hw3 ex14

February 20, 2020

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[1]: #!/usr/bin/env python3
     # -*- coding: utf-8 -*-
     Created on Mon Feb 17 17:01:26 2020
     @author: aswart
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     # python 3 version 2/15
     import numpy as np
     import matplotlib.pyplot as plt
     def rk4(x,t,tau,derivsRK):
     #% Runge-Kutta integrator (4th order)
     #% Input arguments -
     \#\% x = current value of dependent variable
     \#\% t = independent variable (usually time)
        tau = step size (usually timestep)
     #%
         derivsRK = right hand side of the ODE; derivsRK is the
     #%
     #%
                    name of the function which returns dx/dt
                    Calling format derivsRK(x,t).
     #%
     #% Output arguments -
     #% xout = new value of x after a step of size tau
         half_tau = 0.5*tau
         F1 = derivsRK(x,t)
         t_half = t + half_tau
         xtemp = x + half tau*F1
         F2 = derivsRK(xtemp,t_half)
         xtemp = x + half_tau*F2
         F3 = derivsRK(xtemp,t_half)
         t_full = t + tau
         xtemp = x + tau*F3
         F4 = derivsRK(xtemp,t_full)
         xout = x + tau/6.*(F1 + F4 + 2.*(F2+F3))
         return xout
     def rka(x,t,tau,err,derivsRK):
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#% Adaptive Runge-Kutta routine
#% Inputs
#%
    \boldsymbol{x}
              Current value of the dependent variable
#%
              Independent variable (usually time)
#%
   tau
              Step size (usually time step)
              Desired fractional local truncation error
#%
   err
#% derivsRK Right hand side of the ODE; derivsRK is the
#%
              name of the function which returns dx/dt
#%
               Calling format derivsRK(x,t).
#% Outputs
            New value of the dependent variable
    xSmall
#%
#%
              New value of the independent variable
#%
   tau
               Suggested step size for next call to rka
#%* Set initial variables
   tSave = t; xSave = x # Save initial values
    safe1 = .9; safe2 = 4. # Safety factors
    eps = np.spacing(1) # smallest value
#%* Loop over maximum number of attempts to satisfy error bound
   maxTry = 100
   for iTry in range(1,maxTry):
#%* Take the two small time steps
       half tau = 0.5 * tau
       xTemp = rk4(xSave,tSave,half_tau,derivsRK)
       t = tSave + half tau
       xSmall = rk4(xTemp,t,half_tau,derivsRK)
  #%* Take the single big time step
       t = tSave + tau
        xBig = rk4(xSave,tSave,tau,derivsRK)
  #%* Compute the estimated truncation error
        scale = err * (np.abs(xSmall) + np.abs(xBig))/2.
        xDiff = xSmall - xBig
        errorRatio = np.max( [np.abs(xDiff)/(scale + eps)] )
        #print safe1, tau, errorRatio
  #%* Estimate news tau value (including safety factors)
       tau old = tau
       tau = safe1*tau_old*errorRatio**(-0.20)
        tau = np.max([tau,tau_old/safe2])
        tau = np.min([tau,safe2*tau_old])
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#%* If error is acceptable, return computed values
        if errorRatio < 1 :</pre>
          \# xSmall = xSmall \#\% + (xDiff)/15
          \# xSmall = (16.*xSmall - xBiq)/15. <math>\# correction
            return xSmall, t, tau
#%* Issue error message if error bound never satisfied
    print ('ERROR: Adaptive Runge-Kutta routine failed')
    return
def gravrk(s,t):
#% Returns right-hand side of Kepler ODE; used by Runge-Kutta routines
#% Inputs
        State vector [r(1) \ r(2) \ v(1) \ v(2)]
#% s
#%
    t
            Time (not used)
#% Output
#% deriv Derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]
    GM = 4*np.pi**2
#%* Compute acceleration
    r = np.array([s[0], s[1]]) # Unravel the vector s into position and
\rightarrow velocity
    v = np.array([s[2],s[3]])
    accel = -GM*r/np.linalg.norm(r)**3 # Gravitational acceleration
#%* Return derivatives [dr(1)/dt \ dr(2)/dt \ dv(1)/dt \ dv(2)/dt]
    derivs = np.array([v[0], v[1], accel[0], accel[1]])
    return derivs
def gravrk_ex14(s,t):
#% Returns right-hand side of Kepler ODE; used by Runge-Kutta routines
#% Inputs
#%
    S
            State vector [r(1) \ r(2) \ v(1) \ v(2)]
#% t
            Time (not used)
#% Output
    deriv Derivatives [dr(1)/dt dr(2)/dt dv(1)/dt dv(2)/dt]
#%
    GM = 4*np.pi**2
    # Hardcoded in vars since
    # Im not sure on an elegant way to add them to function
    alpha = .1
#%* Compute acceleration
    r = np.array([s[0], s[1]]) # Unravel the vector s into position and
\rightarrow velocity
   v = np.array([s[2], s[3]])
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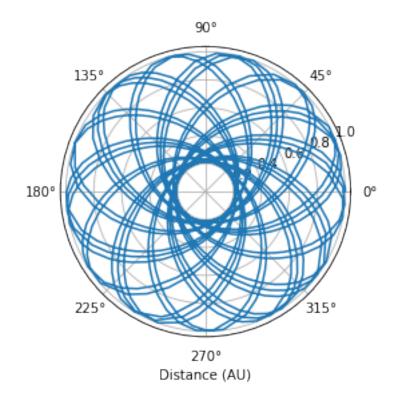
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# Gravitational acceleration
    Grav_mult = 1-(alpha/np.linalg.norm(r))
    accel = (-GM*r/np.linalg.norm(r)**3)*Grav_mult
#%* Return derivatives [dr(1)/dt \ dr(2)/dt \ dv(1)/dt \ dv(2)/dt]
    derivs = np.array([v[0], v[1], accel[0], accel[1]])
    return derivs
# orbit - Program to compute the orbit of a comet.
#clear all; help orbit; % Clear memory and print header
def orbit(input_dict = {}, calc_info = False, plot_momentum = False,
          inter_input = False, plot_traj = True, plot_energy = False):
    if inter input:
        # Set initial position and velocity of the comet.
        r0 = float(input("Enter initial radial distance (AU): "))
        # v0 = float(input("Enter initial tangential velocity (AU/yr): "))
        vstr = str(input("Enter initial tangential velocity (AU/yr) as a number ⊔
→or multiple of Pi (e.g.,2*pi): "))
        # modify input to allow 'pi'
        vinp = vstr.split('*')
        if (vinp[-1].lower() == 'pi'):
          v0 = float(vinp[0])*np.pi
        else:
          v0 = float(vinp[0])
        nStep = int(input("Enter number of steps: "))
        tau = float(input("Enter time step (yr): "))
        NumericalMethod=0
        while(NumericalMethod not in np.array([1,2,3,4,5,6,7])):
            NumericalMethod = int(input("Choose a number for a numerical method:
\hookrightarrow \n
         1-Euler, 2-Euler-Cromer, 3-Runge-Kutta 4-Adaptive R-K: "))
    elif input dict:
        r0 = input_dict['r0']
        v0 = input_dict['v0']
        nStep = input_dict['nStep']
        tau = input_dict['tau']
        NumericalMethod = input_dict['NumericalMethod']
    else:
        r0 = 1
        v0 = 2*np.pi
        nStep = 1000
        tau = .01
        NumericalMethod = 3
```

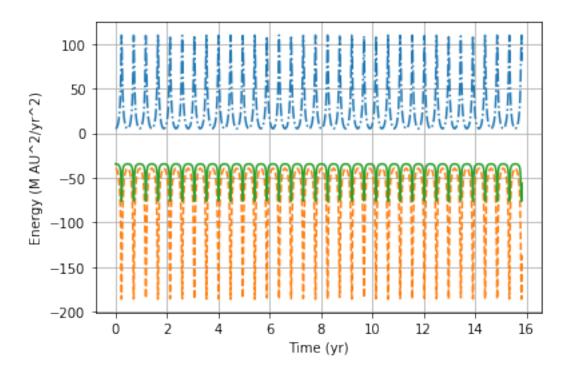
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r = np.array([r0, 0.])
   v = np.array([0., v0])
   state = np.array([ r[0], r[1], v[0], v[1] ]) # Used by R-K routines
   #Set physical parameters (mass, G*M)
   GM = 4*np.pi**2  # Grav. const. * Mass of Sun (au^3/yr^2)
   mass = 1. # Mass of comet
   adaptErr = 1.e-4 # Error parameter used by adaptive Runge-Kutta
   time = 0.0
   #%* Loop over desired number of steps using specified
   #% numerical method.
   for istep in range(0,nStep):
     #%* Record position and energy for plotting.
     # Initially set the arrays for the first step
     if istep == 0:
         rplot = np.linalg.norm(r)
         thplot = np.arctan2(r[1],r[0])
         tplot = time
         kinetic = .5*mass*np.linalg.norm(v)**2
         potential= - GM*mass/np.linalg.norm(r)
         momentum = [np.linalg.norm(np.cross(r,mass*v))]
         perihelion = []
         aphelion = []
     else:
         rplot = np.append(rplot,np.linalg.norm(r)) #Record position_
→ for polar plot
         thplot = np.append(thplot,np.arctan2(r[1],r[0]))
         tplot = np.append(tplot,time)
         kinetic = np.append(kinetic,0.5*mass*np.linalg.norm(v)**2) # Recordu
\rightarrow energies
         potential= np.append(potential, - GM*mass/np.linalg.norm(r))
         momentum.append(np.linalg.norm(np.cross(r, mass*v)))
     #%* Calculate new position and velocity using desired method.
     if NumericalMethod == 1 :
      accel = -GM*r/np.linalg.norm(r)**3
      r = r + tau*v
                                # Euler step
      v = v + tau*accel
      time = time + tau
     elif NumericalMethod == 2:
      accel = -GM*r/np.linalg.norm(r)**3
      v = v + tau*accel
      r = r + tau*v
                                # Euler-Cromer step
      time = time + tau
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elif NumericalMethod == 3:
       state = rk4(state,time,tau,gravrk_ex14)
                                          # 4th order Runge-Kutta
      r = np.array([state[0], state[1]])
      v = np.array([state[2], state[3]])
      time = time + tau
     else:
       [state, time, tau] = rka(state,time,tau,adaptErr,gravrk_ex14)
      r = np.array([state[0], state[1]])
                                          # Adaptive Runge-Kutta
      v = np.array([state[2], state[3]])
     # Find perihelion, aphelion
     if istep >= 2:
         if rplot[istep-1] > rplot[istep-2] and rplot[istep-1] > rplot[istep]:
             \#if\ rplot[istep-1] >= min(rplot)-tau\ and\ rplot[istep-1] <=
\rightarrow min(rplot)+tau:
                perihelion.append( (rplot[istep-1], tplot[istep-1]) )
            if rplot[istep-1] >= max(rplot)-tau and rplot[istep-1] <=__
→max(rplot)+tau:
                aphelion.append(_
elif rplot[istep-1] < rplot[istep-2] and rplot[istep-1] <
→rplot[istep]:
            perihelion.append(
→[rplot[istep-1],thplot[istep-1]],tplot[istep-1]])
   # Given alpha=.1, calculate a
  alpha = .1
  a = np.sqrt(1 + (GM*mass**2*alpha / (momentum[0]**2)))
  shift rad = (360*(1-a)/a)*(np.pi/180)
   # Calculate aphelion, perihelion
  aphelion = np.array(aphelion)
  perihelion = np.array(perihelion)
  ap thetas = aphelion[:,1]
  peri_thetas = perihelion[:,1]
  a1 = ap_{thetas}[0:-1]
  a2 = ap_thetas[1:]
  adelta_norm = np.array([i+2*np.pi if i<0 else i for i in (a1-a2)])
  p1 = peri_thetas[0:-1]
  p2 = peri_thetas[1:]
  pdelta_norm = np.array([i+2*np.pi if i<0 else i for i in (p1-p2)])</pre>
  print("Difference between predicted and average observed orbit precession⊔
→in aphelion:")
  print(np.mean(shift_rad+adelta_norm))
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if plot_traj:
        #%* Graph the trajectory of the comet.
       plt.figure(1); plt.clf() #Clear figure 1 window and bring forward
       plt.polar(thplot,rplot,'-') # Use polar plot for graphing orbit
       plt.xlabel('Distance (AU)')
       plt.grid(True)
    if plot_energy:
        #%* Graph the energy of the comet versus time.
       plt.figure(2); plt.clf() # Clear figure 2 window and bring forward
       totalE = kinetic + potential # Total energy
       plt.plot(tplot,kinetic,'-.',tplot,potential,'--',tplot,totalE,'-')
       #plt.legend('Kinetic', 'Potential', 'Total')
       plt.xlabel('Time (yr)'); plt.ylabel('Energy (M AU^2/yr^2)')
       plt.grid(True)
       plt.show()
   if plot_momentum:
        # Plots angular momentum as a function of time
       plt.figure(3)
       plt.plot(tplot,momentum)
   return rplot, thplot
if __name__ == "__main__":
    # elliptical
   input_dict = {
        'r0': 1,
        'v0': 1*np.pi,
        'nStep': 1000,
        'tau': .01,
        'NumericalMethod': 4
        }
   rplot, thplot = orbit(input_dict,plot_energy = True)
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Difference between predicted and average observed orbit precession in aphelion: -0.002481749038696069





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