

Physics 303
Classical Mechanics II

Continuum Mechanics

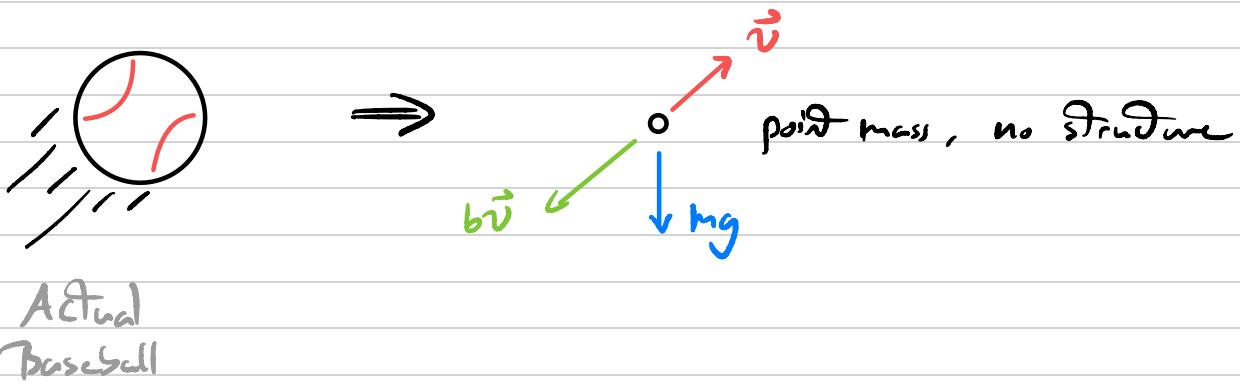
A.W. Jackura — William & Mary

Continuum Mechanics

Classical Mechanics can be generally divided into three main areas, with increasing complexity

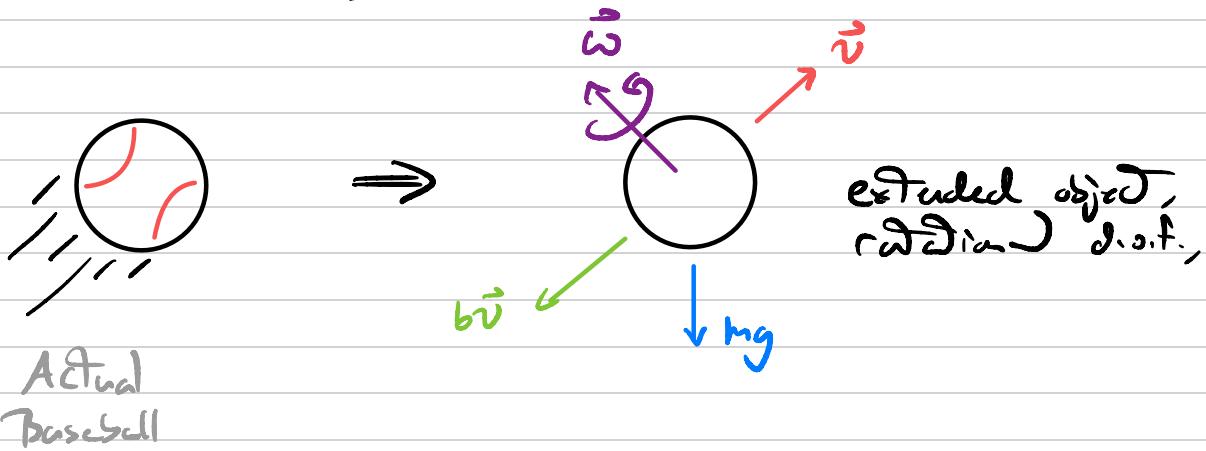
1. Mechanics of point particles

e.g., flight of baseball



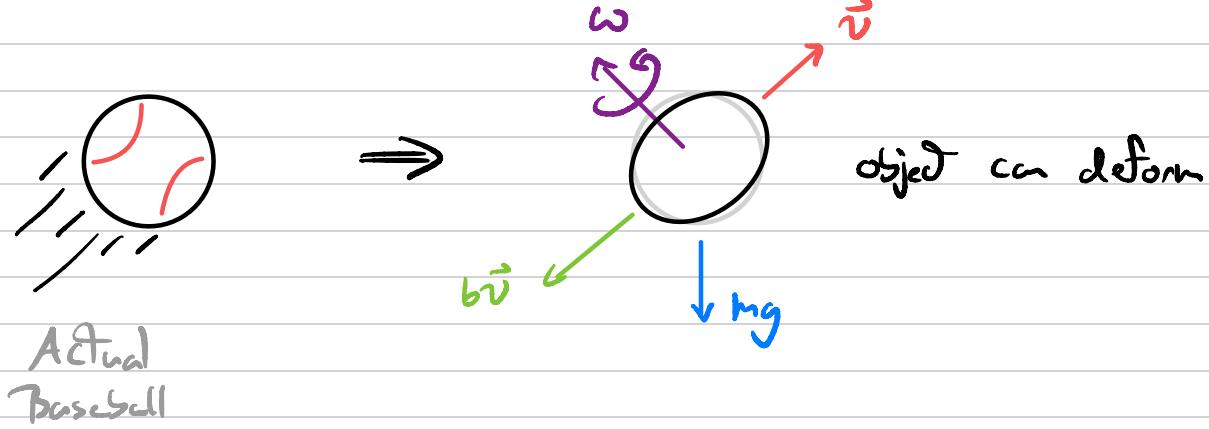
Actual
Baseball

2. Mechanics of rigid bodies



Actual
Baseball

3. Mechanics of continua



Actual
Baseball

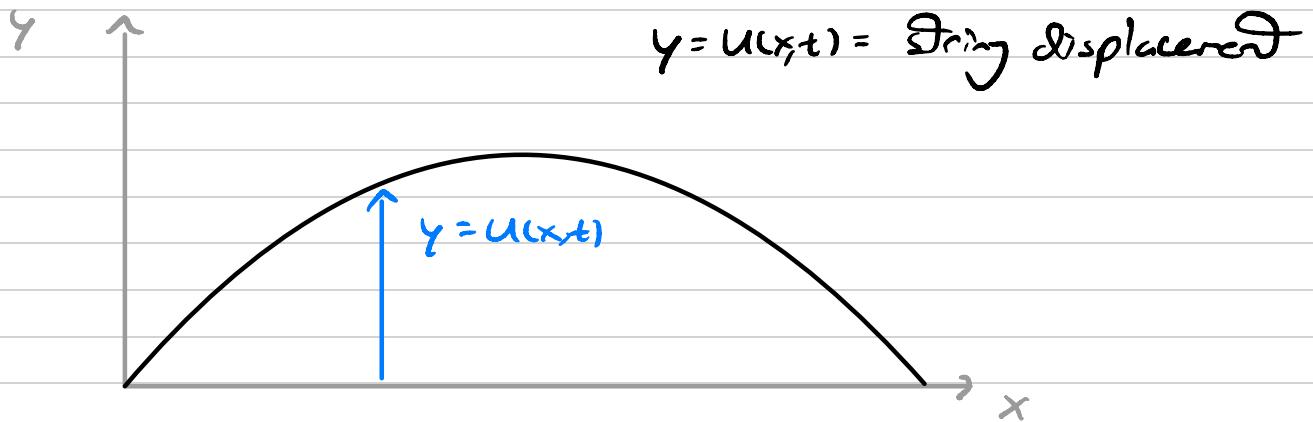
Continuum mechanics can be divided into

- Solid mechanics (our focus)
- Fluid mechanics (see Phys. 302)

In this study, the ordinary differential equations generated from Newton's laws or Euler-Lagrange become partial differential equations.

Wave Motion on a Taut String

As our first example, let's consider the wave motion on an one-dimensional string.



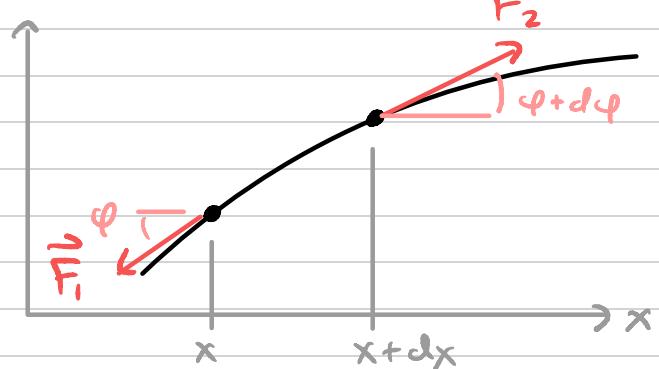
In equilibrium, $y = u(x,t) = 0$

Pick a small segment from

$x \rightarrow x + dx$

Assume small displacements

$$\varphi, \varphi + d\varphi \ll 1$$



The net force in x is

$$\begin{aligned}
 F_x^{\text{net}} &= T \cos(\varphi + d\varphi) - T \cos \varphi \\
 &\approx T \cos \varphi - T d\varphi \sin \varphi - T \cos \varphi \\
 &= -T d\varphi \sin \varphi \approx -T \varphi d\varphi = \mathcal{O}(\varphi^2)
 \end{aligned}$$

$$\begin{aligned}
 F_y^{\text{net}} &= T \sin(\varphi + d\varphi) - T \sin \varphi \\
 &\approx T \sin \varphi + T d\varphi \cos \varphi - T \sin \varphi \\
 &= T d\varphi \cos \varphi \approx T d\varphi
 \end{aligned}$$

Since $\varphi \ll 1 \Rightarrow \sin \varphi \approx \varphi$ & $\cos \varphi \approx 1$

Notice that $\tan \varphi \approx \varphi \approx \frac{\partial y}{\partial x} = \frac{\partial u}{\partial x}$

Therefore, $F_y^{\text{net}} \approx T d\varphi \approx T \frac{\partial \varphi}{\partial x} dx = T \frac{\partial^2 u}{\partial x^2} dx$

$\nabla \vec{F} \Rightarrow \vec{F} = m \vec{a} \Rightarrow F_y^{\text{net}} = dm a_y \leftarrow \text{acceleration in } y \text{ direction}$

↳ mass element of string

$$\begin{aligned}
 \Rightarrow T \frac{\partial^2 u}{\partial x^2} dx &= a_y dm \\
 &= \frac{\partial^2 u}{\partial t^2} (\mu dx) \\
 &\quad \text{↳ linear mass density}
 \end{aligned}$$

$$\Rightarrow \frac{\partial^2 u}{\partial t^2} = \frac{T}{\mu} \frac{\partial^2 u}{\partial x^2}$$

Define $c = \sqrt{\frac{T}{\mu}}$ as the speed of the wave.

Notice, more taut string has higher speed!

So,

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$$

This is the wave equation!

Gauge solution to wave Equation

Introduce two variables $\xi = x - ct$ & $\eta = x + ct$

$$\Rightarrow x = \frac{1}{2}(\xi + \eta), \quad t = \frac{1}{2c}(\eta - \xi)$$

so,

$$\frac{\partial}{\partial \xi} = \frac{\partial x}{\partial \xi} \frac{\partial}{\partial x} + \frac{\partial t}{\partial \xi} \frac{\partial}{\partial t} = \frac{1}{2} \left(\frac{\partial}{\partial x} - \frac{1}{c} \frac{\partial}{\partial t} \right)$$

$$\frac{\partial}{\partial \eta} = \frac{\partial x}{\partial \eta} \frac{\partial}{\partial x} + \frac{\partial t}{\partial \eta} \frac{\partial}{\partial t} = \frac{1}{2} \left(\frac{\partial}{\partial x} + \frac{1}{c} \frac{\partial}{\partial t} \right)$$

$$\Rightarrow \frac{\partial^2}{\partial \xi \partial \eta} = -\frac{1}{4c^2} \left(\frac{\partial^2}{\partial t^2} - c^2 \frac{\partial^2}{\partial x^2} \right)$$

$$\text{So, wave eqn. } \frac{\partial^2 u}{\partial t^2} - c^2 \frac{\partial^2 u}{\partial x^2} = 0$$

becomes

$$\frac{\partial}{\partial \xi} \frac{\partial}{\partial \eta} = 0$$

To solve this eqn., let $h = \frac{\partial u}{\partial \eta}$

$\Rightarrow \frac{\partial h}{\partial \xi} = 0 \Rightarrow h$ is independent of ξ , but
it can depend on η ,
 $\Rightarrow h = h(\eta)$

so, for a given ξ ,

$$\frac{\partial u}{\partial \eta} = h(\eta) \Rightarrow u = \int d\eta h(\eta) + \text{const.}$$

$$\text{Since } h \neq h(\xi) \Rightarrow \int d\eta h(\eta) = g(\eta)$$

Also, the const is for a given $\xi \Rightarrow \text{const} \rightarrow f(\xi)$

so, general solution is

$$u(\xi, \eta) = f(\xi) + g(\eta)$$

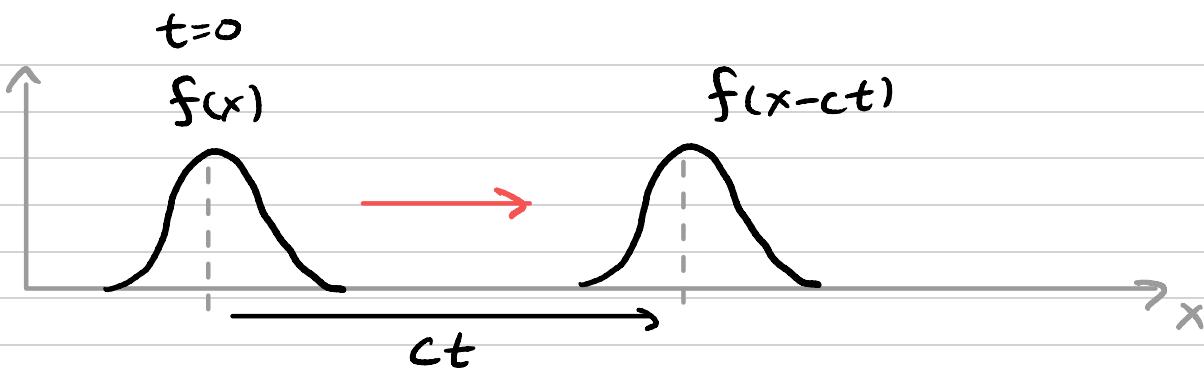
or,

$$u(x, t) = f(x - ct) + g(x + ct)$$



wave moving
to right

wave moving
to left



Consider solution $u(x,t) = f(x-ct)$

At $t=0$, $f(x)$ has a maximum $\nexists x=0$

At t , $f(x-ct)$ has a maximum $\nexists x-ct=0$

$$\Rightarrow x=ct$$

A special example is the standing wave

consider $f(x-ct) = A \sin(kx - \omega t)$

where $\omega = kc$ & A, k are arbitrary constants.

A is called the amplitude,

k is wave number $\Rightarrow \lambda = \frac{2\pi}{k}$ is wave length

ω is angular frequency $\Rightarrow T = \frac{2\pi}{\omega}$ is period

If $g(x+ct) = A \sin(kx + \omega t)$

then,

$$u(x,t) = A \sin(kx - \omega t) + A \sin(kx + \omega t)$$

$$= 2A \sin(kx) \cos(\omega t)$$

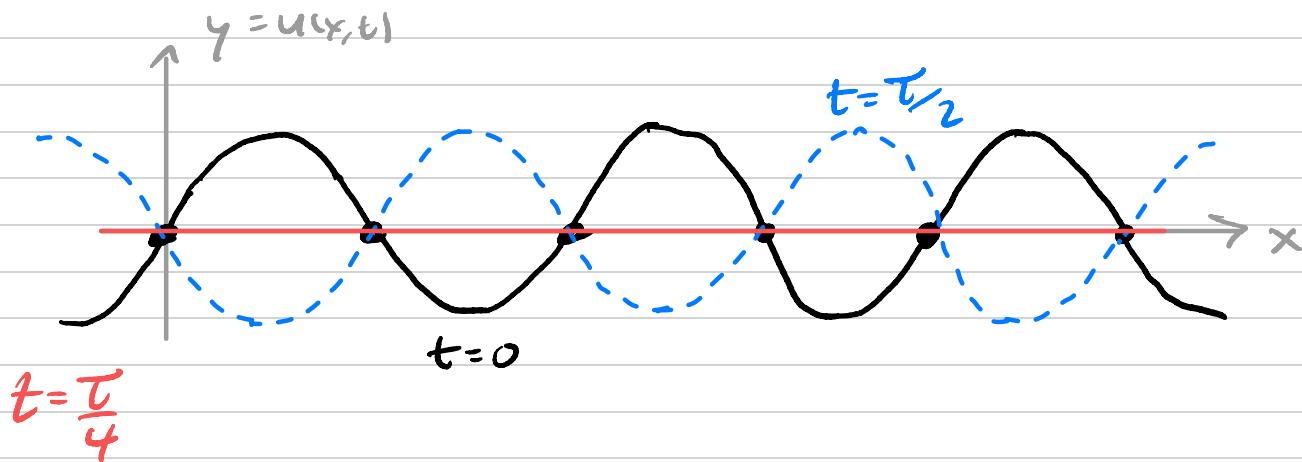
Notice that the wave does not travel, it merely oscillates up and down.

$$u(x,t) = [2A \sin(kx)] \cos(\omega t)$$

amplitude \hookrightarrow oscillatory time dependence

Notice that the zeros of the amplitude are fixed

$$\Rightarrow kx = n\pi \Rightarrow x = \frac{n\pi}{k} \text{ are } \underline{\text{nodes}}$$



We will see that these standing waves are the continuum analogue of normal modes in coupled oscillators.

Boundary Conditions on Finite String

The wave equation requires initial and/or boundary conditions to completely specify a solution.

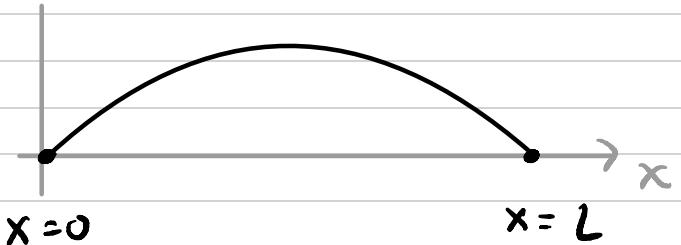
Consider a wave on a finite string subject to Dirichlet BCs.

$$u(0, t) = u(L, t) = 0$$

for all t .

We want a solution to

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$$



$$\text{Try } u(x, t) = X(x) \cos(\omega t - \delta)$$

↑ "Separation of Variables"

$$\Rightarrow -\omega^2 X(x) \cos(\omega t - \delta) = c^2 \frac{d^2 X}{dx^2} \cos(\omega t - \delta)$$

$$\Rightarrow \frac{d^2 X}{dx^2} = -k^2 X \quad \text{w/ } k = \frac{\omega}{c}$$

$$\text{Solution is } X(x) = A \sin kx + B \cos kx$$

Now, this solution is subject to Dirichlet BCs

$$\Rightarrow X(0) = X(L) = 0$$

$$X(0) = 0 \Rightarrow 0 = A \cdot 0 + B(1) \Rightarrow B = 0$$

$$X(L) = 0 \Rightarrow 0 = A \sin(kL)$$

$A=0$ is trivial solution, so find $kL = n\pi$, $n \in \mathbb{N}$

$$\Rightarrow k_n = \frac{n\pi}{L}, n = 1, 2, \dots$$

The wave vector is quantized!

$$\Rightarrow \omega_n = \frac{n\pi}{L} c$$

So,

$$u(x,t) = \sum_n A_n \sin(k_n x) \cos(\omega_n t - \delta_n)$$

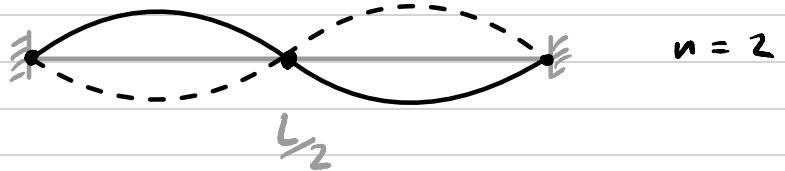
↳ an infinite # of standing waves

⇒ Normal modes

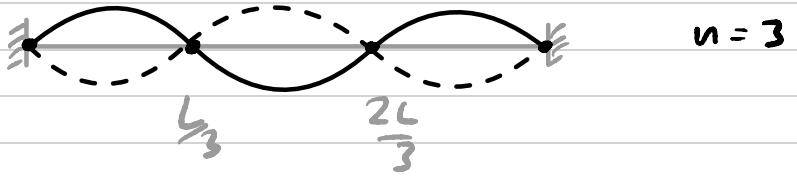
$$\sin\left(\frac{\pi}{L}x\right)$$



$$\sin\left(\frac{2\pi}{L}x\right)$$



$$\sin\left(\frac{3\pi}{L}x\right)$$



:

Coefficients A_n & δ_n are fixed by initial configuration

$$\text{Let } A_n \cos(\omega_n t - \delta_n) = B_n \cos \omega_n t + C_n \sin \omega_n t$$

$$\Rightarrow u(x,t) = \sum_n \sin \omega_n x (B_n \cos \omega_n t + C_n \sin \omega_n t)$$

At $t=0$, we are given $u(x,0)$ & $\dot{u}(x,0)$

$$\text{we find } u(x,0) = \sum_n B_n \sin \omega_n x$$

$$\dot{u}(x,0) = \sum_n C_n \omega_n \sin \omega_n x$$

To get B_n & C_n , use Fourier's trick from Fourier Series

$$u(x,0) = \sum_n B_n \sin\left(\frac{n\pi x}{L}\right)$$

$$\Rightarrow \int_0^L dx u(x,0) \sin\left(\frac{m\pi x}{L}\right) = \sum_n B_n \int_0^L dx \sin\left(\frac{m\pi x}{L}\right) \sin\left(\frac{n\pi x}{L}\right)$$

Can show that

$$\int_0^L dx \sin\left(\frac{m\pi x}{L}\right) \sin\left(\frac{n\pi x}{L}\right) = \frac{L}{2} \delta_{mn}$$

$$\Rightarrow B_n = \frac{2}{L} \int_0^L dx u(x,0) \sin\left(\frac{n\pi x}{L}\right)$$

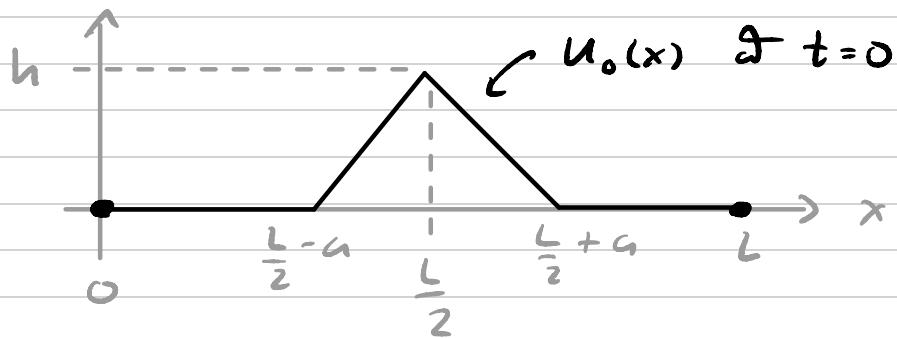
$$\text{Similarly, for } C_n \text{ find } C_n = \frac{2}{L} \frac{1}{W_n} \int_0^L dx i(x,0) \sin\left(\frac{n\pi x}{L}\right)$$

Let's look at a particular example.

Example: Triangular wave on finite string

for example,

$$h = L = a = 1$$



So,

$$u_0(x) = u(x, 0) = \begin{cases} 0 & 0 \leq x < L/2 - a \\ \frac{h}{a}(x - (\frac{L}{2} - a)) & L/2 - a \leq x < L/2 \\ \frac{h}{a}(\frac{L}{2} + a - x) & L/2 \leq x < L/2 + a \\ 0 & L/2 + a \leq x \leq L \end{cases}$$

in our example,

$$u_0(x) = u(x, 0) = \begin{cases} 0 & 0 \leq x < 3 \\ (x - 3) & 3 \leq x < 4 \\ (5 - x) & 4 \leq x < 5 \\ 0 & 5 \leq x < 8 \end{cases}$$

We are also given $\dot{u}(x, 0) = 0 \quad \forall x \in [0, L]$

Notice that the wave is symmetric about $x = \frac{L}{2} (= 4)$

Define $x' = x - \frac{L}{2}$, so $u_0(x') = u_0(-x')$

$$\Rightarrow B_1 = \frac{2}{L} \int_0^L dx u_0(x) \sin\left(\frac{n\pi}{L}x\right)$$

check sign

$$= \frac{2}{L} \int_{-\frac{L}{2}}^{\frac{L}{2}} dx' u_0(x') \sin\left(\frac{n\pi}{L}x' - \frac{n\pi}{2}\right)$$

Since $u_o(x')$ is even

$$\Rightarrow B_{2n} = 0 \quad \forall n \in \mathbb{N}$$

Why?

$$\sin(\alpha - n\pi) = (-1)^n \sin(\alpha)$$

$$B_{2n} = \frac{2}{L} \int_{-L/2}^{L/2} dx' u_o(x') \sin\left(\frac{2\pi n}{L} x' - n\pi\right)$$

$$= (-1)^n \frac{2}{L} \int_{-L/2}^{L/2} dx' u_o(x') \sin\left(\frac{2\pi n}{L} x'\right) = 0$$

↑ even ↑ even ↑ odd

So, look at odd modes

$$B_{2n+1} = \frac{2}{L} \int_{-L/2}^{L/2} dx' u_o(x') \sin\left(\frac{(2n+1)\pi}{L} x' - \frac{(2n+1)\pi}{2}\right)$$

$$= (-1)^n \frac{2}{L} \int_{-L/2}^{L/2} dx' u_o(x') \sin\left(\frac{(2n+1)\pi}{L} x' - \frac{\pi}{2}\right)$$

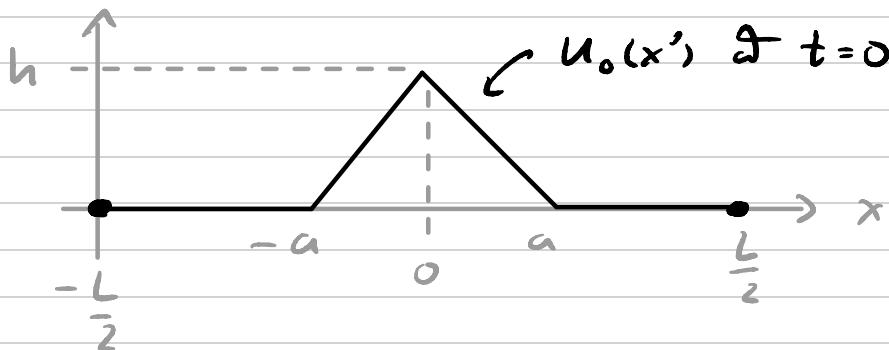
$$\hookrightarrow \sin(\alpha - \pi/2) = -\cos\alpha$$

$$= (-1)^{n+1} \frac{2}{L} \int_{-L/2}^{L/2} dx' u_o(x') \cos\left(\frac{(2n+1)\pi}{L} x'\right)$$

↑ even ↑ even ↑ even

$$\Rightarrow B_{2n+1} \neq 0$$

To exclude further, $u_0(x')$ is given by $x = x' + \frac{L}{2}$



$$u_0(x') = u_0(x', 0) = \begin{cases} 0 & 0 \leq x' < -a \\ h(x'+a)/a & -a \leq x' < 0 \\ h(-x'+a)/a & 0 \leq x' < a \\ 0 & a \leq x' < L \end{cases}$$

So,

$$\mathcal{B}_{2n+1} = (-1)^{n+1} \frac{2}{L} \int_{-L/2}^{L/2} dx' u_0(x') \cos\left(\frac{(2n+1)\pi}{L} x'\right)$$

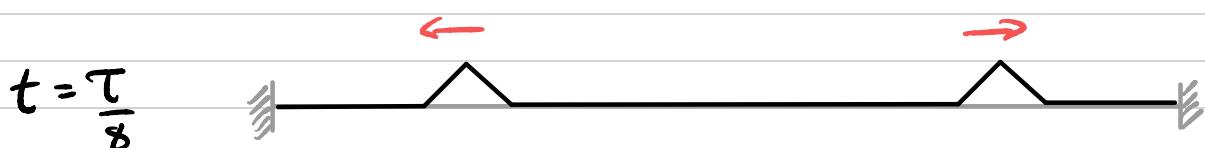
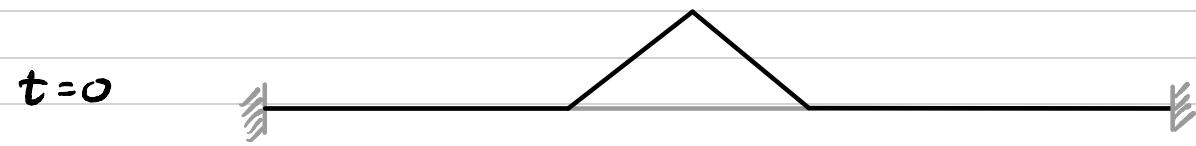
$$= (-1)^{n+1} \frac{4}{L} \cdot \frac{h}{a} \int_0^a dx' (a-x') \cos\left(\frac{(2n+1)\pi}{L} x'\right)$$

$$= (-1)^{n+1} \frac{4h}{La} \cdot \frac{L^2}{(2n+1)^2 \pi^2} \left[1 - \cos\left((2n+1)\pi \frac{a}{L}\right) \right]$$

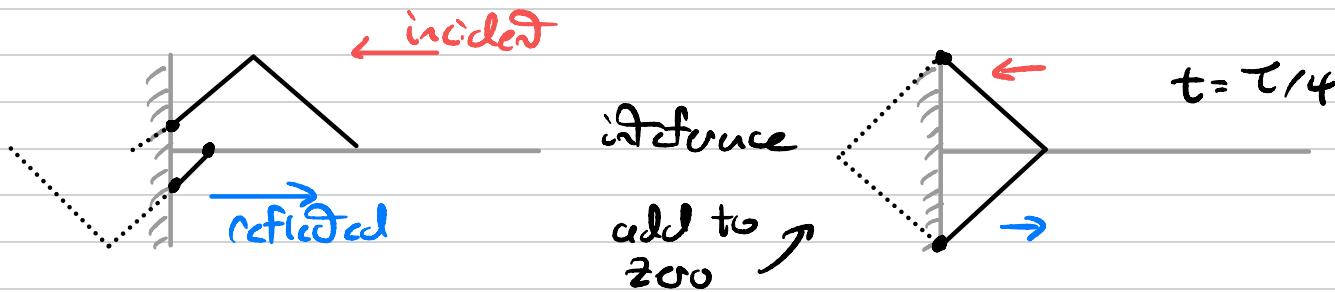
$$= \frac{4hL}{a} \frac{(-1)^{n+1}}{(2n+1)^2 \pi^2} \left[1 - \cos\left((2n+1)\pi \frac{a}{L}\right) \right]$$

* $(-1)^n$ from wrong sign as $\cos \alpha = \sin(\alpha + \frac{\pi}{2})$

Let's look at the time evolution for the fundamental frequency $\omega_1 = \frac{\pi}{L} c \Rightarrow \tau = \frac{2\pi}{\omega_1}$



At the boundary



Wave Equation in 3D

We can generalize the 1D wave eqn, $\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$, to 3D in the expected way.

Let $\rho(\vec{r}, t) = \rho(x, y, z, t)$ denote some disturbance of a 3D system (e.g., pressure in sound wave through air), then the wave equation is

$$\frac{\partial^2 \rho}{\partial t^2} = c^2 \left(\frac{\partial^2 \rho}{\partial x^2} + \frac{\partial^2 \rho}{\partial y^2} + \frac{\partial^2 \rho}{\partial z^2} \right)$$

c = speed of wave

For sound in air, $c^2 = \frac{B M}{\rho_0}$ → Bulk modulus (see later)

↳ equilibrium density

Define the vector operator

$$\vec{\nabla} = \left(\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

so $\vec{\nabla}^2$

$$\vec{\nabla}^2 = \vec{\nabla} \cdot \vec{\nabla} = \left(\frac{\partial}{\partial x} \right)^2 + \left(\frac{\partial}{\partial y} \right)^2 + \left(\frac{\partial}{\partial z} \right)^2$$

Laplacian

Therefore, the 3D wave eqn. is

$$\boxed{\frac{\partial^2 p}{\partial t^2} = c^2 \vec{\nabla}^2 p}$$

Plane Wave Solution

If the wave front is propagating in the \hat{n} direction,
then

$$p(\vec{r}, t) = f(\hat{n} \cdot \vec{r} - ct)$$

Verify: $\vec{\nabla} p = \frac{\partial f}{\partial (\hat{n} \cdot \vec{r})} \hat{n} = -\frac{1}{c} \frac{\partial f}{\partial t} \hat{n}$

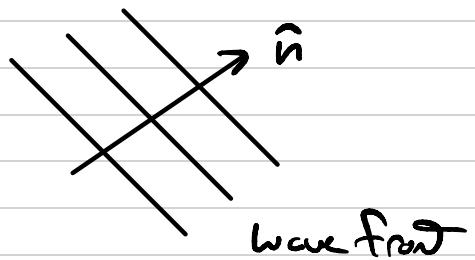
$$\Rightarrow \vec{\nabla}^2 p = -\frac{1}{c} \frac{\partial}{\partial t} \left(\frac{\partial f}{\partial t} \right) \hat{n} \quad \xrightarrow{\text{Using trick as before}} \frac{\partial f(x-y)}{\partial x} = -\frac{\partial f(x-y)}{\partial y}$$

$$= -\frac{1}{c} \frac{\partial}{\partial t} \left(-\frac{1}{c} \frac{\partial f}{\partial t} \hat{n} \right) \cdot \hat{n}$$

$$= \frac{1}{c^2} \frac{\partial^2 f}{\partial t^2}$$

If wave is in free space, no boundary conditions,
then

$$p(\vec{r}, t) \propto \cos(k(\hat{n} \cdot \vec{r} - ct))$$



Spherical Wave Solutions

Another important example is spherical wave solutions, i.e., a disturbance traveling radially outward.

$$\rho = \rho(r, t)$$

Can show $\vec{\nabla}^2 \rho = \frac{1}{r} \frac{\partial^2 (r\rho)}{\partial r^2}$

So, wave eqn. is

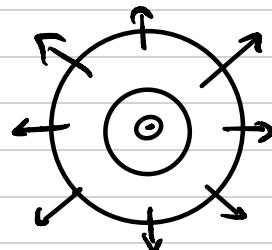
$$\frac{\partial^2 \rho}{\partial t^2} = c^2 \frac{1}{r} \frac{\partial^2 (r\rho)}{\partial r^2}$$

Can see that $\frac{\partial^2 (r\rho)}{\partial t^2} = c^2 \frac{\partial^2 (r\rho)}{\partial r^2}$

has a solution $r\rho(r, t) = f(r - ct) + g(r + ct)$

If disturbance is radially outward, $g=0$

$$\Rightarrow \rho(r, t) = \frac{1}{r} f(r - ct)$$

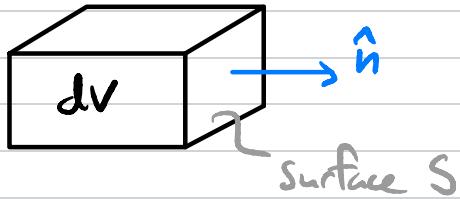
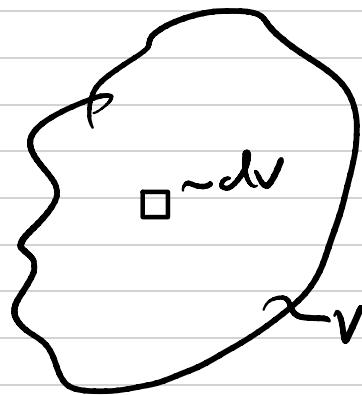


Volume & Surface Forces

We now aim to construct the Equations of motion
for a continuous 3D system.

→ Apply NII to small
mass element

Consider small element



Surface area is specified
by normal vector \hat{n}
pointed "outward"

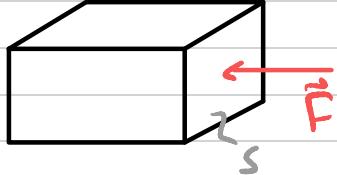
Two types of forces on dV

- Volume forces ($F \propto dV$)

e.g., gravity $\vec{F} = \rho g dV$

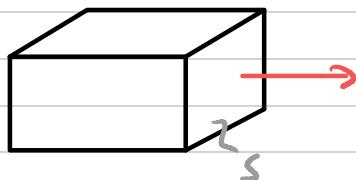
\uparrow mass density

- Surface forces ($F \propto dA$)

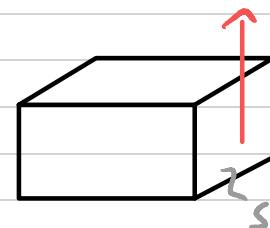


pressure

$$\vec{F} = -\rho \hat{n} dA$$



tension



Shear

Ideal fluids have no shear modulus

(Real fluids have small shear modulus \Rightarrow viscosity)

Isotropic Pressure of Fluids

Let S_1 & S_2 be two surfaces w/ normal vectors \hat{n}_1 & \hat{n}_2 . Construct third surface S_3 w/ \hat{n}_3 to form isoscles prism

NII gives $\vec{F}_2 = -\rho_3 \hat{n}_3 dA_3$

$$\vec{F}_1 = -\rho_1 \hat{n}_1 dA_1$$
$$\vec{F}_3 = -\rho_3 \hat{n}_3 dA_3$$
$$\vec{F}_1 + \vec{F}_2 + \vec{F}_3 + \vec{F}_{vol} = m\vec{a}$$
$$\Rightarrow \underbrace{\vec{F}_1 + \vec{F}_2 + \vec{F}_3}_{\text{Surface forces}} = \underbrace{m\vec{a} - \vec{F}_{vol}}_{\text{Volume forces}}$$

Shrink size by λ factor

$$\Rightarrow \lambda^2 (\vec{F}_1 + \vec{F}_2 + \vec{F}_3) = \lambda^3 (m\vec{a} - \vec{F}_{vol})$$

$$\text{as } \lambda \rightarrow 0 \Rightarrow \vec{F}_1 + \vec{F}_2 + \vec{F}_3 = \lambda (m\vec{a} - \vec{F}_{vol}) = \vec{0}$$

Since isosoles, $\hat{n}_3 \parallel \vec{F}_1 \parallel \hat{n}_1 \parallel \vec{F}_2 \Rightarrow F_1 = F_2 \Rightarrow \rho_1 = \rho_2$

\Rightarrow isotropic pressure (Direct result of no shear modulus)

Stress & Strain

Stress is the ratio of surface force F to the applied area

examples: Stress = $\frac{F}{A}$ = pressure P for static fluid

$$\text{Stress} = \frac{T}{A} \quad \text{for wire in tension}$$

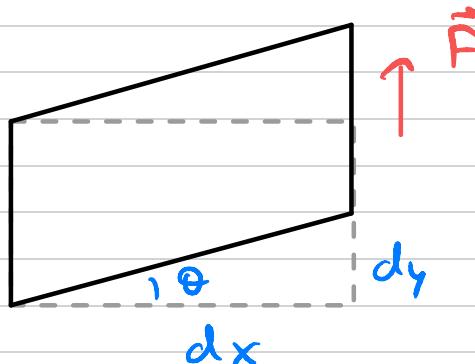
$$\text{Stress} = \frac{\text{Shear Force}}{A} = \text{shear stress}$$

Strain is the deformation of object as the result of stress. (fractional deformation)

examples: Strain = $\frac{dV}{V}$ for static fluid

$$\text{Strain} = \frac{dl}{l} \quad \text{for wire in tension}$$

$$\text{Strain} = \frac{dy}{dx} \quad \text{for shear}$$



Stress & Strain are related by properties of matter. For stresses on a medium which is not too large, expect strain to be linear to stress

$$\text{Stress} \propto \text{Strain}$$

The proportionality factor is called the Elastic modulus

For a stretched wire,

$$\frac{dF}{A} = YM \frac{dl}{l}$$

↳ Young's modulus

For hydrostatic pressure,

$$dp = -BM \frac{dV}{V}$$

↳ Bulk modulus

For shearing forces,

$$\frac{F}{A} = SM \frac{dy}{dx}$$

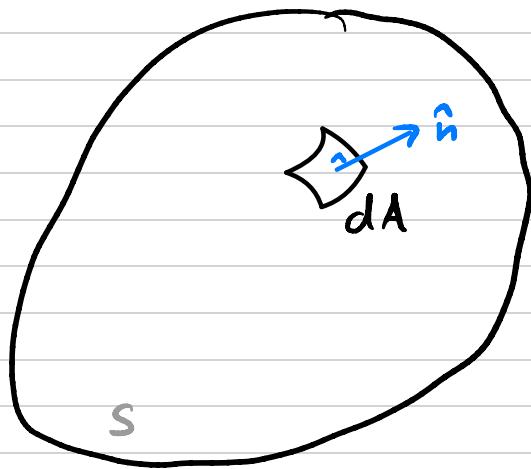
↳ Shear modulus

The Stress Tensor

Now we will make the concept of stress more rigorous. Consider a surface force on a small area $d\vec{A}$ of a closed surface S of some continuous medium.

Define oriented vector

$$d\vec{A} = \hat{n} dA$$



The surface force acting on the area $d\vec{A}$ is $\vec{F}(d\vec{A})$.

It is a linear function of $d\vec{A}$, i.e.,

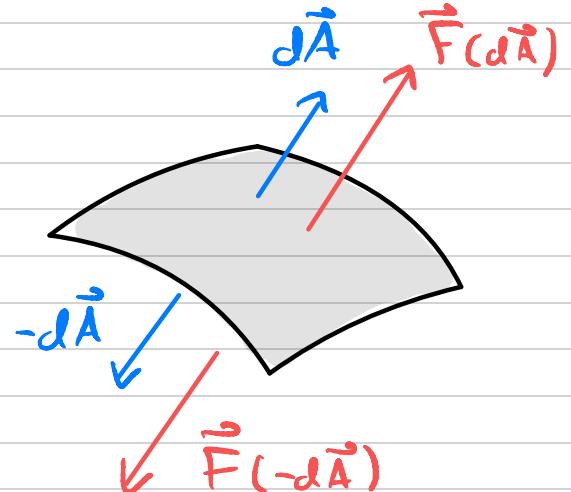
$$\vec{F}(\lambda_1 d\vec{A}_1 + \lambda_2 d\vec{A}_2) = \lambda_1 \vec{F}(d\vec{A}_1) + \lambda_2 \vec{F}(d\vec{A}_2)$$

Proof: First note that as long as $d\vec{A}$ small,

$$\vec{F}(\lambda d\vec{A}) = \lambda \vec{F}(d\vec{A})$$

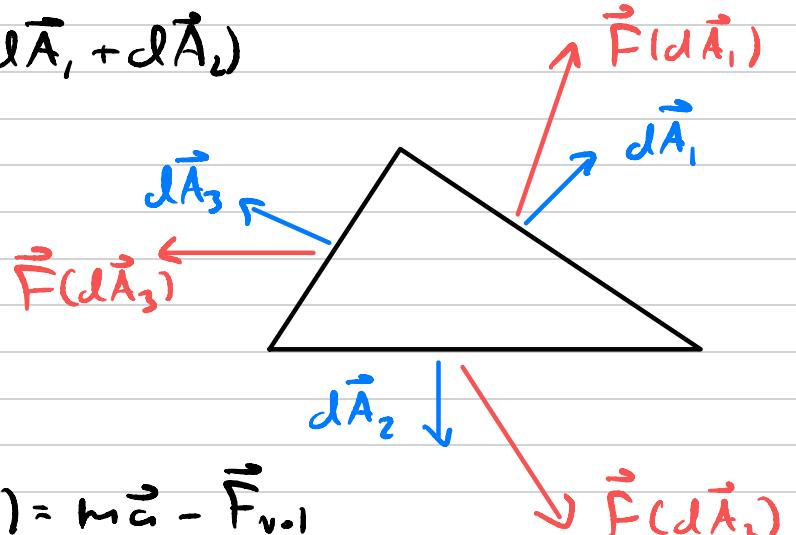
like wise, by NIII

$$\vec{F}(-d\vec{A}) = -\vec{F}(d\vec{A})$$



Next, consider two elements $d\vec{A}_1$ & $d\vec{A}_2$

Find a third $d\vec{A}_3 = -(d\vec{A}_1 + d\vec{A}_2)$



so, NII gives

$$\vec{F}(d\vec{A}_1) + \vec{F}(d\vec{A}_2) + \vec{F}(d\vec{A}_3) = m\vec{a} - \vec{F}_{v,0}$$

As Lfor, if surface size $\rightarrow 0$, then

$$\vec{F}(d\vec{A}_1) + \vec{F}(d\vec{A}_2) + \vec{F}(d\vec{A}_3) = \vec{0}$$

$$\Rightarrow \vec{F}(-d\vec{A}_3) = -\vec{F}(d\vec{A}_3) = \vec{F}(d\vec{A}_1) + \vec{F}(d\vec{A}_2)$$

$$T3D, d\vec{A}_3 = -(d\vec{A}_1 + d\vec{A}_2)$$

$$\Rightarrow \vec{F}(d\vec{A}_1 + d\vec{A}_2) = \vec{F}(d\vec{A}_1) + \vec{F}(d\vec{A}_2)$$

Combine with $\vec{F}(\lambda d\vec{A}) = \lambda \vec{F}(d\vec{A})$, and yield

$$\vec{F}(\lambda_1 d\vec{A}_1 + \lambda_2 d\vec{A}_2) = \lambda_1 \vec{F}(d\vec{A}_1) + \lambda_2 \vec{F}(d\vec{A}_2) \quad \blacksquare$$

So, $\vec{F}(d\vec{A})$ is linear in $d\vec{A}$.

For a fluid, $F(d\vec{A}) = -\rho d\vec{A}$. BT, the most general relation is

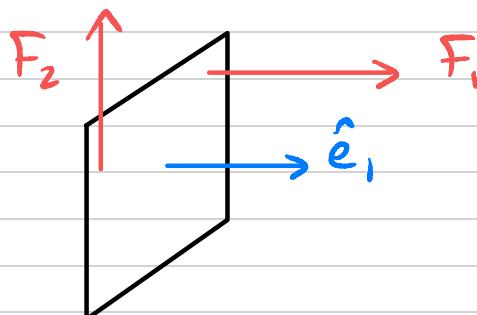
$$F_j(d\vec{A}) = \sum_{k=1}^3 \sigma_{jk} dA_k \quad (j, k = 1, 2, 3 = x, y, z)$$

This defines the 3×3 stress tensor Σ , with elements σ_{jk} . In matrix form,

$$\vec{F}(d\vec{A}) = \Sigma \cdot d\vec{A}$$

Notice - \vec{F} does not need to be \parallel or \perp to surface

Consider the surface element $d\vec{A} = dA \hat{e}_1$



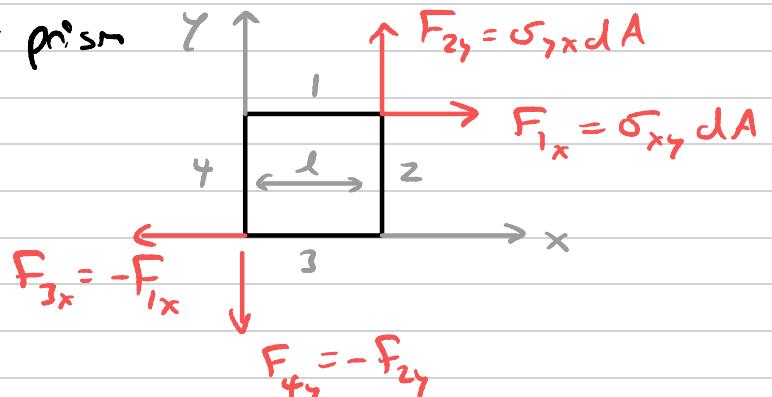
$$\begin{aligned} F_j &= \sum_k \sigma_{jk} dA e_{1,k} \\ &= \sigma_{j1} dA \end{aligned}$$

$$\text{so, } F_1 = \sigma_{11} dA \quad \text{normal force}$$

$$\begin{aligned} F_2 &= \sigma_{21} dA \\ F_3 &= \sigma_{31} dA \end{aligned} \quad \left. \begin{array}{l} \} \\ \} \end{array} \right. \text{shear forces}$$

The stress tensor is symmetric. Consider

the square element of prism



The torque on the element is

$$\Gamma_z = F_{2y} l - F_{1x} l$$

$$= (\sigma_{yx} - \sigma_{xy}) l dA$$

$$= \frac{dL_z}{dt}$$

Now, rescale 3D prism by λ

$$\Rightarrow \Gamma_z \rightarrow \lambda^3 \Gamma_z$$

$$\text{but, } L_z \propto I \propto m l^2 \propto \rho l^4 \rightarrow \lambda^4 L_z$$

$$\Rightarrow \text{as } \lambda \rightarrow 0 \Rightarrow \Gamma_z \rightarrow 0 \Rightarrow \sigma_{xy} = \sigma_{yx}$$

Similar arguments hold for other edges

$$\Rightarrow \boxed{\sigma_{jk} = \sigma_{kj}} \Rightarrow \mathbb{I}' \text{ is symmetric!}$$

So, \mathbb{I}' has 6 independent components

Simple example - hydrostatic fluid

In this case, $\vec{F}(d\vec{A}) = -p d\vec{A}$

↳ contact

$$\Rightarrow \sigma_{jk} = -p \delta_{jk}$$

$$\Rightarrow \Sigma' = -p \mathbf{1}$$

$$= \begin{pmatrix} -p & 0 & 0 \\ 0 & -p & 0 \\ 0 & 0 & -p \end{pmatrix}$$