

## hw3\_ex12

February 20, 2020

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[1]: #!/usr/bin/env python3
# -*- coding: utf-8 -*-
"""

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"""

# python 3 version 2/15
import numpy as np
import matplotlib.pyplot as plt

def rk4(x,t,tau,derivsRK):
    ## Runge-Kutta integrator (4th order)
    ## Input arguments -
    ## x = current value of dependent variable
    ## t = independent variable (usually time)
    ## tau = step size (usually timestep)
    ## derivsRK = right hand side of the ODE; derivsRK is the
    ## name of the function which returns dx/dt
    ## Calling format derivsRK(x,t).
    ## Output arguments -
    ## xout = new value of x after a step of size tau
    half_tau = 0.5*tau
    F1 = derivsRK(x,t)
    t_half = t + half_tau
    xtemp = x + half_tau*F1
    F2 = derivsRK(xtemp,t_half)
    xtemp = x + half_tau*F2
    F3 = derivsRK(xtemp,t_half)
    t_full = t + tau
    xtemp = x + tau*F3
    F4 = derivsRK(xtemp,t_full)
    xout = x + tau/6.*(F1 + F4 + 2.*(F2+F3))
    return xout
def rka(x,t,tau,err,derivsRK):
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%% Adaptive Runge-Kutta routine
%% Inputs
%% x          Current value of the dependent variable
%% t          Independent variable (usually time)
%% tau        Step size (usually time step)
%% err        Desired fractional local truncation error
%% derivsRK   Right hand side of the ODE; derivsRK is the
%%            name of the function which returns dx/dt
%%            Calling format derivsRK(x,t).
%% Outputs
%% xSmall     New value of the dependent variable
%% t          New value of the independent variable
%% tau        Suggested step size for next call to rka

%%* Set initial variables
tSave = t; xSave = x    # Save initial values
safe1 = .9; safe2 = 4.  # Safety factors
eps = np.spacing(1) # smallest value

%%* Loop over maximum number of attempts to satisfy error bound
maxTry = 100

    for iTry in range(1,maxTry):

%%* Take the two small time steps
        half_tau = 0.5 * tau
        xTemp = rk4(xSave,tSave,half_tau,derivsRK)
        t = tSave + half_tau
        xSmall = rk4(xTemp,t,half_tau,derivsRK)

%%* Take the single big time step
        t = tSave + tau
        xBig = rk4(xSave,tSave,tau,derivsRK)

%%* Compute the estimated truncation error
        scale = err * (np.abs(xSmall) + np.abs(xBig))/2.
        xDiff = xSmall - xBig
        errorRatio = np.max( [np.abs(xDiff)/(scale + eps)] )

        #print safe1,tau,errorRatio

%%* Estimate news tau value (including safety factors)
        tau_old = tau

        tau = safe1*tau_old*errorRatio**(-0.20)
        tau = np.max([tau,tau_old/safe2])
        tau = np.min([tau,safe2*tau_old])

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    ##* If error is acceptable, return computed values
    if errorRatio < 1 :
        # xSmall = xSmall ## + (xDiff)/15
        # xSmall = (16.*xSmall - xBig)/15. # correction
        return xSmall, t, tau

##* Issue error message if error bound never satisfied
print ('ERROR: Adaptive Runge-Kutta routine failed')
return

def gravrk(s,t):
    ## Returns right-hand side of Kepler ODE; used by Runge-Kutta routines
    ## Inputs
    ## s State vector [r(1) r(2) v(1) v(2)]
    ## t Time (not used)
    ## Output
    ## deriv Derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]

    GM = 4*np.pi**2

    ##* Compute acceleration
    r = np.array([s[0], s[1]]) # Unravel the vector s into position and
    →velocity
    v = np.array([s[2], s[3]])
    accel = -GM*r/np.linalg.norm(r)**3 # Gravitational acceleration

    ##* Return derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]
    derivs = np.array([v[0], v[1], accel[0], accel[1]])
    return derivs

def gravrk_ex12(s,t):
    ## Returns right-hand side of Kepler ODE; used by Runge-Kutta routines
    ## Inputs
    ## s State vector [r(1) r(2) v(1) v(2)]
    ## t Time (not used)
    ## Output
    ## deriv Derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]

    GM = 4*np.pi**2
    # Hardcoded in vars since
    # Im not sure on an elegant way to add them to function
    r0 = np.array([1.0,0])
    init_Grav_accel = .01*-GM*r0/np.linalg.norm(r0)**3

    ##* Compute acceleration
    r = np.array([s[0], s[1]]) # Unravel the vector s into position and
    →velocity

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v = np.array([s[2] ,s[3]])
# Gravitational acceleration
accel = -GM*r/np.linalg.norm(r)**3 + init_Grav_accel

#%% Return derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]
derivs = np.array([v[0], v[1], accel[0], accel[1]])
return derivs

# orbit - Program to compute the orbit of a comet.
#clear all; help orbit; % Clear memory and print header

def orbit(input_dict = {},calc_info = False, plot_momentum = False,
          inter_input = False, plot_traj = True, plot_energy = False):
    if inter_input:
        # Set initial position and velocity of the comet.
        r0 = float(input("Enter initial radial distance (AU): "))
        # v0 = float(input("Enter initial tangential velocity (AU/yr): "))
        vstr = str(input("Enter initial tangential velocity (AU/yr) as a number,
→or multiple of Pi (e.g.,2*pi): "))

        # modify input to allow 'pi'
        vinp = vstr.split('*')
        if (vinp[-1].lower() == 'pi'):
            v0 = float(vinp[0])*np.pi
        else:
            v0 = float(vinp[0])

        nStep = int(input("Enter number of steps: "))
        tau = float(input("Enter time step (yr): "))
        NumericalMethod=0
        while(NumericalMethod not in np.array([1,2,3,4,5,6,7])):
            NumericalMethod = int(input("Choose a number for a numerical method:
→\n\
            1-Euler, 2-Euler-Cromer, 3-Runge-Kutta 4-Adaptive R-K: "))
    elif input_dict:
        r0 = input_dict['r0']
        v0 = input_dict['v0']
        nStep = input_dict['nStep']
        tau = input_dict['tau']
        NumericalMethod = input_dict['NumericalMethod']
    else:
        r0 = 1
        v0 = 2*np.pi
        nStep = 1000
        tau = .01
        NumericalMethod = 3

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r = np.array([r0, 0.])
v = np.array([0., v0])
state = np.array([ r[0], r[1], v[0], v[1] ])    # Used by R-K routines

#Set physical parameters (mass, G*M)
GM = 4*np.pi**2      # Grav. const. * Mass of Sun (au3/yr2)
mass = 1.              # Mass of comet
adaptErr = 1.e-4      # Error parameter used by adaptive Runge-Kutta
time = 0.0

#%% Loop over desired number of steps using specified
    numerical method.
for istep in range(0,nStep):

    #%% Record position and energy for plotting.
    # Initially set the arrays for the first step
    if istep == 0:
        rplot = np.linalg.norm(r)
        thplot = np.arctan2(r[1],r[0])
        tplot = time
        kinetic = .5*mass*np.linalg.norm(v)**2
        potential= - GM*mass/np.linalg.norm(r)
        momentum = [np.linalg.norm(np.cross(r,mass*v))]
    else:
        rplot = np.append(rplot,np.linalg.norm(r))          #Record position
    for polar plot
        thplot = np.append(thplot,np.arctan2(r[1],r[0]))
        tplot = np.append(tplot,time)
        kinetic = np.append(kinetic,0.5*mass*np.linalg.norm(v)**2)    # Record
    energies
        potential= np.append(potential,- GM*mass/np.linalg.norm(r))
        momentum.append(np.linalg.norm(np.cross(r, mass*v)))

    #%% Calculate new position and velocity using desired method.
    if NumericalMethod == 1 :
        accel = -GM*r/np.linalg.norm(r)**3
        r = r + tau*v          # Euler step
        v = v + tau*accel
        time = time + tau
    elif NumericalMethod == 2:
        accel = -GM*r/np.linalg.norm(r)**3
        v = v + tau*accel
        r = r + tau*v          # Euler-Cromer step
        time = time + tau
    elif NumericalMethod == 3:
        state = rk4(state,time,tau,gravrk_ex12)

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    r = np.array([state[0], state[1]])    # 4th order Runge-Kutta
    v = np.array([state[2], state[3]])
    time = time + tau
else:
    [state, time, tau] = rka(state,time,tau,adaptErr,gravrk)
    r = np.array([state[0], state[1]])    # Adaptive Runge-Kutta
    v = np.array([state[2], state[3]])
    # If sometime after first step and theta goes from neg to pos, then we
    → know we've completed an orbit
#     if istep != 0 and (thplot[-2:]*[-1,1] > 0).all():
#         break

# print(thplot[-5:])

if plot_traj:
    ##* Graph the trajectory of the comet.
    plt.figure(1); plt.clf()    # Clear figure 1 window and bring forward
    plt.polar(thplot,rplot,'-')    # Use polar plot for graphing orbit
    plt.xlabel('Distance (AU)')
    plt.grid(True)

if plot_energy:
    ##* Graph the energy of the comet versus time.
    plt.figure(2); plt.clf()    # Clear figure 2 window and bring forward
    totalE = kinetic + potential    # Total energy
    plt.plot(tplot,kinetic,'-.',tplot,potential,'--',tplot,totalE,'-')
    # plt.legend('Kinetic','Potential','Total')
    plt.xlabel('Time (yr)'); plt.ylabel('Energy (M AU^2/yr^2)')
    plt.grid(True)
    plt.show()

if plot_momentum:
    # Plots angular momentum as a function of time
    plt.figure(3)
    plt.plot(tplot,momentum)

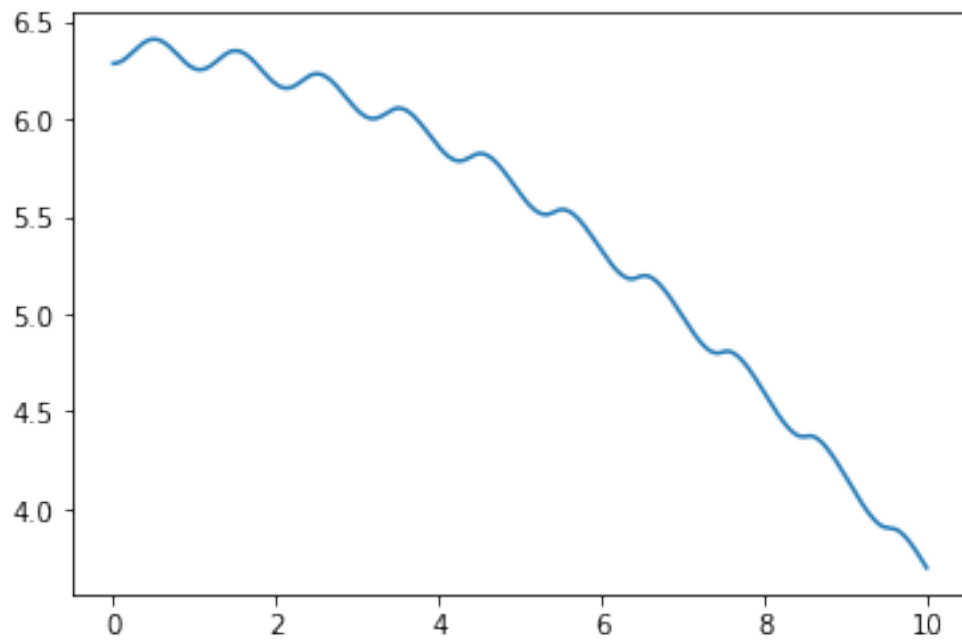
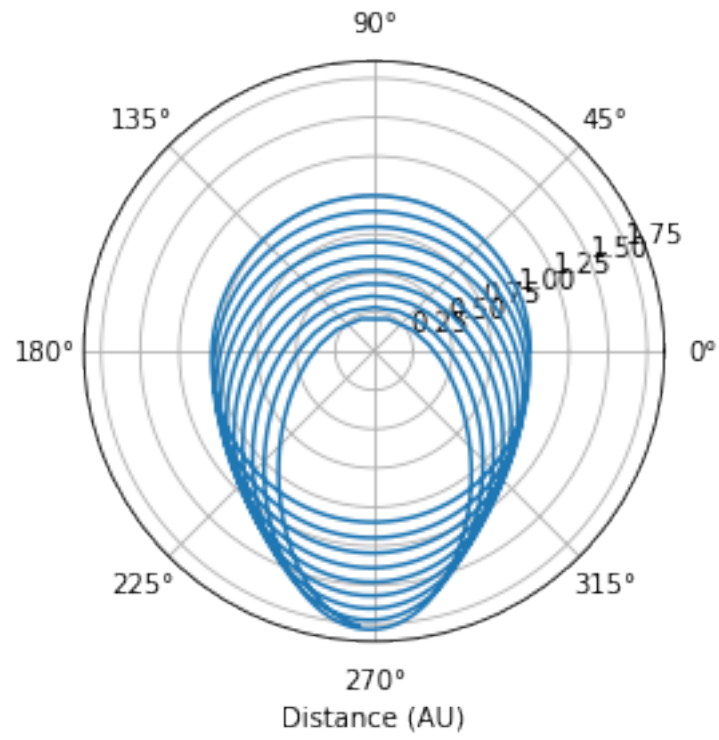
return rplot, thplot

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[2]: # non-elliptical
input_dict = {
    'r0': 1,
    'v0': 2*np.pi,
    'nStep': 1000,
    'tau': .01,
    'NumericalMethod': 3
}
rplot, thplot = orbit(input_dict,plot_momentum = True)

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