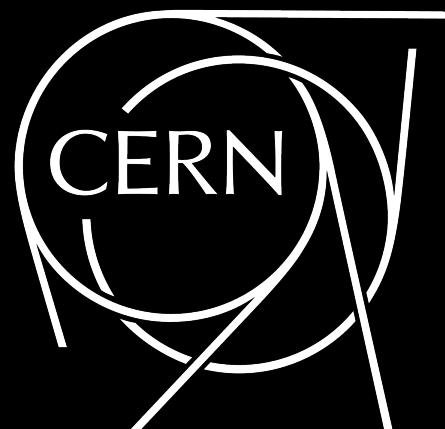
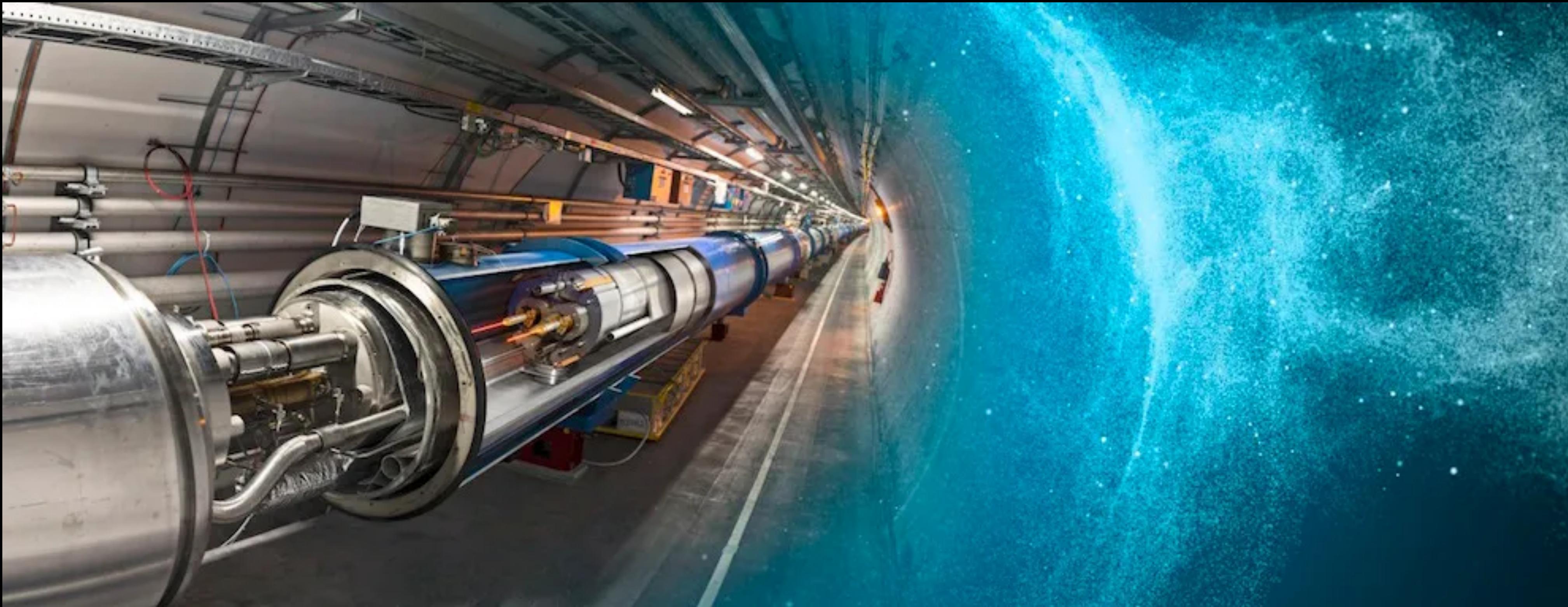


Estudio de la materia en condiciones extremas: del cosmos a los aceleradores de partículas



Alba Soto Ontoso
Seminario de Invitados – Máster en Física
Granada, 8-9th January, 2024

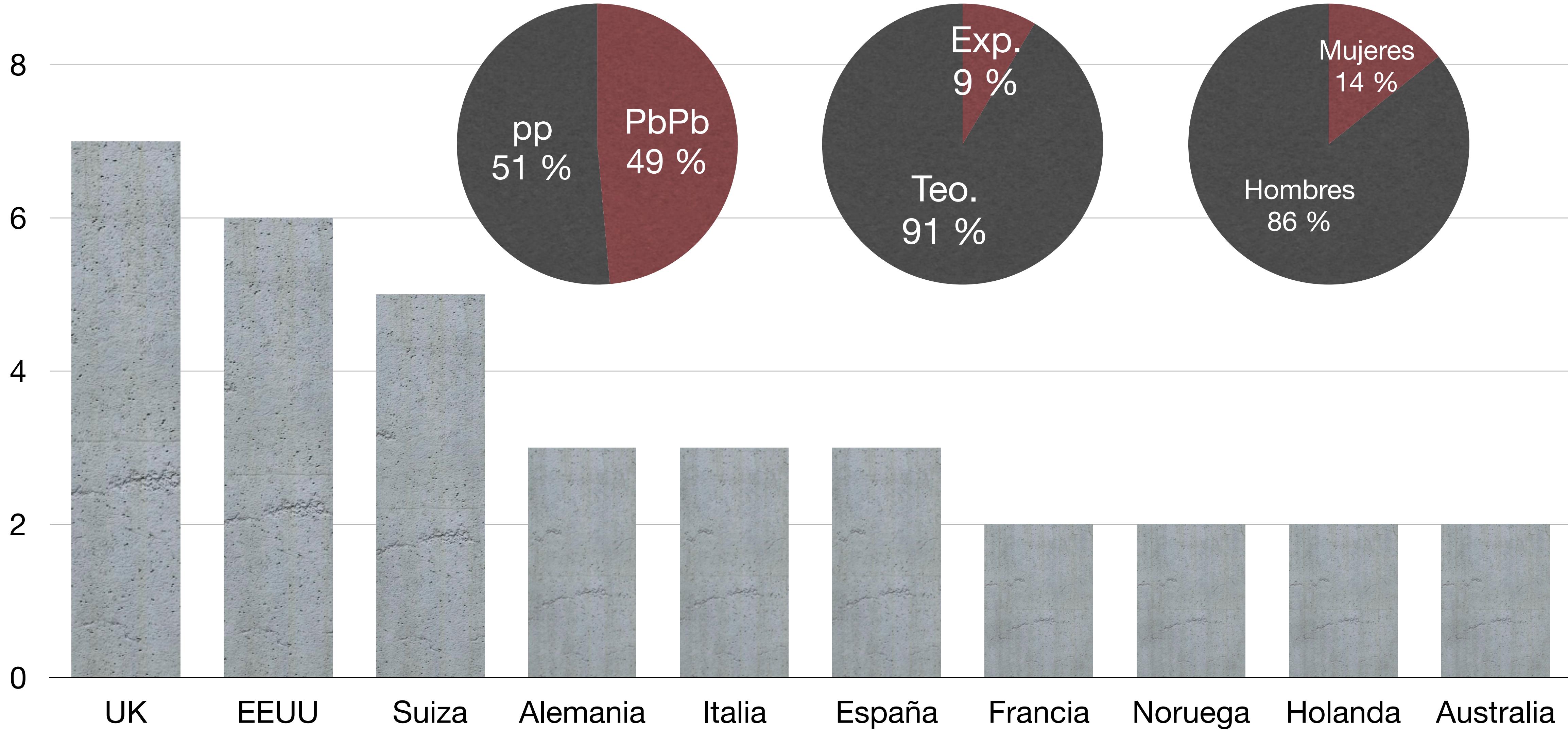
PRELUDIO

Quién soy yo?

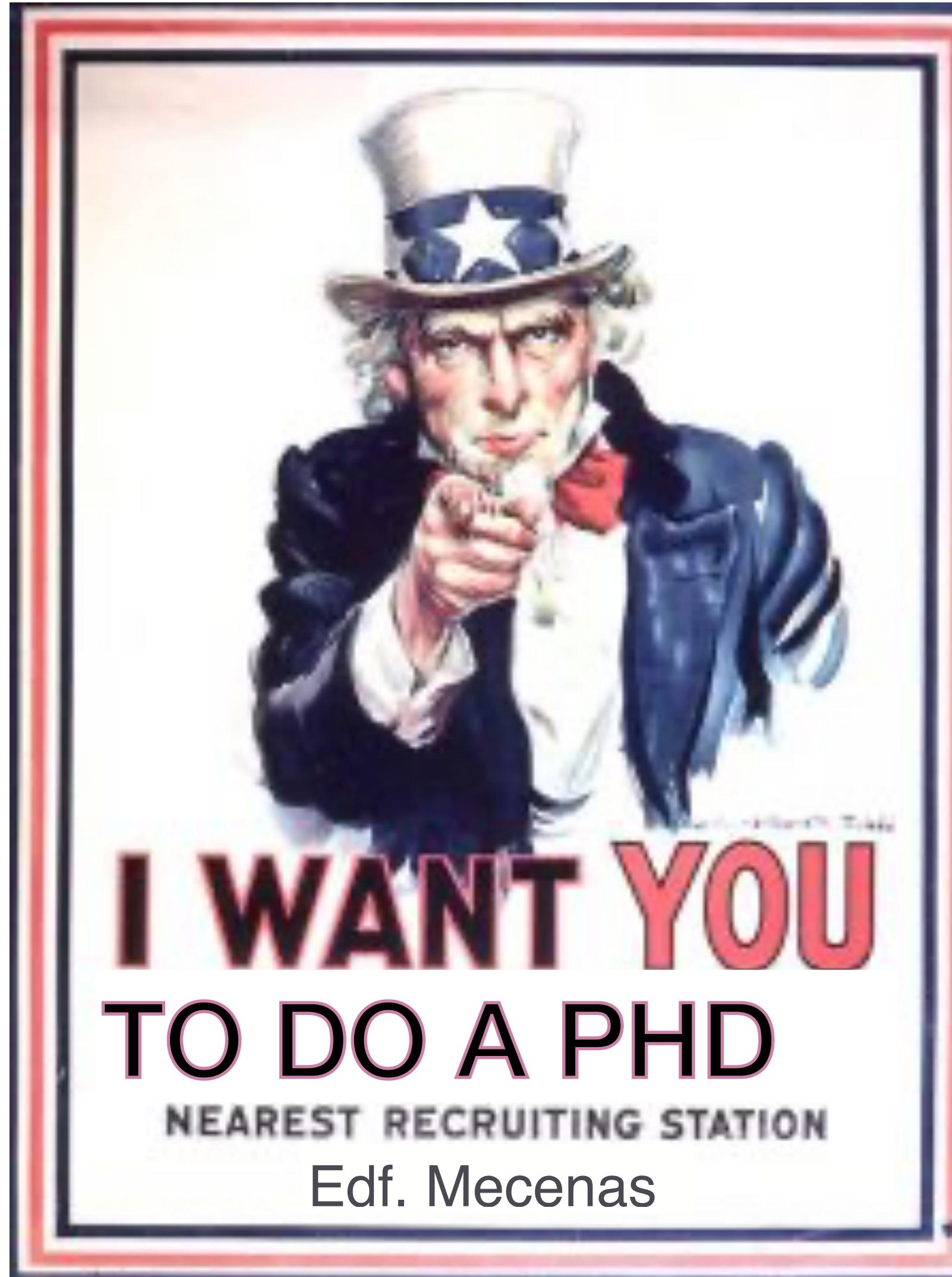


- Sep 2010-June 2015, **BSc. + MS.c Physics** @University of Sevilla + Granada (Spain)
- Sep 2015-June 2018, **Ph.D** @University of Granada (Spain) + Goethe University (Germany) w/ Javier L. Albacete and Hannah Elfner
- Oct 2018-Oct 2020, **Research Associate** @Brookhaven National Lab (USA) w/ Yacine Mehtar-Tani
- Oct 2020-Oct 2022, **Postdoc** @Institute de Physique Theorique (France) w/ Gregory Soyez
- Oct 2022-Sep 2024, **Senior fellow** @CERN (Switzerland)
- Sep 2024-∞(?), **Ramón y Cajal** @University of Granada (Spain)

Mi red de colaboradores/as



Únete

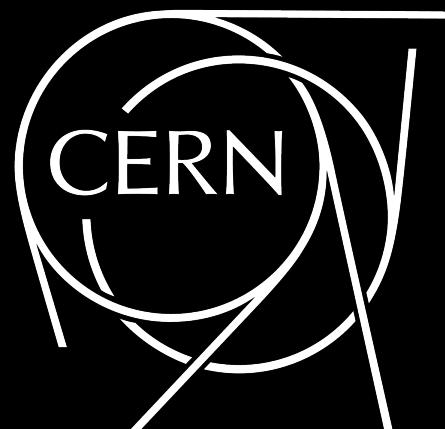
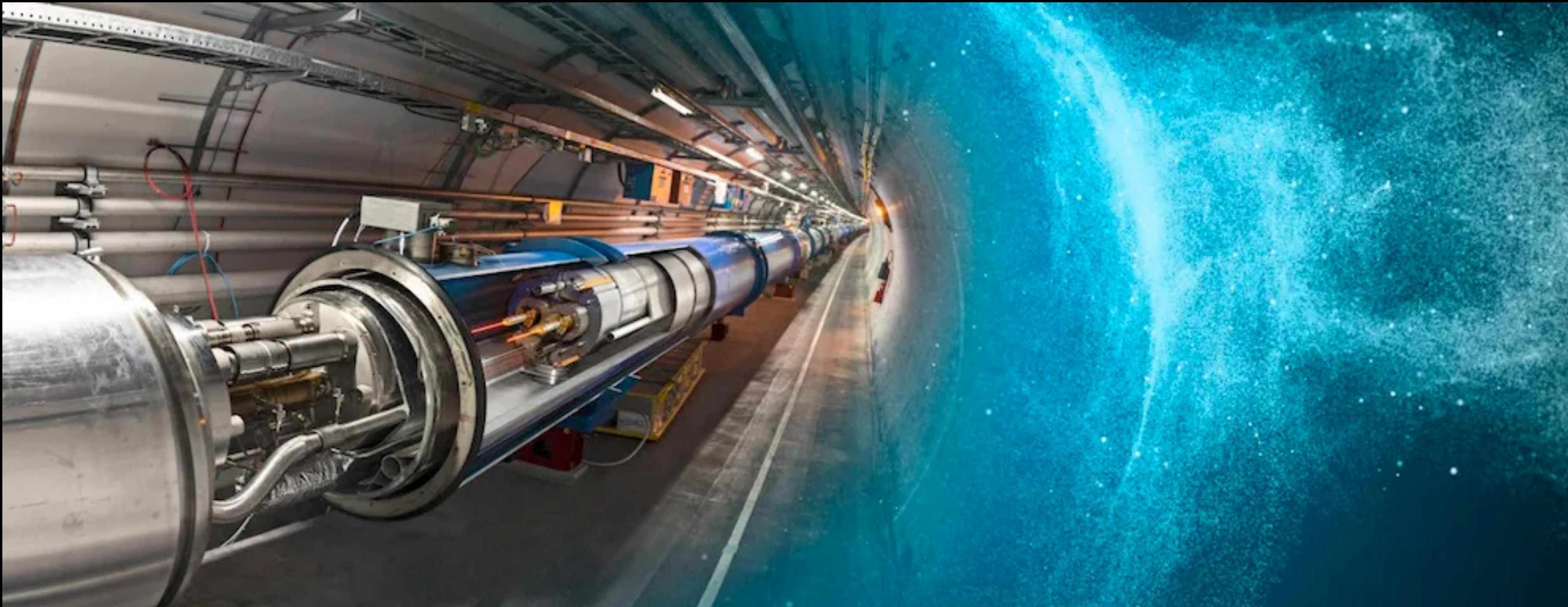


Si estás interesado/a en trabajar
conmigo, escríbeme un correo:

alba.soto.ontoso@cern.ch

y hablamos!

Estudio de la materia en condiciones extremas: del cosmos a los aceleradores de partículas



Alba Soto Ontoso
Seminario de Invitados – Máster en Física
Granada, 8-9th January, 2024

Plan for the course

Lecture 1: big picture

- The Strong force
- QCD Thermodynamics
- The Quark-Gluon Plasma
- Little and Big Bangs

Lecture 3: fluid evolution

- Hydrodynamic paradigm
- Connection to nuclear structure
- Small systems puzzle

Lecture 2: early stages

- Spatio-temporal evolution of a heavy-ion collision
- Nuclear PDFs
- Geometry of the collision

Lecture 4: jets in the QGP

- Parton shower in vacuum
- Interleaving of vacuum and medium-induced emissions
- Color randomization in the QGP

Plan for the course

Besides these slides, some extra material that we will use can be found in

<https://github.com/albaontoso/2024-Granada-heavyion-lectures>

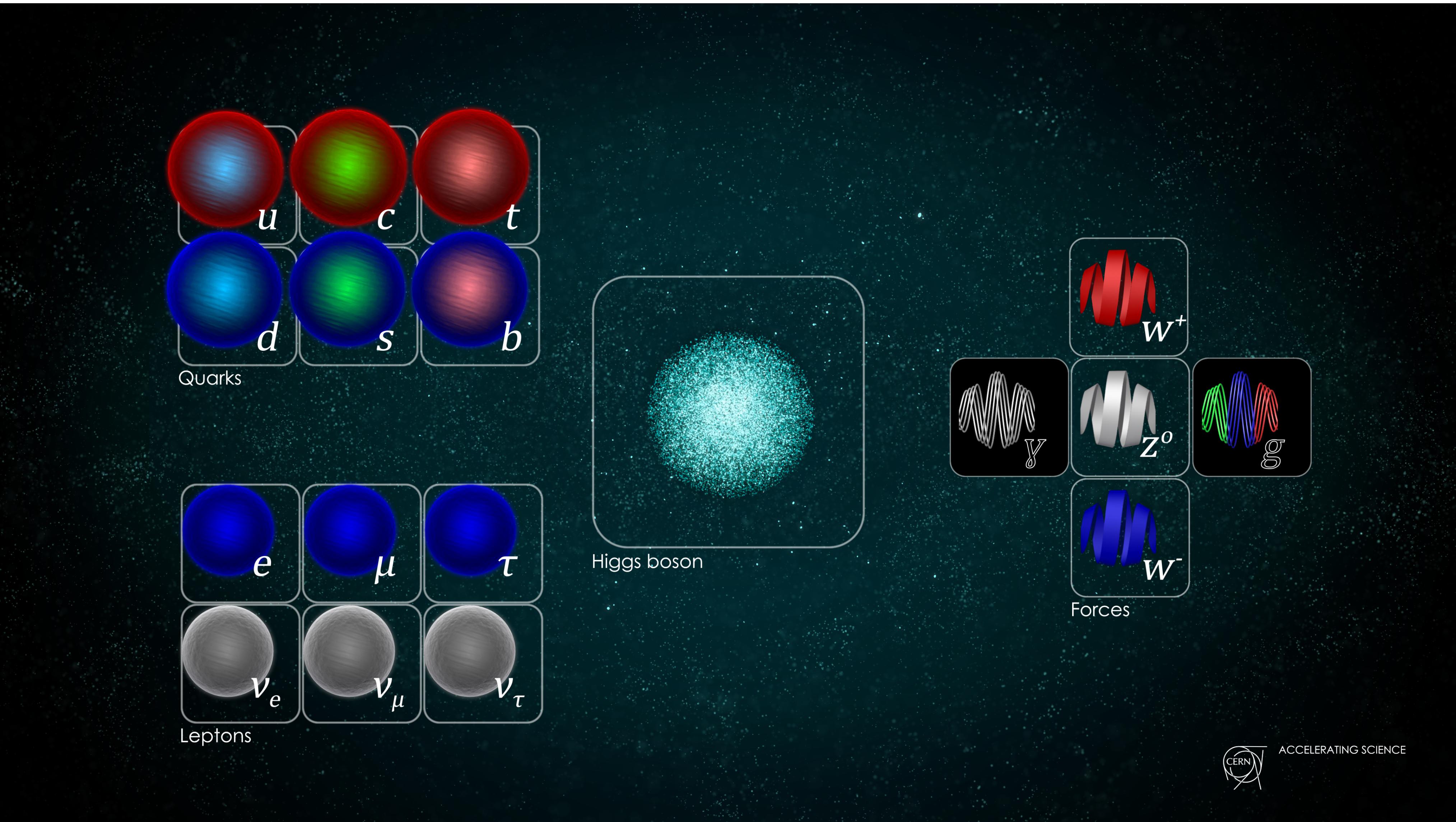
To download it in your computer do

```
git clone git@github.com:albaontoso/2024-Granada-heavyion-lectures.git
```

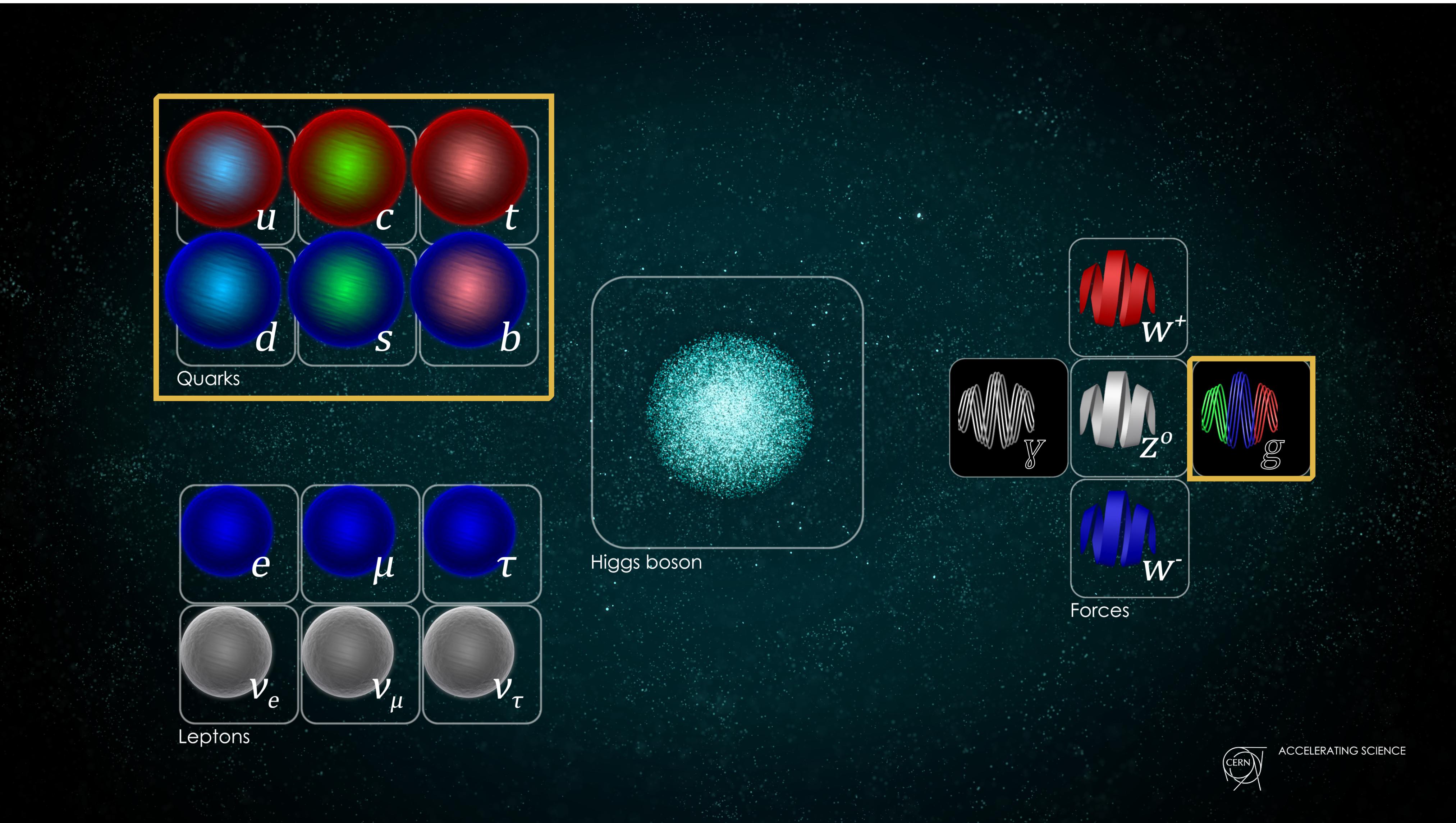
Everything written in PYTHON. Few tools to use Jupyter notebooks:

- Visual studio code (best editor in my opinion)
- <https://colab.research.google.com/>
- <https://jupyter.org/try>

Matter at the femtometer scale

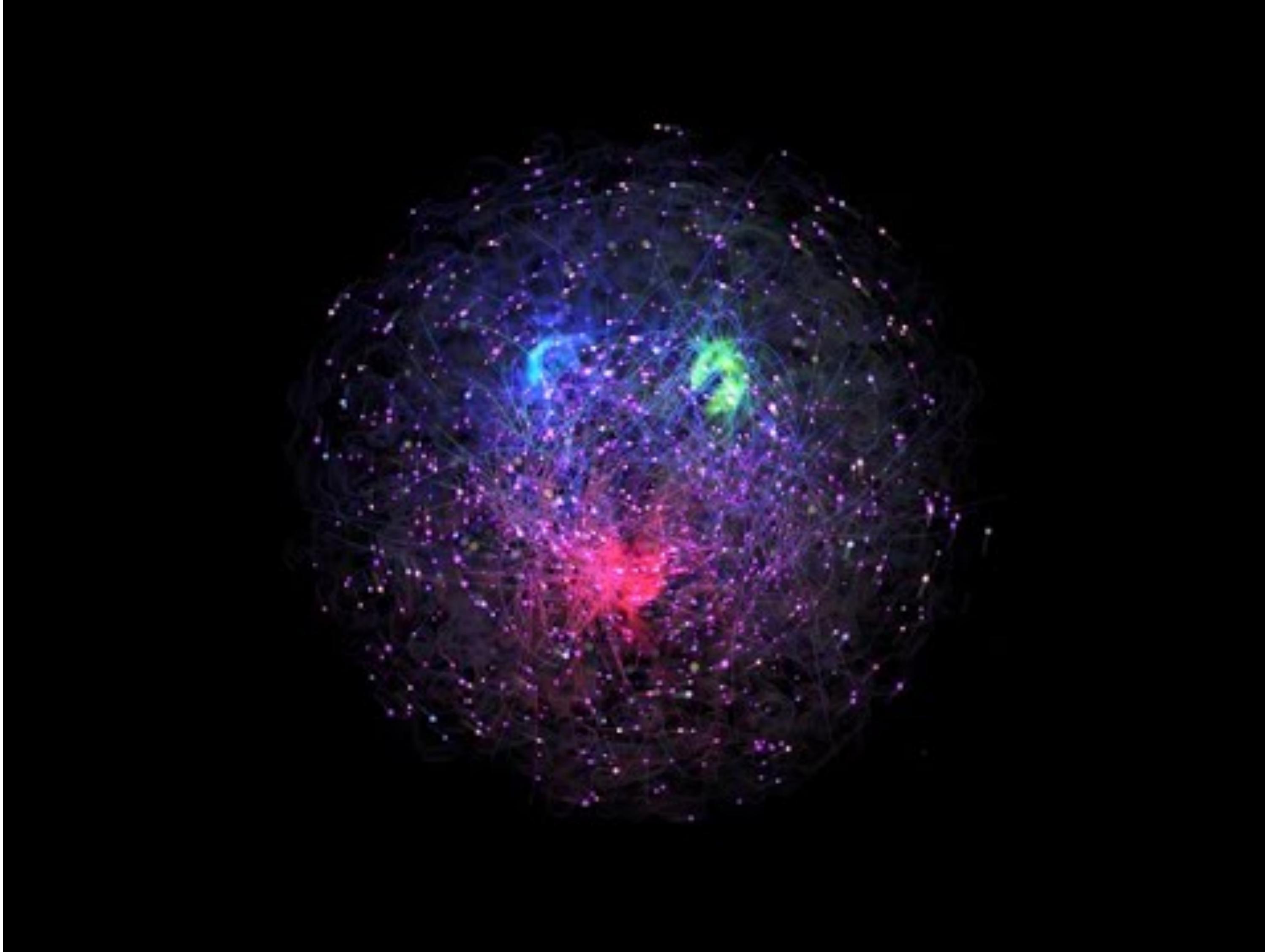


Focus of these lectures: quarks and gluons



Fascinating dynamics inside the proton

[Arts at MIT, <https://youtube.com/watch?v=Dt8FZ4ksWiY>]



Fascinating dynamics inside the proton

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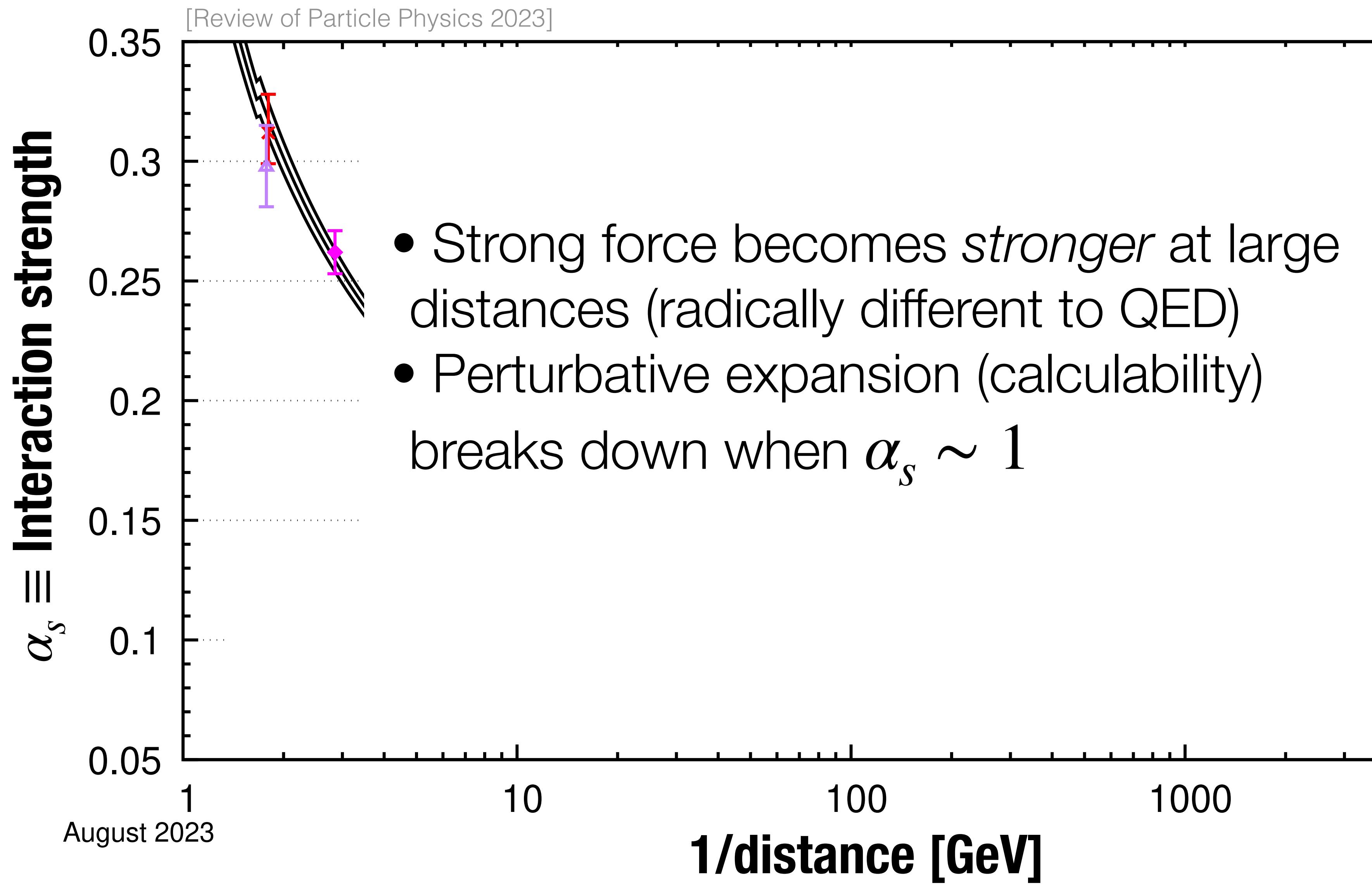
Quarks and gluons (i.e. color-charged particles) cannot be isolated

Key aspects of QCD: confinement and asymptotic freedom

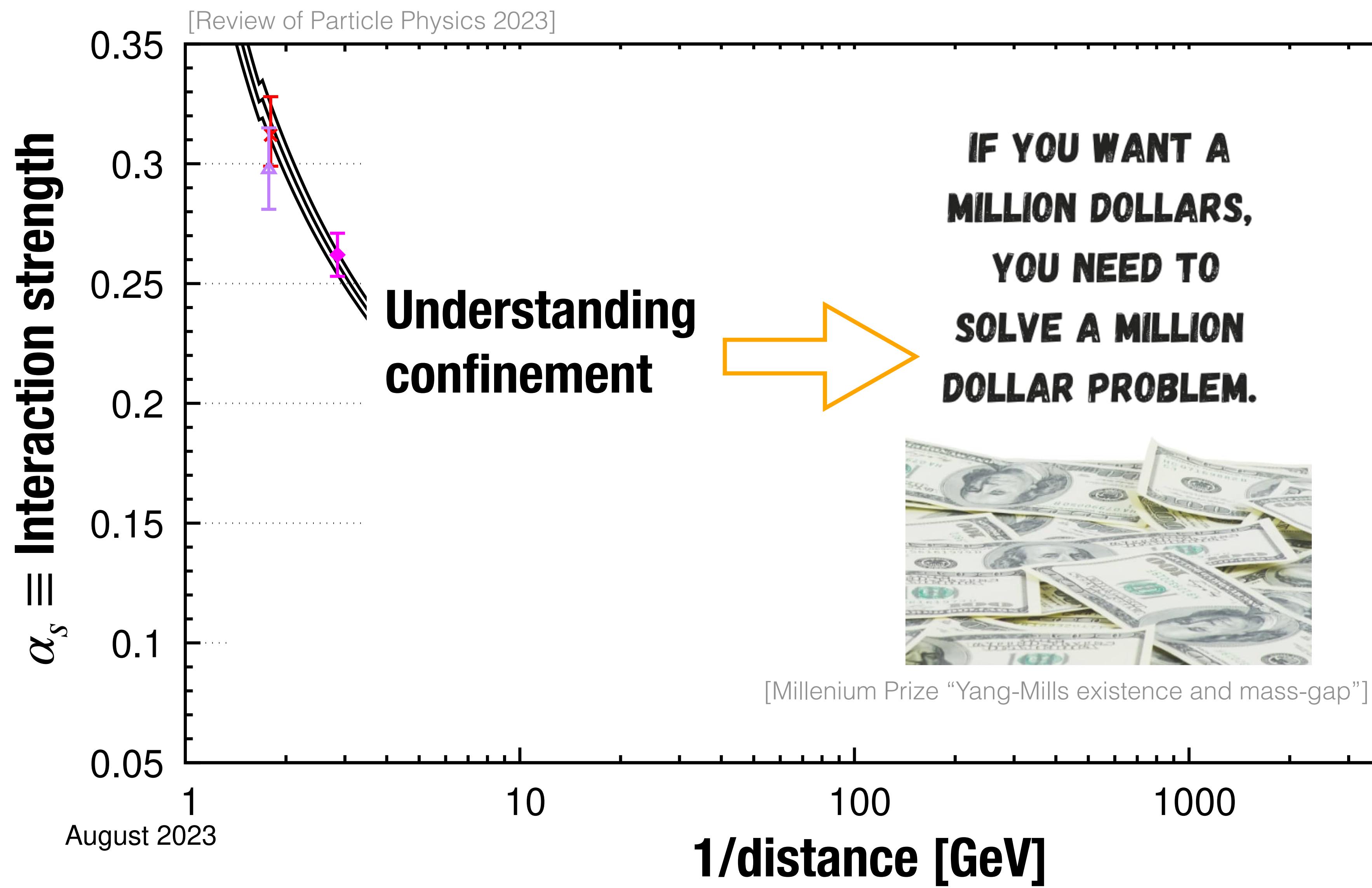
$$\alpha_s \equiv \text{Interaction strength}$$

1/distance [GeV]

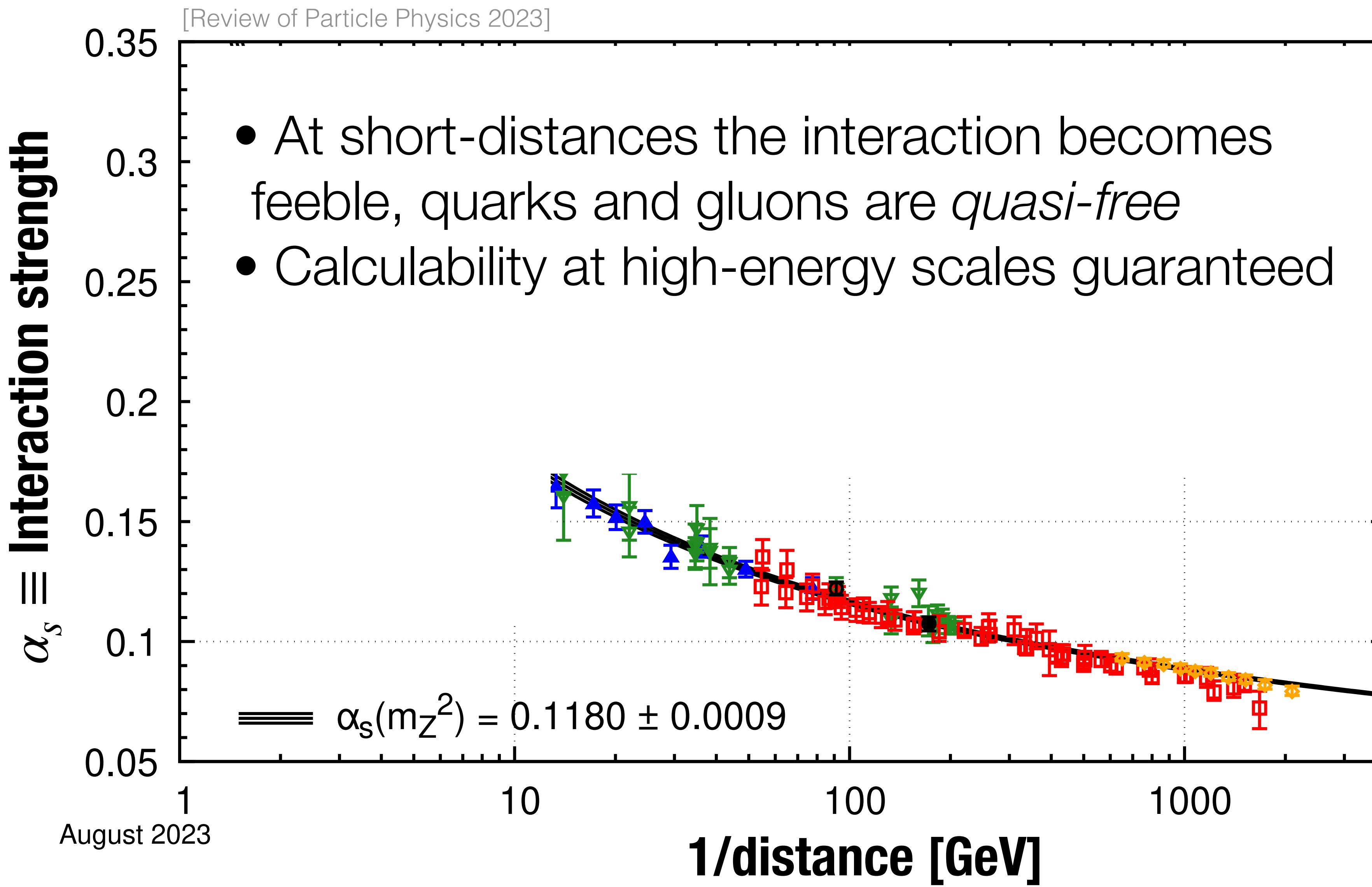
Key aspects of QCD: confinement



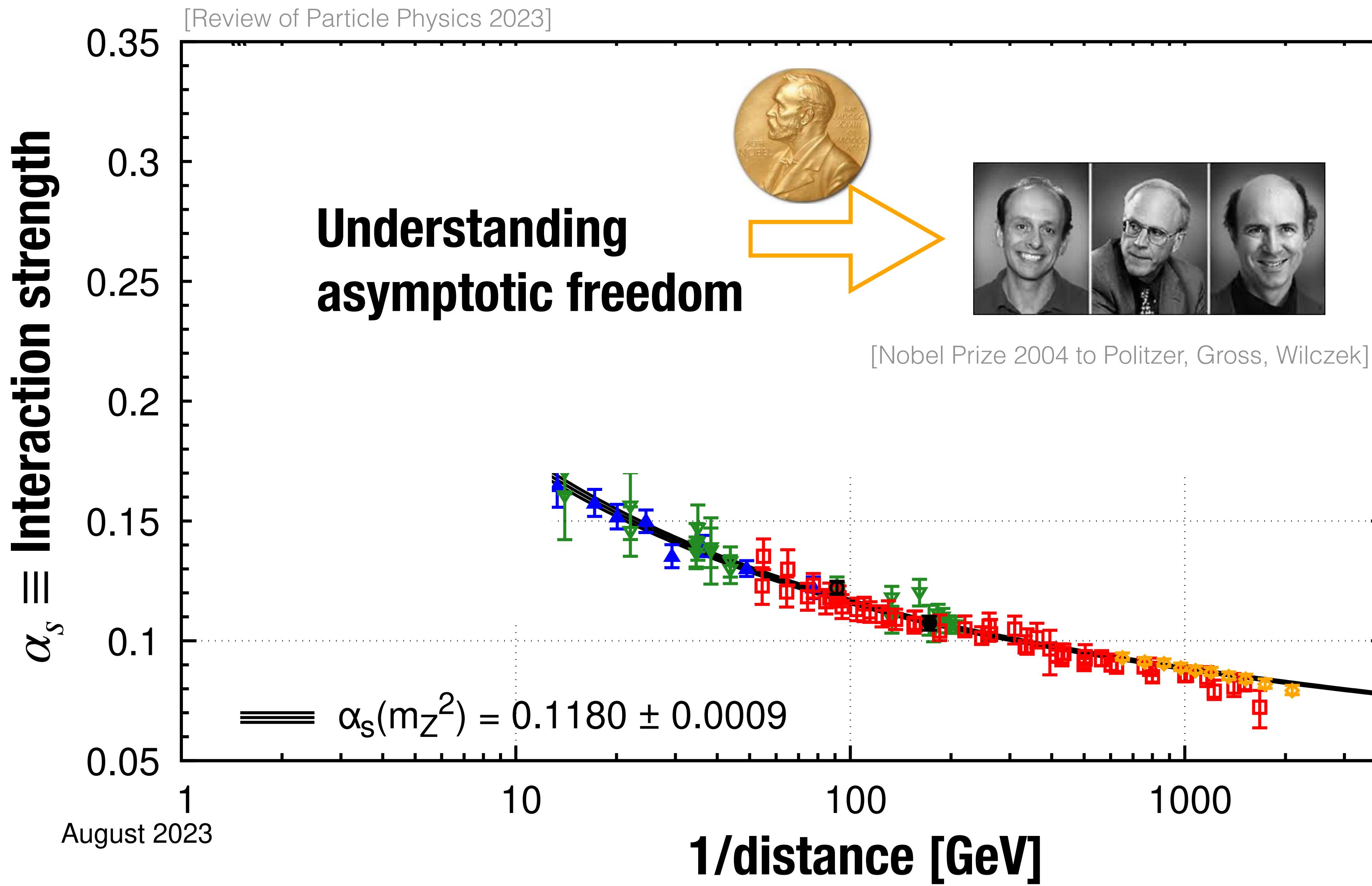
Key aspects of QCD: confinement



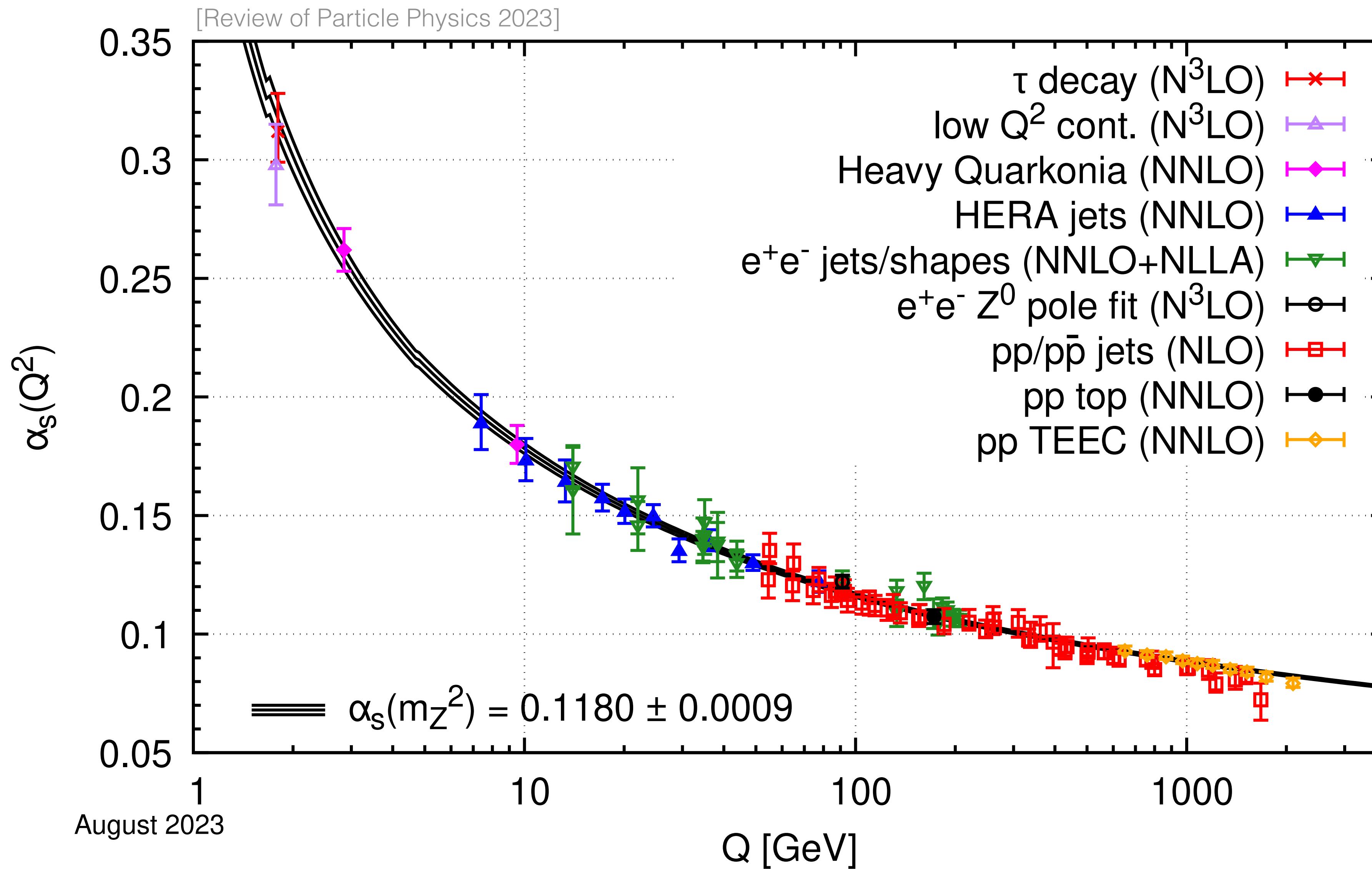
Key aspects of QCD: asymptotic freedom



Key aspects of QCD: asymptotic freedom

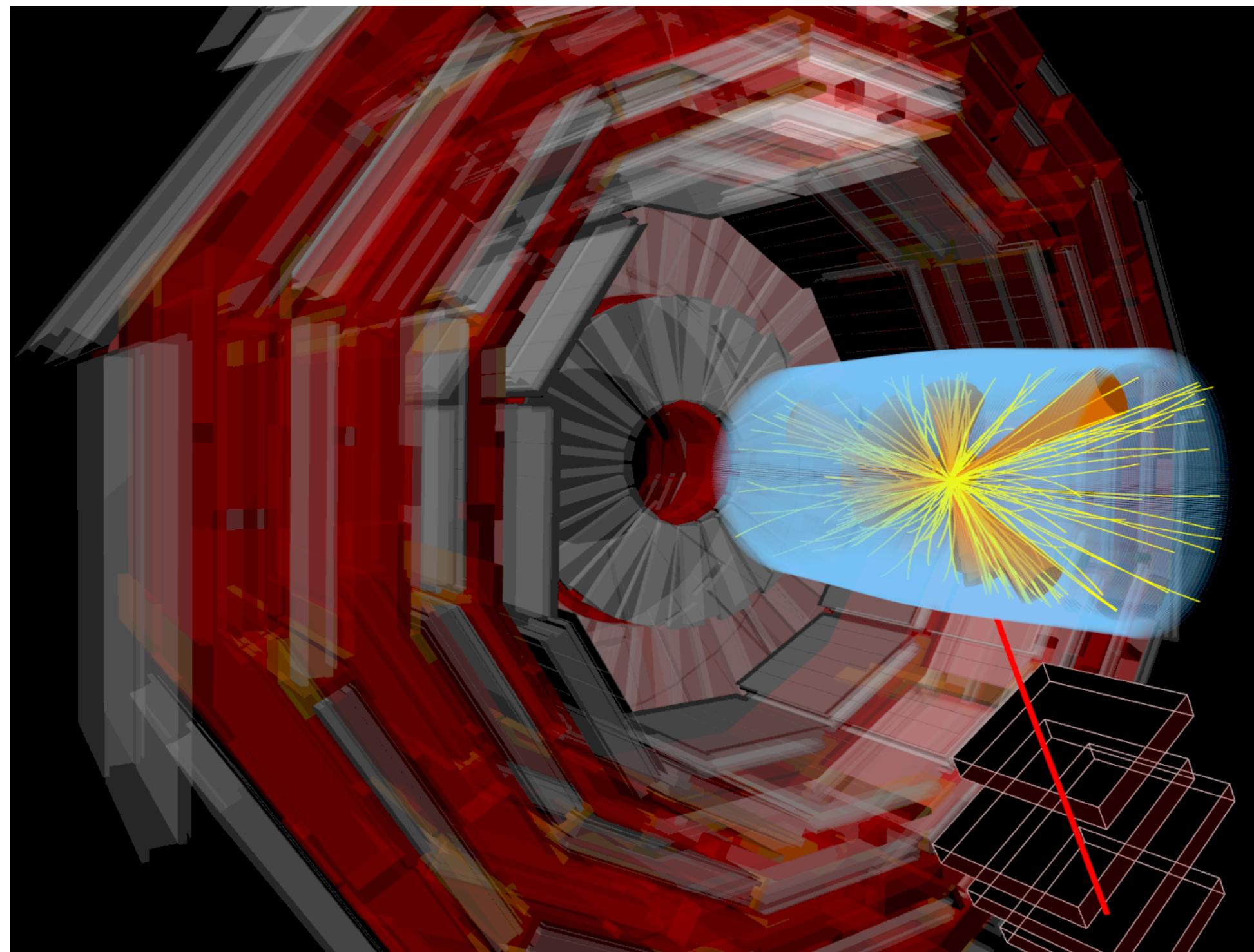


Key aspect of QCD: running of the coupling



Technical (but amazing) point: this behaviour is connected to the gluon self-interaction

Asymptotic freedom opens up the LHC program



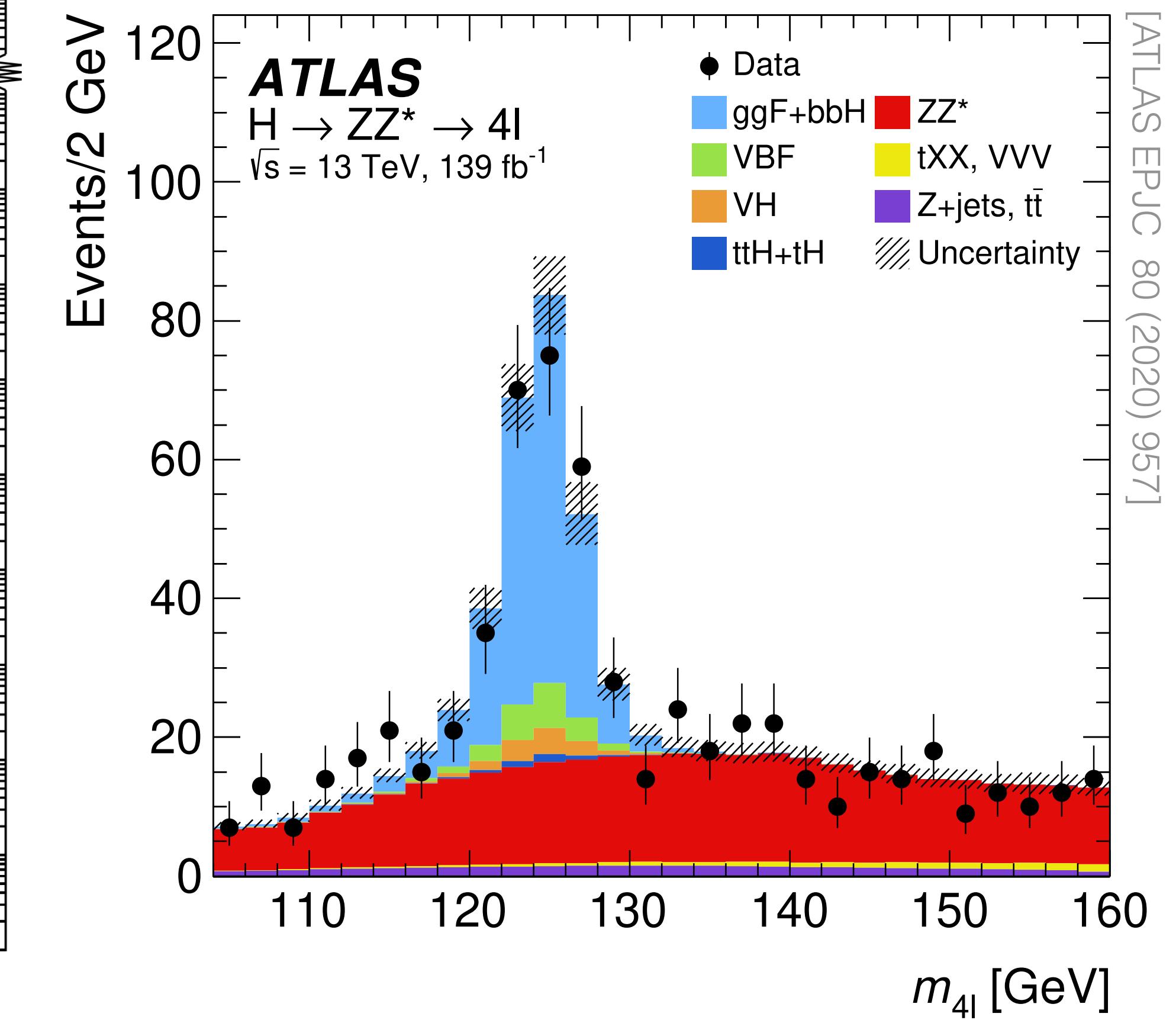
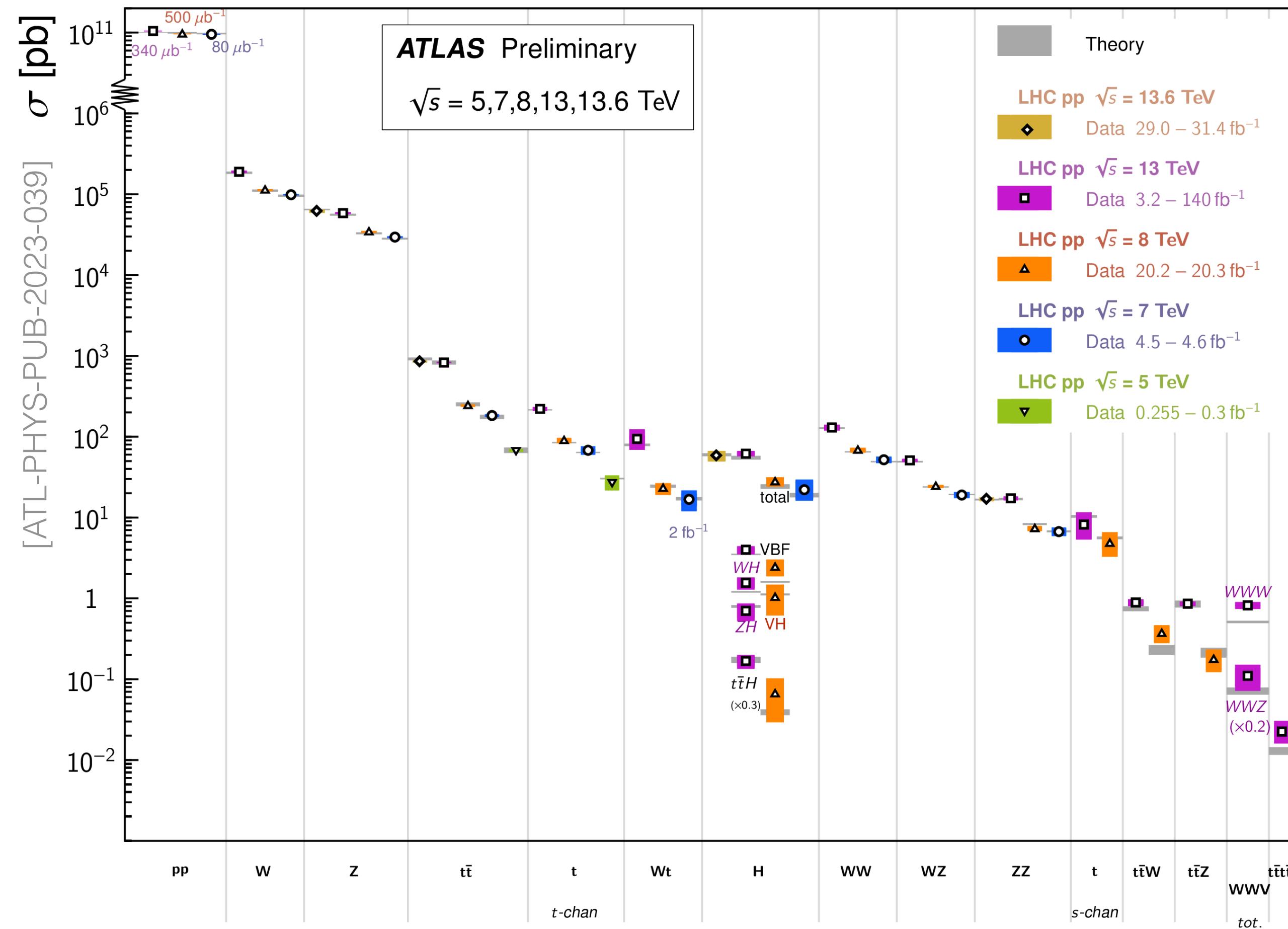
$$= \left| \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \right| + \left| \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \right| - \left| \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \right| \times \left| \begin{array}{c} \text{---} \\ | \\ \text{---} \\ | \\ \text{---} \end{array} \right| + \mathcal{O}(\alpha_s^3)$$

Proton-proton collision computed in terms of Feynman diagrams
with quarks and gluons as degrees of freedom

QCD in the real world

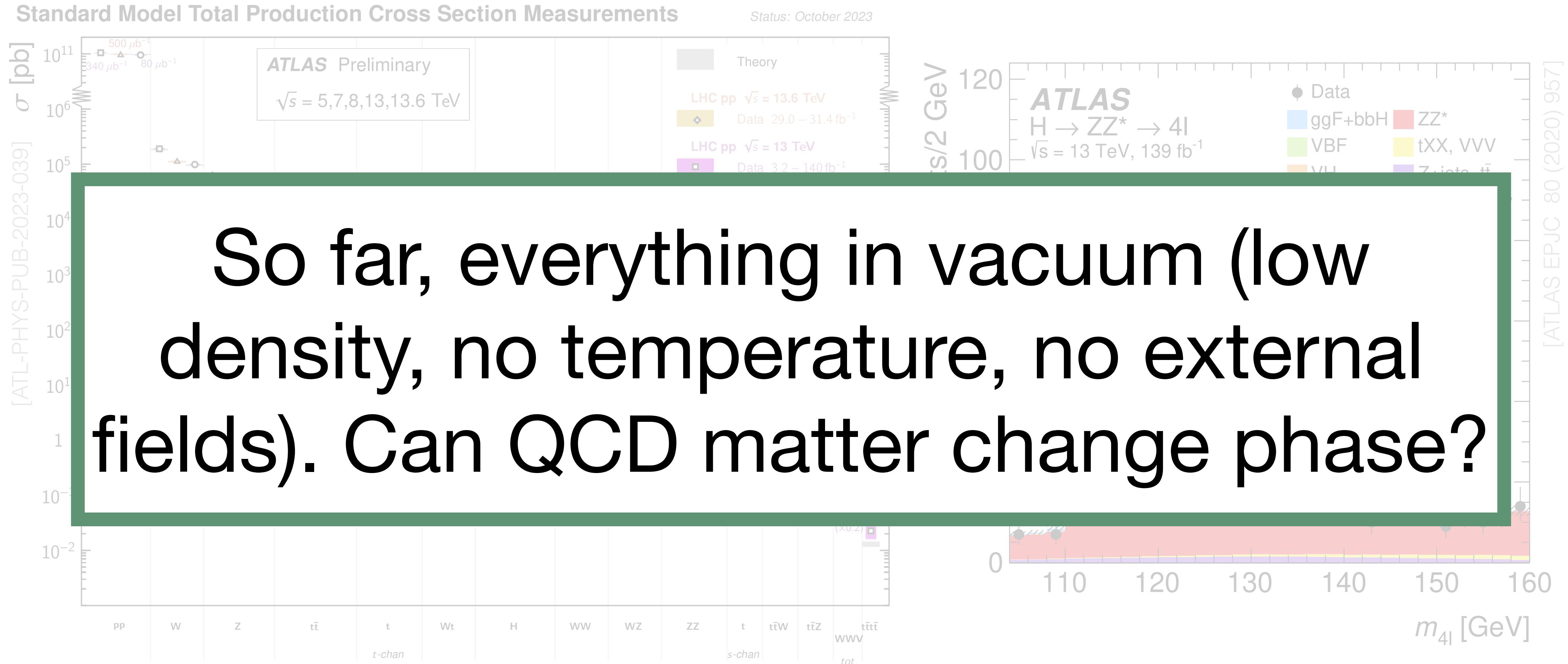
Standard Model Total Production Cross Section Measurements

Status: October 2023



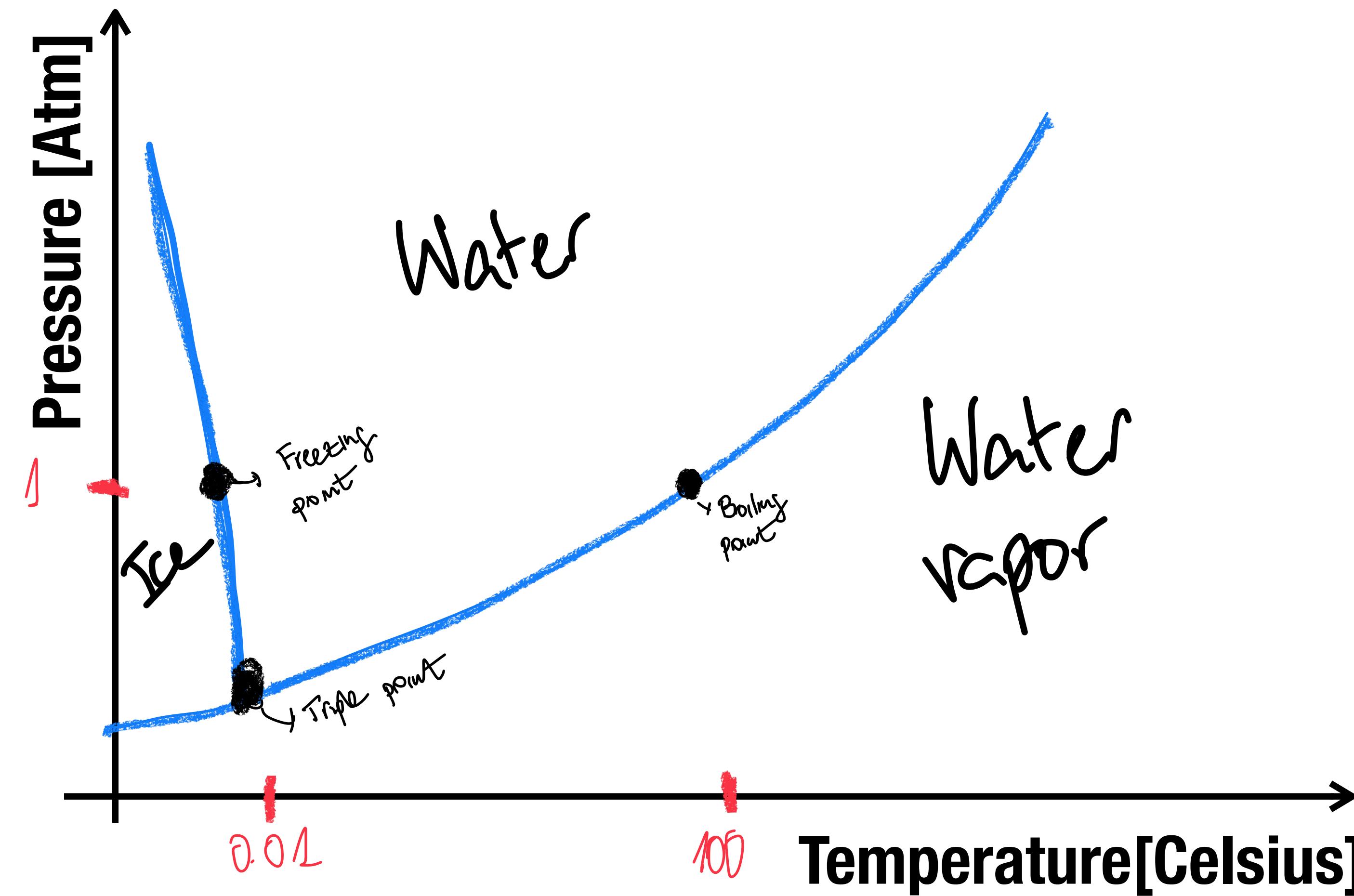
The Higgs wouldn't have been discovered without precise QCD

QCD in the real world

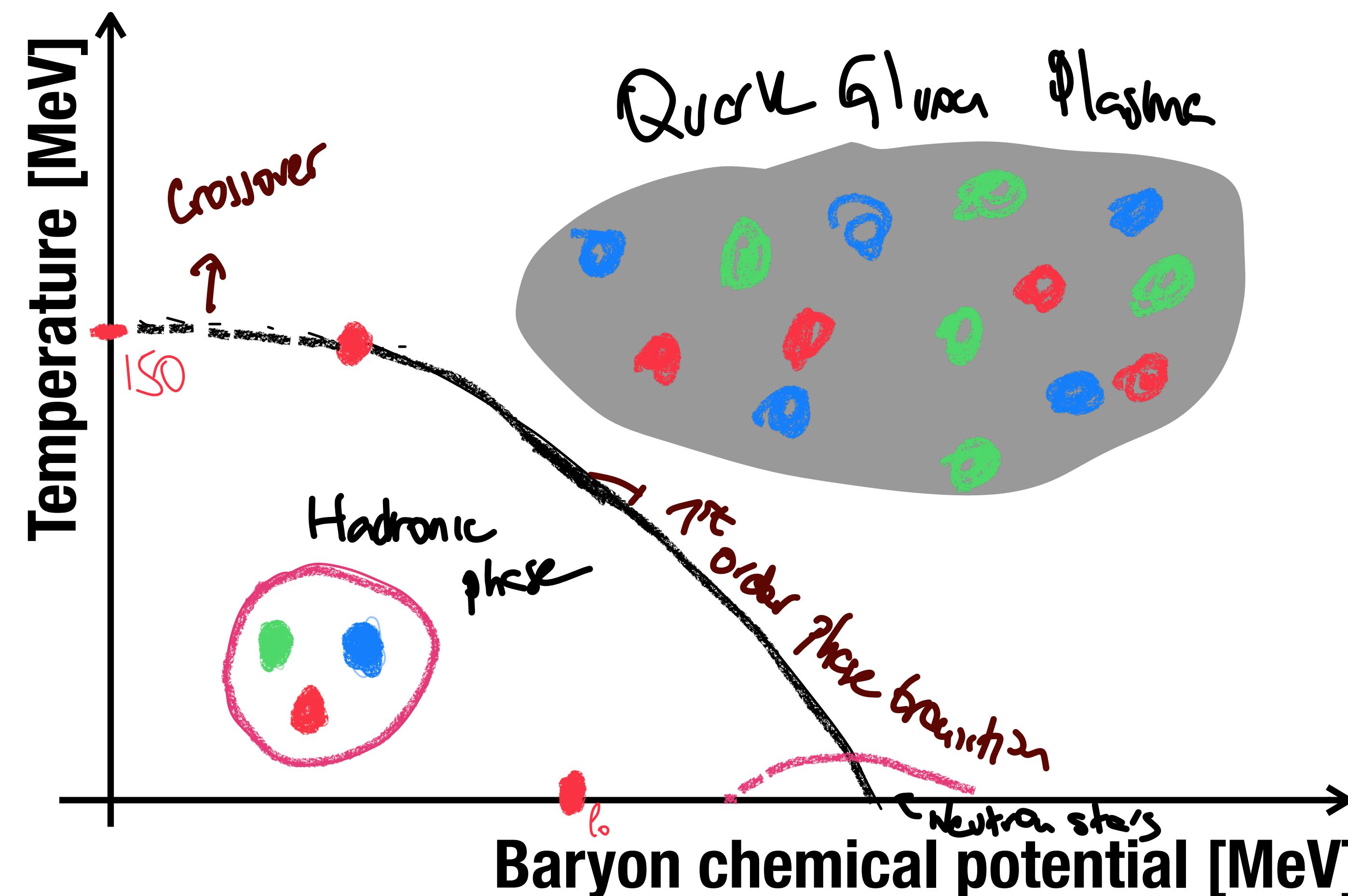


The Higgs wouldn't have been discovered without precise QCD

Reminder: phase diagram of water

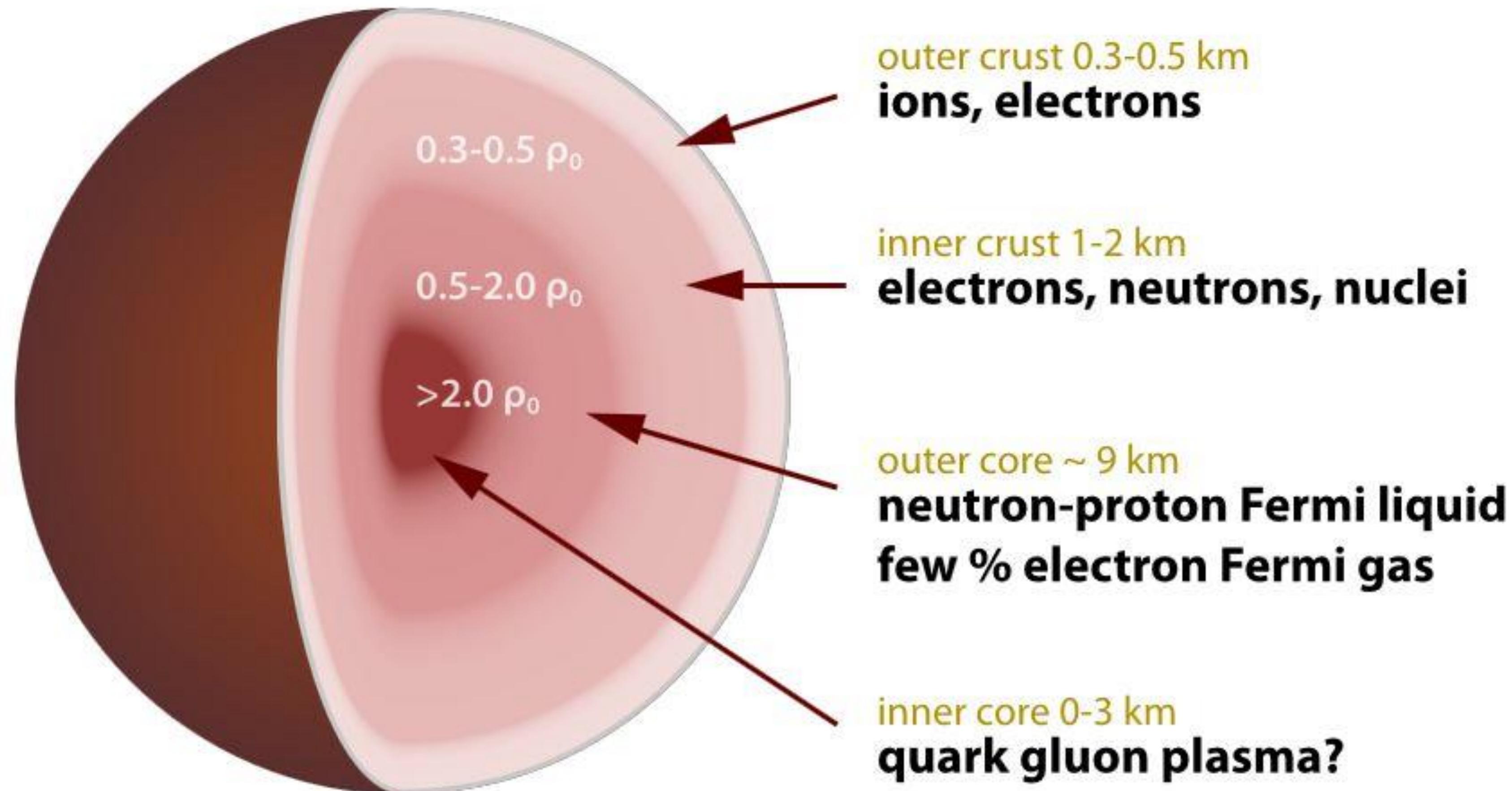


Phase diagram of QCD matter



ρ_0 : nuclear saturation mass density $\rho_0 \sim m_n n_0 = m_n \frac{A}{\frac{4}{3} \pi R^3} \sim 900 \text{ MeV}$

Neutron stars: as dense as it gets

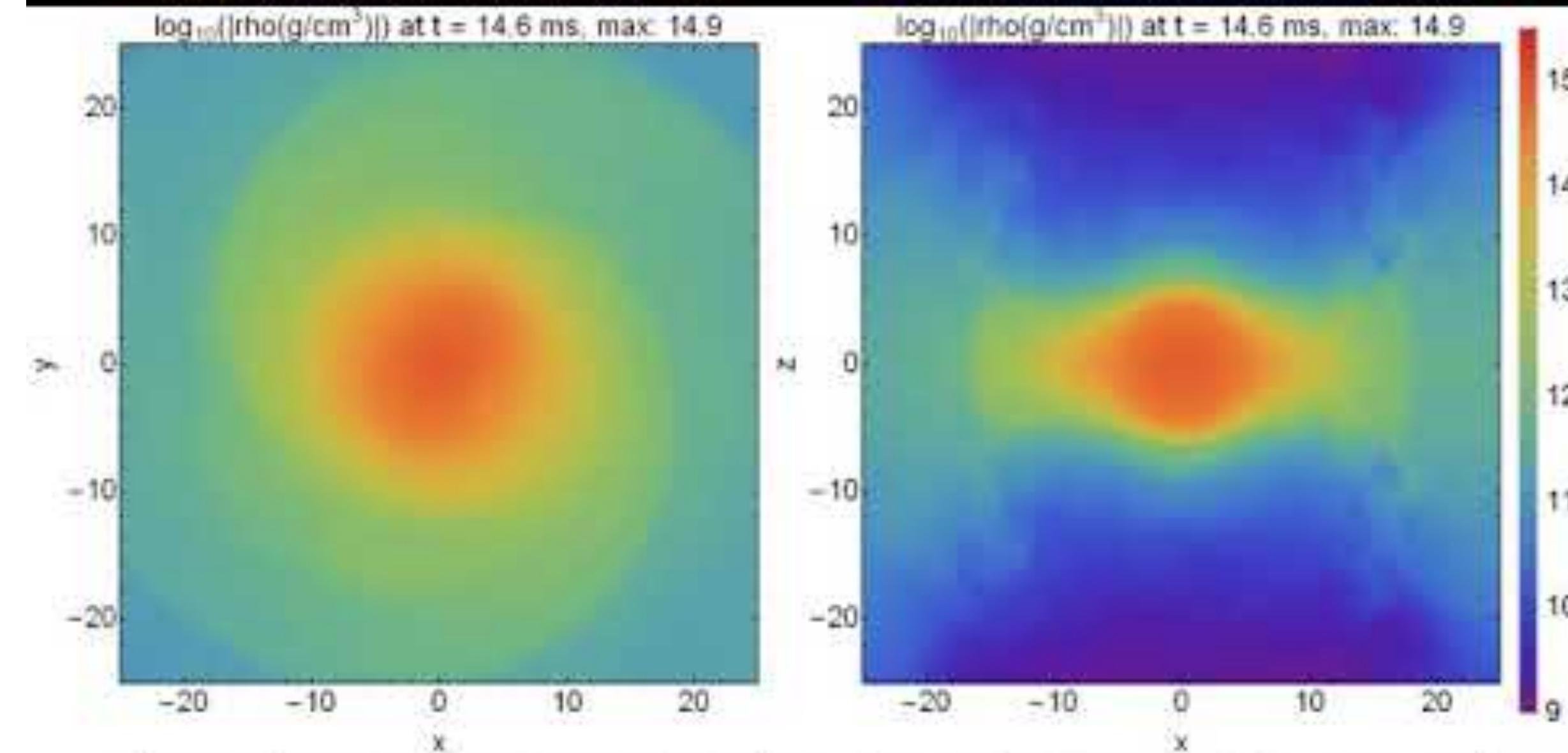


Critical density around 10^{17} kg/m³ ~ put planet Earth into Los Carmenes

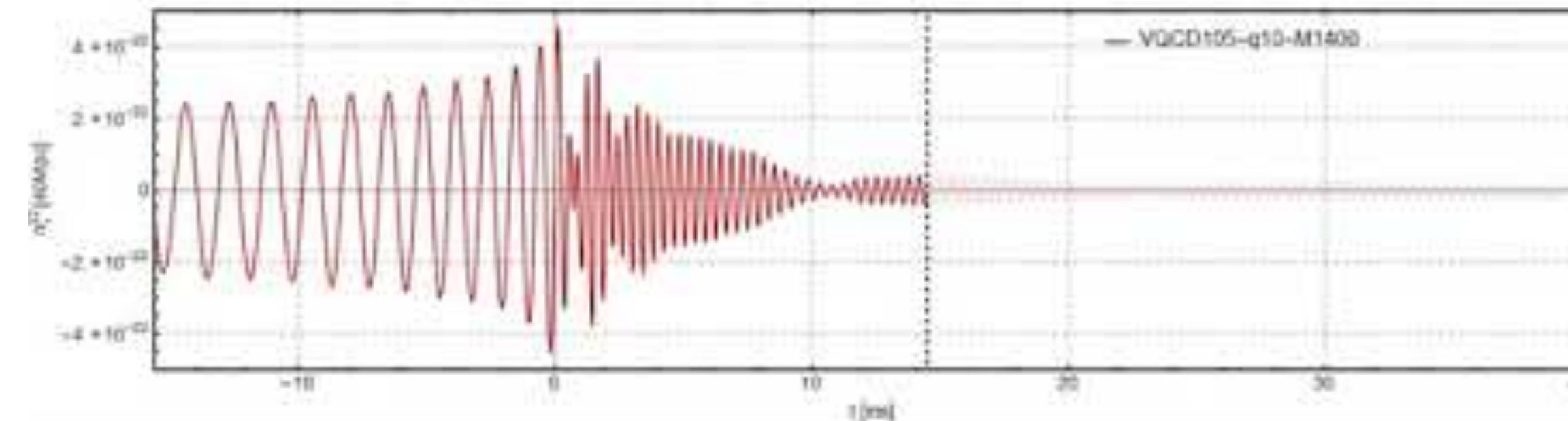
Neutron stars produce gravitational waves

[Eckert et al PRD 101 (2020) 10, 103006] <https://youtube.com/watch?v=p51XSfYcSb8>

Mass density



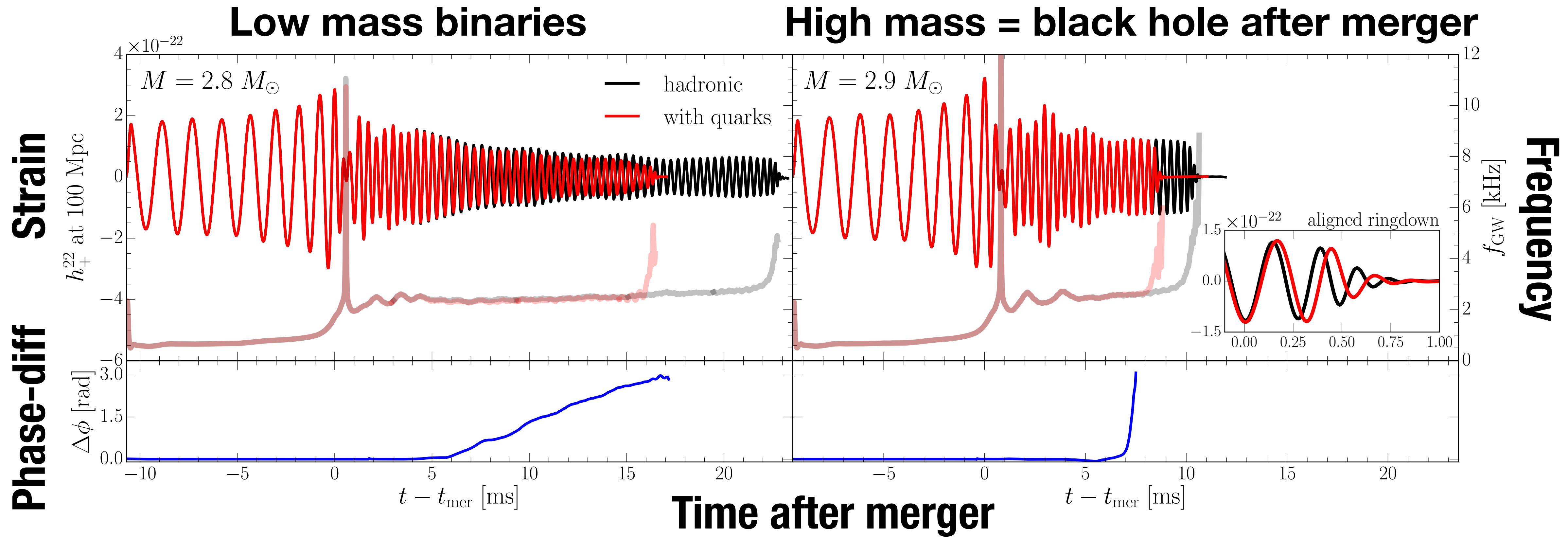
**Strain
gravitational
wave**



The merger of 2 neutron stars create gravitational waves measurable at LIGO

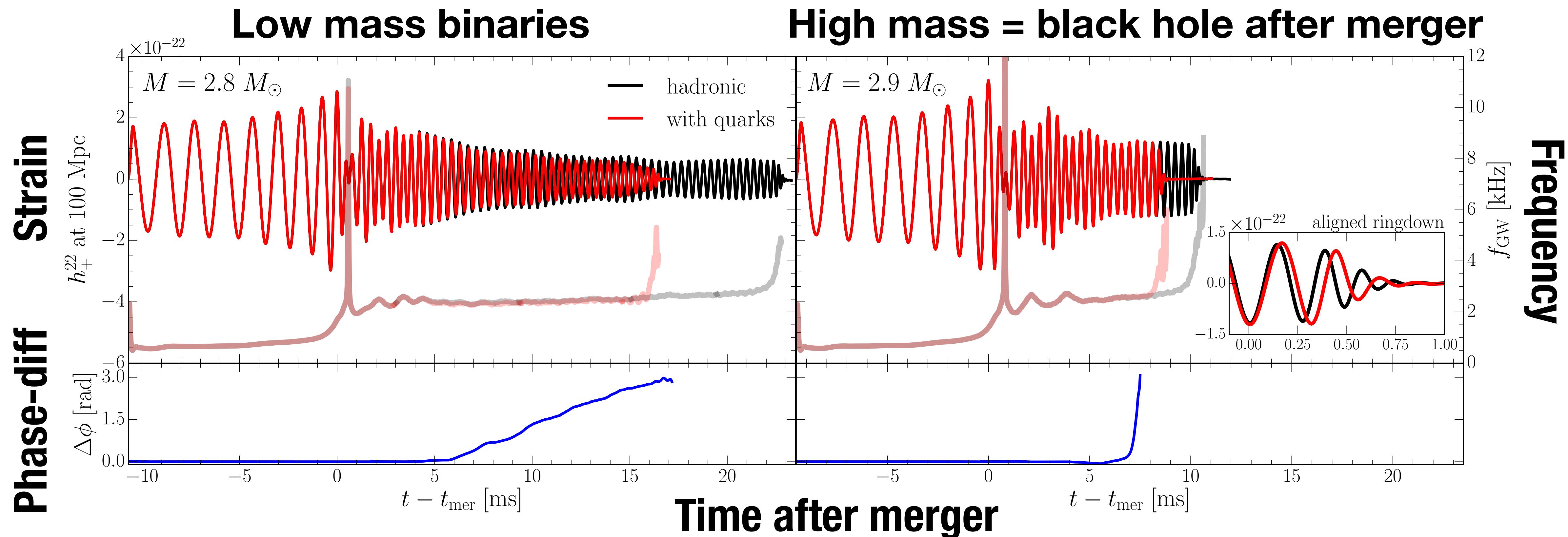
QCD phase-transition imprint into gravitational waves

[Most et al PRL 122 (2019) 6, 061101]



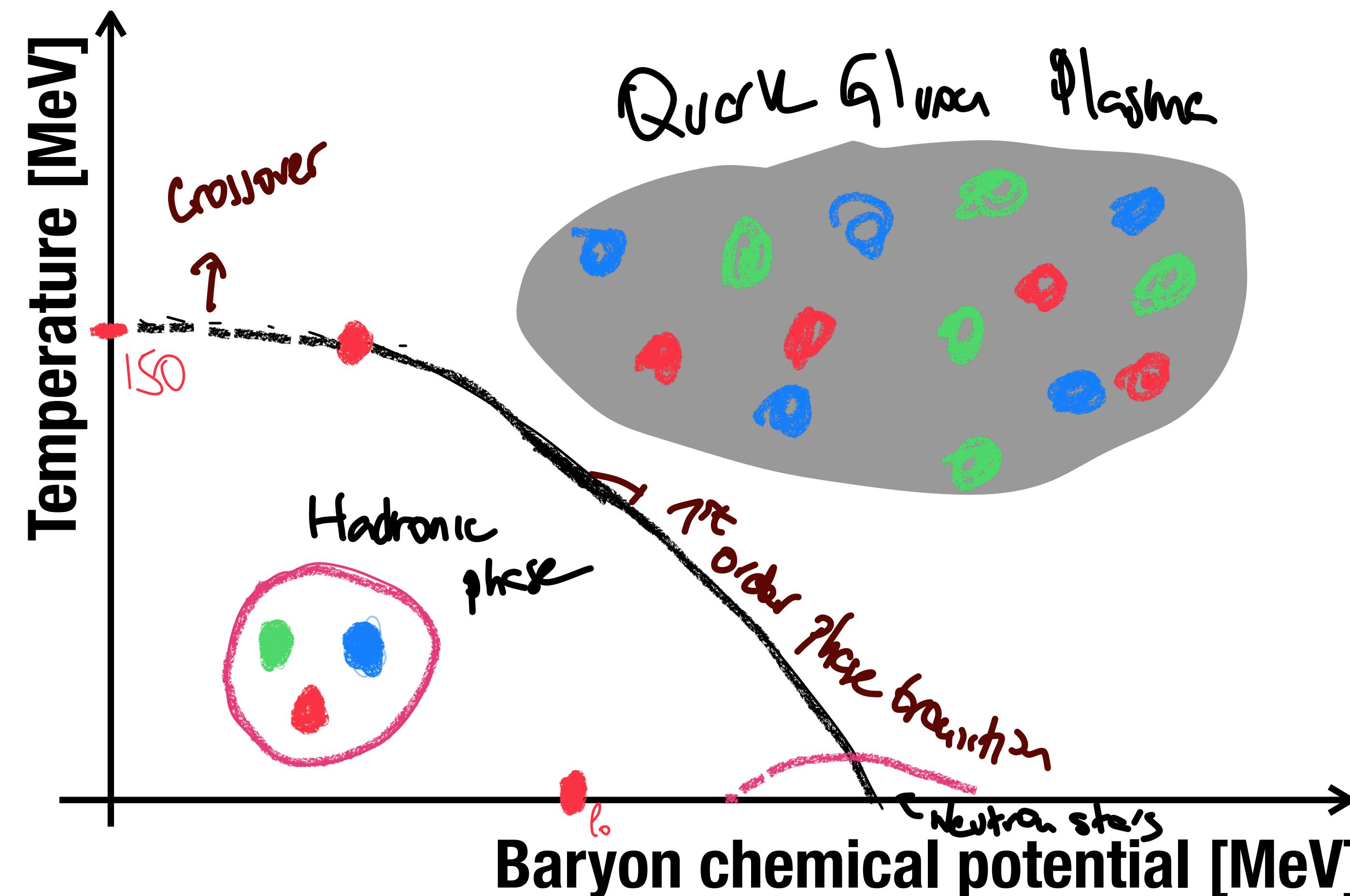
QCD phase-transition imprint into gravitational waves

[Most et al PRL 122 (2019) 6, 061101]



- During inspiral phase no quarks are formed: identical GW signal
- Low mass: post merger spectra sensitive to the presence of quarks
- High mass: signal quickly suppressed by the collapse into a BH

Phase diagram of QCD matter



ρ_0 : nuclear saturation mass density $\rho_0 \sim m_n n_0 = m_n \frac{A}{\frac{4}{3} \pi R^3} \sim 900 \text{ MeV}$

QCD thermodynamics: simple estimates

Consider an ideal gas composed of N_F massless quarks and N_B gluons. The so-called Stefan-Boltzmann pressure is given by:

$$p(T) = \frac{\pi^2}{90} \left(N_B + \frac{7}{8} N_F \right) T^4$$

where

$$N_B = 2 \times 8 \qquad \qquad N_F = 2 \times 2 \times 3 \times 3$$

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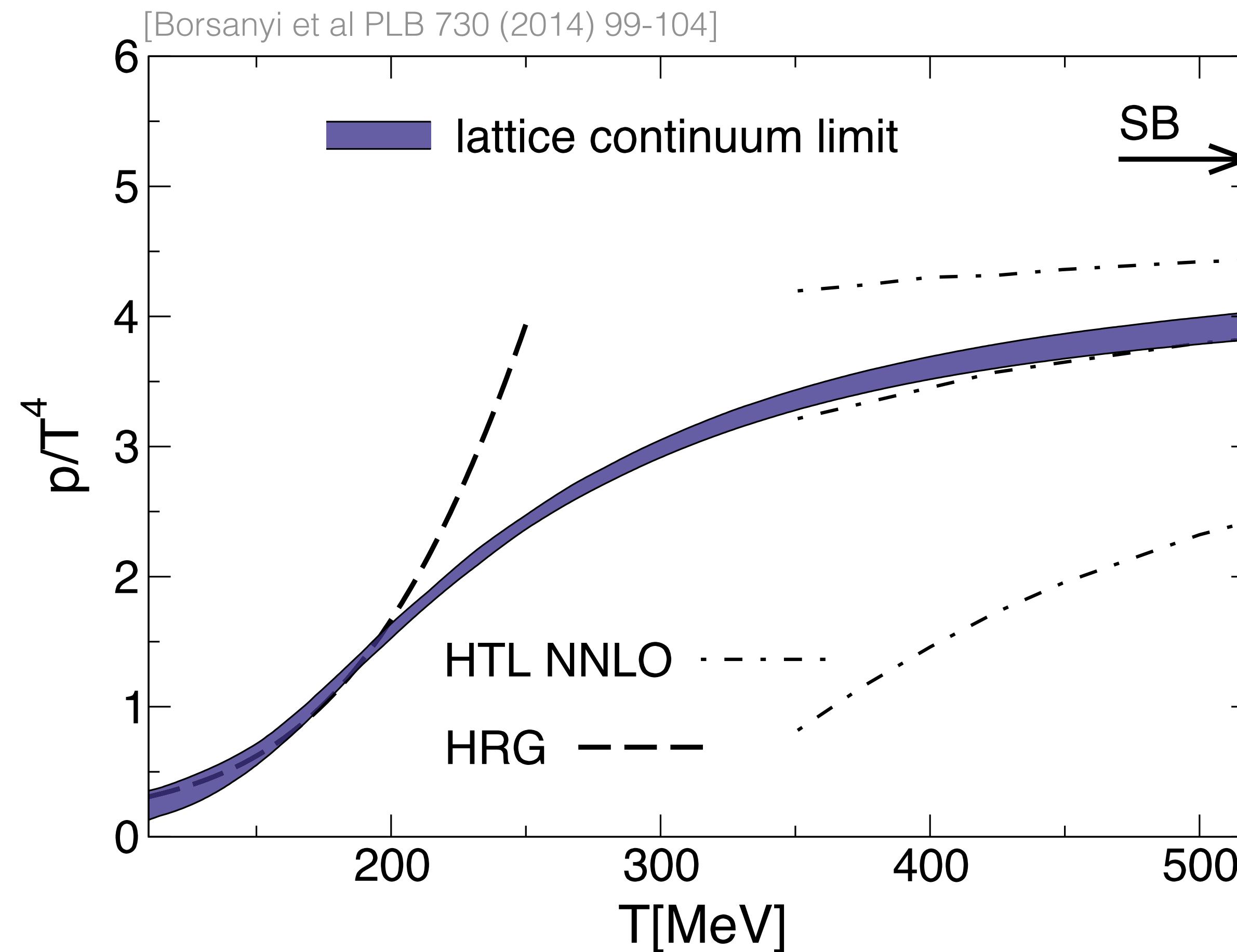
The diagram shows the expression $N_B = 2 \times 8$ at the top. Two arrows point downwards from the numbers 2 and 8 to the words "spin" and "colour" respectively, indicating that the product 2 times 8 is being broken down into these two components.

$$N_F = 2 \times 2 \times 3 \times 3$$

↑ ↑
colour flavour (u/d/s)

↓ ↓
spin particle/antiparticle

QCD thermodynamics: using lattice techniques

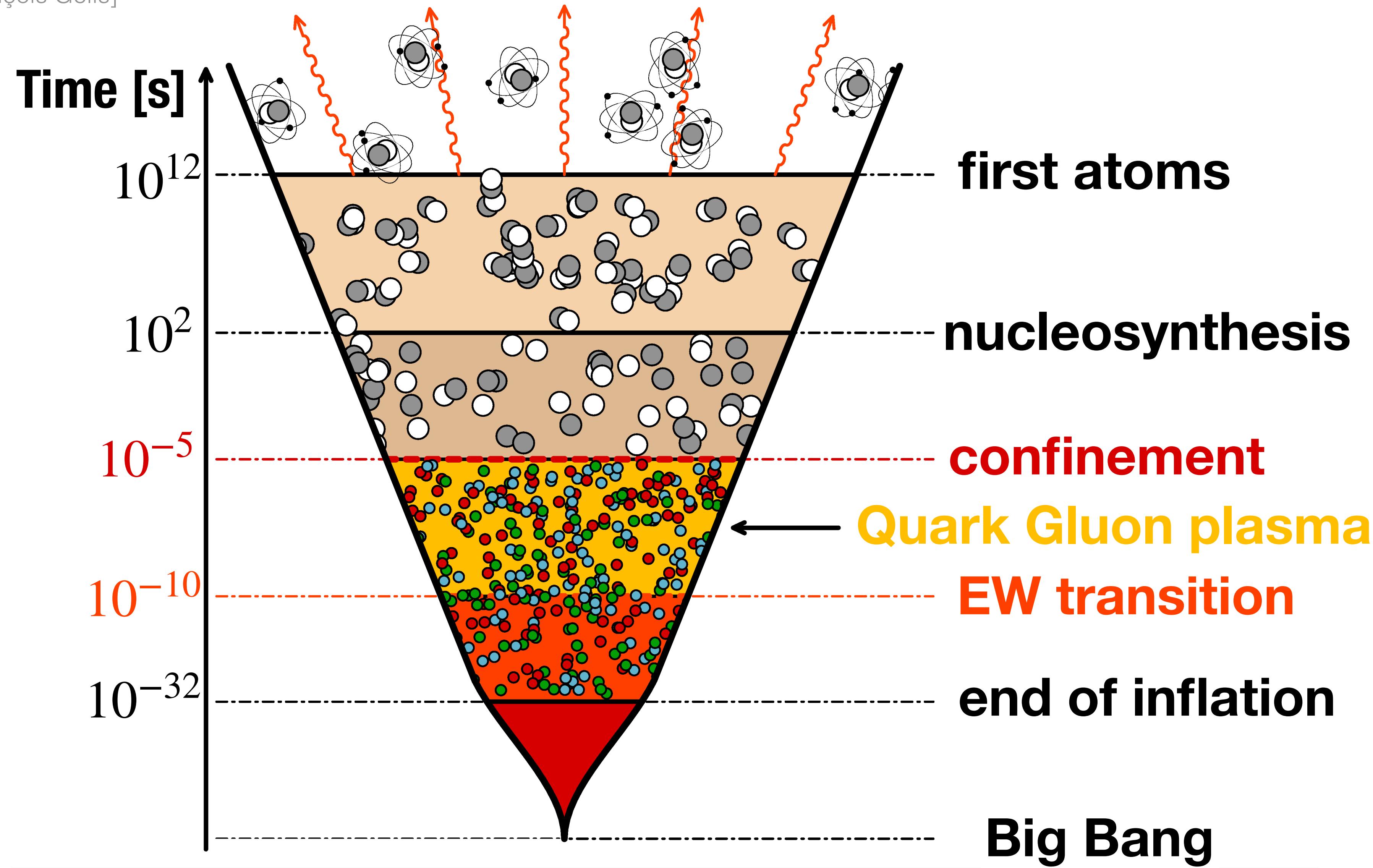


- QGP far from Stefan-Boltzmann limit
- Cross over transition (no sharp rise) from hadrons to quarks and gluons
- Curiosity: lattice result approaches strongly coupled N=4 SYM at high T

Critical temperature around $150 \text{ MeV} = 10^5 T_\odot = 2 \text{ billion degrees}$

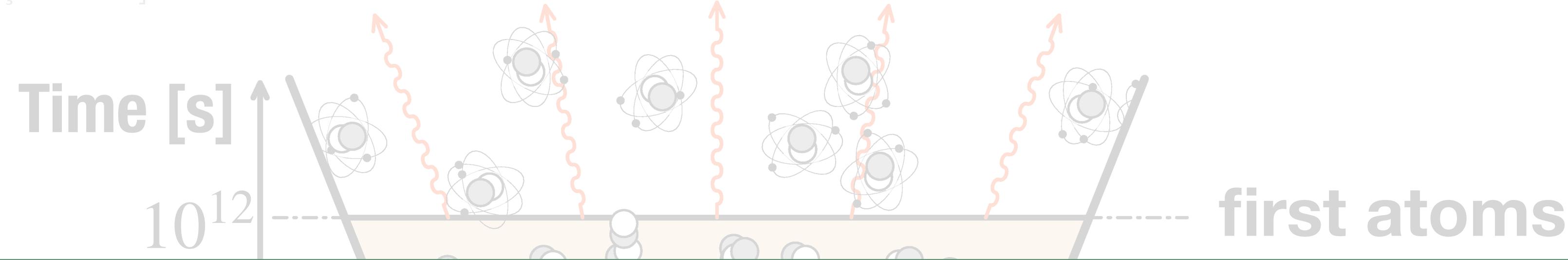
QGP in the early universe

[Adapted from François Gelis]

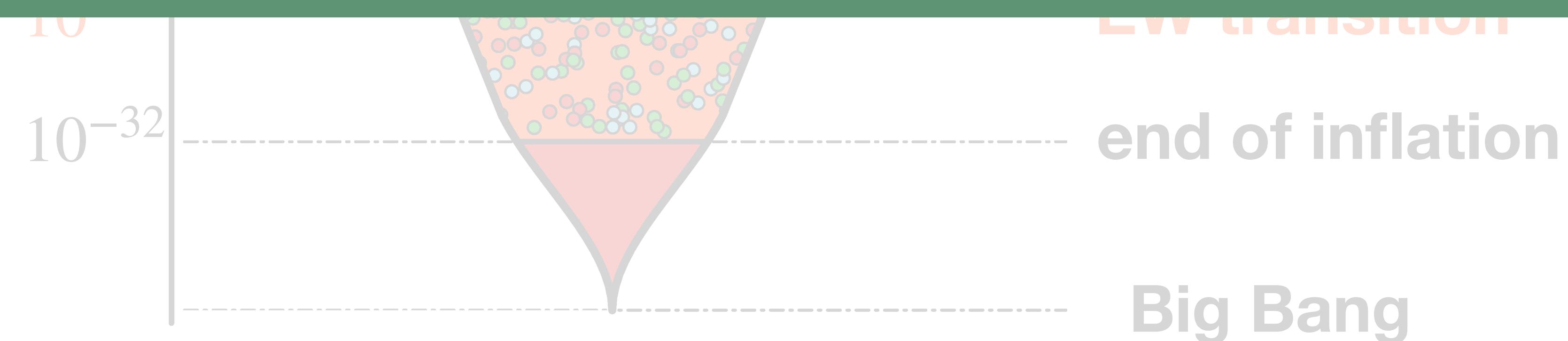


QGP in the early universe

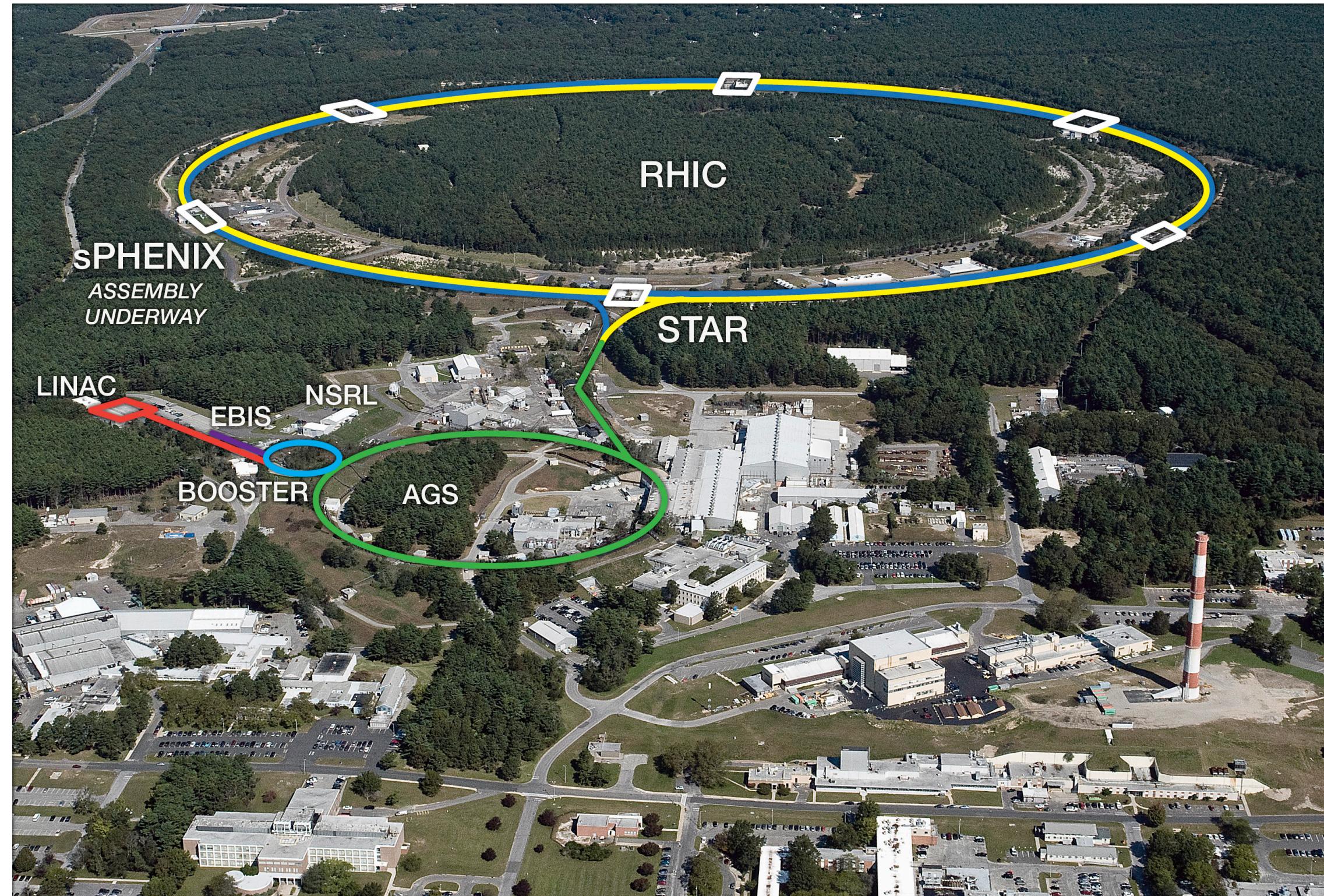
[Adapted from François Gelis]



Is it possible to recreate these high-temperatures in the lab? Yes!



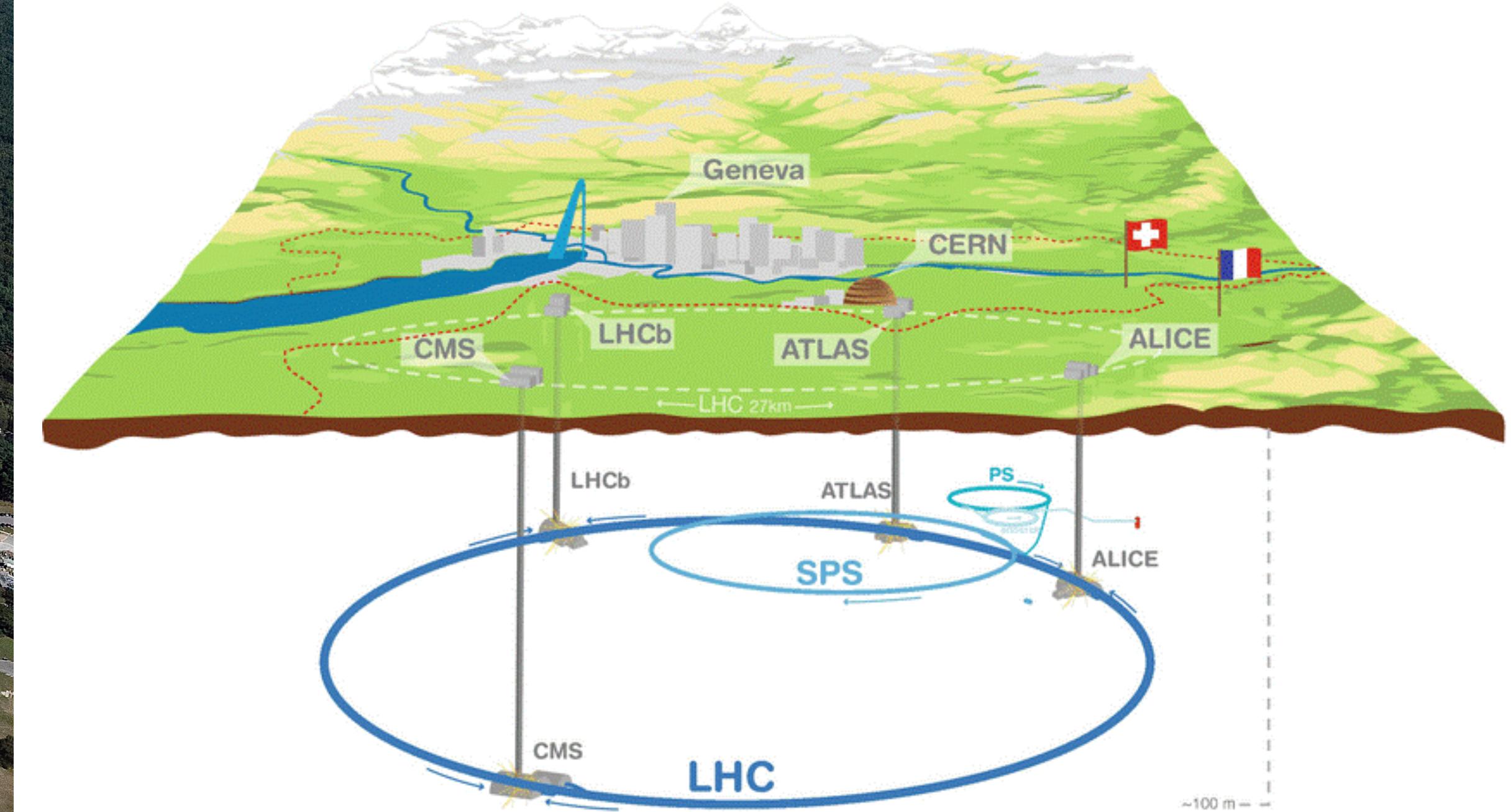
Ultra-relativistic heavy-ion colliders



Relativistic Heavy-Ion Collider

@ Brookhaven National Lab (New York)

p, Au, U, Zr, Ru @ $7.7 \leq \sqrt{s_{NN}} \leq 200$ GeV

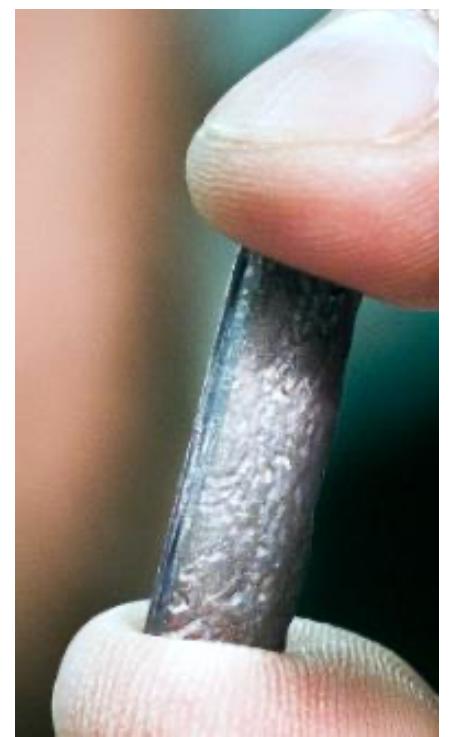


Large Hadron Collider

@ CERN (Geneva)

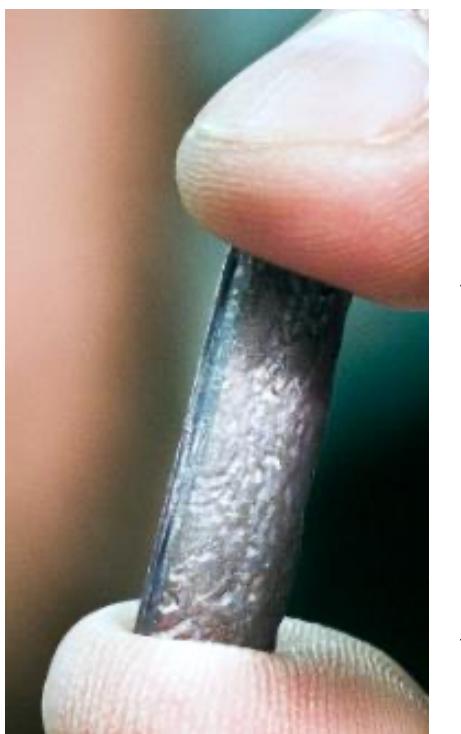
p, Pb, Xe, O @ $2.76 \leq \sqrt{s_{NN}} \leq 6$ TeV

How does a heavy-ion collider work in a nutshell



↑ 2 cm, 500 mg
of pure Lead
1\$/mg

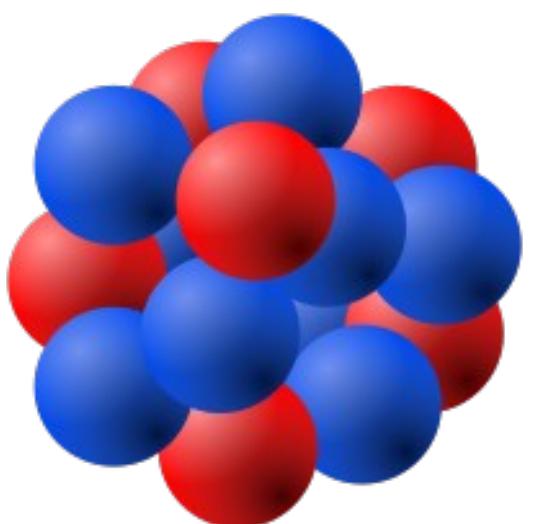
How does a heavy-ion collider work in a nutshell



2 cm, 500 mg
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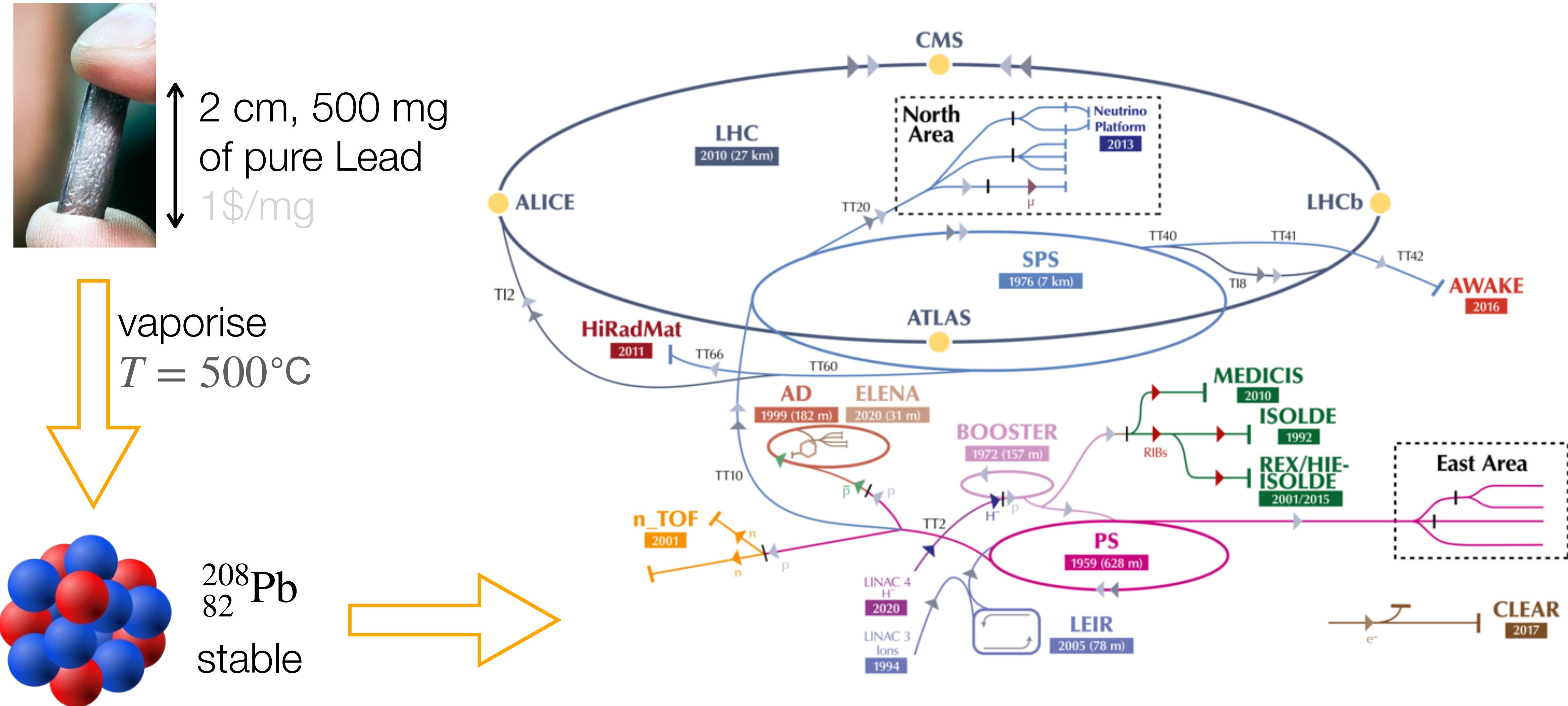


vaporise
 $T = 500^\circ\text{C}$



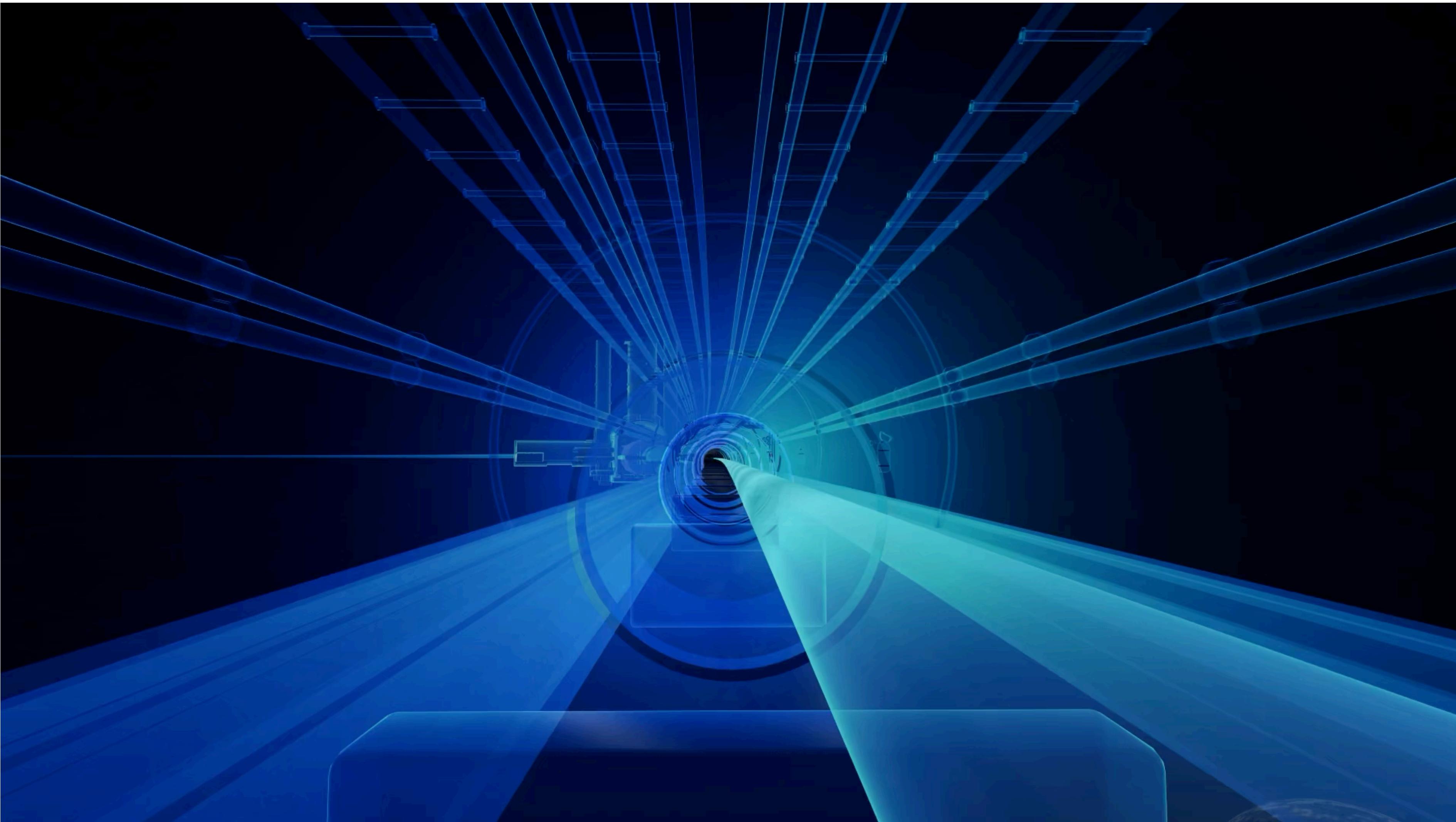
$^{208}\text{Pb}_{82}$
stable

How does a heavy-ion collider work in a nutshell



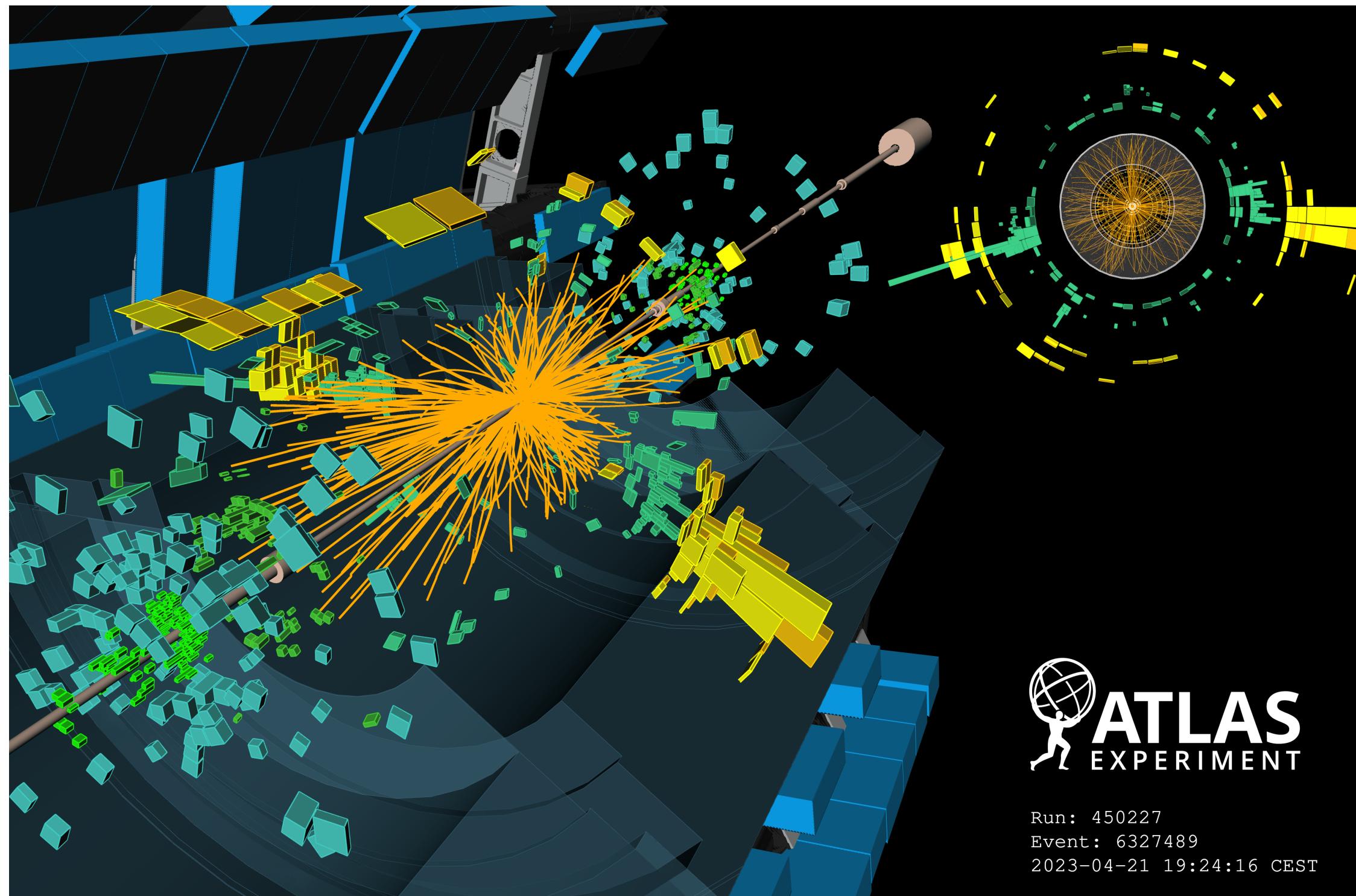
From the Big Bang to the Little Bang

<https://home.cern/resources/video/physics/heavy-ion-collision-event-animation>

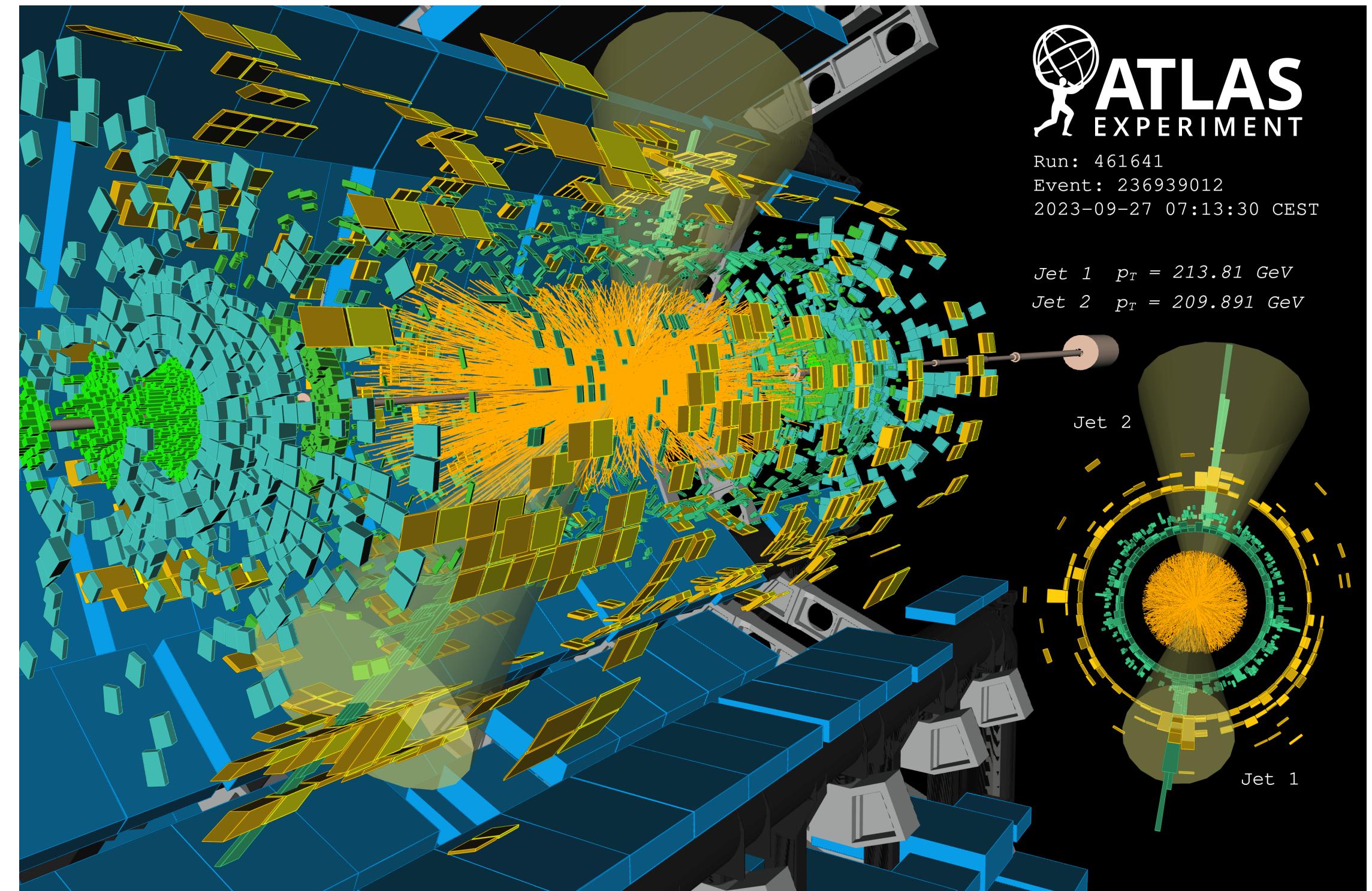


From the Big Bang to the Little Bang

pp@13.6 TeV

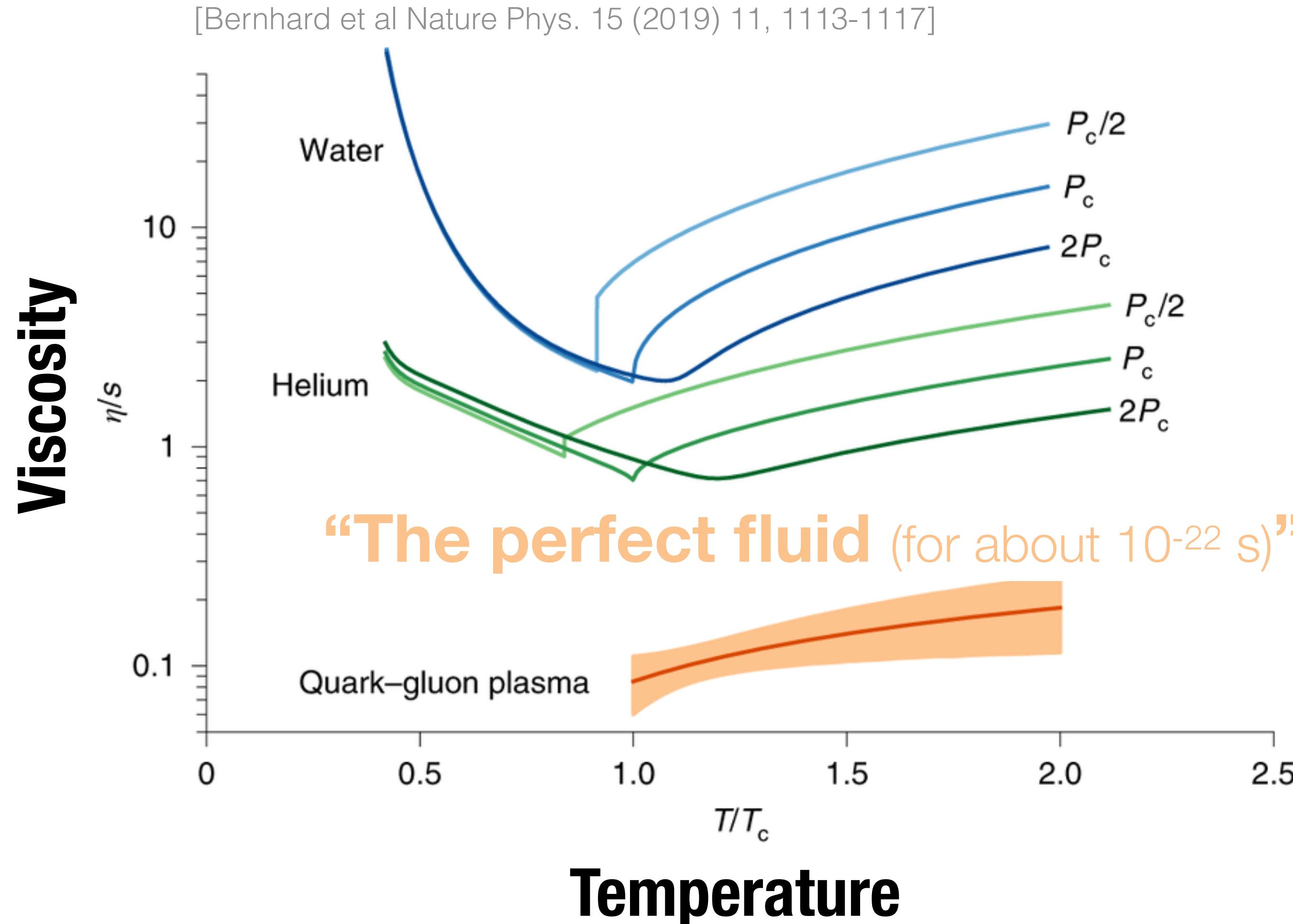


PbPb@5.36 TeV



Challenge: extract QGP properties out of the $\mathcal{O}(10^3)$ detected hadrons

Sneak peek into some of the most fascinating results



[STAR Collab. Nature 548 (2017) 62-65]



Sneak peek into some of the most fascinating unknowns

- How does the QGP emerge from an asymptotically free theory?
- What is the mechanism behind QGP thermalisation in yoctoseconds?
<https://igfae.usc.es/yoctolhc/project/>
- How small can a QGP droplet be?
- Why are many bulk phenomena so similar in pp, pPb and PbPb?
- Is there a critical point in the region explored by heavy-ion collisions?
- Which modifications does the QGP induce into a parton shower?