

CSIR

Defence, Peace, Safety and Security

Pyradi: an open-source toolkit for infrared calculation and data processing

SPIE S+D 8543-19

Willers, Willers, Santos,
van der Merwe, Calitz, de Waal & Mudau





Infrared Data Analysis

Who has done infrared measurement data analysis?

You will understand the next slide....



Infrared Data Analysis



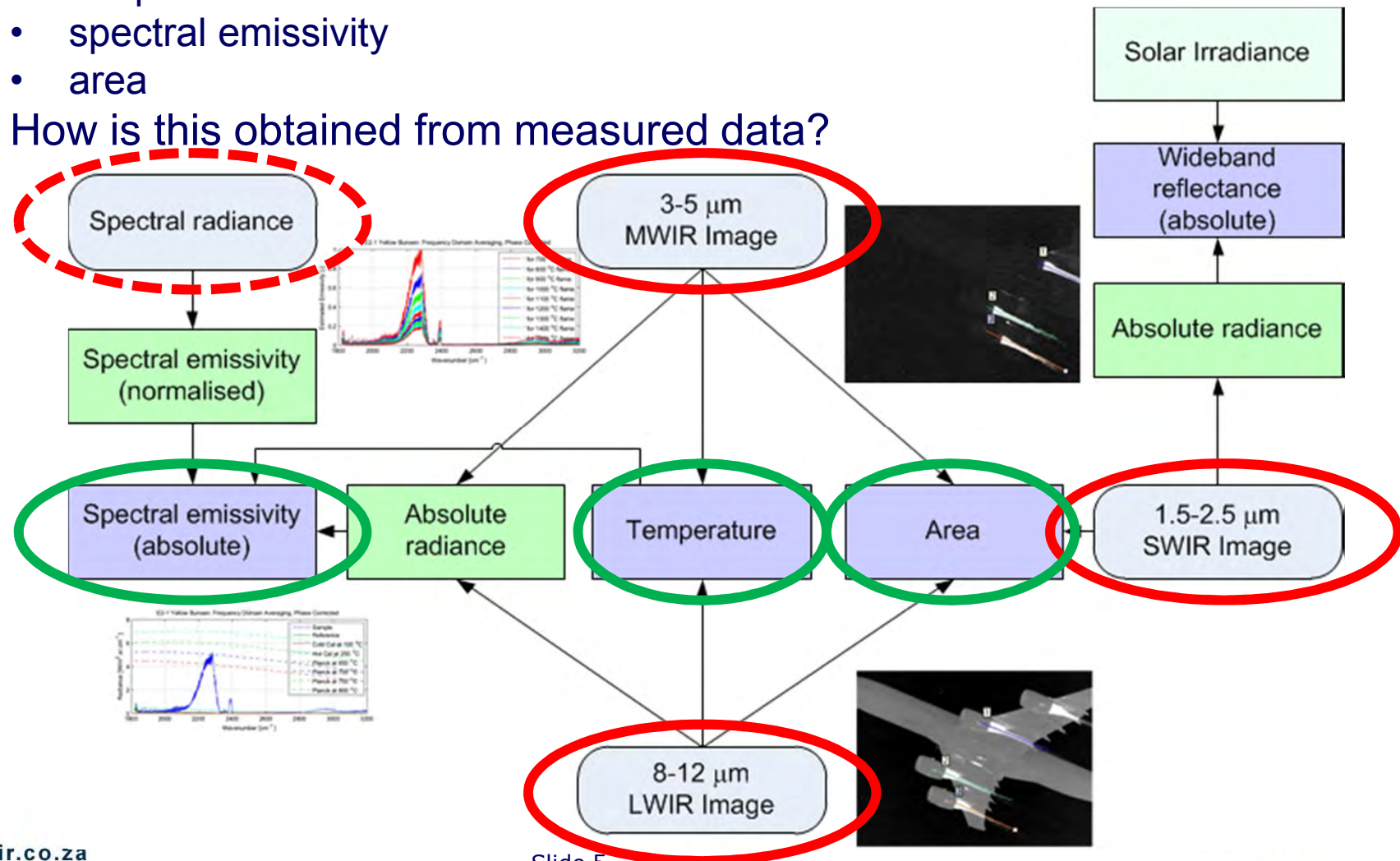


Signatures From Measurements

Main focus of infrared target modelling:

- temperature
- spectral emissivity
- area

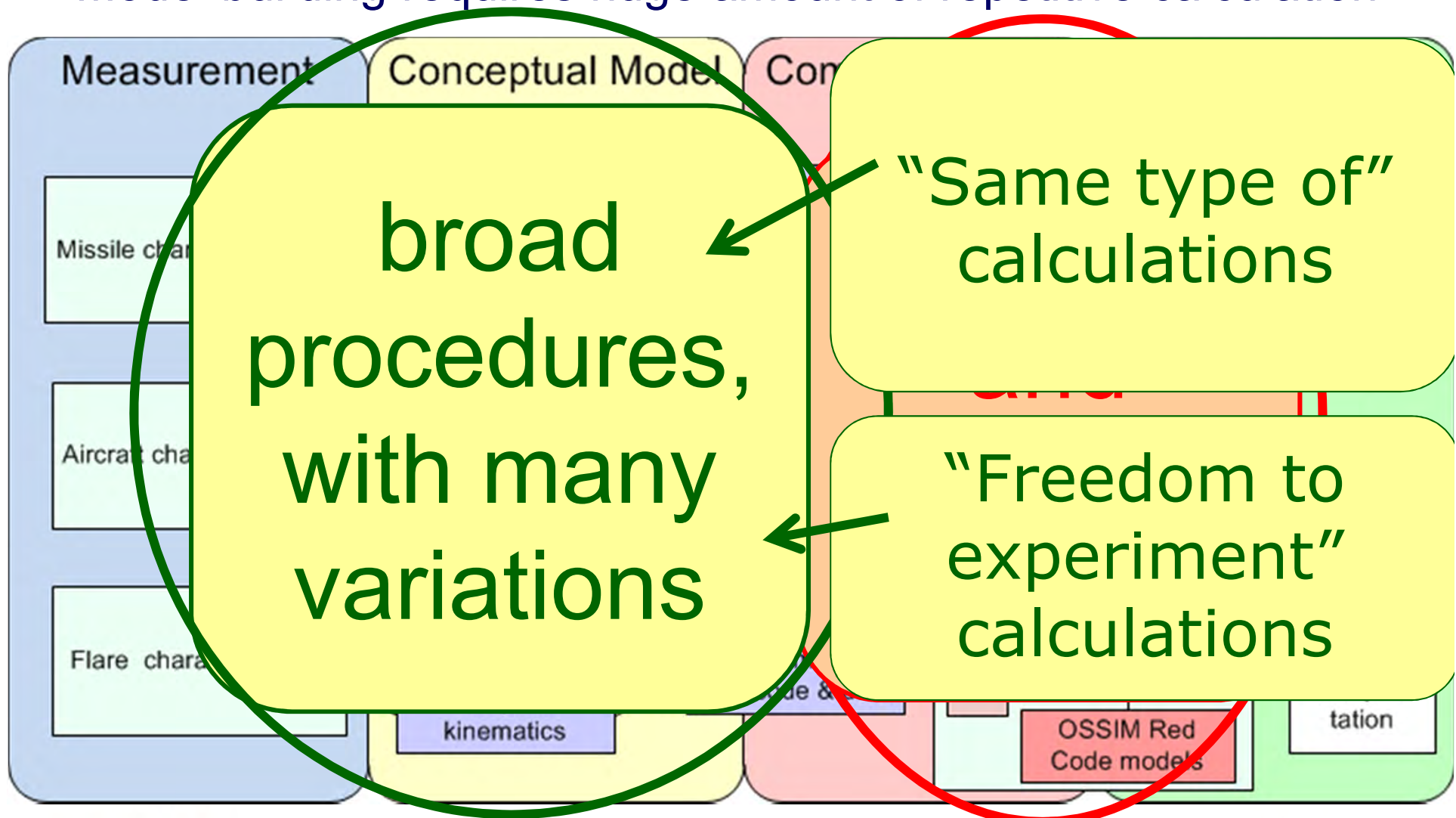
How is this obtained from measured data?





From Measurements to Results

Model building requires huge amount of repetitive calculation





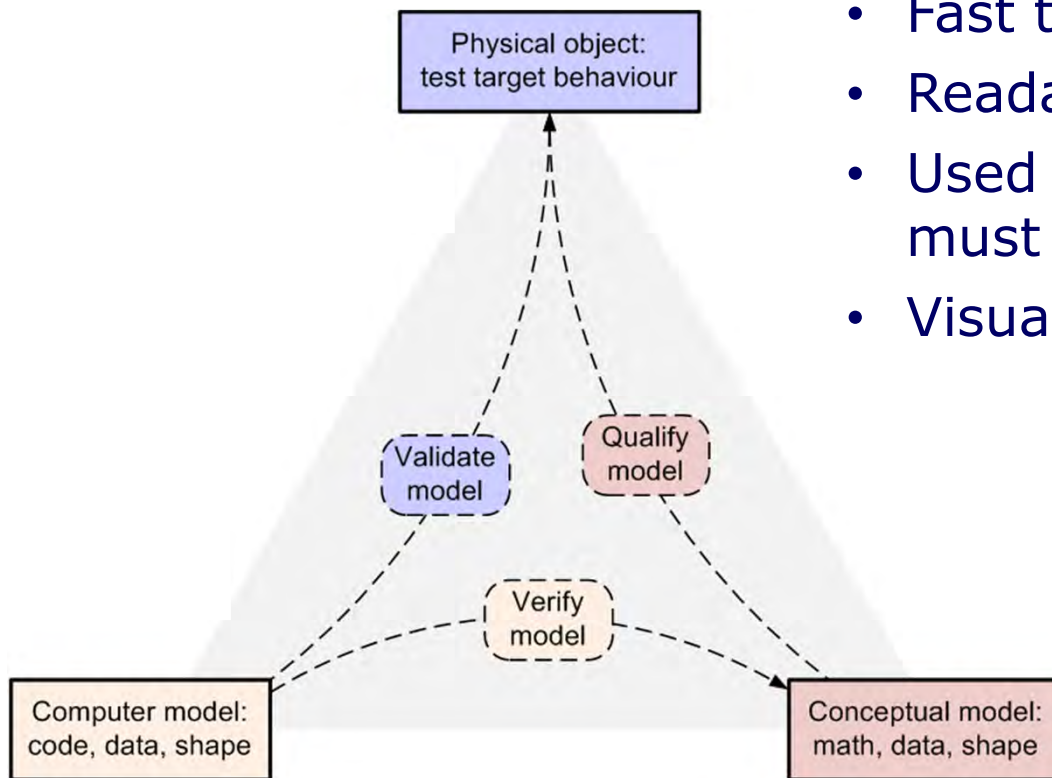
Vital to Insight!





Key Requirements

- Easy to use.
- Fast to prototype.
- Readable, uncluttered code.
- Used in validation/verification: must be a well-tested library.
- Visualisation is critical!





pyradi

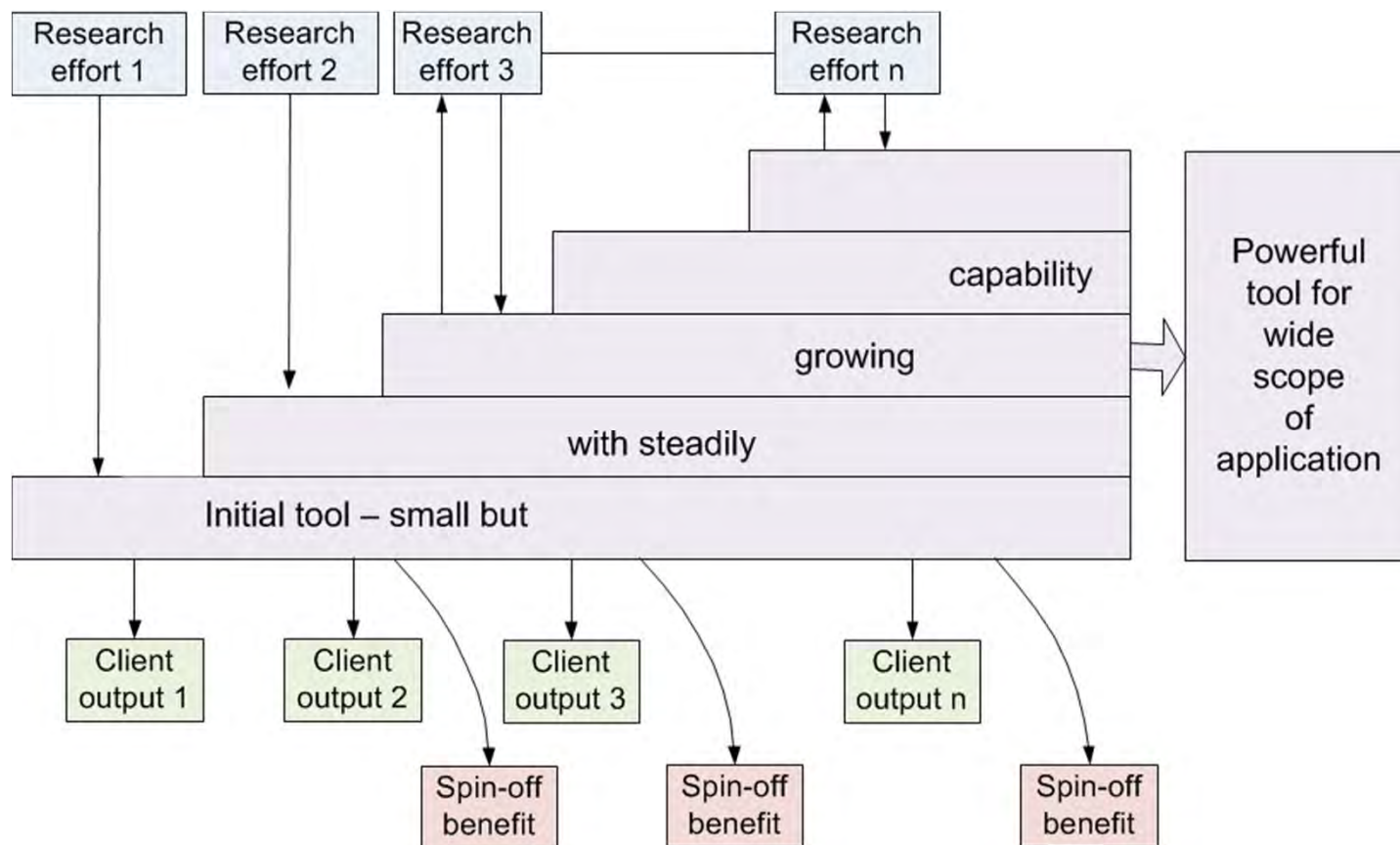
Focussed on electro-optics calculations.

The toolkit covers

- models of physical concepts,
- mathematical operation,
- data manipulation, and
- graphical visualisation.



Cumulative Development





Scripting Language

Matlab vs Python.

Extensive experience with both languages.

Python won:

- General purpose, mainstream, language.
- A more powerful and productive language.
- Many extensions in very many application areas.
- Numpy & Scipy equal to Matlab for scientific work.
- Superior visualisation & graphics tools.
- Easy to learn & very strong open source support.
- Zero acquisition cost & low usage cost.
- Negative: no Simulink, fewer toolboxes.





Unix Tool Philosophy

Write programs that do one thing and do it well.

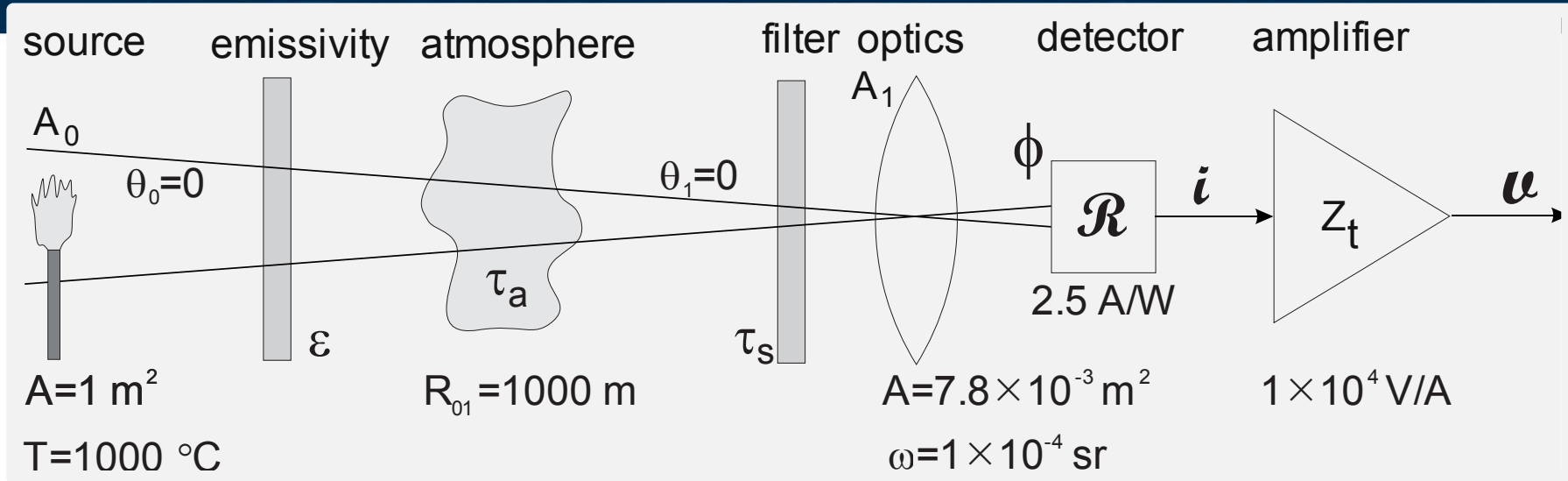
Write programs to work together.

- Simple parts connected by clean interfaces.
- Clarity in coding.
- Write big programs putting together building blocks.
- Programmer time is expensive; conserve it.
- Quick prototyping before polishing.
- Design for the future, because it will be here sooner than you think.





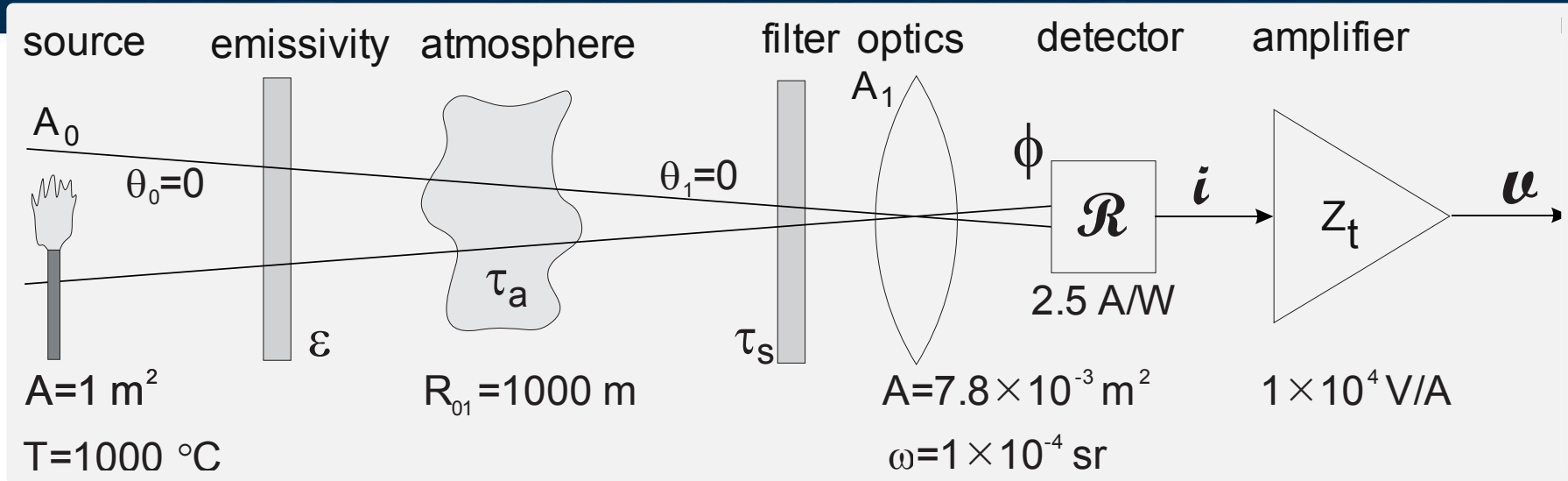
Example: Flame Detection



Detect the presence or absence of a flame,
through the atmosphere in $3\text{-}5 \text{ }\mu\text{m}$ band.



Example: Flame Detection



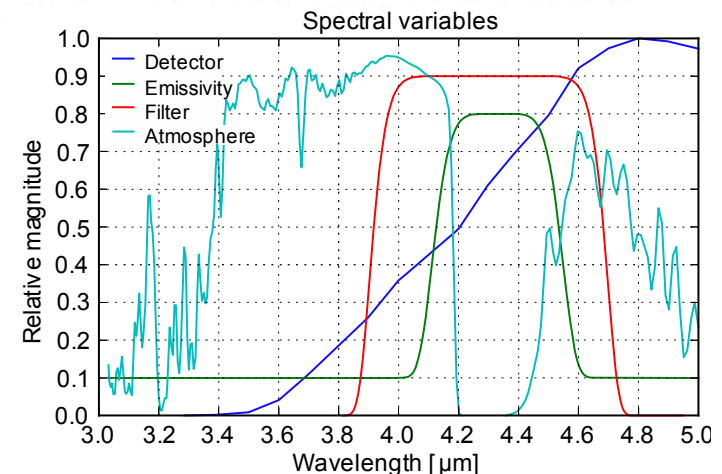
Signal from the flame:

$$v = Z_t \int_{A_0} \int_{A_1} \frac{1}{r_{01}^2} \int_0^\infty \epsilon_\lambda L_\lambda(T, A_0) \tau_{a\lambda} \tau_{s\lambda}(A_1) \mathcal{R}_\lambda d\lambda d(\cos \theta_0 A_0) d(\cos \theta_1 A_1).$$

Signal from the atmospheric radiance:

$$v = Z_t \omega_{\text{optics}} A_{\text{optics}} \int_0^\infty L_{\text{path}\lambda} \tau_{s\lambda} \mathcal{R}_\lambda d\lambda,$$

Modtran spectral transmittance & path radiance





Solution Page 1 of 2

```
#load atmospheric transmittance from file created in Modtran in wavenumbers
# the transmittance is specified in the wavenumber domain with
# 5 cm-1 intervals, but we want to work in wavelength with 2.5 cm-1
waven = numpy.arange(2000.0, 3300.0, 2.5).reshape(-1, 1)
wavel= ryutils.convertSpectralDomain(waven, type='nw')

#remove comment lines, and scale path radiance from W/cm2.sr.cm-1 to W/m2.sr.cm-1
tauA = ryfiles.loadColumnTextFile('data/path1kmflamesensor.txt',
    [1],abscissaOut=waven, comment='%')
lpathwn = ryfiles.loadColumnTextFile('data/pathspaceflamesensor.txt',
    [9],abscissaOut=waven, ordinateScale=1.0e4, comment='%')

#convert path radiance spectral density from 1/cm^-1 to 1/um, at the sample
#wavenumber points
(dum, lpathwl) = ryutils.convertSpectralDensity(waven, lpathwn, type='nw')

#load the detector file in wavelengths, and interpolate on required values
detR = ryfiles.loadColumnTextFile('data/detectorflamesensor.txt',
    [1],abscissaOut=wavel, comment='%')

#construct the flame emissivity from parameters
emis = ryutils.sfilter(wavel,center=4.33, width=0.45, exponent=6, taupass=0.8,
    taustop=0.1 )
#construct the sensor filter from parameters
sfilter = ryutils.sfilter(wavel,center=4.3, width=0.8, exponent=12,
    taupass=0.9, taustop=0.0001)
```

Only two
pages





Solution Page 2 of 2

define sensor scalar parameters

```
opticsArea=7.8e-3 # optical aperture area [m2]
opticsFOV=1.0e-4 # sensor field of view [sr]
transZ=1.0e4 # amplifier transimpedance gain [V/A]
responsivity=2.5 # detector peak responsivity =A/W]
```

define the flame properties

```
flameTemperature = 1000+273.16 # temperature in [K]
flameArea = 1 # in [m2]
distance = 1000 # [m]
fill = (flameArea /distance**2) / opticsFOV # how much of FOV is filled
fill = 1 if fill > 1 else fill # limit target solid angle to sensor FOV
```

flame get spectral radiance in W/m².sr.cm-1

```
radianceFlame = ryplanck.planck(waven, flameTemperature, type='en')\
    .reshape(-1, 1)/numpy.pi
inbandirradianceFlame = radianceFlame * detR * tauA * emis * sfilter *\
    fill * opticsFOV
totalirradianceFlame = numpy.trapz(inbandirradianceFlame.reshape(-1, 1),
    waven, axis=0)[0]
signalFlame = totalirradianceFlame *transZ*responsivity *opticsArea
```

now do path

```
inbandirradiancePath = lpathwn * detR * sfilter * opticsFOV
totalirradiancePath = numpy.trapz(inbandirradiancePath.reshape(-1, 1),waven, axis=0)[0]
signalPath = totalirradiancePath * transZ*responsivity *opticsArea
```

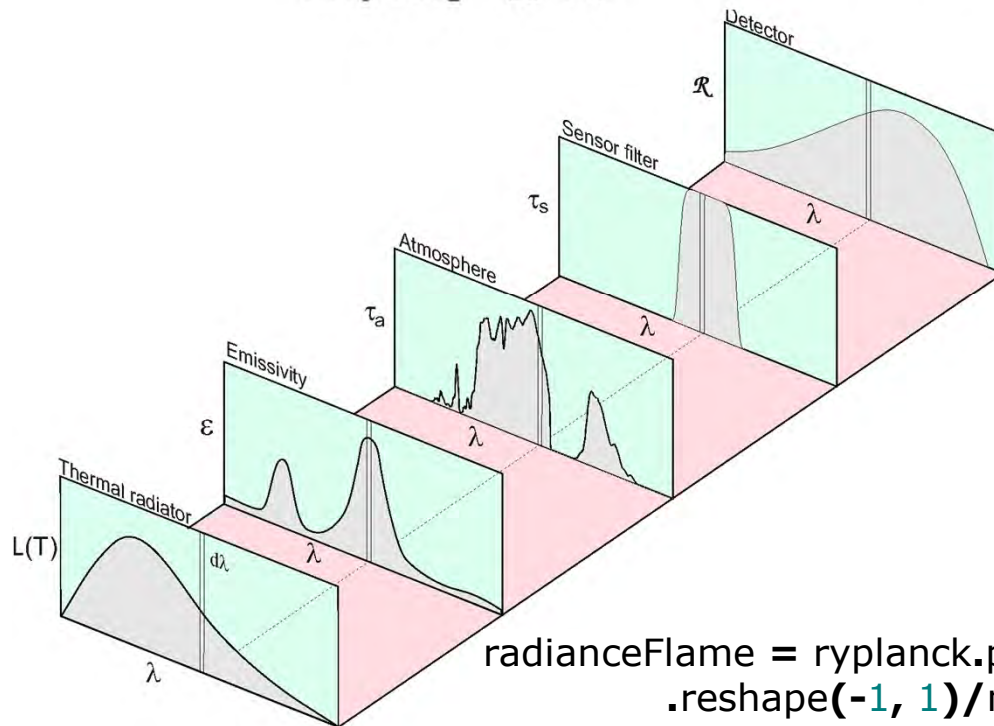
The code is:

- Compact
- Clean
- Intuitive



Multi-spectral Calculations

$$v = Z_t \int_{A_0} \int_{A_1} \frac{1}{r_{01}^2} \int_0^\infty \epsilon_\lambda L_\lambda(T, A_0) \tau_{a\lambda} \tau_{s\lambda}(A_1) \mathcal{R}_\lambda d\lambda d(\cos \theta_0 A_0) d(\cos \theta_1 A_1).$$



- Spectral variables are vectors.
- Calculate irradiance at each wavelength 'bin' by vector element-wise multiplication.
- Add flux at all wavelengths.

```
radianceFlame = ryplanck.planck(waven, flameTemperature, type='en')\
    .reshape(-1, 1)/numpy.pi
inbandirradianceFlame = radianceFlame* emis* tauA * sfilter * detR *\
    opticsFOV
totalirradianceFlame = numpy.trapz(inbandirradianceFlame.reshape(-1, 1),
    waven, axis=0)[0]
signalFlame = totalirradianceFlame *transZ*responsivity *opticsArea
```




Planck's Law

$$M_{e\lambda}(T) = \frac{2\pi hc^2}{\lambda^5 \left(e^{\frac{hc}{\lambda kT}} - 1 \right)}$$

Pyradi provides functions for Planck Law emittance, as well as Planck Law temperature derivative in terms of: [W] or [q/s] and [μm , cm^{-1} , Hz].

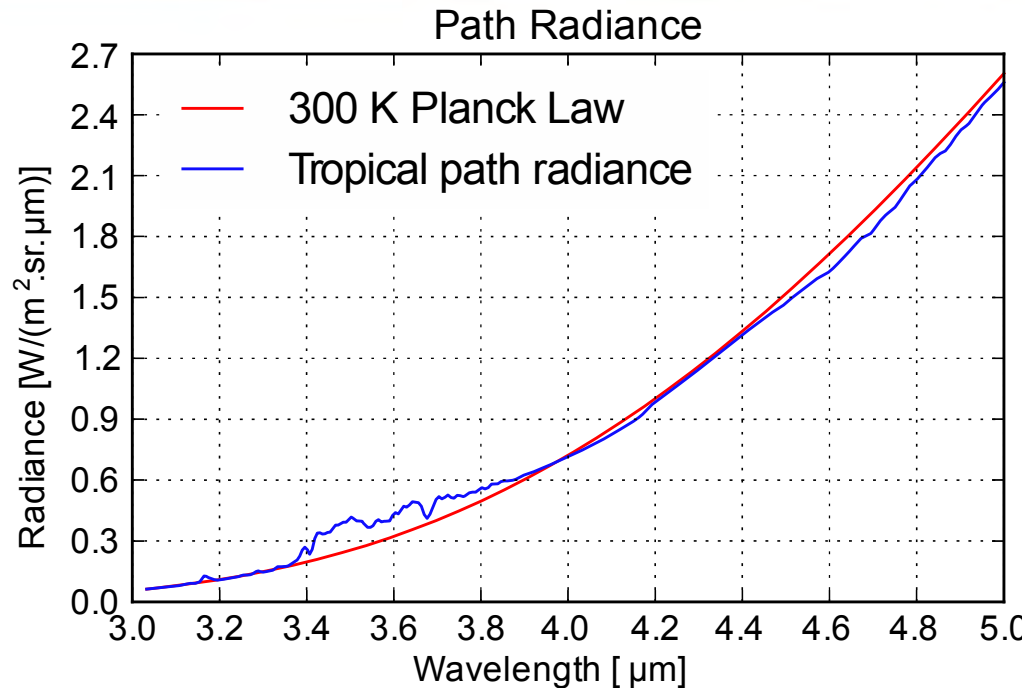


```
Mel = planck(wl, temperature[0], type='el') # [W/(m$^2$. $\mu$m)]  
Mql = planck(wl, temperature[0], type='ql') # [q/(s.m$^2$. $\mu$m)]  
Men = planck(n, temperature[0], type='en') # [W/(m$^2$.cm$^{-1}$)]
```

```
Mqn = dplanck(n, temperature[0], type='qn') # [q/(s.m$^2$.cm$^{-1}$.K)]  
Mef = dplanck(f, temperature[0], type='ef') # [W/(m$^2$.Hz.K)]  
Mqf = dplanck(f, temperature[0], type='qf') # [q/(s.m$^2$.Hz.K)]
```



Visualising Path Radiance



Flame sensor path:

- Horizontal path to space.
- Notice correspondence.
- Visualisation important!

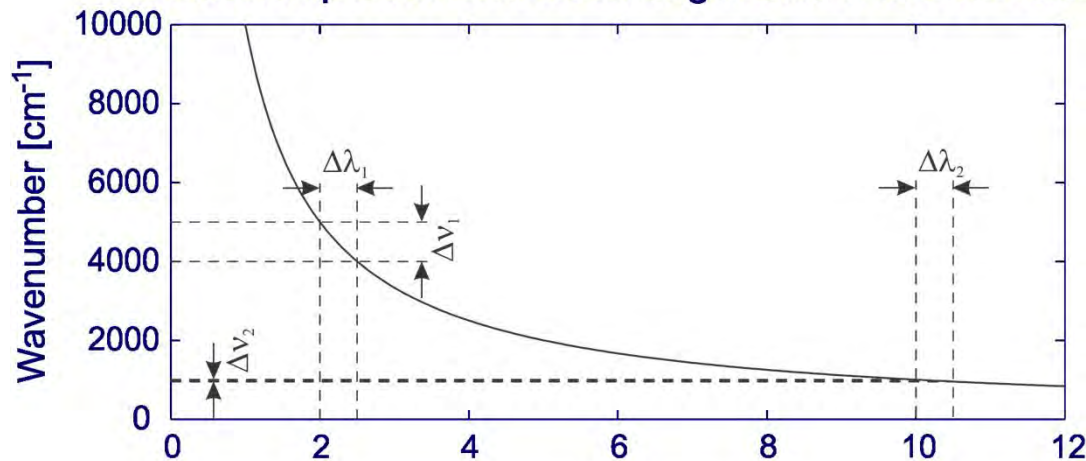
#check path radiance against Planck's Law for atmo temperature

```
LbbTropical = ryplanck.planck(wavel, 273+27, type='el').reshape(-1, 1)/numpy.pi
plot1.plot(2, wavel, LbbTropical, plotCol=['r'], label=['300 K Planck Law'])
plot1.plot(2, wavel, lpathwl, plotCol=['b'], label=['Tropical path radiance'])
currentP = plot1.getSubPlot(2)
currentP.set_xlabel('Wavelength [ $\mu\text{m}$ ']
currentP.set_ylabel('Radiance [ $\text{W}/(\text{m}^2\text{sr}.\mu\text{m})$ '])
currentP.set_title('Path Radiance')
```

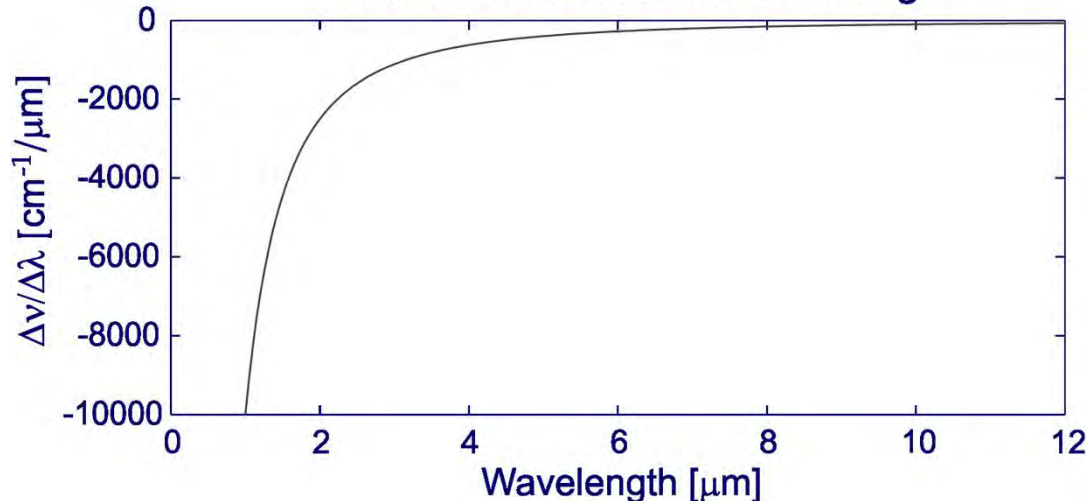


Wavelength and Wavenumber

Relationship between wavelength and wavenumber



$\Delta\nu/\Delta\lambda$ as function of wavelength



Wavenumber is given by

$$\nu = 10^4 / \lambda$$

where λ is in units of μm .

The conversion of a spectral quantity requires the derivative, as follows:

$$d\nu = \frac{-10^4}{\lambda^2} d\lambda = \frac{\nu^2}{-10^4} d\lambda$$

$$dL_\nu = dL_\lambda \frac{\lambda^2}{10^4} = dL_\lambda \frac{10^4}{\nu^2}$$

Pyradi converts between spectral variables:

$[\mu\text{m}, \text{cm}^{-1}, \text{Hz}]$,

as well as spectral densities:

$[\text{W}/(\text{m}^2 \cdot \mu\text{m})]$,

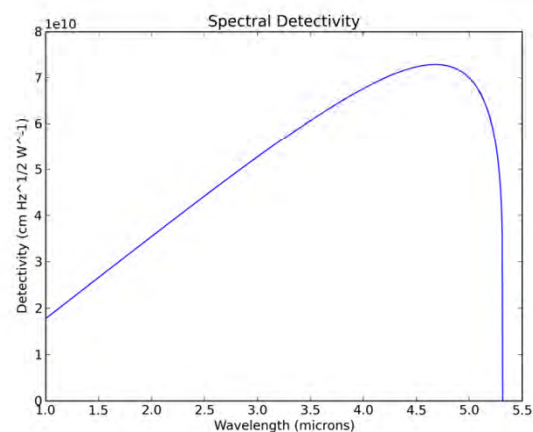
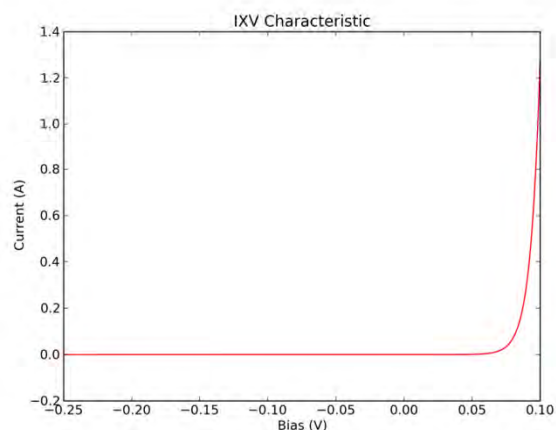
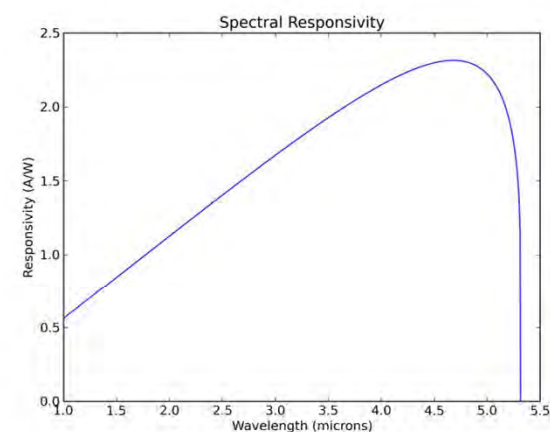
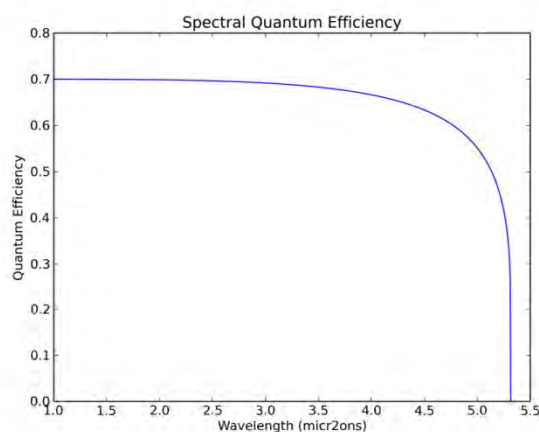
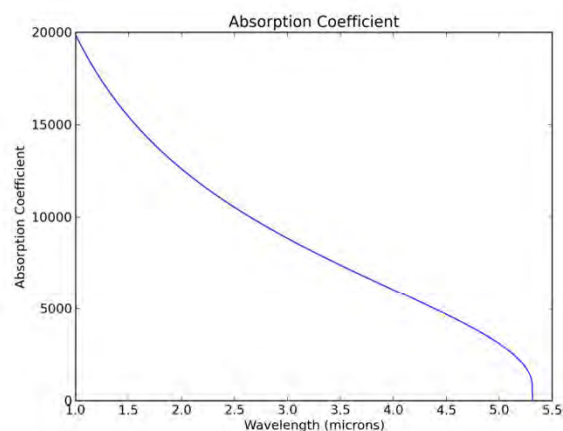
$[\text{W}/(\text{m}^2 \cdot \text{cm}^{-1})]$,

$[\text{W}/(\text{m}^2 \cdot \text{Hz})]$.



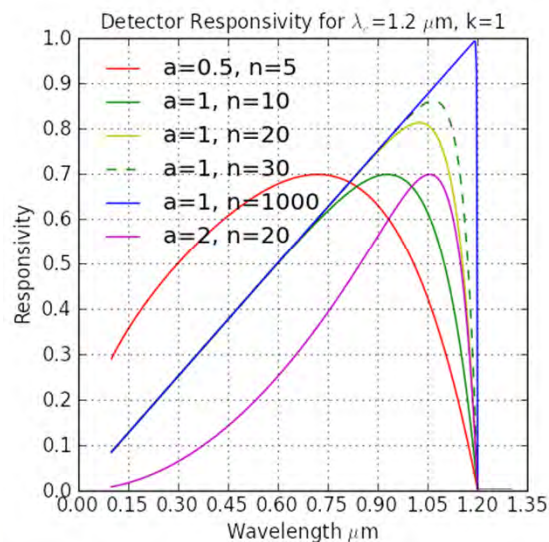
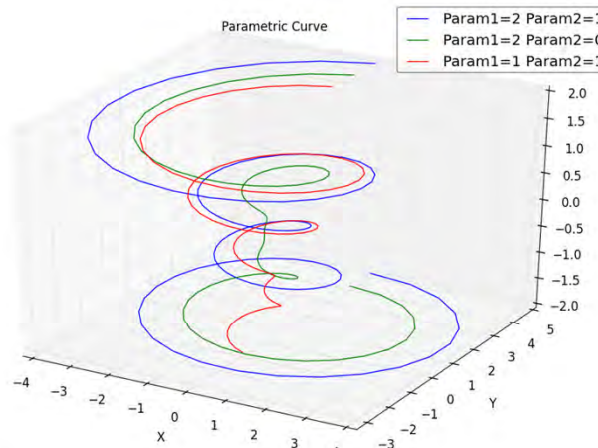
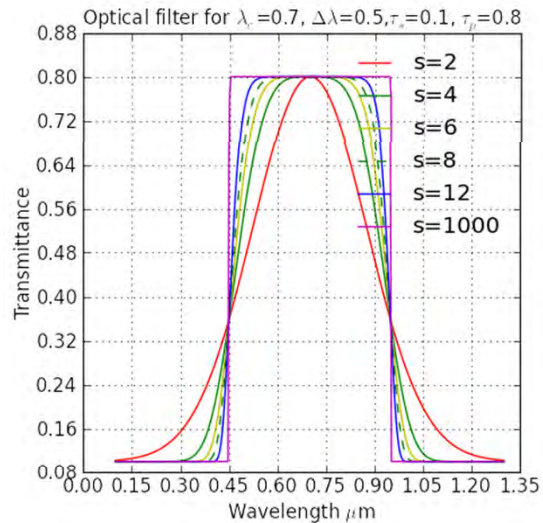
Detector Modelling

Detector model, based on first principles physics.
Example shown is for InSb detector.



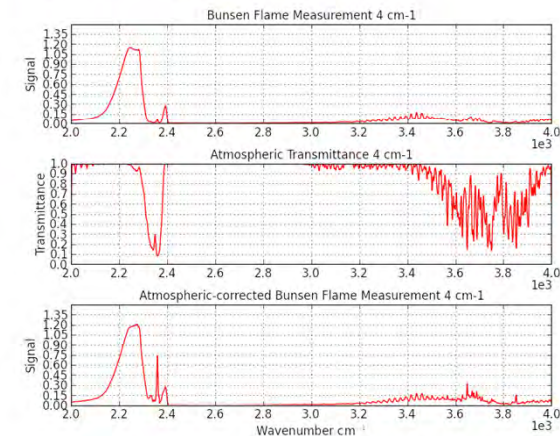
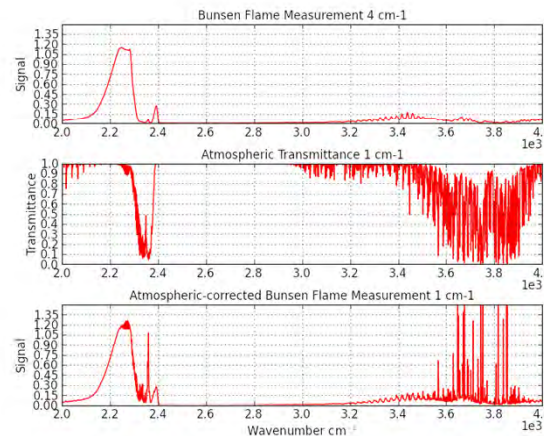


Utility Functions



General purpose models:

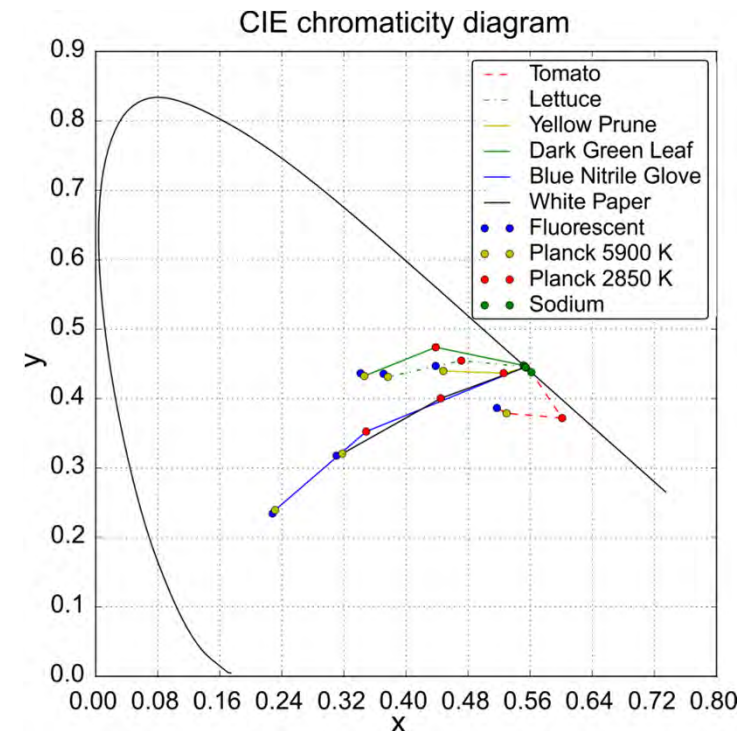
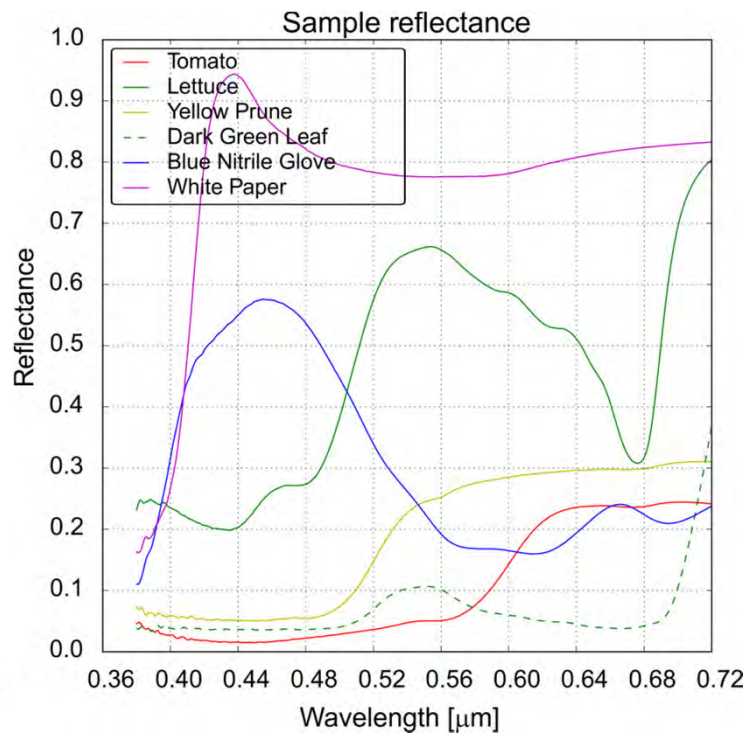
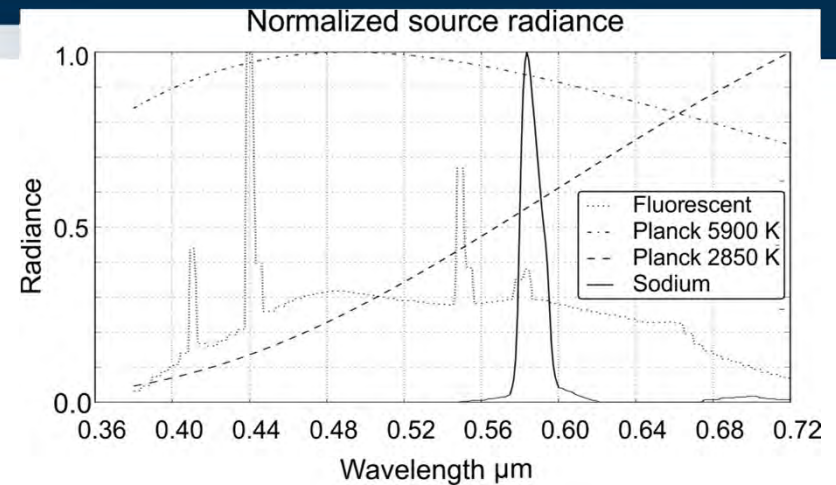
- Spectral filter shapes.
- Photon detector shapes.
- Spectral convolution.
- Effective value normalisation.
- Various 2-D & 3-D plotting functions.





Chromaticity Coordinates

Basic functionality for CIE-1931 coordinate calculation for spectral source radiance and sample reflectance

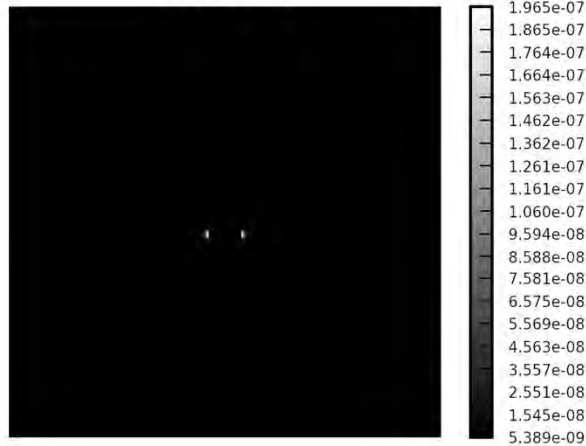




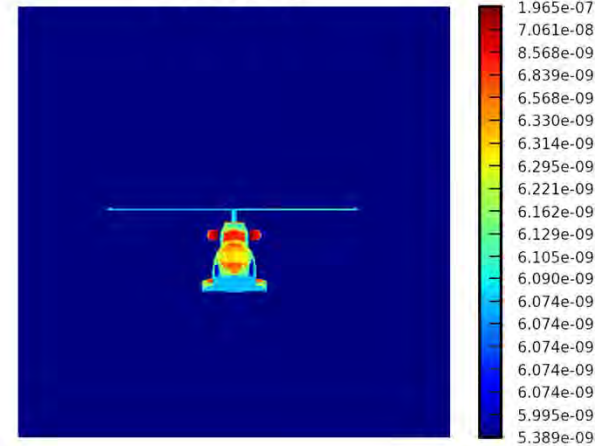
Histogram Equalisation

Image Irradiance (Square Root then Equalised) [W/m^2]

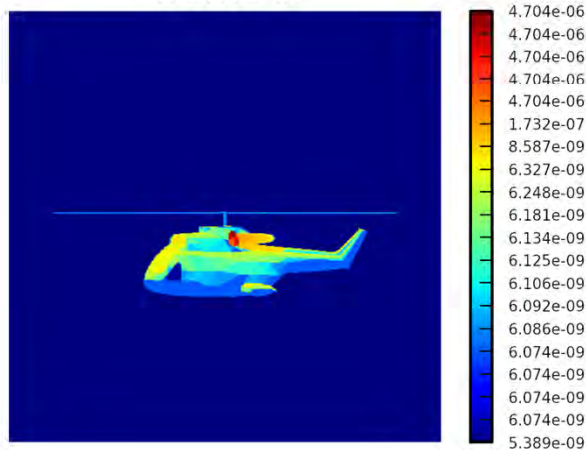
frame 0



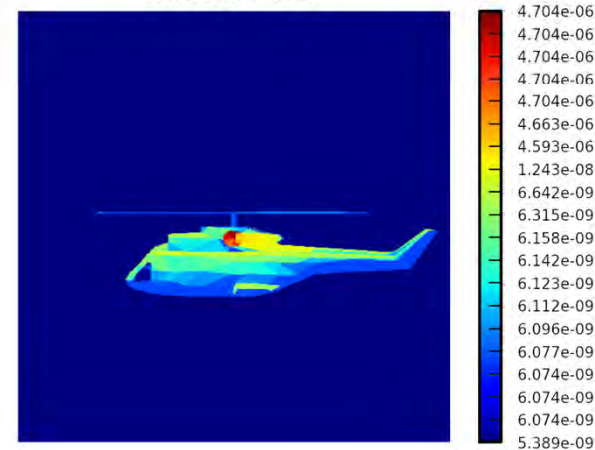
frame 0



frame 9



frame 19





Reading Data Files

- Binary raw image files: select frames, size & data type.

```
imagefile = 'data/sample.ulong'  
rows = 100  
cols = 100  
vartype = numpy.uint32  
framesToLoad = [1, 3, 5, 7]  
frames, img = readRawFrames(imagefile, rows, cols, vartype, framesToLoad)
```

- FLIR Inc (CEDIP) *.ptw image files: select frames.

```
ptwfile = 'data/PyradiSampleMWIR.ptw'  
header = readPTWHeader(ptwfile)  
#loading sequence of frames  
framesToLoad = [3,4,10]  
data = getPTWFrames (header, framesToLoad)
```

(Instrument calibration scaling still to follow).



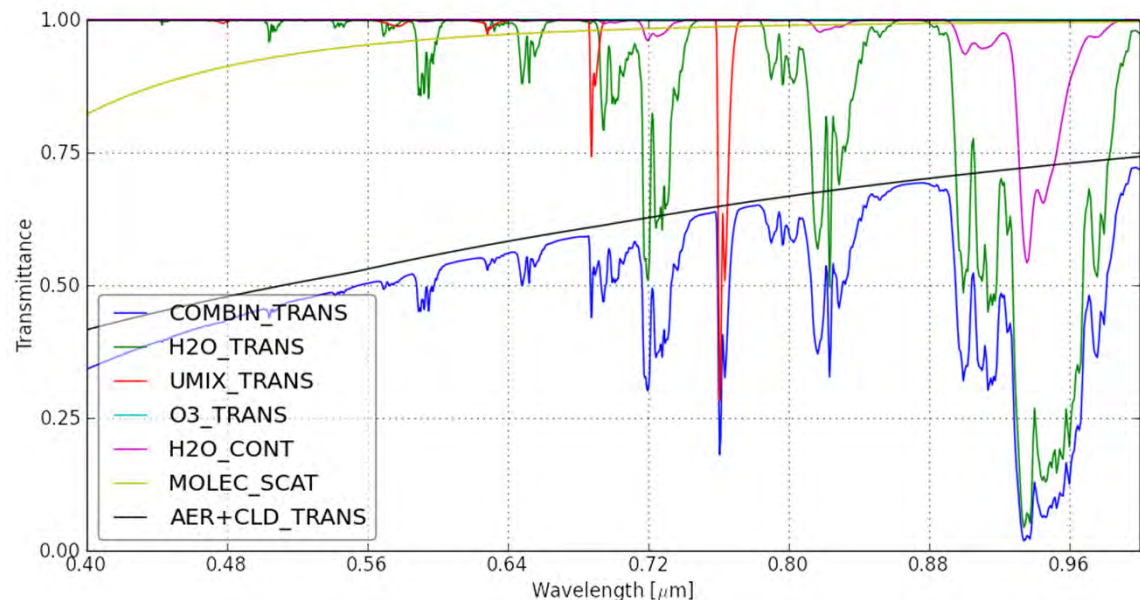
Read Modtran tape7 Files

```
colSelect = ['FREQ_CM-1', 'COMBIN_TRANS', 'H2O_TRANS', 'UMIX_TRANS', \
            'O3_TRANS', 'H2O_CONT', 'MOLEC_SCAT', 'AER+CLD_TRANS']
tape7= loadtape7("data/tape7VISNIR5kmTrop23Vis", colSelect )

wavelen = ryutils.convertSpectralDomain(tape7[:,0], type='nl')
mT = ryplot.Plotter(1, 1, 1, "Modtran Tropical, 23 km Visibility (Rural)" \
    + ", 5 km Path Length", figsize=(12,6))
mT.plot(1, wavelen, tape7[:,1:], "", "Wavelength [ $\mu$ m]", "Transmittance",
    label=colSelect[1:], legendAlpha=0.5, pltaxis=[0.4, 1, 0, 1],
    maxNX=10, maxNY=4, powerLimits = [-4, 4, -5, 5])
mT.saveFig('ModtranPlot.png')
```

- Select columns by header title.
- Supports all four IEMSCT file variants.

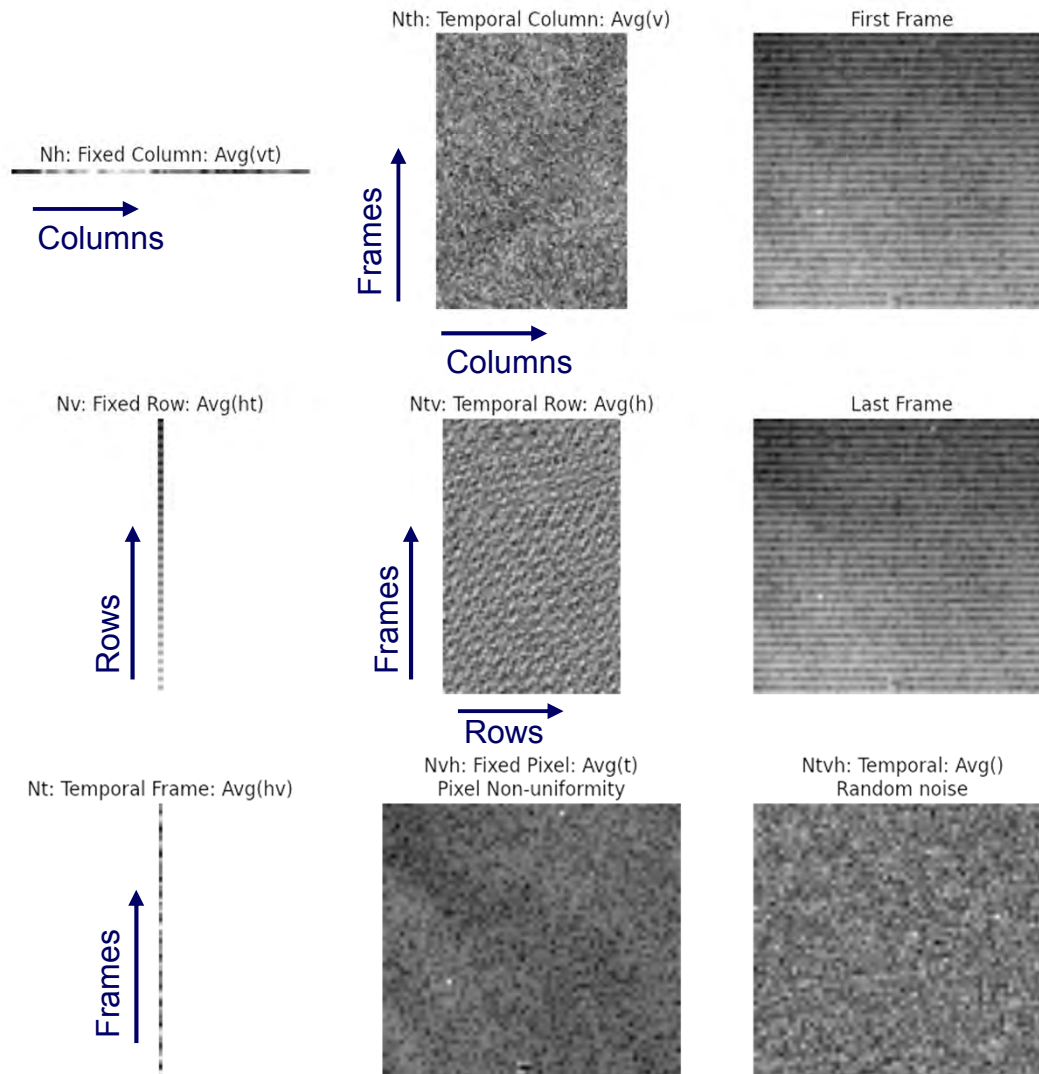
Modtran Tropical, 23 km Visibility (Rural), 5 km Path Length





3-D Noise Analysis

PTW Image Sequence 3-D Noise Components



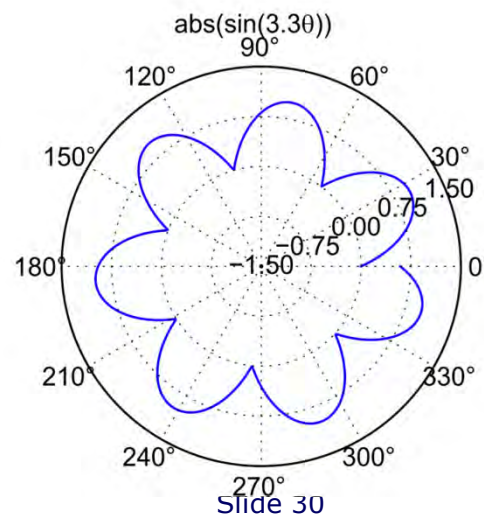
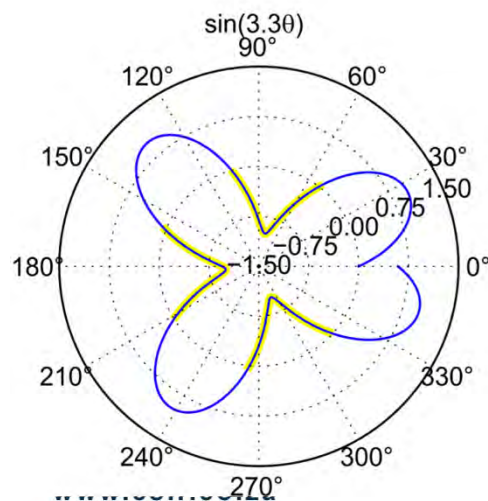
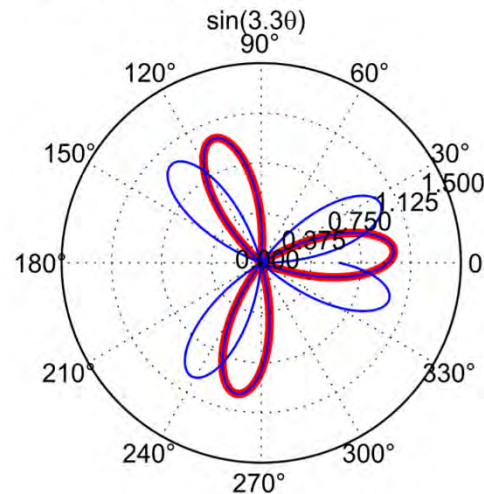
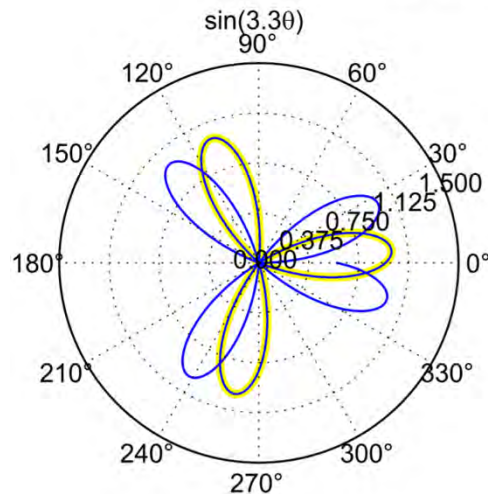
3-D noise analysis of image sequences.

- Averages and filters along frames, rows and columns.
- Isolates individual noise causes.



Polar Plots

Polar Plots of Negative Data



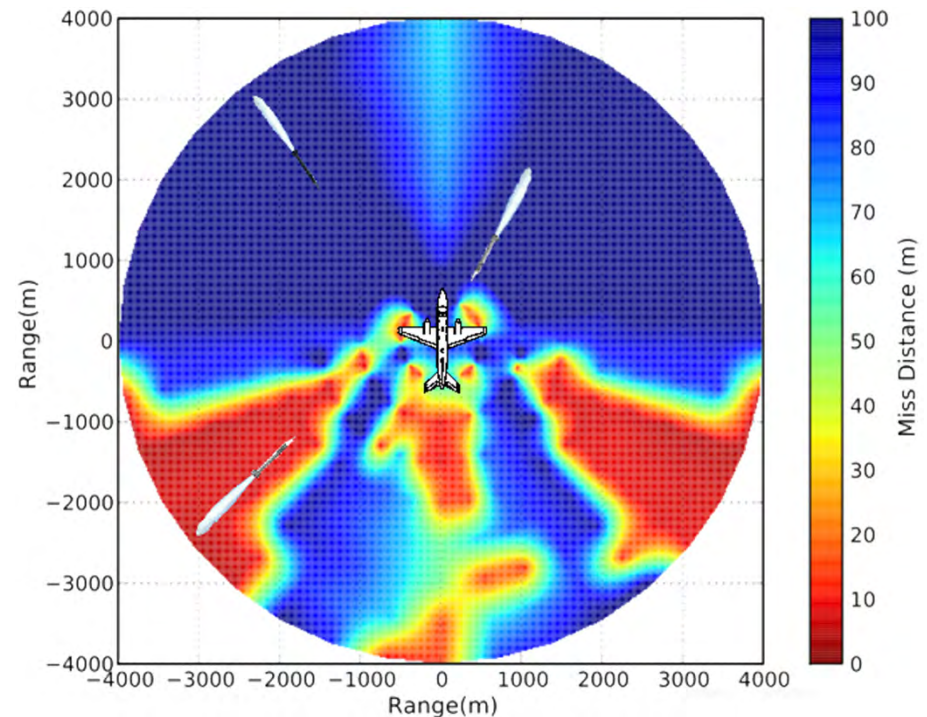
- -1 is a π phase shift: plot on 'other' side.
- Confuse direction with magnitude!
- Two options:
 - Highlight negative.
 - Zero not in centre.



Aircraft Vulnerability

Missiles launched from red position are kills.

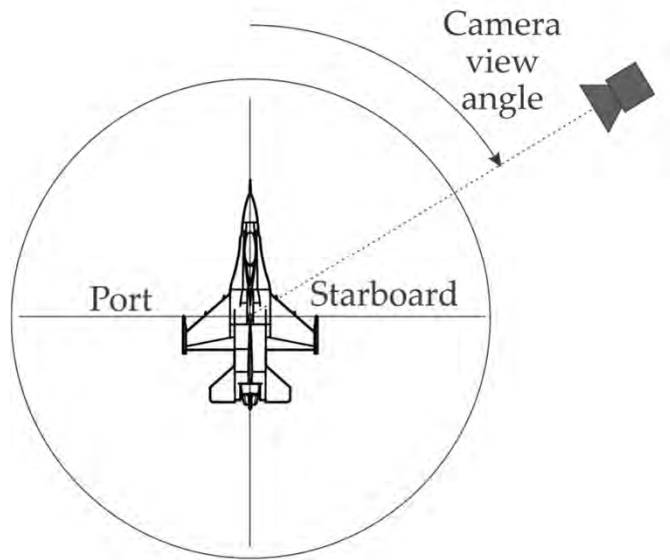
- Why are countermeasures not effective?
- Investigate Jam / Signal ratios.
- Investigate geometry: aircraft fuselage obscuration.
- Investigate angular 3-D signatures.



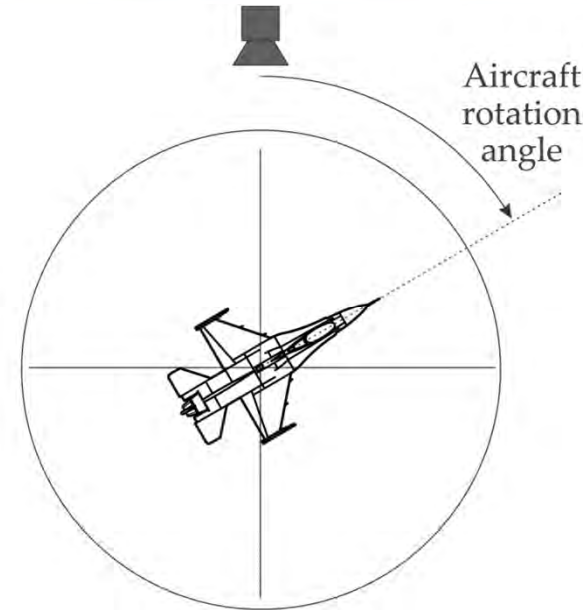
Welcome to the haunted house of 3-D signatures!



Two 3-D Scenarios



- Stationary target.
- Orbiting camera.
- Varying background.
- Stable sun position.

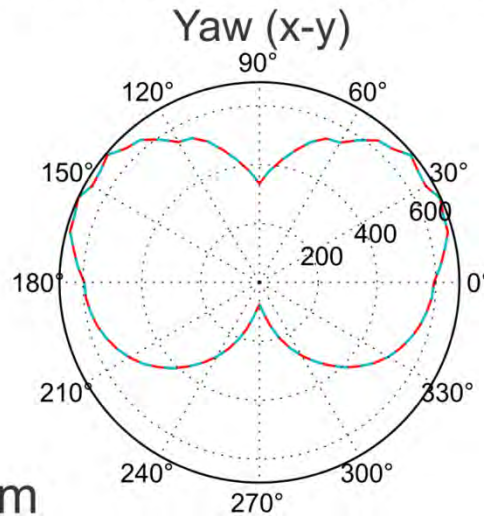
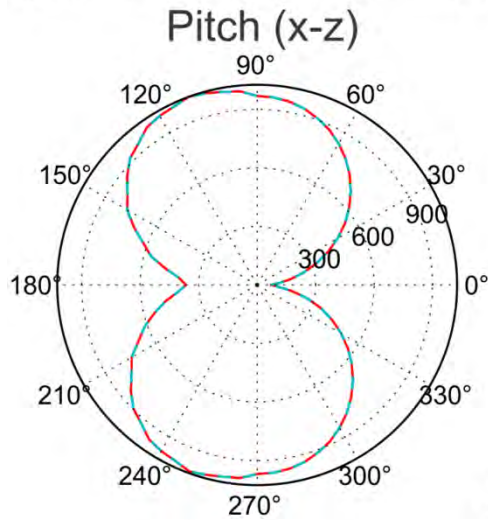


- Stationary camera.
- Rotating target.
- Varying sun position.
- Stable background.

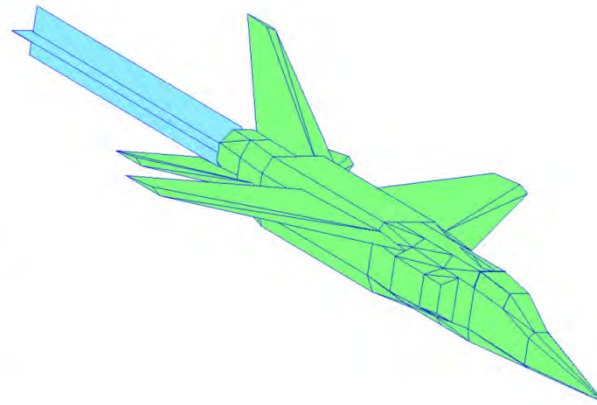
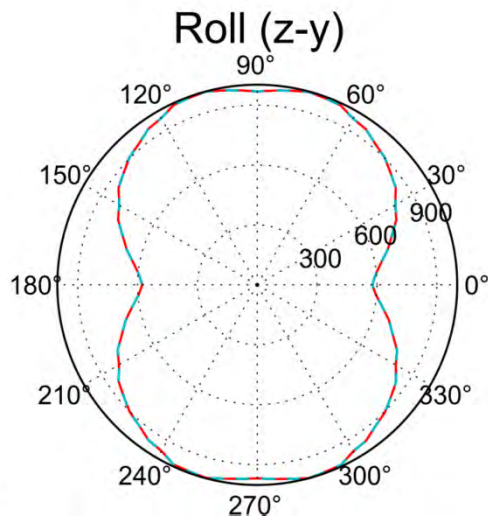


Intensity Polar Plots 8-12 μm

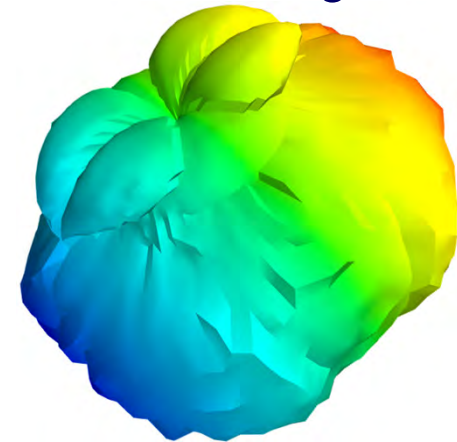
tp01a/m/n/o Fighter, 100 m ASL, 100% dry, 300 m/s, MLS, Intensity [W/sr]



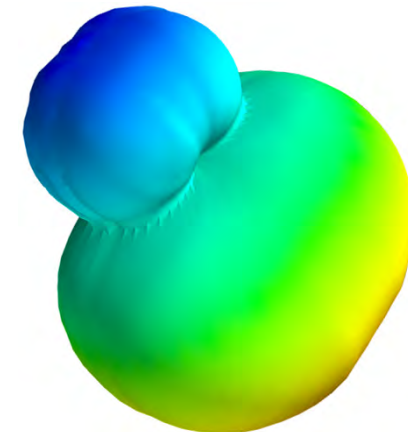
8-12 μm



Rotating target:
constant background



Orbiting sensor:
varying background

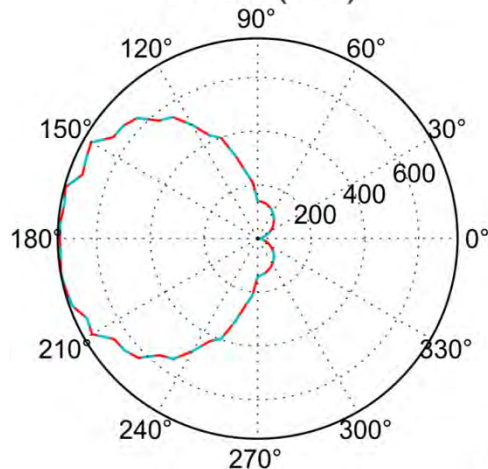




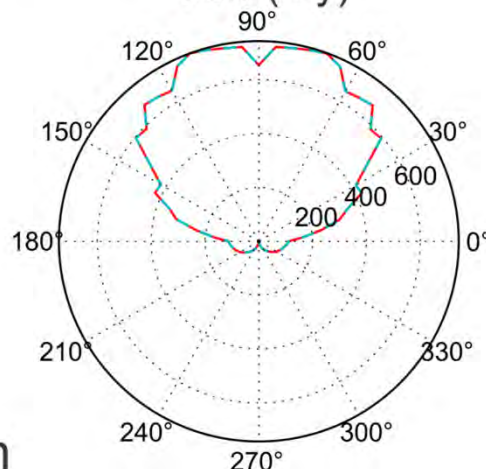
Intensity Polar Plots 3-5 μm

tp01a/m/n/o Fighter, 100 m ASL, 100% dry, 300 m/s, MLS, Intensity [W/sr]

Pitch (x-z)

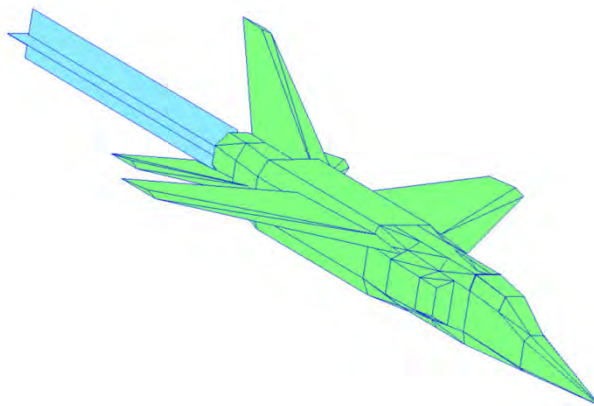
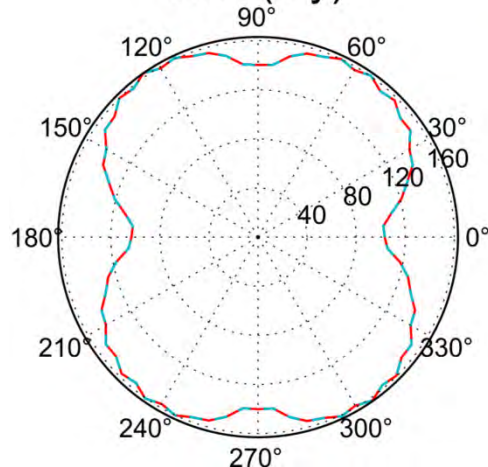


Yaw (x-y)

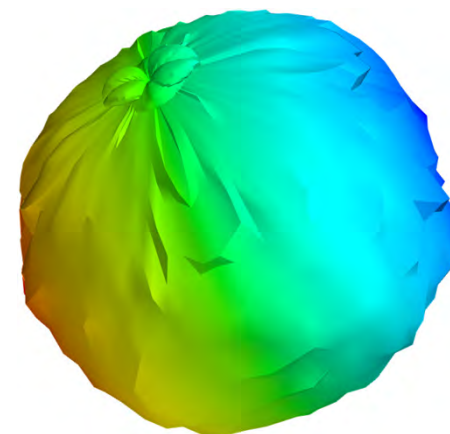


3-5 μm

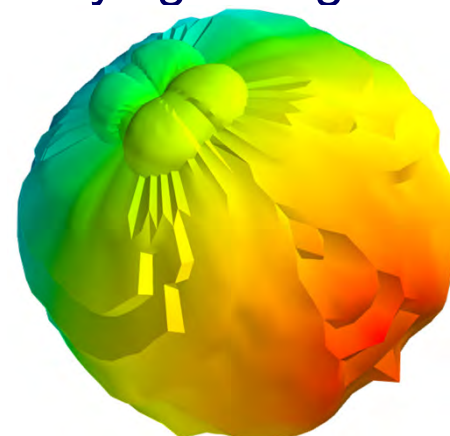
Roll (z-y)



Rotating target:
constant background



Orbiting sensor:
varying background

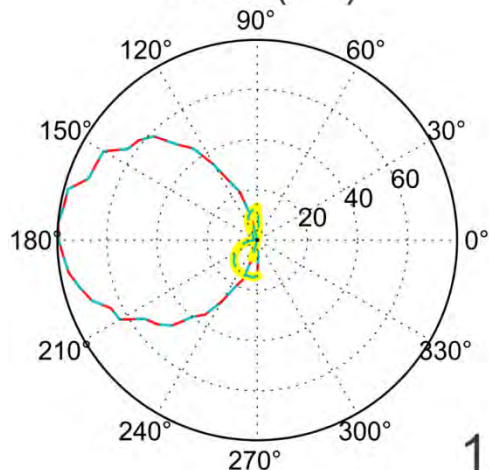




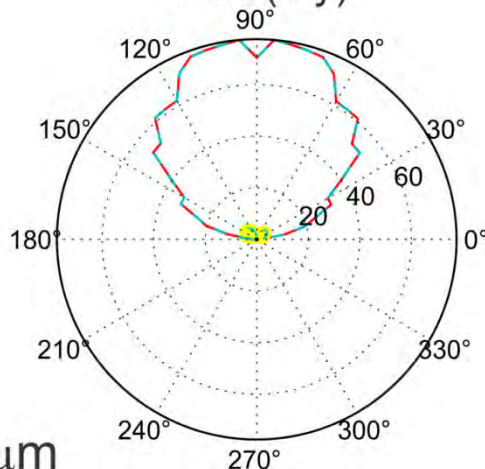
Intensity Polar Plots 1.5-2.5 μm

tp01a/m/n/o Fighter, 100 m ASL, 100% dry, 300 m/s, MLS, Intensity [W/sr]

Pitch (x-z)

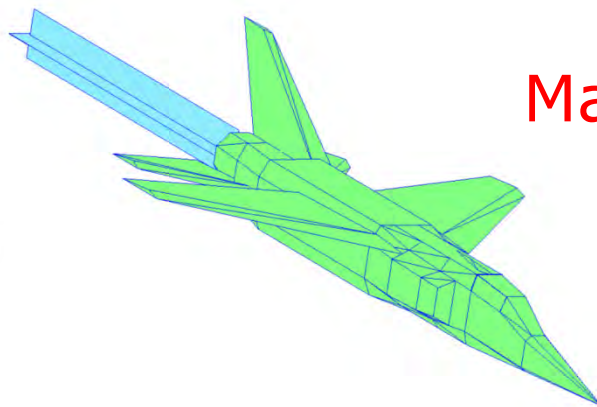
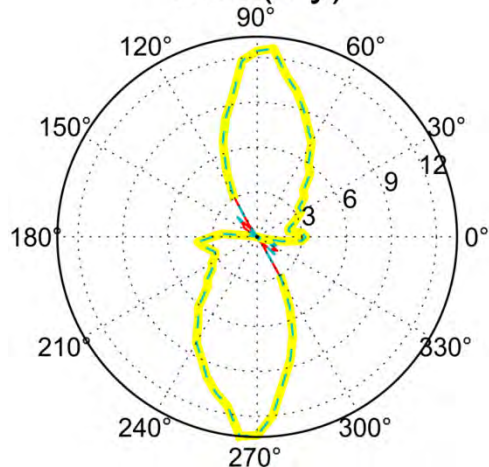


Yaw (x-y)



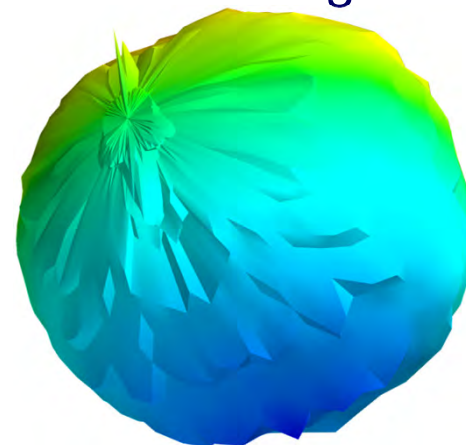
1.5-2.5 μm

Roll (z-y)

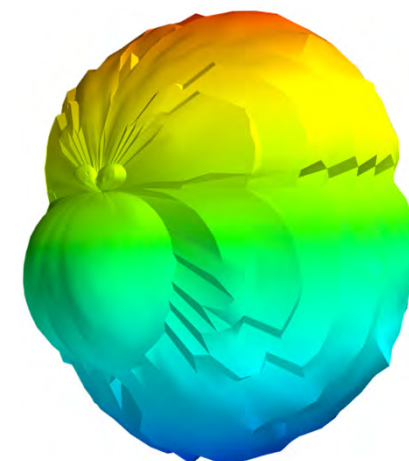


Mario!

Rotating target:
constant background



Orbiting sensor:
varying background

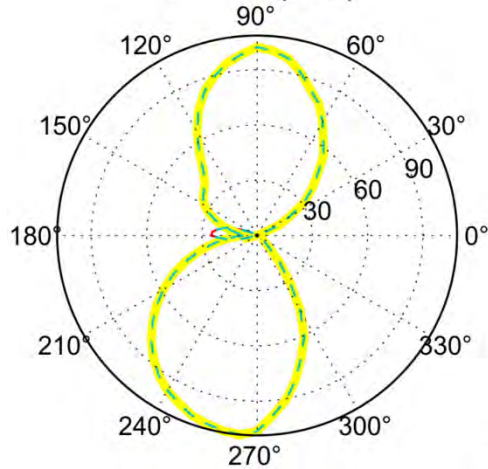




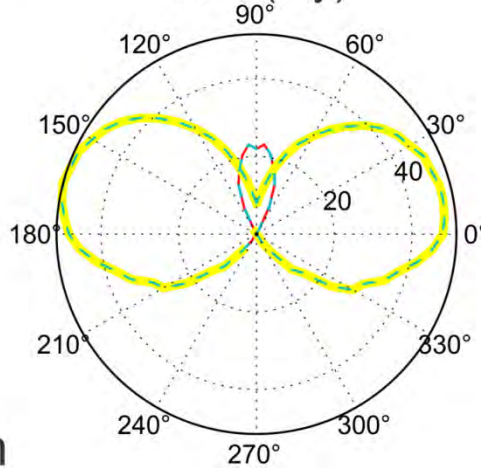
Intensity Polar Plots 1-2 μm

tp01a/m/n/o Fighter, 100 m ASL, 100% dry, 300 m/s, MLS, Intensity [W/sr]

Pitch (x-z)

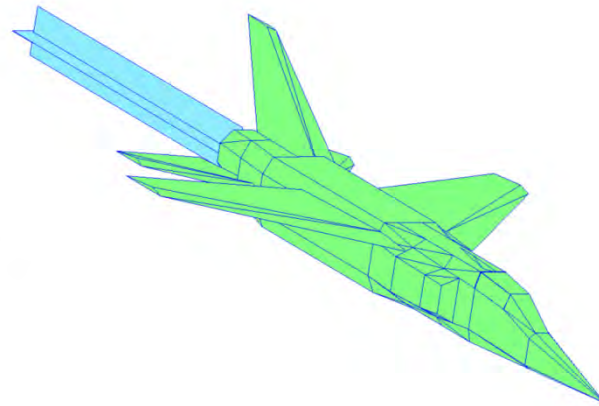
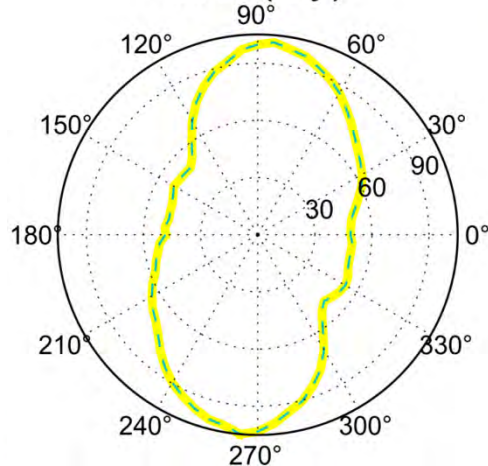


Yaw (x-y)

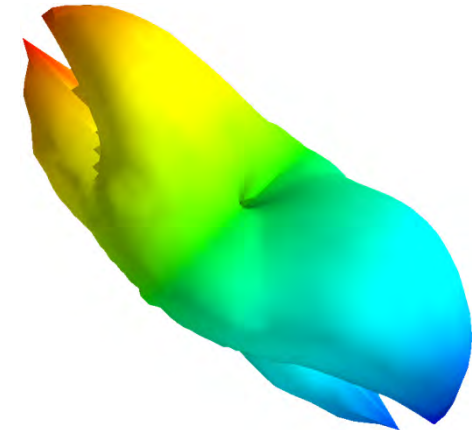


1-2 μm

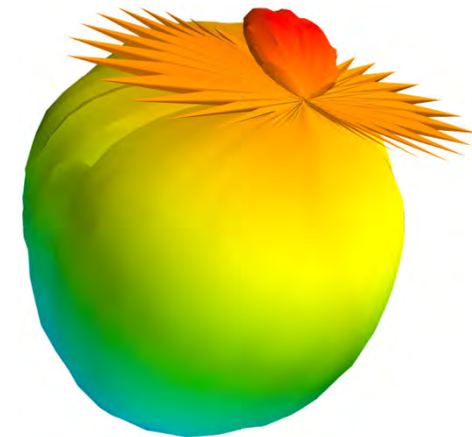
Roll (z-y)



Rotating target:
constant background



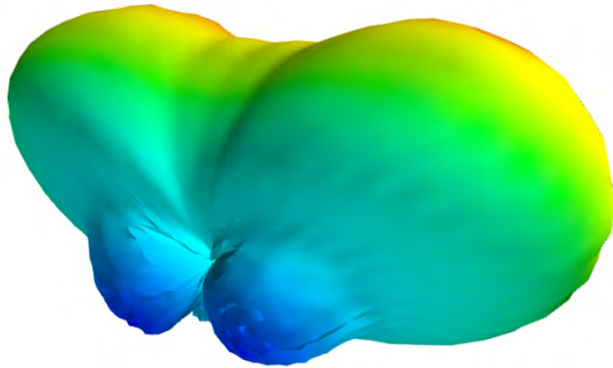
Orbiting sensor:
varying background



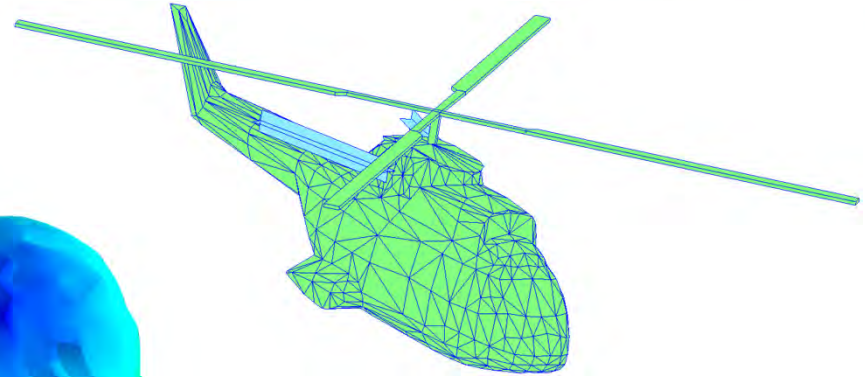
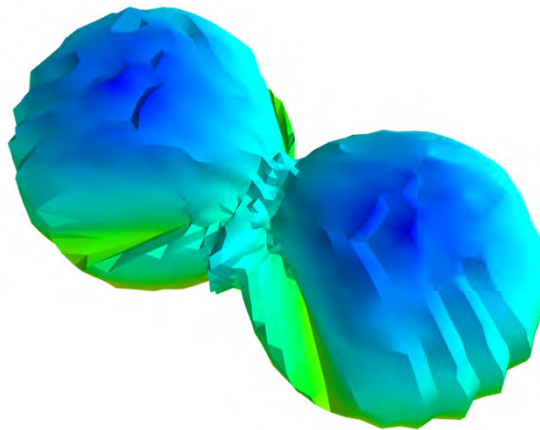


Puma Helicopter Model

8-12 μm

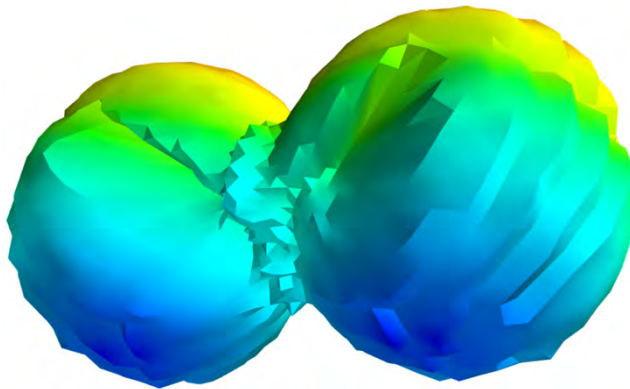


3-5 μm

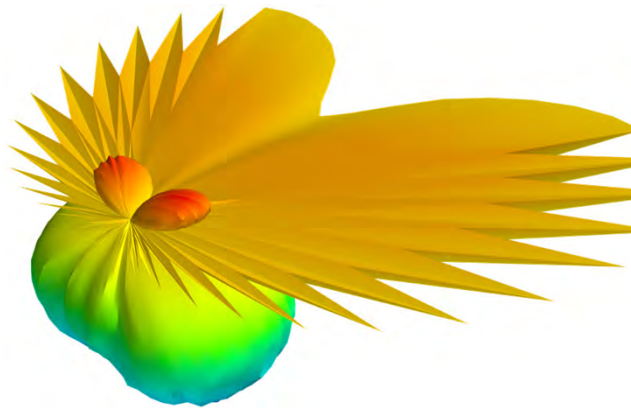


Orbiting sensor:
varying background

1.5-2.5 μm



1-2 μm



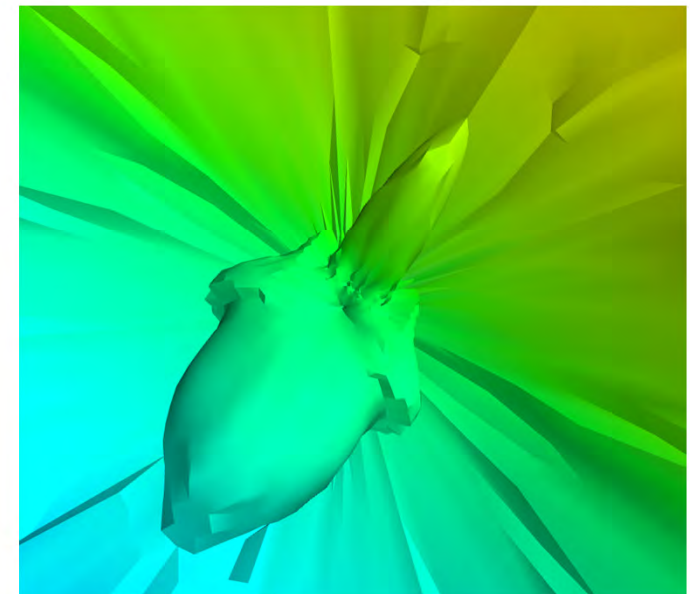
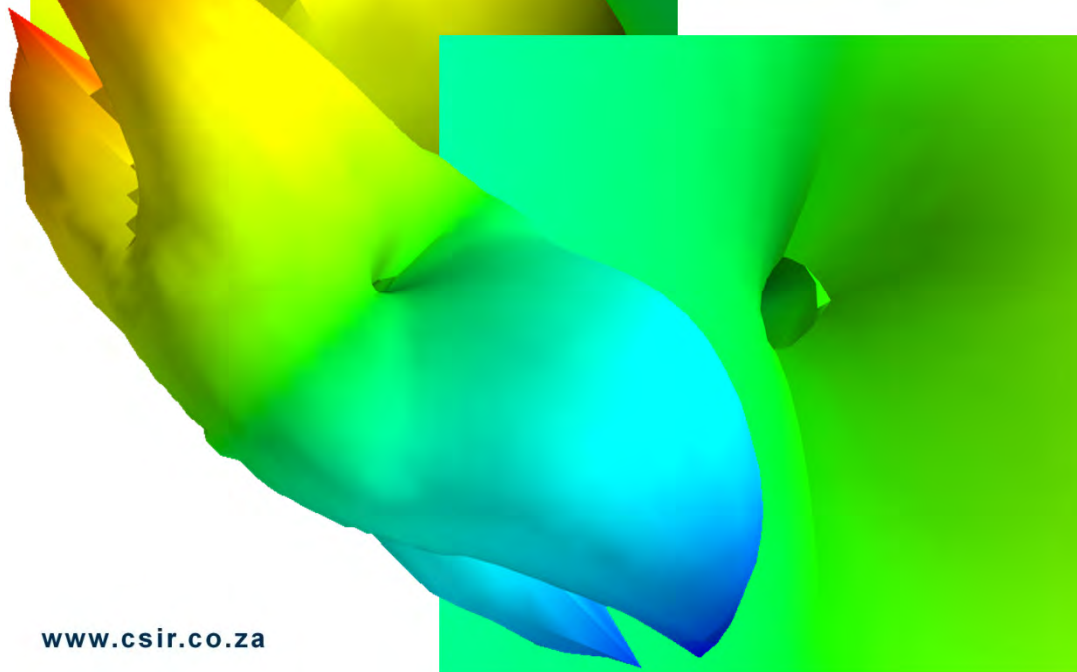
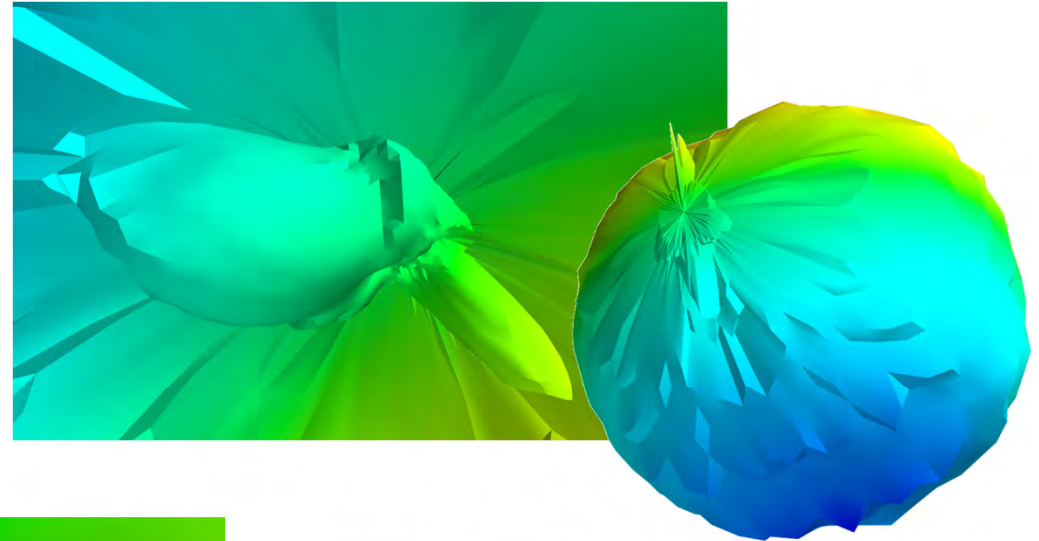
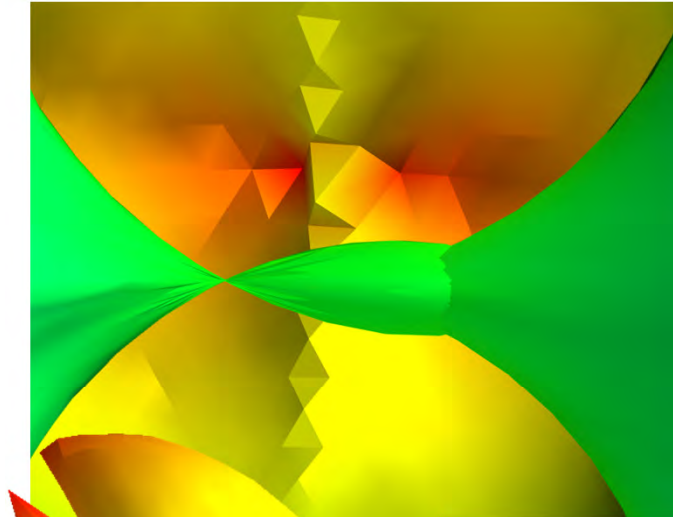


Going Inside!

You can actually go inside these shapes!



Going Inside!





Requirements

- Moderate hardware requirements.
- Supported on Linux & Windows.
- Software requirements (all are free):
 - Python 2.7.3.
 - Numpy 1.6.1.
 - Scipy 0.10.1.
 - Matplotlib 1.1.
 - Mayavi 4.1.



Authors, Sponsors & Users

- Collaborative effort between
 - CSIR, South Africa.
 - Denel Dynamics, South Africa.
 - Instituto Tecnológico de Aeronáutica, Brazil.
- Used in CSIR and Denel for R&D, simulation model development and analysis.
- Targeting academic environment.
- Working towards the tool of choice in infrared laboratories.





Open Source Considerations

Pyradi is available as open source software.

Mozilla Public License 1.1:

- Authors retain ownership of their respective contributions.
- Modifications by users must be released under compatible terms.

Potential users are encouraged to join in, and commit changes, updates and new contributions.



Hosted at Google Code: pyradi

<http://code.google.com/p/pyradi>

http://pyradi.googlecode.com/svn/trunk/doc/_build/html/index.html

The image shows two overlapping browser windows. The background window displays the Google Code project page for 'pyradi'. The page header includes the pyradi logo and the text 'Python toolkit to do optical radiometry calculations'. Navigation links for 'Project Home', 'Issues', 'Source', and 'Administer' are visible. The 'Project Information' sidebar on the left lists 'Stared by 0 users', 'Code license: Mozilla Public License 1.1', 'Labels: Radiometry, Flux, Optical, Python, Toolkit', and 'Members: neliswillers (8 committers)'. The main content area has a 'Tip' about discussing project duties, followed by an 'Objective' section describing the toolkit's purpose and a list of six features: 1. Models of physical concepts (e.g. Planck's Law), 2. Mathematical operations (e.g. spectral integrals, spatial integration), 3. Data manipulation (e.g. file input/output, interpolation, spectral fitting), 4. Detector modelling, 5. Graphical depiction (2-D and 3-D graphs), and 6. All these in an interactive development environment. Below this is a 'Prerequisites' section mentioning 'Matplotlib (1.1)' and 'Mayavi (4.1)', a 'Status' section noting the project is in 'early alpha', and a 'Documentation' section with a 'here' link. The foreground window shows the 'pyradi documentation' page, which has a dark blue header with 'next | modules | index' links. It features a 'Table Of Contents' sidebar with links to 'Welcome to pyradi's documentation!', 'Contents:', 'Examples of code use', and 'Indices and tables'. The main content area is titled 'Welcome to pyradi's documentation!' and includes the pyradi logo, a description of the toolkit, its availability at <http://code.google.com/p/pyradi>, documentation at http://pyradi.googlecode.com/svn/trunk/doc/_build/html/index.html, and a Google Group at <http://groups.google.com/group/pyradi-dev>. A 'Contents:' section lists: Planck and thermal radiation (with sub-items Overview and Module functions), File reading/writing utility (with sub-items Overview and Module functions), and Plotting utility.



Closing

- Pyradi only recently started – already paying dividends.
- Visualisation proved to be most valuable.
- All are welcome to use and contribute.