



Machine Learning for Optic Correction in the LHC

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Introduction

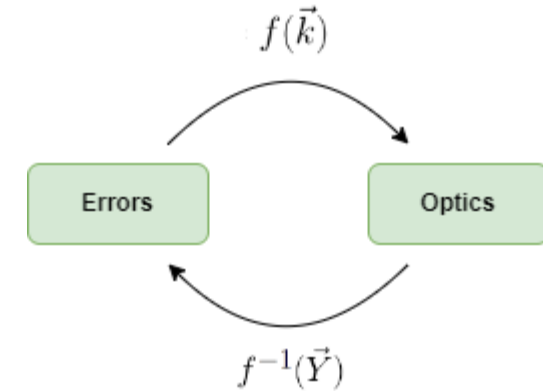
Elena's work continuation **"Supervised learning-based reconstruction of magnet errors in circular accelerators"** [1]

Actual methods for quadrupole magnetic error correction consist on knob tuning, response matrix ...

- Problems, no information on the actual magnet errors
- Correcting the optics, but not the actual magnet errors!

Machine learning promises multiple new ways to manage quadrupole errors

- The effect of magnet errors on optics can be calculated using simulation software
- ML can be used to model the relation between optics and errors regardless on how complex



$$\vec{Y} = f(\vec{k})$$

$$\vec{k} = f^{-1}(\vec{Y})$$

$$\vec{Y} = (\beta_x, \beta_y, \mu_x \dots)$$

Methods: Data pipeline

Generating random quadrupole strength errors and calculating the corresponding twiss parameters for data generation

Creating a ML model with:

- Input: Optic measurements, 3346 features
- Output: Quadrupole Magnet strength error

Possible improvements include:

- Using **MAD-NG** for non linear errors or model training
- Adding noise to simulated optic data, more realistic training data

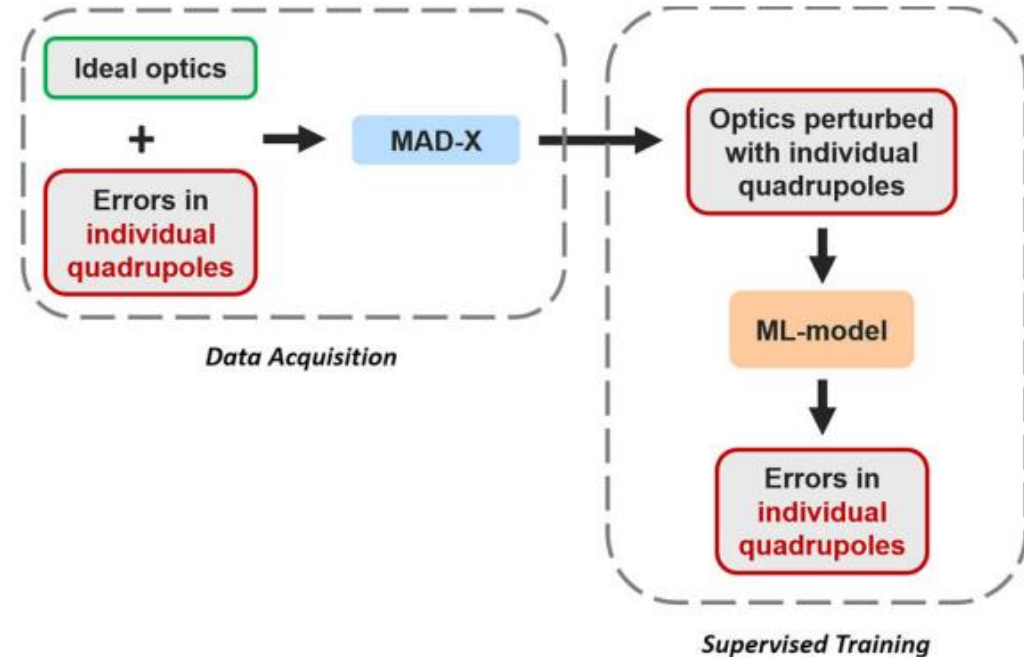


Fig 1. Data pipeline [1]

Methods: ML Model

Ridge and Linear Regression:

$$Loss = Error(Y - \hat{Y}) + \lambda \sum_1^n w_i^2$$

Least squares regression with L2 regularization

Bagging:

Training using ten different subsets of data and averaging the results.

This methods decreases variance of the model and overfitting

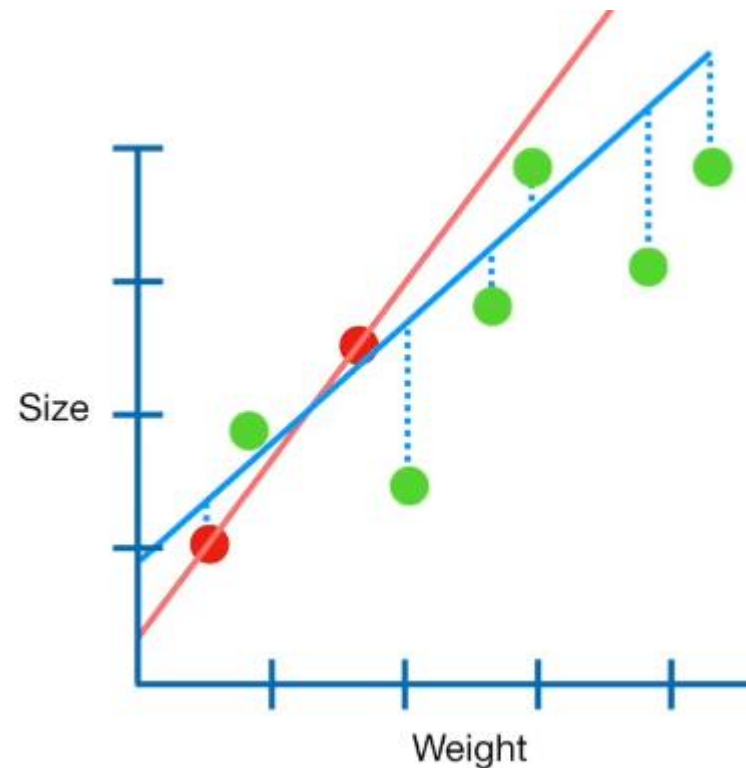


Fig 2. Example Ridge regression [2]

Results: Error Generation and Simulation

Previous setup, Elenas:

- 2016 40 CM optics
- MADX script with older python MADX wrapper
- Error generation Tab 1. and dipole errors according to best knowledge model but not used as input
- Matching tunes

Magnet	$\sigma_K / K_1 [10^{-4}]$	$\sigma_s [\text{mm}]$
MQ	19	...
MQX	4	6
MQY	11	...
MQM	12	...
MQW	15	...
MQT	75	...

New setup:

- 2023 45CM and 30CM optics
- CPYMAD instead of MADX
- Same error generation
- Matching tunes

*Tab 1. Error Generation
parameters obtained from WISE
[\[1\]](#) [\[3\]](#)*

Results: Error Generation and Simulation

- **70k Samples using 2023 30 CM optics**

Errors seem to be too big for this optics 15% of twiss failed!

- **70k Samples with 2023 45 CM optics**

In this case 1-2% of twiss failed, indicating that errors are probably too big for 30CM optics also with the “new” ATS arcs

- **All results shown are for 2023 30 CM optics**

Results: Error Simulation

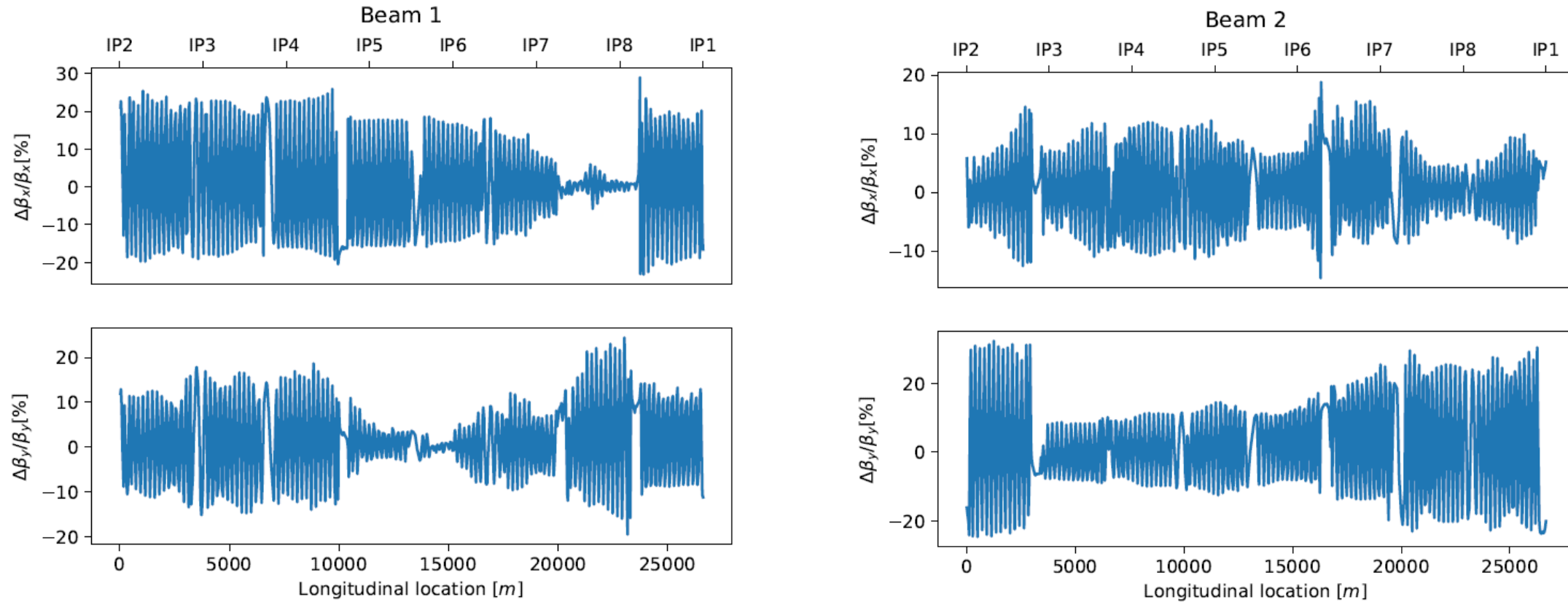


Fig 3. Example Beta Beating

Results: Model Training

Algorithm	Correlation Coefficient: R2	Mean Absolute Error: MAE
Ridge Regression	Train: 0.853	Train: 4.67e-06
	Test: 0.843	Test: 4.83e-06
Linear Regression	Train: 0.888	Train: 3.45e-06
	Test: 0.872	Test: 3.69e-06

Decision tree regression: Worse results

For 45CM optics results are similar

$$R^2(\mathbf{y}, \hat{\mathbf{y}}) = 1 - \frac{\text{Var}\{\mathbf{y} - \hat{\mathbf{y}}\}}{\text{Var}\{\mathbf{y}\}}$$

$$\text{MAE}(\mathbf{y}, \hat{\mathbf{y}}) = \frac{1}{n} \sum_{i=1}^n |y_i - \hat{y}_i|$$

Results: Model Training

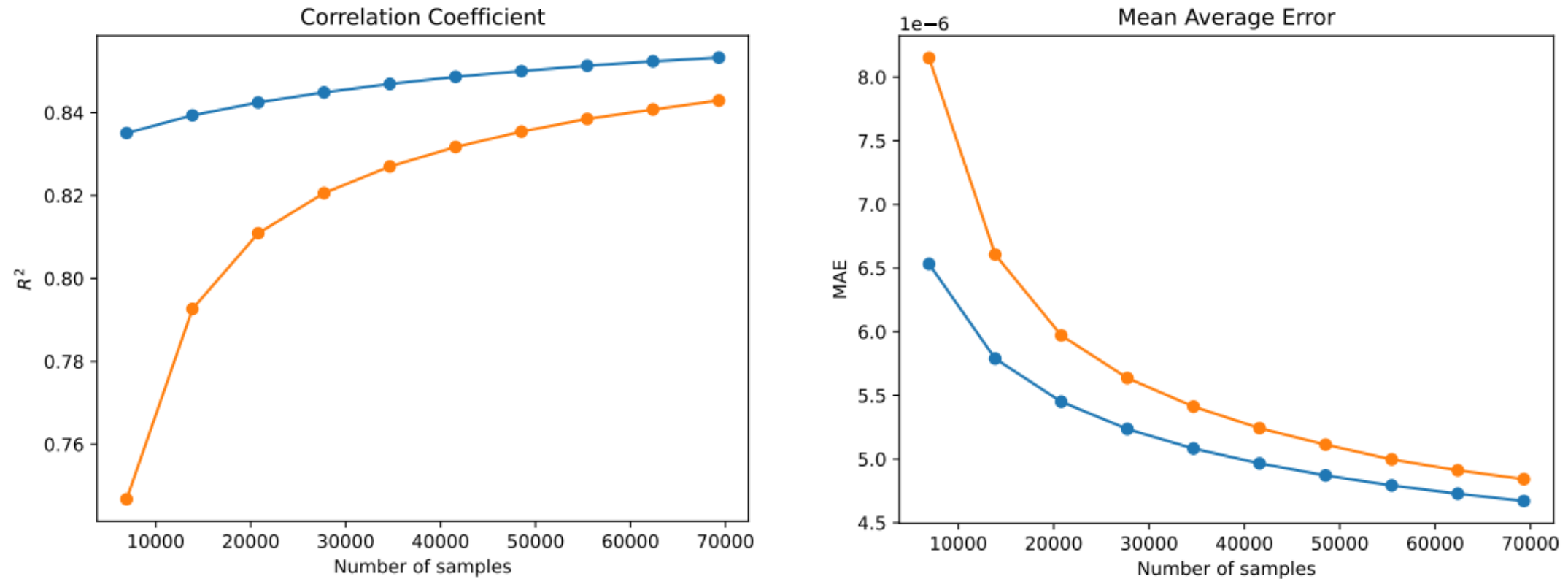


Fig 4. Ridge regression training for different number of samples

Results: Model Training

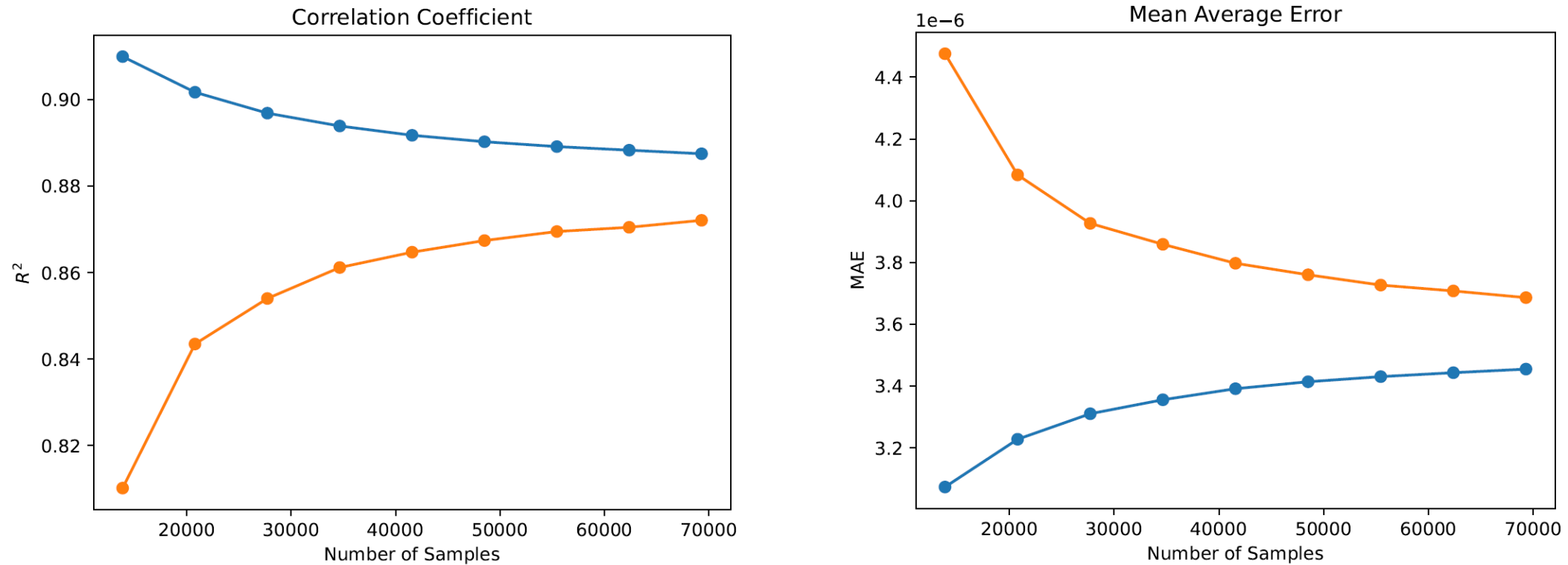


Fig 5. Linear regression training for different number of samples

Results: Example prediction

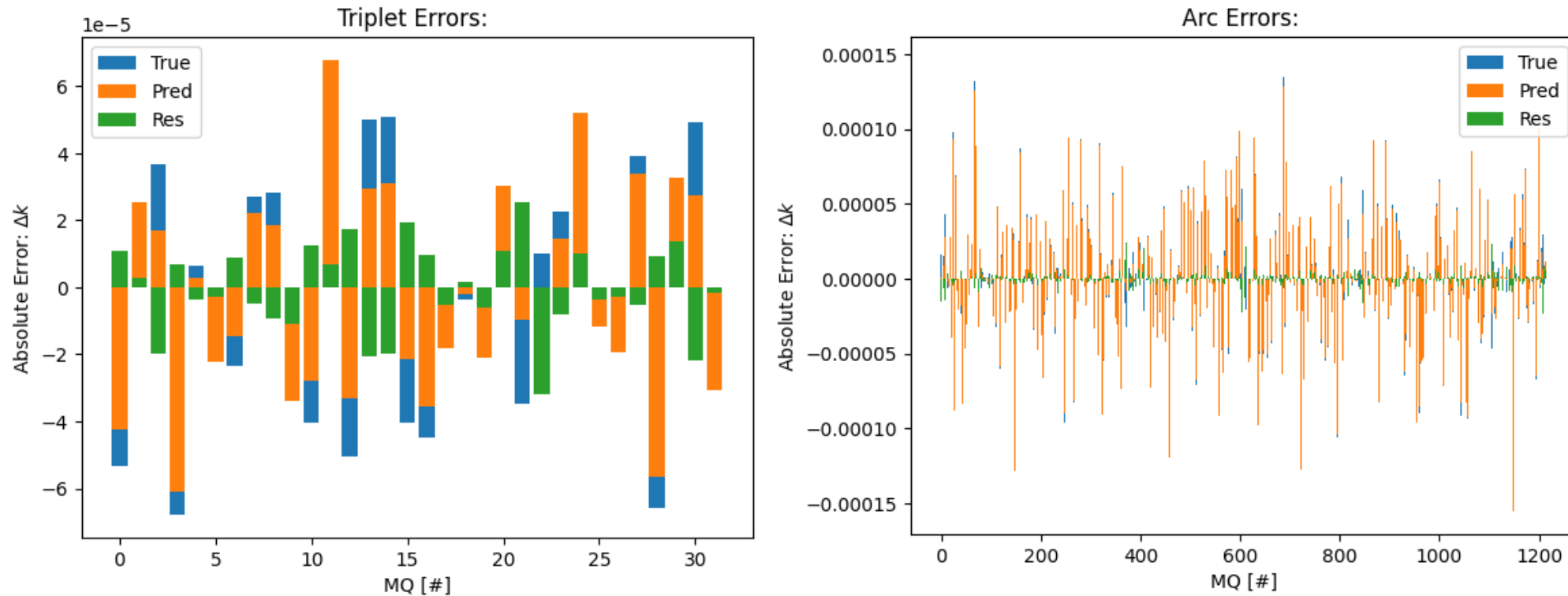
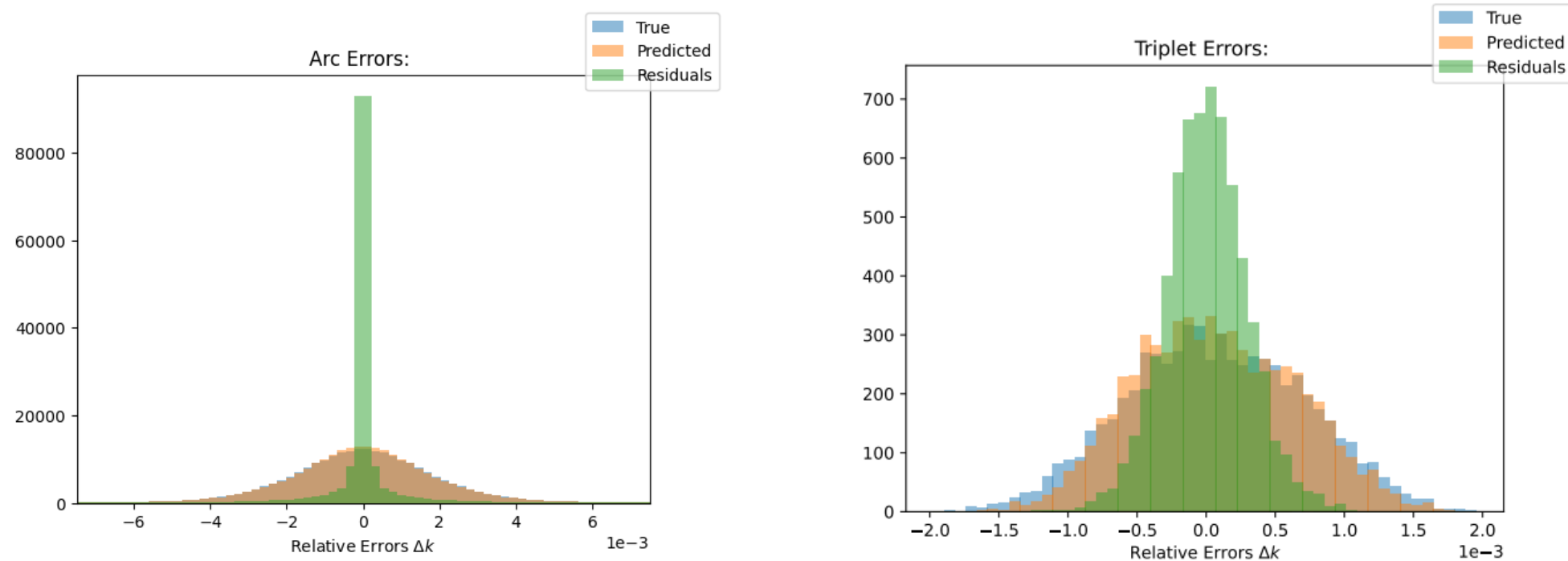


Fig 6. Quadrupole error prediction using linear reg. (random sample)

Results: Error Histogram Prediction



*Fig 7. Relative error histograms
for 200 test samples*

Results: Worse performing samples

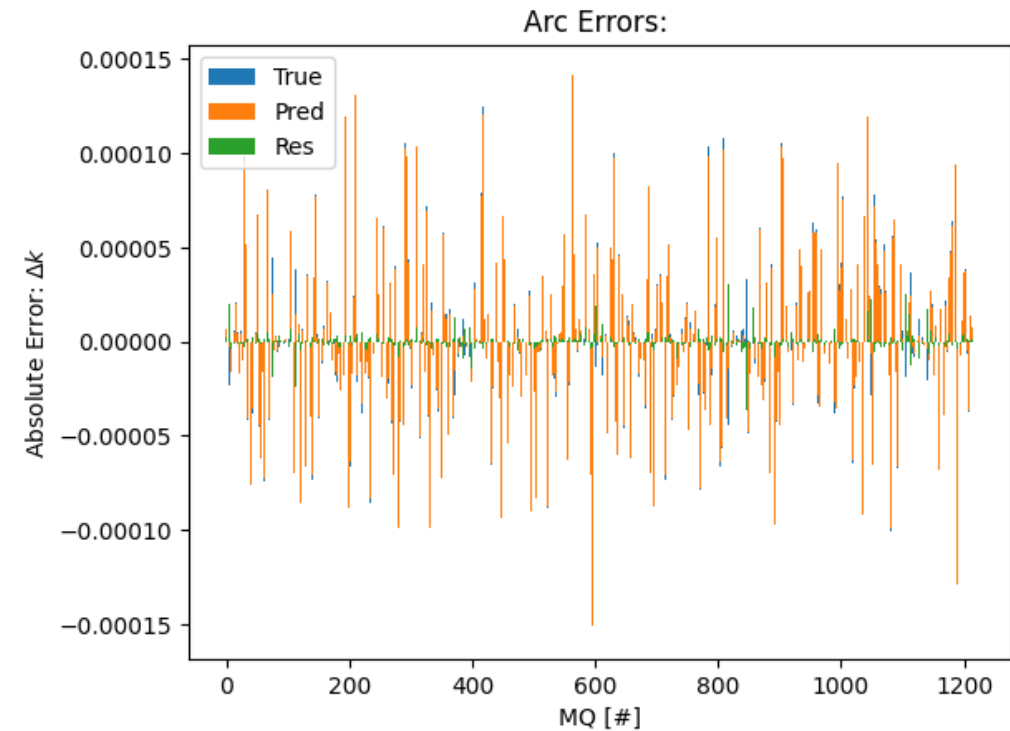
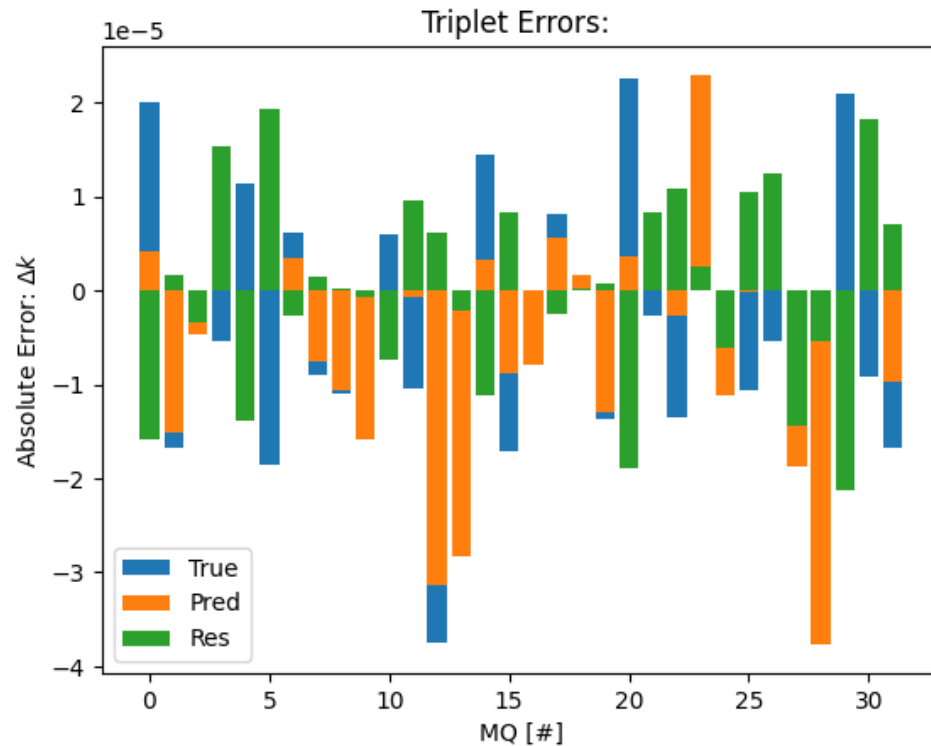


Fig 8. Example of worse performing sample (Triplets!)

Results: Performance for different magnets

	Magnets	Correlation Coefficient: R2	Mean Absolute Error: MAE
Triplet		Train: 0.853	Train: 4.67e-06
		Test: 0.843	Test: 4.83e-06
Arc + Triplet		Train: 0.888	Train: 3.45e-06
		Test: 0.872	Test: 3.69e-06

Unexpected! R2 slightly worse or similar for triplet errors, expected better R2

MAE Is not a great indicator since error generation is different for triplet and arc

Results: 45CM Optics

Triplet quadrupole performance is worse than in arc. Maybe the errors generated for 40CM in 2016 are too big for 30CM. Testing for 45 CM optics.

Train: $R^2 = 0.895$ $MAE = 3.118e-06$
Test: $R^2 = 0.884$ $MAE = 3.292e-06$

R^2 Test 0.872 for 30CM

Similar results, hypothesis is wrong.

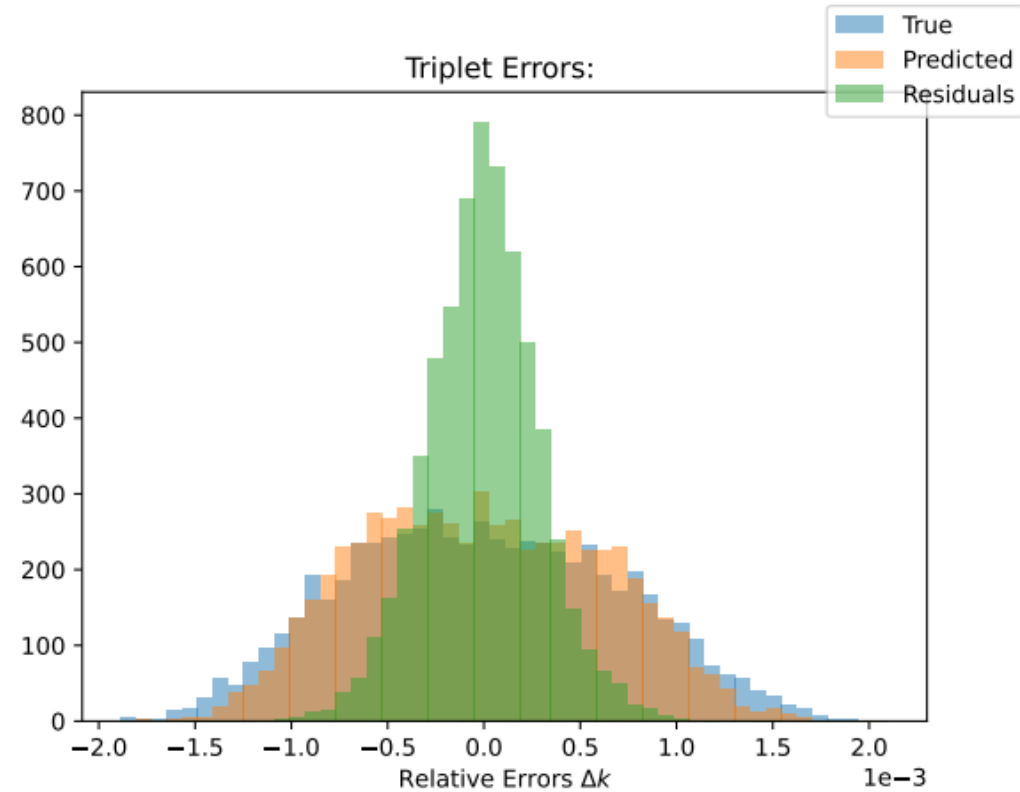


Fig 9. Quadrupole error histograms for 200 test samples

Conclusion

- **Improvement in simulation with respect to 2016, maybe because of the decrease in degeneracy in the arc magnets due to change to ATS optics**
- **Linear regression shows better results than ridge**
- **Triplet errors prediction is more challenging than arc magnet error prediction in our case obtaining worse samples and worse R^2 , this might be because of the “new” ATS optics, bigger effect and less degeneracy for arc magnets could explain better performance**
- **Testing the model on real world data is the next step and most important!**

References

- [1] Fig 1. Data pipeline. **"Supervised learning-based reconstruction of magnet errors in circular accelerators"** by **E. Fol**, 2021, <https://doi.org/10.1140/epjp/s13360-021-01348-5>
- [2] Fig 2. Example Ridge regression. **"Regularization Part 1: Ridge (L2) Regression"** by **J. Stamer**, 2018, <https://www.youtube.com/watch?v=Q81RR3yKn30>
- [3] Tab 1. P.Hagen, M.Giovannozzi, J.-P.Koutchouk, T.Risselada, F.Schmidt, E.Todesco, and E.Wildner, **"WISE: A Simulation of the LHC Optics including Magnet Geometrical Data, LHC-Project-Report-1123"**, 2008, <https://cds.cern.ch/record/1123714>

Backup Slides: Error generation macros

"All quadrupoles in the lattice are assigned a random relative gradient error obtained from uniform distribution with the same rms error σ per magnets family."

Rr = 0.017
GCUTR = 3
B2r = Tab. 1

```
SetEfcomp_Q: macro = {  
  Efcomp, radius = Rr, order= 1,  
  dknr:={0,  
    1E-4*( B2s *ON_B2S + B2r *ON_B2R * TGAUSS(GCUTR)),  
    1E-4*(-B3s *ON_B3S + B3r *ON_B3R * TGAUSS(GCUTR)),  
    1E-4*( B4s *ON_B4S + B4r *ON_B4R * TGAUSS(GCUTR)),  
    1E-4*(-B5s *ON_B5S + B5r *ON_B5R * TGAUSS(GCUTR)),  
    1E-4*( B6s *ON_B6S + B6r *ON_B6R * TGAUSS(GCUTR)),  
    1E-4*(-B7s *ON_B7S + B7r *ON_B7R * TGAUSS(GCUTR)),  
    1E-4*( B8s *ON_B8S + B8r *ON_B8R * TGAUSS(GCUTR)),  
    1E-4*(-B9s *ON_B9S + B9r *ON_B9R * TGAUSS(GCUTR)),  
    1E-4*( B10s *ON_B10S + B10r *ON_B10R * TGAUSS(GCUTR)),  
    1E-4*(-B11s *ON_B11S + B11r *ON_B11R * TGAUSS(GCUTR))},  
  dksr:={0,  
    1E-4*(-A2s *ON_A2S + A2r *ON_A2R * TGAUSS(GCUTR)),  
    1E-4*( A3s *ON_A3S + A3r *ON_A3R * TGAUSS(GCUTR)),  
    1E-4*(-A4s *ON_A4S + A4r *ON_A4R * TGAUSS(GCUTR)),  
    1E-4*( A5s *ON_A5S + A5r *ON_A5R * TGAUSS(GCUTR)),  
    1E-4*(-A6s *ON_A6S + A6r *ON_A6R * TGAUSS(GCUTR)),  
    1E-4*( A7s *ON_A7S + A7r *ON_A7R * TGAUSS(GCUTR)),  
    1E-4*(-A8s *ON_A8S + A8r *ON_A8R * TGAUSS(GCUTR)),  
    1E-4*( A9s *ON_A9S + A9r *ON_A9R * TGAUSS(GCUTR)),  
    1E-4*(-A10s *ON_A10S + A10r *ON_A10R * TGAUSS(GCUTR)),  
    1E-4*( A11s *ON_A11S + A11r *ON_A11R * TGAUSS(GCUTR))};  
}
```

Backup Slides: Tuning for Ridge reg

