

FIG. 1. Momentum distribution of photons from positrons annihilating in metal-ammonia solutions, pure ammonia, and pure cesium. The distributions have been normalized to the same amplitude at 5.5 mrad. MNH_3 is a typical liquid-metal-ammonia solution. All other distributions are for solids and were measured close to liquid-nitrogen temperature. The insert shows the difference between the solution and solid ammonia. In the insert all distributions have been renormalized to the same peak height.

while the CsNH_3 solution was frozen close to the eutectic (about 17%).² Typical data runs accumulated about 20 000 counts per point in the vicinity of the peak. The drastic change produced by removing the solvated electrons is apparent when one realizes that the distribution for solid NaNH_3 (0.7%) was identical with that marked MNH_3 prior to freezing. No such change occurs when CsNH_3 is frozen. The angular distribution for pure cesium is included for comparison. The central portion of the cesium distribution is closely parabolic indicative of free electron behavior. At angles greater than 5.5 mrad the metal-ammonia curves are coincident with pure ammonia. The pure cesium distribution never completely matches the other curves. The relative heights of the peaks are not very accurate due to normalization at 5.5 mrad where the amplitude is quite low. For better comparison, the background from positrons annihilating in solid ammonia has been subtracted and the difference curves (narrow component in the customary terminology) renormalized to the same height. Note the very close similarity of solid CsNH_3 and solid LiNH_3 .

¹ R. A. Levy, Phys. Rev. **102**, 31 (1956).

² R. L. Schroeder, J. C. Thompson, and P. L. Oertel, Phys. Rev. **178**, 298 (1969).

³ J. A. Arias-Limonta and P. G. Varlashkin, J. Chem. Phys. **52**, 581 (1970).

⁴ J. A. Arias-Limonta and P. G. Varlashkin, Phys. Rev. B **1**, 142 (1970).

⁵ See, for example, *Positron Annihilation*, edited by A. T. Stewart and L. O. Roellig (Academic, New York, 1967).

⁶ P. G. Varlashkin and A. T. Stewart, Phys. Rev. **148**, 459 (1966).

⁷ P. G. Varlashkin, J. Chem. Phys. **49**, 3088 (1968).

Methane as a Numerical Experiment for Polarization Basis Function Selection

STEPHEN ROTHENBERG

Information Systems Design, 7817 Oakport Street, Oakland, California, 94621

AND

H. F. SCHAEFER III

Department of Chemistry, University of California, Berkeley, California 94720

(Received 9 November 1970)

The selection of basis functions for *ab initio* molecular calculations still seems to be more of an art than a science. This is due primarily to the fact that very complete work has only been done for diatomic molecules.¹ The significantly higher cost of calculations for polyatomic molecules has meant a relative scarcity of data.² Furthermore, the optimization of basis sets for polyatomic molecules *in situ* is too prohibitive to be done routinely and therefore, rules from a few representative cases appear to be the only realistic alternative.

A study was carried out on methane using the MOLE Quantum Chemistry System³ in order to examine the effect of basis function choice as a precursor to an SCF study of second-row hydrides.⁴ The data is of interest in itself, in particular, since it illustrates the comparative value of the various alternatives for SCF-type calculations and probably is valid for extended treatments as well.

Polarization functions are those functions which are necessary to represent the distortion (polarization) of the atom within the molecular environment; they are equivalent to higher angular momentum orbitals than are nominally occupied in the isolated atomic constituents. In the case of methane, the carbon 1s, 2s, and

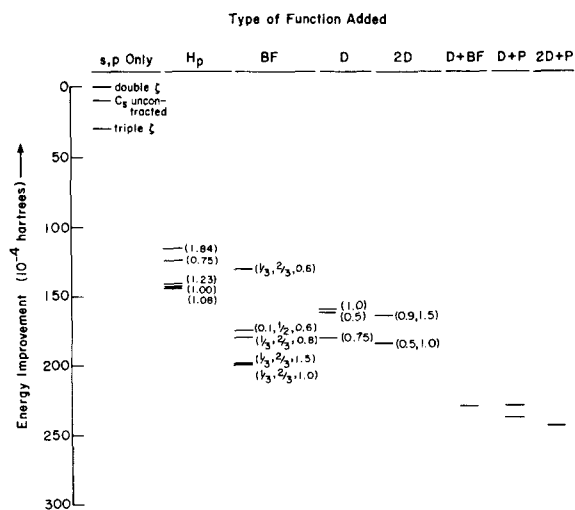


FIG. 1. Effect of polarization function selection for methane

TABLE I. Methane^a calculations.

| | Energy | Improvement ×10 ⁴ | Basis ^b | Comments |
|---|----------------------|---------------------------------|--|--|
| Reference calculations | | | | |
| 1. | 40.1282 | | (2s ² 1p ¹ /1s ¹) | Minimal STO optimization |
| 2. | 40.1672 | | (3s ¹⁰ 1p ⁵ /1s ⁵) | Whitten ^c double-zeta AO basis |
| 3. | 40.1801 | | X: (4s ¹⁰ 2p ⁵ /2s ⁵) | Double-zeta MO basis ^c |
| 4. | 40.1833 | 30 | (5s ¹⁰ 3p ⁵ /3s ⁵) | Dunning ^d contraction (triple-zeta) |
| | 40.1811 | 10 | (10s ¹⁰ 2p ⁵ /2s ⁵) | Carbon s completely free |
| Carbon d functions | | | | |
| 5. | 40.1963 | 162 | X+d _c (0.5) | Exponent values for d orbital shown in parentheses; Optimum carbon d exponent is 0.74. |
| 6. | 40.1982 | 181 | (0.75) | |
| 7. | 40.1961 | 160 | (1.00) | |
| 8. | 40.1975 | 164 | X+d _c (0.9, 1.5) | |
| 9. | 40.1986 | 185 | (0.5, 1.0) | |
| C-H bond functions | | | | |
| 10. | 40.1931 | 130 | X+BF(1/3, 2/3; 0.6) | Bond functions placed 1/3 and 2/3 of the distance down the C-H bond from the carbon with identical exponents of 0.6. |
| 11. | 40.1981 | 180 | 0.8 | |
| 12. | 40.2000 | 199 | 1.0 | |
| 13. | 40.1999 | 198 | 1.5 | Bond functions placed 0.1 bohr from carbon and 1/2 the distance down the C-H bond with identical exponents of 0.6, etc. |
| 14. | 40.1976 | 175 | X+BF(0.1, 1/2; 0.6) | |
| Hydrogen p functions | | | | |
| 15. | 40.1928 | 127 | x+p _H (0.75) | Optimum hydrogen p exponent 1.08 |
| 16. | 40.1943 | 142 | 1.00 | |
| 17. | 40.1944 | 143 | 1.08 | |
| 18. | 40.1942 | 141 | 1.23 | |
| 19. | 40.1916 | 115 | 1.84 | |
| 20. | 40.1983 ^e | 182 | (5s ¹⁰ 5p ⁵ /4s ⁴ 1p ¹) | |
| Carbon d functions and C-H bond functions | | | | |
| 21. | 40.2031 | 230 | X+d(1.0)+BF(1/3, 2/3; 1.0) | |
| Carbon d functions and hydrogen p functions | | | | |
| 22. | 40.2031 | 230 | X+d _c (1.0)+p _H (1.84) | |
| 23. | 40.2039 | 238 | X+d _c (0.75)+P _H (1.08) | |
| 24. | 40.2045 | 244 | X+d _c (0.5, 1.0)+P _H (1.0) | |
| Estimated ^f Hartree-Fock limit | | | | |
| 25. | 40.225±0.005 | 400-500 | | |

^a All methane calculations use a tetrahedral geometry with a C-H distance of 2.0665.

^b The notation for the basis functions given as $a s^b$ means *b* *s*-type primitive functions have been grouped to provide *a* basis orbitals.

^c Reference 2(b).

^d Reference 2(l).

^e Reference 2(c).

^f Previous estimate of the methane Hartree-Fock limit was given by C. Hollister and O. Sinanoglu [J. Am. Chem. Soc. **88**, 13 (1966)] as 40.22. Our analysis agrees with this value. We would estimate the main deficiency in our best calculation (#24) lies in the carbon *s*, *p*, and hydrogen *s* functions.

2*p* orbitals and the hydrogen 1*s* orbital are nominally the occupied atomic orbitals. Polarization functions available are *d*, *f*, *g*, etc., on carbon and *p*, *d*, *f*, etc., on hydrogen as well as any function located off one of the nuclear centers. This study is restricted to those

functions shown most significant¹ in diatomic studies, i.e., (a) 3*d* functions on carbon, (b) 2*p* functions on hydrogen, and finally to (c) bond functions, which in the present work are 1*s*-type functions along the C-H bond axis. The reference point for the measurement of

the effectiveness of the various polarization terms was chosen to be a basis set^{2(a),2(b),2(d)} that has double-zeta flexibility in the molecule and gave ^{2(b)} essentially double-zeta results in the carbon atom.

The results are shown in Table I as absolute energies and energies relative to the double-zeta reference calculation. Figure 1 provides a visual illustration of these results. The main conclusions from this data are:

(1) Bond functions provide a reasonable alternative to the more conventional type of polarization functions. In particular, the effect of eight bond functions (calculation No. 12) was not improved on by using fewer than 18 conventional polarization functions (calculation No. 23);

(2) a single *d* function accomplishes practically all that numerous *d* functions would;

(3) polarization is added effectively by all three abovementioned techniques [(a), (b), (c)] with hydrogen *p*'s being somewhat less effective (and less economical!);

(4) the use of a second kind of polarization in addition to any one of types [(a), (b), (c)] has about half⁵ the energetic contribution it would have had by itself;

(5) the results of various placements of the bond-functions along the bond axis (certainly somewhat arbitrary here) appears to be fairly insensitive to the exact location along the axis;

(6) polarization functions are energetically more important than relaxing the contraction of the *s* and *p* functions⁶;

We appreciate a number of helpful discussions with Dr. Ralph Young and Dr. Thom Dunning.

¹ (a) P. E. Cade, K. D. Sales, and A. C. Wahl, *J. Chem. Phys.* **44**, 1973 (1966); (b) P. E. Cade and W. H. Huo, *ibid.* **47**, 614, 649 (1967).

² (a) S. Huzinaga, *J. Chem. Phys.* **42**, 1293 (1965); (b) J. Whitten, *ibid.* **44**, 359 (1966); (c) C. D. Ritchie and H. F. King, *ibid.* **47**, 564 (1967); (d) H. Basch, M. B. Robin, and N. A. Kuebler, *ibid.* **47**, 1201 (1967); **49**, 5007 (1968); (e) G. P. Arrighini, C. Guidotto, M. Maestro, R. Moccia, and O. Slavetti, *ibid.* **49**, 224 (1968); (f) A. Veillard, *Theoret. Chim. Acta* **12**, 405 (1968); (g) H. Basch, C. J. Hornback, and J. W. Moskowitz, *J. Chem. Phys.* **41**, 1311 (1969); (h) W. J. Hehre, R. F. Stewart, and J. A. Pople, *ibid.* **51**, 2657 (1969); (i) S. Huzinaga and Y. Sakai, *ibid.* **50**, 1371 (1969); (j) J. D. Petke, J. L. Whitten, and A. W. Douglas, *ibid.* **51**, 256 (1969); (k) D. R. Whitman and C. J. Hornback, *ibid.* **51**, 398 (1969); (l) T. Dunning, *ibid.* **53**, 2823 (1970).

³ S. Rothenberg with P. Kollman, M. E. Schwartz, E. F. Hayes, and L. C. Allen, *Intern. J. Quantum Chem.* **3s**, 715 (1970).

⁴ S. Rothenberg, R. A. Young, and H. F. Schaefer, *J. Am. Chem. Soc.* **92**, 3243 (1970).

⁵ Arrighini *et al.* (Ref. 2(c)) have carried out an SCF study of methane with (5s3p1d/2s1p) STO basis. Their total energy is almost exactly the same as our (4s¹⁰2p⁵1d¹/2s⁵1p¹) Gaussian calculation. Their calculation showed an improvement of 167×10⁻⁴ due to the carbon *d* functions and 84×10⁻⁴ due to the hydrogen *p* functions, and in combination a contribution of 259×10⁻⁴ (i.e., an almost additive effect). The additive effect is probably due to their orbital exponents being somewhat different from more optimal values (compare with Table I nonoptimal calculations 7, 19, and 22).

⁶ Two calculations were done (see Table I) to see if additional flexibility in the *s* and *p* basis were of the same order of magnitude as the addition of polarization functions. Dunning's contractions (5s¹⁰3p⁵/3s⁵) [essentially triple-zeta flexibility and an almost saturated carbon *s* set (10s¹⁰2p⁵/2s⁵)] only gave energy improvements an order of magnitude less than use of polarization functions.

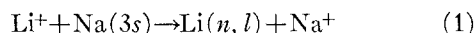
Simple Approximate Wavefunctions for Alkali Atoms

DONALD RAPP AND JEAN R. WARD

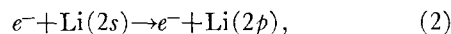
The University of Texas at Dallas, Dallas, Texas 75230

(Received 1 September 1970)

In calculations of cross sections for inelastic collision processes, it is often useful to have simple approximate one-electron wavefunctions for atoms. Slater orbitals,¹ for example, are simple nodeless functions that have been used in a variety of applications.² For more accurate work, self-consistent field functions are sometimes applied. However, these have several disadvantages, namely that they are not solutions of simple one-electron Schrödinger equations, and they do not give accurate values for the energy levels of the excited states. In treating the dynamics of charge transfer³ involving reactions such as



and in excitation by electron impact



great simplification results from representing Li and Na as (Li⁺+e⁻) and (Na⁺+e⁻), i.e., as one-electron atoms. The dynamical equations require that the wavefunction for the single electron on the ion satisfy a one-electron Schrödinger equation,

$$H\psi_n = E_n\psi_n, \quad (3)$$

where

$$H = (-\hbar^2/2m)\nabla^2 + V(r). \quad (4)$$

Furthermore, it is important³ that the energy levels of the functions be fairly accurate.

A simple way⁴ of doing this is to choose an effective potential for the single electron which depends on the shielding by the inner core electrons as an exponential function of distance:

$$\begin{aligned} V(r) &= -Z_{\text{eff}}(\epsilon^2/r) \\ &= -(\epsilon^2/r)[(Z-1)\exp(-r/\alpha)+1], \end{aligned} \quad (5)$$

where the effective nuclear charge,

$$Z_{\text{eff}}(r) = (Z-1)\exp(-r/\alpha)+1, \quad (6)$$

varies from the true nuclear charge *Z* at *r*=0, to 1 as