





## Simulation of Natural Convection under the Influence of Magnetic Field by Explicit Local Radial Basis Function Collocation Method

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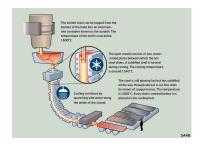
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#### Outline

- ► Introduction
- ► Numerical method
- ► Problem description
- ► Results
- ► Conclusions

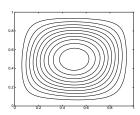
# Introduction Focus and motivation



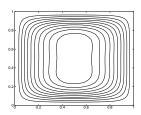


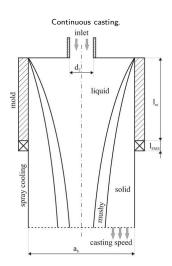
#### Focus and motivation

#### Natural convection.



Natural convection with applied magnetic field.



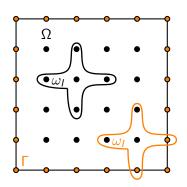


#### Numerical method

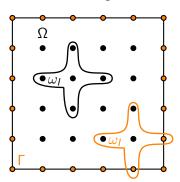
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Used for solving PDF's.

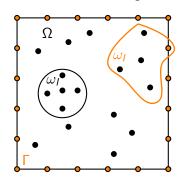
- ► Local influence domain
- Radial Basis Function are used as basis
- ► Collocation (interpolation) of scattered data



#### Uniform node arrangement



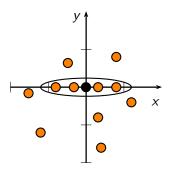
#### Non-uniform node arrangement



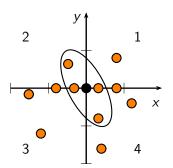
Influence domain for non-uniform node arrangement:

- ► spatial position
- ▶ range

#### Random neighbours



#### Optimal neighbours



General approximation function:

$$\Phi(\vec{p}) \approx \sum_{i=1}^{N} \alpha_i \psi_i(\vec{p})$$

$$\underline{\Psi}\vec{\alpha} = \vec{\phi}$$

Collocation condition:

$$\Phi(\vec{p}_i) = \phi_i$$

$$\frac{\partial^{i}}{\partial p_{j}^{i}}\Phi(\vec{p}) = \sum_{n=1}^{N} \alpha_{n} \frac{\partial^{i}}{\partial p_{j}^{i}} \psi_{n}(\vec{p})$$

Multiquadric RBFs with normalized subdomain and shape parameter c:

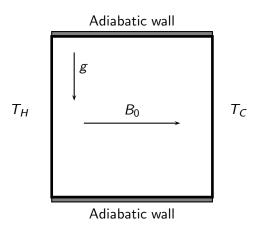
$$\psi_n(\vec{p}) = \sqrt{r_n^2(\vec{p}) + c^2}$$

$$r_n = \sqrt{\left(\frac{p_x - p_{xn}}{p_{x_{max}}}\right)^2 + \left(\frac{p_y - p_{yn}}{p_{y_{max}}}\right)^2 + \left(\frac{p_z - p_{zn}}{p_{z_{max}}}\right)^2}$$

### Problem description

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## Problem description



## Assumptions

- ► Laminar, steady fluid flow
- ► Incompressible fluid flow:  $\nabla \vec{v} = 0$ .
- ▶ Magnetic Raynold's number:  $Re_m = VL\mu_0\sigma \ll 1$ 
  - induced magnetic field is negligible compared to applied magnetic field.
- ▶ Boussinesq approximation:  $\rho = \rho_b (1 \beta_T (T T_C))$ .

## Governing equations

Momentum eq.:

$$\rho\left(\frac{\partial \vec{v}}{\partial t} + (\vec{v} \cdot \nabla)\vec{v}\right) = -\nabla p + \mu \nabla^2 \vec{v} + \vec{F}_m - \vec{g}\rho\beta_T(T - T_C)$$

Continuity eq.:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0 \qquad \Rightarrow \qquad \nabla \cdot \vec{v} = 0$$

Energy eq.:

$$\frac{\partial T}{\partial t} + (\vec{v} \cdot \nabla)T = \alpha_T \nabla^2 T$$

#### Lorentz force term

Equations for Electromagnetic field

Lorentz force:

$$\vec{F}_m = \vec{j} \times \vec{B}$$

Ohm's law:

$$\vec{j} = \sigma(-\nabla\phi + \vec{v} \times \vec{B}) = \vec{j}_i + \vec{j}_d$$

$$\nabla \cdot \vec{j} = 0$$

$$\nabla^2 \phi = \nabla (\vec{v} \times \vec{B})$$

## Boundary conditions

Velocities on walls:

$$u = 0, v = 0.$$

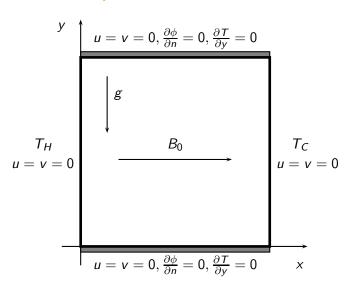
Temperatures:

- ▶ at top and bottom wall (adiabatic):  $\frac{\partial T}{\partial y} = 0$
- ▶ on the left and right walls have predetermined values:  $T_H$ ,  $T_C$  Electric potential (insulating boundary):

$$\frac{\partial \phi}{\partial n} = (\vec{v} \times \vec{B})_{boundary} \cdot \vec{n} \qquad \Rightarrow \qquad \frac{\partial \phi}{\partial n} = 0$$

## Problem description

Domain scheme with boundary conditions



## Governing equations

Dimensionless governing equations

Momentum eq.:

$$\left(rac{\partial ec{V}}{\partial au} + (ec{V}\cdot
abla)ec{V}
ight) = -rac{1}{
ho}
abla P + Pr
abla^2 Pr ec{V}\hat{j} - Gr\Theta$$

Continuity eq.:

$$\nabla \cdot \vec{V} = 0$$

Energy eq.:

$$\frac{\partial \Theta}{\partial au} + (\vec{V} \cdot \nabla)\Theta = \nabla^2 \Theta$$

## Governing equations

Dimensionless numbers and nondimensionalization

#### Dimensionless numbers:

$$Ra = \frac{g\beta_{T}(T_{H} - T_{C})L^{3}}{\nu\alpha_{T}}$$

$$Pr = \frac{\nu}{\alpha_{T}}$$

$$Ha = BL\sqrt{\frac{\sigma}{\mu}}$$

$$Gr = \frac{g\beta_{T}(T_{H} - T_{C})L^{3}}{\alpha_{T}^{2}}$$

#### Non-dimensional variables:

$$X = \frac{x}{L}, \qquad Y = \frac{y}{L}$$

$$U = \frac{uL}{\alpha_T}, \qquad V = \frac{vL}{\alpha_T}$$

$$P = \frac{pL^2}{\alpha_T^2}, \qquad \Theta = \frac{T - T_C}{T_H - T_C}$$

$$\tau = \frac{t\alpha_T}{L^2}$$

#### Results

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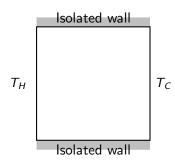
#### Test cases

Three different test cases were devised in order to test the numerical method:

- ► De Vahl Davis benchmark test (TC1)
- ► Natural convection under the influence of magnetic field (TC2)
- Natural convection under the influence of magnetic field for low Pr numbers (TC3)

# Natural convection De Vahl Davis benchmark test (TC1)

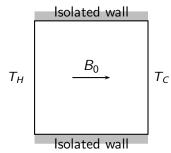
- closed, differentially heated square cavity
- ► Pr = 0.71
- $Arr Ra = 10^4 10^6$
- ► *Ha* = 0



#### Natural convection

Natural convection under the influence of magnetic field (TC2)

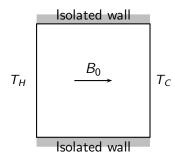
- ► closed, differentially heated square cavity
- ► Pr = 0.71
- $Gr = 10^4 10^6$
- ► Ha = 0 100



#### Natural convection

Natural convection under the influence of magnetic field for low Pr numbers (TC3)

- ► closed, differentially heated square cavity
- ► Pr = 0.14
- $Gr = 10^4 10^6$
- ► Ha = 0 100



#### Natural convection

Natural convection under the influence of magnetic field for low Pr numbers (TC3)

#### Material properties of molten steel:

▶ 
$$\rho = 7200 \text{ kg/m}^3$$

► 
$$c_p = 700 \text{ J/(kg K)}$$

$$ightharpoonup$$
  $\alpha_T = 30 \text{ W/(mK)}$ 

$$\beta_T = 1.0 \cdot 10^{-4} / \text{K}$$

• 
$$\mu = 0.006 \text{ kg/(ms)}$$

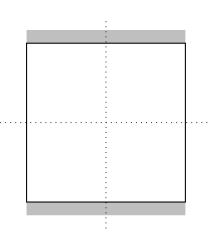
$$Pr = 0.14$$

#### Results

► Nusselt number:

$$Nu = \int_0^L \frac{\partial T}{\partial x} \bigg|_{x=0} dy$$

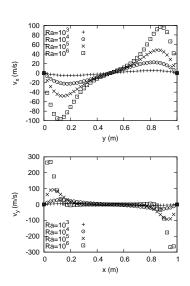
- ▶ u along y = 0.5
- $\triangleright$  v along x = 0.5



# Results De Vahl Davis benchmark test (TC1)

Ra	Nu(a)	Nu(b)	Nu(c)
$10^{3}$	1.108	1.116	1.101
10 <sup>4</sup>	2.223	2.234	2.075
$10^{5}$	4.497	4.510	4.624
$10^{6}$	8.685	8.798	8.97*

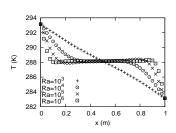
(a) present; (b) De Vahl Davis, 1983; (c) Kosec et al., 2007



# Results De Vahl Davis benchmark test (TC1)

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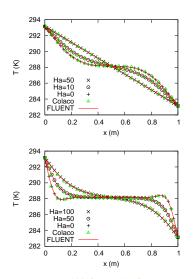


#### Results

#### Natural convection under the influence of magnetic field (TC2)

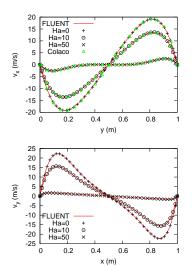
На	Nu(a)	<i>Nu</i> (b)	Nu(c)			
$Gr = 10^4$						
0	2.03	2.02	2.06			
10	1.71	1.70	1.84			
50	1.01	0.97	1.06			
$Gr = 10^6$						
0	8.15	9.21	7.98			
10	7.99	9.04	7.88			
100	3.33	3.54	4.27			

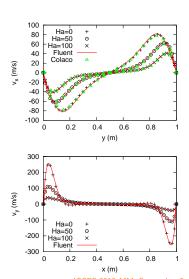
(a) present; (b) Colaço et all, 2009; (c) FVM



#### Results

#### Natural convection under the influence of magnetic field (TC2)

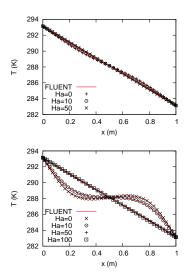


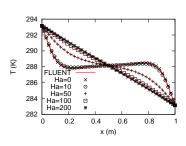


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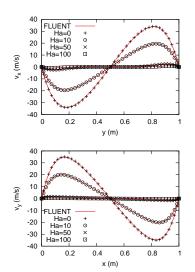
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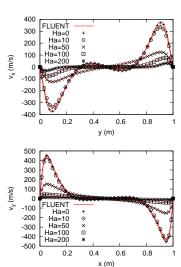
# Natural convection under the influence of magnetic field for low Pr numbers (TC3) Temperatures along the line through the center of the cavity



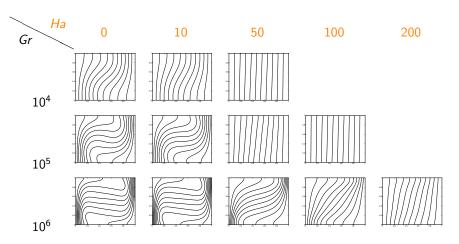


# Natural convection under the influence of magnetic field for low Pr numbers (TC3) Velocities along line through the center of the cavity



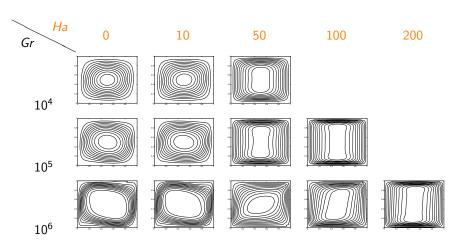


# Natural convection under the influence of magnetic field for low Pr numbers (TC3) | Southerms |



# Natural convection under the influence of magnetic field for low Pr numbers (TC3)

Stream Functions



#### **Conclusions**

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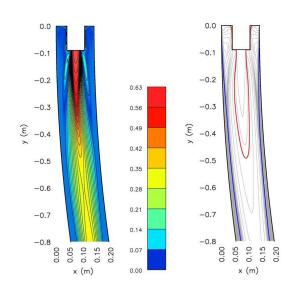
# Summary and Conclusions Summary

- Method is tested first for a natural convection benchmark test and than for natural convection under the influence of external magnetic field.
- ► A comparison between in-house meshless method, FLUENT comercial code and results from published papers (De Vahl Davis, 1983; Colaço et al., 2009) is made.

#### **Conclusions**

#### Future plans:

- complex geometries
- simplified industrial application
- continuous casting real curved geometry



#### **Affiliation**

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#### Nusselt numbers

