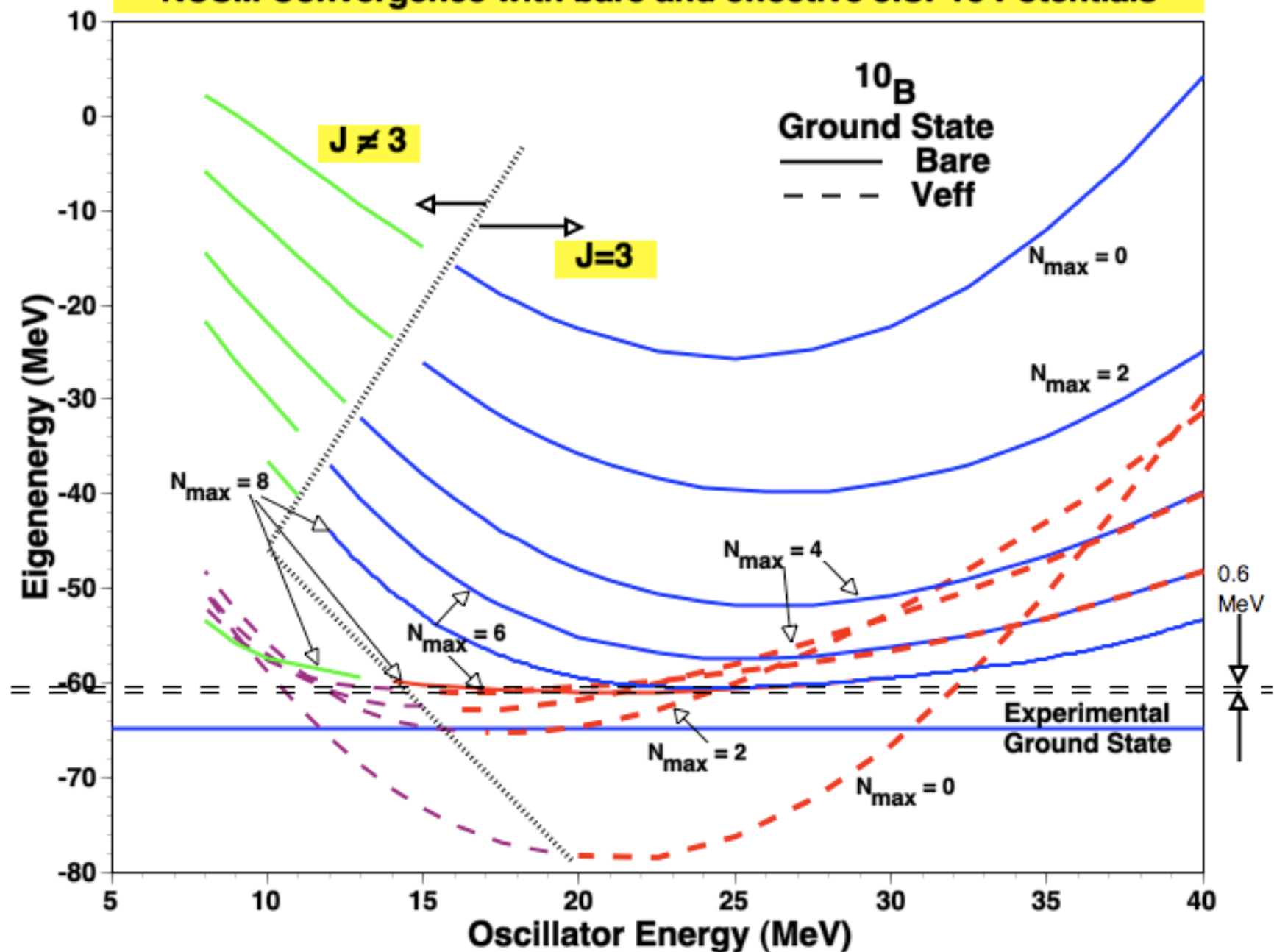
4-He first excited state RMS with JISP16 (Veff)

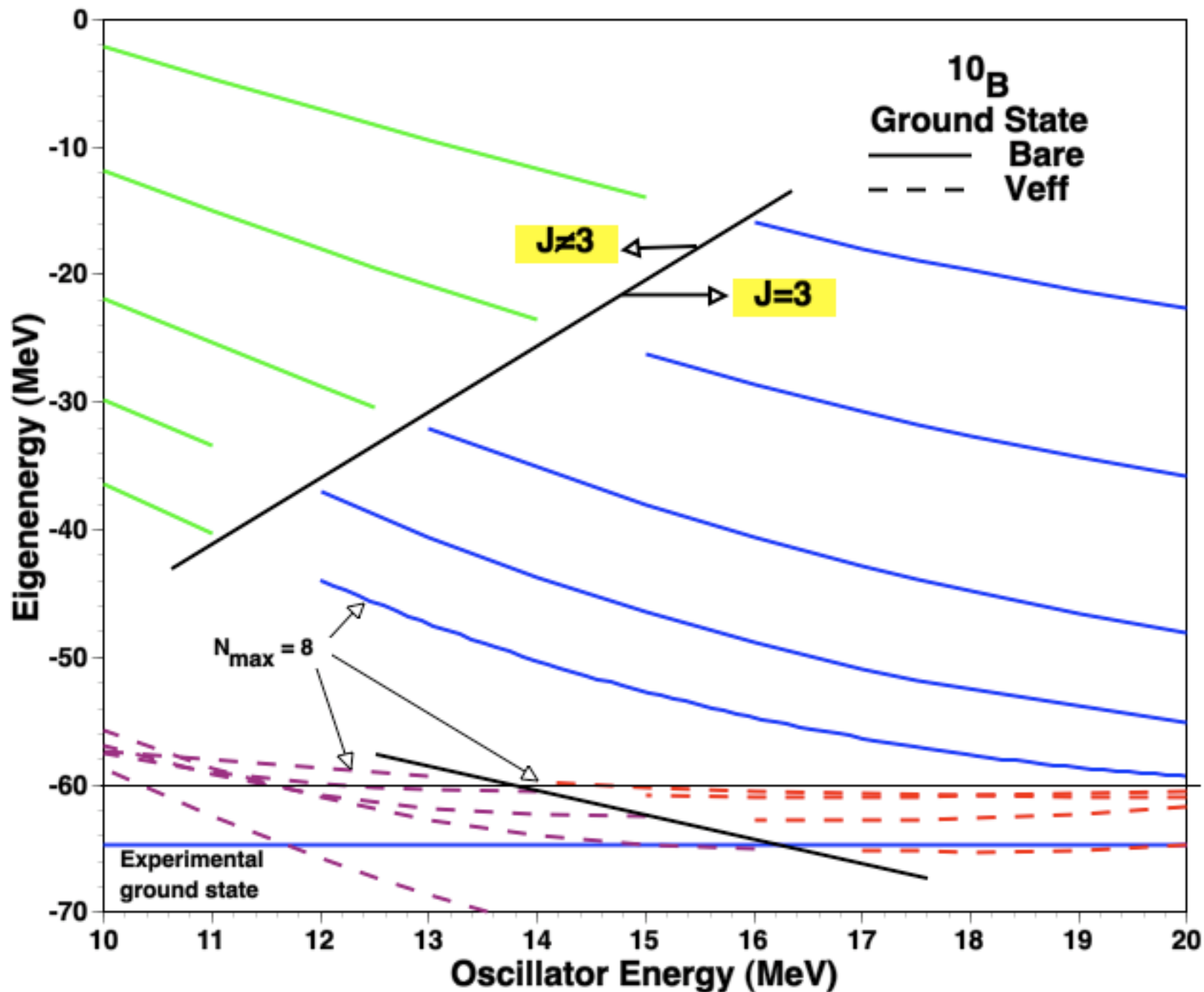


Ground state energy E_{gs} and excitation energies E_x (in MeV), ground state point-proton rms radius r_p (in fm) and quadrupole moment Q (in $e \cdot \text{fm}^2$) of the ${}^6\text{Li}$ nucleus; $\hbar\omega = 17.5$ MeV.

Interaction		JISP6	JISP16	AV8'+TM'	AV18+UIX	AV18+IL2
Method	Nature	NCSM, $10\hbar\omega$ [6]	NCSM, $12\hbar\omega$	NCSM, $6\hbar\omega$ [2]	GFMC [8,15]	GFMC [10,15]
$E_{gs}(1_1^+, 0)$	-31.995	-31.48	-31.00	-31.04	-31.25(8)	-32.0(1)
r_p	2.32(3)	2.083	2.151	2.054	2.46(2)	2.39(1)
Q	-0.082(2)	-0.194	-0.0646	-0.025	-0.33(18)	-0.32(6)
$E_x(3^+, 0)$	2.186	2.102	2.529	2.471	2.8(1)	2.2
$E_x(0^+, 1)$	3.563	3.348	3.701	3.886	3.94(23)	3.4
$E_x(2^+, 0)$	4.312	4.642	5.001	5.010	4.0(1)	4.2
$E_x(2^+, 1)$	5.366	5.820	6.266	6.482		5.5
$E_x(1_2^+, 0)$	5.65	6.86	6.573	7.621	5.1(1)	5.6

NCSM Convergence with bare and effective JISP16 Potentials





Same as in Table 4 but for the ^{10}B nucleus; $\hbar\omega = 15$ MeV.

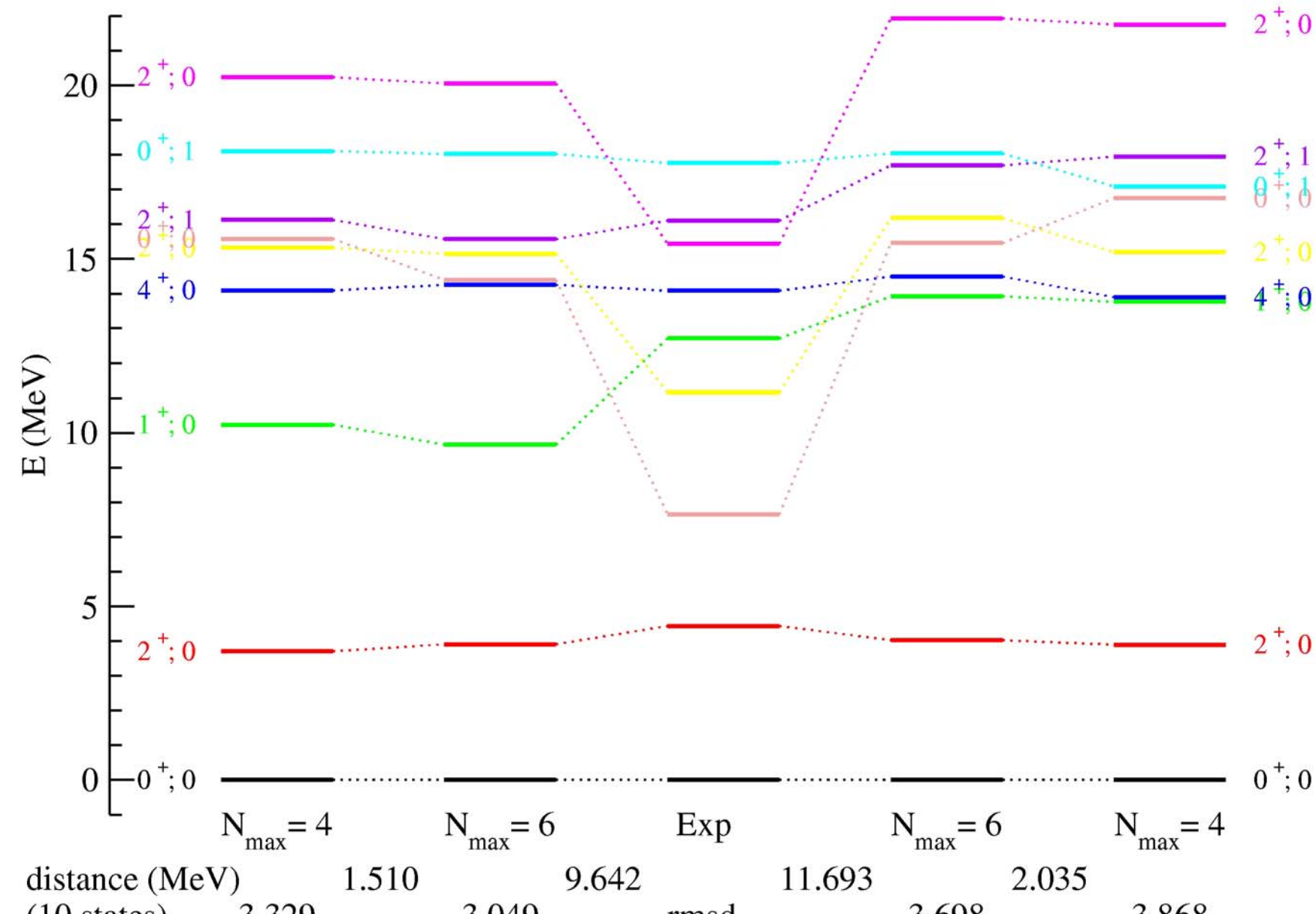
Interaction		JISP16	AV8'+TM'	AV18+IL2
Method	Nature	NCSM, $8\hbar\omega$	NCSM, $4\hbar\omega$ [2]	GFMC [16]
$E_{gs}(3_1^+, 0)$	-64.751	-60.14	-60.57	-65.6(5)
r_p	2.30(12)	2.168	2.168	2.33(1)
Q	+8.472(56)	6.484	+5.682	+9.5(2)
$E_x(1_1^+, 0)$	0.718	0.555	0.340	0.9
$E_x(0^+, 1)$	1.740	1.202	1.259	
$E_x(1_2^+, 0)$	2.154	2.379	1.216	
$E_x(2_1^+, 0)$	3.587	3.721	2.775	3.9
$E_x(3_2^+, 0)$	4.774	6.162	5.971	
$E_x(2_1^+, 1)$	5.164	5.049	5.182	
$E_x(2_2^+, 0)$	5.92	5.548	3.987	
$E_x(4^+, 0)$	6.025	5.775	5.229	5.6
$E_x(2_2^+, 1)$	7.478	7.776	7.491	

^{10}B Basis space	Exp -	JISP16 $8\hbar\Omega$	JISP16 $6\hbar\Omega$	$N3LO + NNN_B$ $6\hbar\Omega$	$N3LO + NNN_A$ $6\hbar\Omega$	$N3LO$ $6\hbar\Omega$
$ E(3^+, 0) $ [MeV]	64.751	60.138	60.800	60.167	64.027	55.613
r_p [fm]	2.30(12)	2.167	2.173	2.248	2.168	2.224
$Q(3_1^+, 0)$ [$e\text{ fm}^2$]	+8.472(56)	6.483	6.239	6.101	6.104	6.665
$\mu(3_1^+, 0)[\mu_N]$	+1.8006	N/A	N/A	N/A	N/A	N/A
$\mu(1_1^+, 0)[\mu_N]$	+0.63(12)	N/A	N/A	N/A	N/A	N/A
$E_x(3_1^+0)$ [MeV]	0.0	0.0	0.0	0.0	0.0	0.0
$E_x(1_1^+0)$ [MeV]	0.718	0.555	0.345	0.728	1.131	-0.877
$E_x(0_1^+1)$ [MeV]	1.740	1.202	1.000	1.662	1.704	1.049
$E_x(1_2^+0)$ [MeV]	2.154	2.379	2.189	2.077	1.529	1.706
$E_x(2_1^+0)$ [MeV]	3.587	3.721	3.323	2.762	3.498	1.797
$E_x(3_2^+0)$ [MeV]	4.774	6.162	5.896	5.093	6.785	4.406
$E_x(2_1^+1)$ [MeV]	5.164	5.049	4.863	4.982	5.518	4.530
$E_x(2_2^+0)$ [MeV]	5.92	5.548	4.992	3.675	4.900	3.732
$E_x(4_1^+0)$ [MeV]	6.025	5.775	5.428	4.398	5.699	4.763
$E_x(2_2^+1)$ [MeV]?	7.478	7.776	7.586	7.998	8.480	5.848
$rms(Exp - Th)$ [MeV]	-	0.539	0.609	0.988	0.875	1.333
$B(E2; 1_1^+0 \rightarrow 3_1^+0)$	4.13(6)	3.317	3.151	0.227	0.356	4.003
$B(E2; 1_2^+0 \rightarrow 3_1^+0)$	1.71(0.26)	0.627	0.540	2.514	2.771	N/A
$B(M1; 2_1^+0 \rightarrow 3_1^+0)$	0.0015(3)	0.0022	0.0022	0.0048	0.0008	N/A
$B(M1; 2_1^+1 \rightarrow 3_1^+0)$	0.041(4)	0.086	0.097	0.002	0.091	N/A
$B(M1; 2_2^+0 \rightarrow 3_1^+0)$	0.050(12)	0.056	0.044	0.031	0.044	N/A
$B(M1; 4_1^+0 \rightarrow 3_1^+0)$	0.043(7)	0.005	0.002	0.002	0.003	N/A
$B(M1; 2_2^+1 \rightarrow 3_1^+0)$	-	3.899	4.113	2.635	3.580	N/A
$B(GT; 2_1^+1 \rightarrow 3_1^+0)$	0.083(3)	0.042	0.041	0.010	0.061	N/A
$B(GT; 2_2^+1 \rightarrow 3_1^+0)$	0.95(13)	1.652	1.745	1.212	1.559	N/A

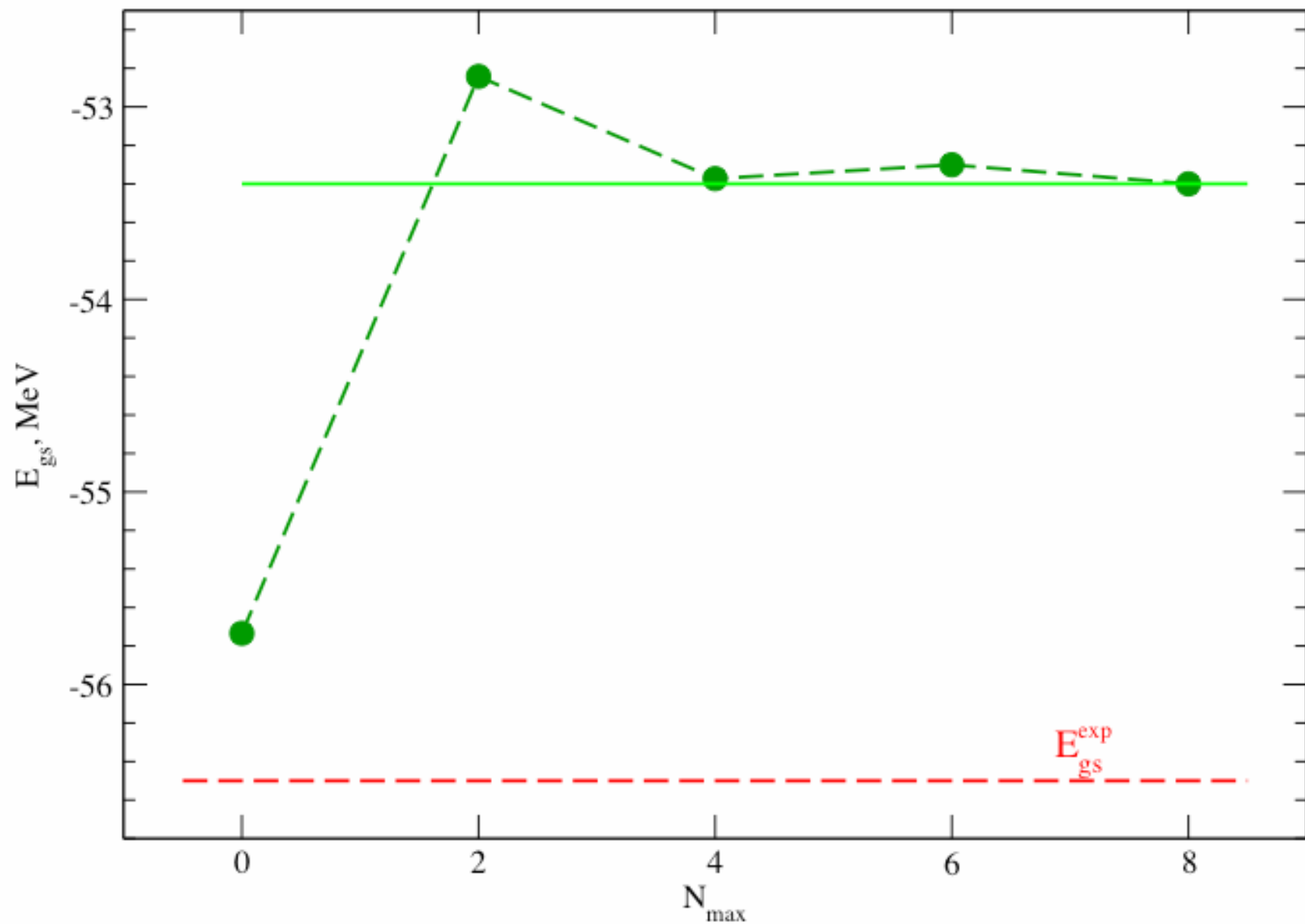
^{12}C Spectral Convergence for $\hbar\Omega = 15$ MeV

N3LO3NFA

JISP16_Veff

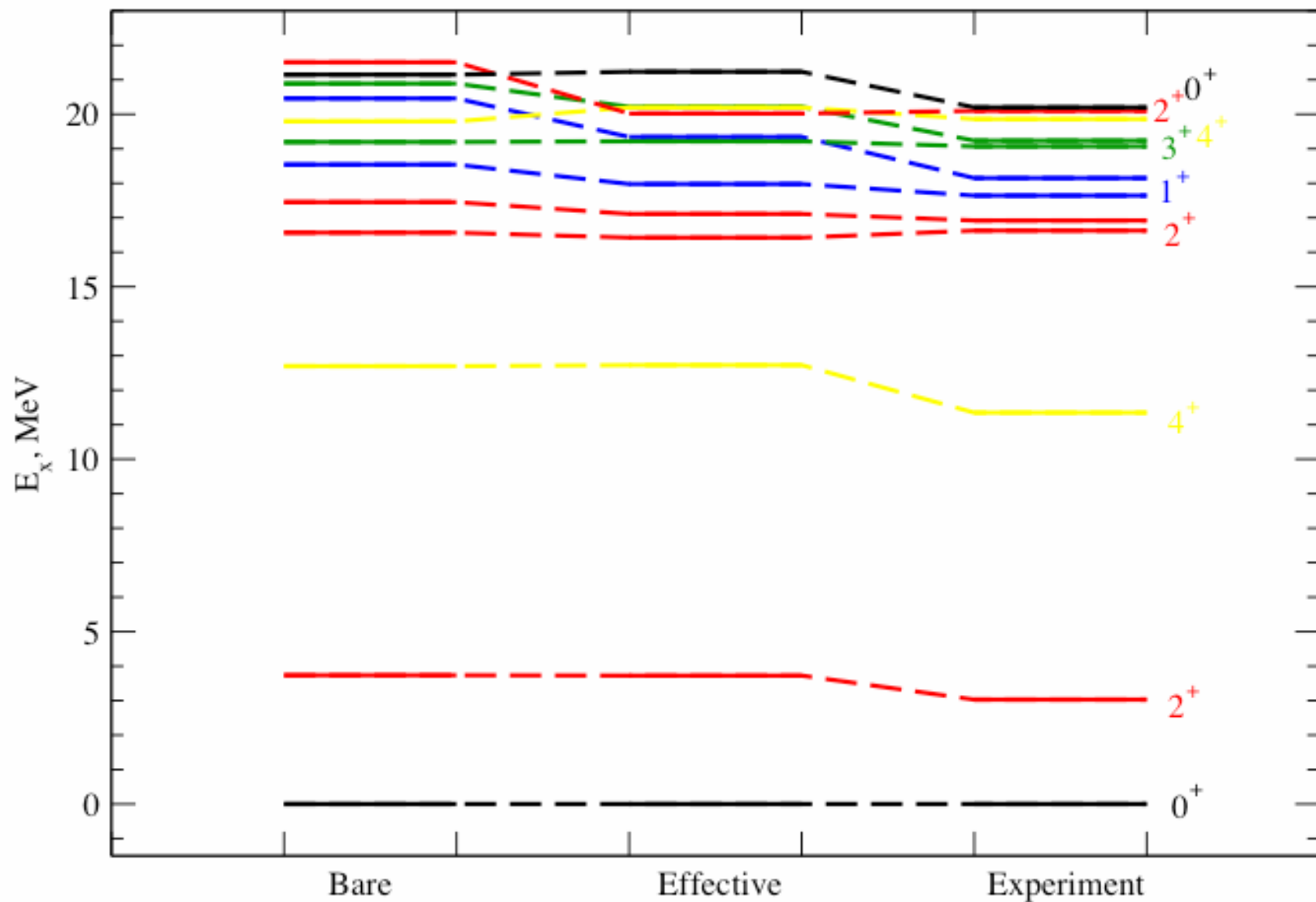


^8Be g.s. convergence with N_{max} $\hbar\Omega$
 $\hbar\Omega = 15 \text{ MeV}$



^8Be spectrum

NCSM, $8h\Omega$ model space



Role of *NNN* force?

- W. Polyzou and W. Glöckle theorem (Few-body Syst. **9**, 97 (1990)):

$$H=T+V_{ij} \Rightarrow H'=T+V'_{ij}+V_{ijk},$$

where V_{ij} and V'_{ij} are phase-equivalent, H and H' are isospectral.

Hope:

$$H'=T+V'_{ij}+V_{ijk} \Rightarrow H=T+V_{ij}$$

with (approximately) isospectral H and H' .

JISP16 seems to be *NN* interaction minimizing *NNN* force.

Without *NNN* force calculations are simpler, calculations are faster, larger model spaces become available.

Conclusions

- JISP16 provides a realistic description of two-body and many-body properties, comparable with modern realistic $NN + NNN$ forces
- Convergence of NCSM calculations with JISP16 is faster, even the bare JISP16 calculation convergence is reasonable, i.e. the results are more reliable. A confidence region of the binding energy predictions can be obtained for many nuclei by comparing the bare and effective interaction results

Plans

- JISP16 improvement by the fit to the same nuclei
- Charge-dependent JISP16
- Extending the calculations to the *sd* shell
- Scattering calculations: NCSM + J-matrix