



Computational Fluid Dynamics

Lecture 10
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The Euler Equations



The Euler equations for 1D flow:

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u(E + p/\rho) \end{pmatrix} = 0 \quad p = (\gamma - 1)\rho e$$

where $E = e + u^2/2$ and we define $H = h + u^2/2$; $h = e + p/\rho$

Ideal Gas: $p = \rho RT$; $e = e(T)$; $c_v = de/dT$

$$h = h(T); \quad c_p = dh/dT; \quad h - e = RT$$

$$R = h/T - e/T = c_p - c_v;$$

$$\gamma = c_p/c_v;$$

$$\text{Pressure and energy are related by } p = \rho(h - e) = (h/e - 1)\rho e = (\gamma - 1)\rho e$$

Speed of sound

$$c^2 = \left(\frac{\partial p}{\partial \rho} \right)_s = \gamma RT = \gamma \frac{p}{\rho}$$



$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u(E + p/\rho) \end{pmatrix} = 0 \quad p = (\gamma - 1)\rho e$$

$$c^2 = \gamma p / \rho$$

Expanding the derivative and rearranging the equations

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ p \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} u & \rho & 0 \\ 0 & u & 1/\rho \\ 0 & \rho c^2 & u \end{pmatrix} \begin{pmatrix} \rho \\ \rho u \\ p \end{pmatrix} = 0 \quad \text{or} \quad \frac{\partial \mathbf{f}}{\partial t} + \mathbf{A} \frac{\partial \mathbf{f}}{\partial x} = 0$$

The Characteristics for the Euler Equations are found by finding the eigenvalues for \mathbf{A}^T

$$\det(\mathbf{A}^T - \lambda \mathbf{I}) = 0$$



Find

$$\det(\mathbf{A}^T - \lambda \mathbf{I}) = 0 \quad \text{or} \quad \det \begin{pmatrix} u - \lambda & 0 & 0 \\ \rho & u - \lambda & \rho c^2 \\ 0 & 1/\rho & u - \lambda \end{pmatrix} = 0$$

$$(u - \lambda)((u - \lambda)^2 - c^2) = 0$$

Therefore

or:

$$(u - \lambda) = 0 \Rightarrow \lambda = u$$

$$\lambda_1 = u - c;$$

$$((u - \lambda)^2 - c^2) = 0$$

$$\lambda_2 = u + c;$$

$$\Rightarrow u - \lambda = \pm c$$

$$\lambda_3 = u$$

$$\Rightarrow \lambda = u \pm c$$



The Rankine-Hugoniot conditions

For a hyperbolic system

$$\frac{\partial \mathbf{f}}{\partial t} + s \frac{\partial \mathbf{F}}{\partial x} = 0$$

The speed of a discontinuity (shock) is found by moving to a frame where a shock moving with speed s is stationary

$$\cancel{\frac{\partial \mathbf{f}}{\partial t}}^0 + s \frac{\partial \mathbf{f}}{\partial x} + \frac{\partial \mathbf{F}}{\partial x} = 0$$

Integrating across the shock yields

$$s[\mathbf{f}] = [\mathbf{F}]$$



Computational Fluid Dynamics Exact Solution:

For the Euler equations:

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u(E + p/\rho) \end{pmatrix} = 0$$

The Rankine-Hugoniot conditions are:

$$s(\rho_L - \rho_R) = ((\rho u)_L - (\rho u)_R)$$

$$s((\rho u)_L - (\rho u)_R) = ((\rho u^2 + p)_L - (\rho u^2 + p)_R)$$

$$s((\rho E)_L - (\rho E)_R) = ((\rho u(E + p/\rho))_L - (\rho u(E + p/\rho))_R)$$

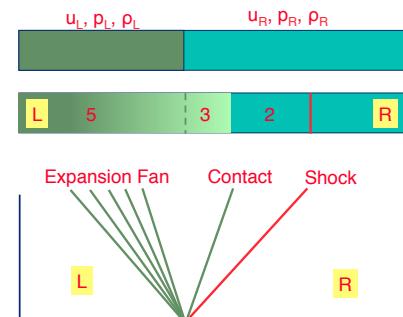


Computational Fluid Dynamics

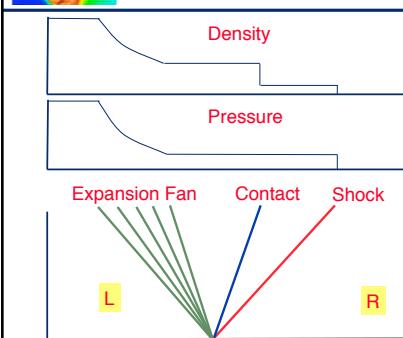
The Shock-Tube Problem Exact Solution



Computational Fluid Dynamics The shock tube problem



Computational Fluid Dynamics



Computational Fluid Dynamics Exact Solution:

$$\text{given: } p_L, \rho_L, u_L, \quad p_R, \rho_R, u_R \quad \alpha = \frac{\gamma+1}{\gamma-1}$$

$$c_L = \sqrt{\gamma p_L / \rho_L} \quad c_R = \sqrt{\gamma p_R / \rho_R}$$

Consider the case $p_L > p_R$:
 Shock separates R and 2
 Contact discontinuity separates 2 and 3
 Expansion fan separates 3 and L



Computational Fluid Dynamics Exact Solution:

$$\text{given: } p_L, \rho_L, u_L, \quad p_R, \rho_R, u_R$$

The Rankine-Hugoniot conditions give a nonlinear relation for the pressure jump across the shock $P = \frac{p_2}{p_R}$

$$P = \frac{p_L}{p_R} \left[1 - \frac{(\gamma-1)(c_R/c_L)(P-1)}{\sqrt{2\gamma(2\gamma+(\gamma+1)(P-1))}} \right]^{2\gamma/(\gamma-1)}$$

which can be solved by iteration

CFD Computational Fluid Dynamics
Exact Solution:

The speed of the shock and velocity behind the shock are found using RH conditions:

$$s_{\text{shock}} = u_R + c_R \left(\frac{\gamma - 1 + (\gamma + 1)P}{2\gamma} \right)^{1/2}$$

$$x_{\text{shock}} = x_0 + s_{\text{shock}} t$$

The speed of the contact is

$$s_{\text{contact}} = u_3 = u_2 = u_L + \frac{2c_L}{\gamma - 1} \left(1 - \left(P \frac{p_R}{p_L} \right)^{\frac{\gamma-1}{2\gamma}} \right)$$

$$x_{\text{contact}} = x_0 + s_{\text{contact}} t$$

The left hand side of the fan moves with speed

$$s_{jL} = -c_L$$

$$x_{jL} = x_0 + s_{jL} t$$

The right hand side of the fan moves with speed

$$s_{jR} = u_2 - c_3$$

$$x_{jR} = x_0 + s_{jR} t$$

CFD Computational Fluid Dynamics
Exact Solution:

$$P = \frac{p_2}{p_R}$$

Left uniform state given: $p_L, \rho_L, u_L,$	In the expansion fan (4) given: $p_L, \rho_L, u_L,$	Behind the contact (3) $p_3 = p_2$ $u_3 = u_2$
$c_L = \sqrt{\gamma p_L / \rho_L}$	$\rho_4 = \rho_4 \left(\frac{1 + (s-x)}{c_L \alpha t} \right)^{\frac{2\gamma}{\gamma-1}}$ $u_4 = u_L + \frac{2}{\gamma+1} \left(\frac{s-x}{t} \right)$	$\rho_3 = \rho_L \left(P \frac{p_R}{p_L} \right)^{1/\gamma}$
	$P_4 = P_4 \left(\frac{1 + (s-x)}{c_L \alpha t} \right)^{\frac{2\gamma}{\gamma-1}}$	

CFD Computational Fluid Dynamics
Exact Solution:

$$\alpha = \frac{\gamma + 1}{\gamma - 1} \quad P = \frac{p_2}{p_R}$$

Behind the contact (3) $p_3 = p_2$ $u_3 = u_2$ $\rho_3 = \rho_L \left(P \frac{p_R}{p_L} \right)^{1/\gamma}$	Behind the shock (2) $\rho_2 = \left(\frac{1 + \alpha P}{\alpha + P} \right) \rho_R$ $p_2 = P p_R$ $u_2 = u_L + \frac{2c_L}{\gamma - 1} \left(1 - \left(P \frac{p_R}{p_L} \right)^{\frac{\gamma-1}{2\gamma}} \right)$	Right uniform state given: P_R, ρ_R, u_R
		$c_R = \sqrt{\gamma p_R / \rho_R}$

CFD Computational Fluid Dynamics
Exact Solution:

The velocity relative to the speed of sound defines the flow regime

$$c = \sqrt{\frac{\gamma p}{\rho}}$$

The Mach number is defined as the ratio of the local velocity over the speed of sound

$$Ma = \frac{u}{c}$$

$Ma < 1$ subsonic	$Ma > 1$ supersonic
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CFD Computational Fluid Dynamics
Exact Solution:

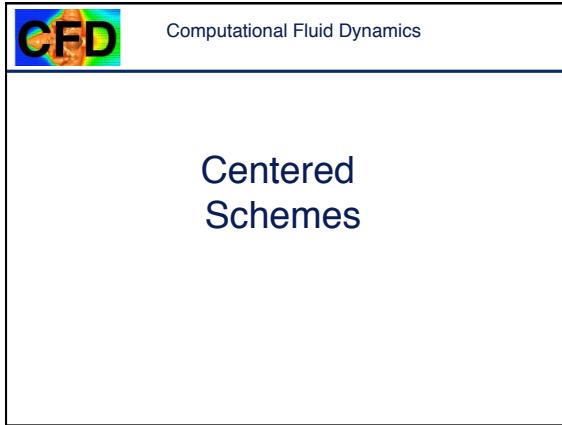
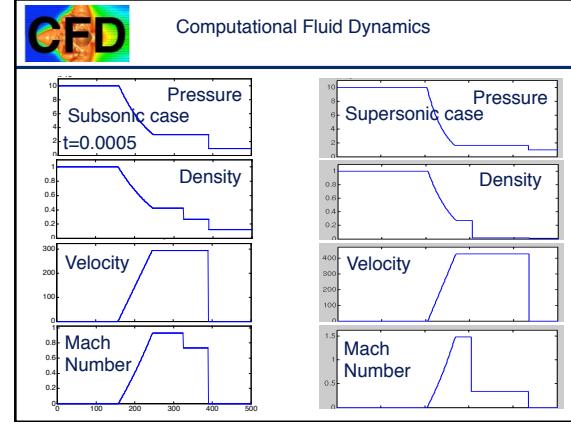
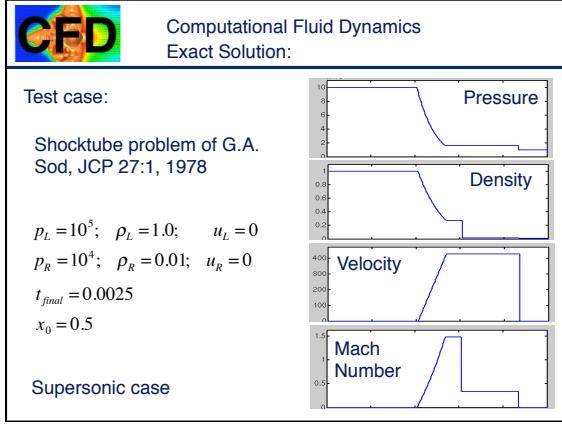
Test case:
Shocktube problem of G.A. Sod, JCP 27:1, 1978

$p_L = 10^5; \rho_L = 1.0; u_L = 0$
 $p_R = 10^4; \rho_R = 0.125; u_R = 0$

$t_{\text{final}} = 0.005$
 $x_0 = 0.5$

Subsonic case

CFD Computational Fluid Dynamics
Exact Solution:



CFD Computational Fluid Dynamics Flux Splitting

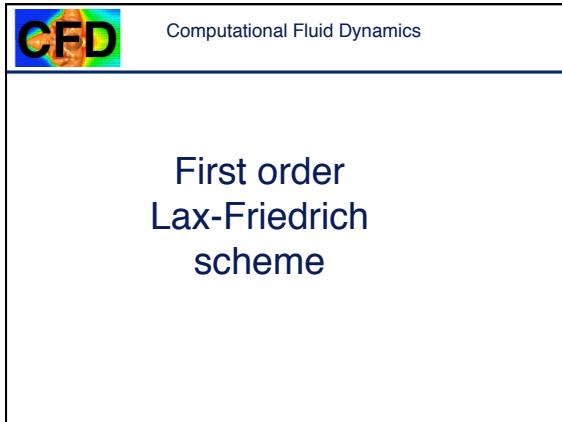
For a system of equations the characteristics often have different slopes and we therefore have different "upwind" directions

The simplest approach to solve the Euler equations is by using centered schemes where it is not necessary to take into account the characteristic direction

We will examine two such schemes:

The first order Lax-Friedrich scheme and

The second order Lax-Wendroff scheme



CFD Computational Fluid Dynamics Exact Solution:

Write the Euler equations as

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}}{\partial x} = 0$$

Update the solution by

$$\mathbf{f}_j^{n+1} = \frac{1}{2}(\mathbf{f}_{j+1}^n + \mathbf{f}_{j-1}^n) - \frac{1}{2} \frac{\Delta t}{h} (\mathbf{F}_{j+1}^n - \mathbf{F}_{j-1}^n)$$

The error term contains an term including $h^2/\Delta t$ but by taking $h \sim \Delta t$ we have a first order scheme

The scheme is often written as

$$\mathbf{f}_j^{n+1} = \mathbf{f}_j^n - \frac{\Delta t}{h} (\mathbf{F}_{j+1/2}^n - \mathbf{F}_{j-1/2}^n)$$

where $\mathbf{F}_{j+1/2} = \frac{1}{2}(\mathbf{F}_{j+1} + \mathbf{F}_j) - \frac{1}{2} \frac{h}{\Delta t} (\mathbf{f}_{j+1}^n - \mathbf{f}_j^n)$

CFD Computational Fluid Dynamics

```
%Euler equation-one-dimensional Lax-Fredrich
nx=256; maxstep=150;
gg=1.4; Jeft=100000*p; right=100000; Jeft=1; right=0.125*u; left=0;
x=10.0*pi*(nx-1); time=0; for i=1:nx;x(i)=i*(t-1);end

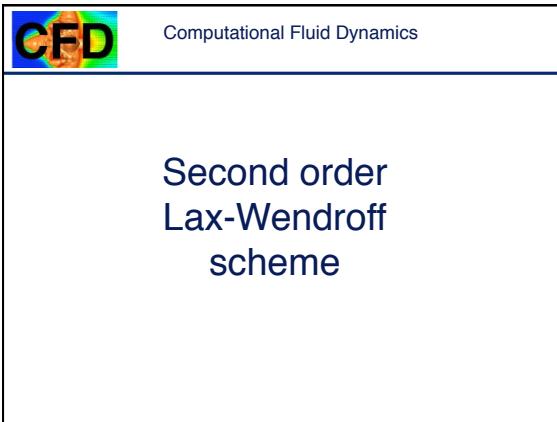
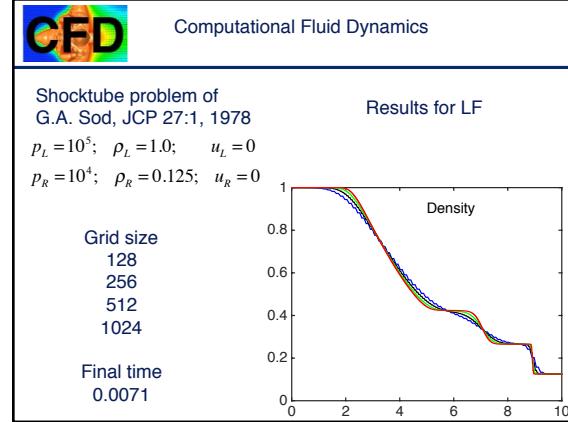
r=zeros(1,nx);ru=zeros(1,nx);re=zeros(1,nx);
rm=zeros(1,nx);ru=zeros(1,nx);rE=zeros(1,nx);
c=zeros(1,nx);ru=zeros(1,nx);m=zeros(1,nx);

for i=1:nx;r(i)=Jeft*right;ru(i)=0.0;re(i)=p*right/(gg-1);end
for i=1:nx/2; r(i)=Jeft;
    rE(i)=(p_left*(gg-1)+0.5*r_left*u_left)/2;
    ru(i)=r_left*u_left;
end

rh=r;ru=ru;rE=rE;

cmax=sqrt(max(gg*p_right*r_right,gg*p_left*u_left)); dt=0.45*h*cmax;
for i=1:nx;
    for j=i-1:maxstep;
        rh(j)=r(j)+(gg-1)*(E(j)-0.5*(r(j)+u(j))); end
        for l=2:nx-1;
            r(l)=0.5*(r(j+1)+r(l-1))-(0.5*dt)*(ru(j+1)-ru(l-1));
            ru(l)=0.5*(r(j+1)+r(l-1))-(0.5*dt)*(rE(j+1)-rE(l-1));
            c(l)=sqrt((r(l)+0.5*dt)*(ru(j+1)^2+rE(j+1)^2)+(p(j+1)-(ru(j-1)^2+rE(j-1))^2));
            (E(l))=0.5*(r(j+1)*E(j+1)+r(l-1)*E(l-1));
            (0.5*dt)*(ru(j+1)*ru(j+1)*(E(j+1)+p(j+1))...
            -(ru(l-1)*ru(l-1)*(E(l-1)+p(l-1))));
            (0.5*dt)*(rE(j+1)*rE(j+1)*(E(j+1)+p(j+1))...
            -(rE(l-1)*rE(l-1)*(E(l-1)+p(l-1))));
        end
    end

    r=rh;ru=ru;rE=rE;
    if mod(j,100000)==0; pause
    time=time+dt; step
    end
end
```



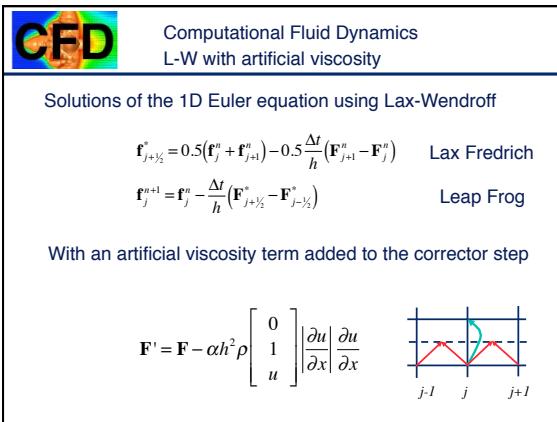
CFD Computational Fluid Dynamics L-W with artificial viscosity

The Euler equations:

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} \rho u \\ \rho u^2 + p \\ \rho u(E + p/\rho) \end{pmatrix} = 0$$

where $E = e + u^2/2; p = (\gamma - 1)\rho e$

Add the artificial viscosity to RHS:

$$= \frac{\partial}{\partial x} \left(\alpha h^3 \rho \begin{vmatrix} 0 \\ \frac{\partial u}{\partial x} \\ \frac{\partial u}{\partial x} \end{vmatrix} \right)$$


CFD Computational Fluid Dynamics L-W with artificial viscosity

Outline of L-W program

```

for istep=1:2000
    for i=1:nx;p(i)=.....; end
    for i=2:nx-1
        rh(i)=.....
        ru(i)=.....
        rEh(i)=.....
    end
    for i=1:nx;ph(i)=.....; end
    for i=2:nx-1
        r(i)=.....
        ru(i)=.....
        rE(i)=.....
    end
    for i=1:nx,u(i)=ru(i)/r(i);end
    for i=2:nx-1
        ru(i)=.....
        rE(i)=.....
    end
    time=time+dt;istep
end

```

%prediction step

%correction step

%artificial viscosity

CFD Computational Fluid Dynamics L-W with artificial viscosity

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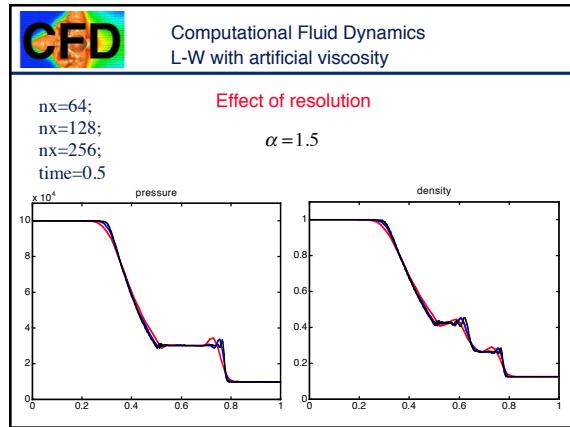
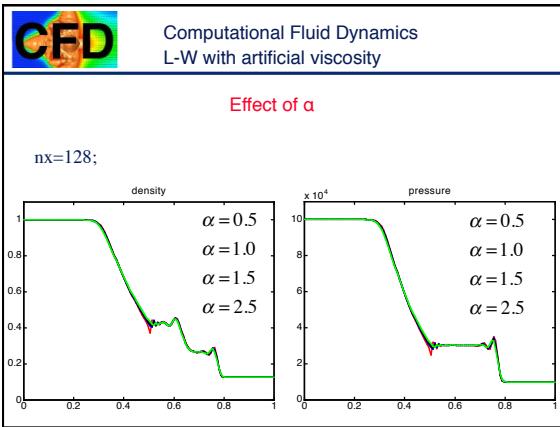
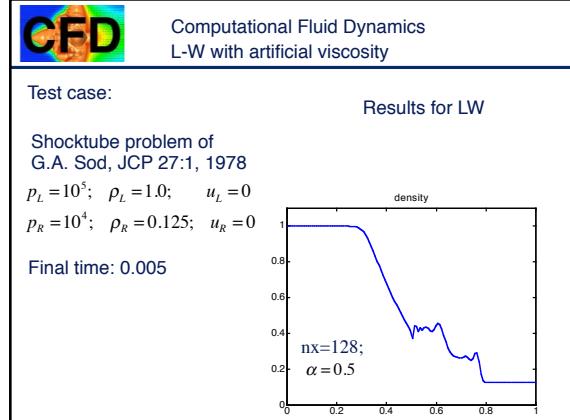
nx=128; artvisc=0.5;
hold off
gg=1; p_left=1000000; p_right=10000; left=-1; right=0.01;
gg=1; p_left=1000000; p_right=10000; left=-1; right=0.125;
dx=10.0; nx=(dx-1)/dx; time=0;

%zeroes(1,nx); u=zeros(1,nx); E=zeros(1,nx); ph=zeros(1,nx);
%zeroes(1,nx); u=zeros(1,nx); E=zeros(1,nx); ph=zeros(1,nx);
for i=1:nx; r(i)=right; u(i)=0; E(i)=0; ph(i)=0; end
for i=1:nx; r(i)=left; u(i)=0; E(i)=0; ph(i)=0; end
r(1)=0;
u(1)=0;
E(1)=0;
ph(1)=0;

for istep=1:2000
    for i=1:nx; p(i)=(gg-1)*(E(i))/((u(i))^2); end
    for i=2:nx-1
        %prediction step
        r(i)=0.5*(r(i)+u(i+1))-0.5*dx/h*((u(i+1)-u(i)));
        u(i)=0.5*(u(i)+u(i+1))-0.5*dx/h*((E(i+1)+2*E(i)+E(i-1))+p(i)-(u(i)^2)*2*dx/p(i));
        E(i)=0.5*(E(i)+E(i+1)+2*E(i+2))-0.5*dx/h*((E(i+1)+u(i+1))*((u(i+1)^2)+(u(i)^2))-...
            (E(i)^2)*(u(i)^2)*dx/p(i));
        ph(i)=0;
    end
    for i=1:nx; r(i)=right; u(i)=0; E(i)=0; ph(i)=0; end
    for i=1:nx; r(i)=left; u(i)=0; E(i)=0; ph(i)=0; end
    r(1)=0;
    u(1)=0;
    E(1)=0;
    ph(1)=0;

    for i=2:nx-1
        %correction step
        r(i)=r(i)-(dx/h)^2*(u(i+1)^2-u(i)^2);
        u(i)=u(i)-(dx/h)^2*(u(i+1)^2-u(i)^2)-(2*dx/h^2)*(1+ph(i)*ph(i-1));
        E(i)=E(i)-(dx/h)^2*(E(i+1)^2-u(i)^2)-(E(i+1)^2-u(i)^2)-(2*dx/h^2)*(1+ph(i)*ph(i-1))...
            -(u(i)^2)*(ph(i)*ph(i));
        ph(i)=0;
    end
    time=time+dt;
    plot(p, 'r', 'LineWidth', 2); title('pressure');
    pause(0.005); break;
end

```



CFD Computational Fluid Dynamics

First order upwinding with flux splitting

For conservation law

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}}{\partial x} = 0$$

Which can be written as

$$\frac{\partial \mathbf{f}}{\partial t} + [A] \frac{\partial \mathbf{f}}{\partial x} = 0; \quad [A] = \frac{\partial \mathbf{F}}{\partial \mathbf{f}}$$

Define $\mathbf{F} = \mathbf{F}^+ + \mathbf{F}^-$ where $[\lambda] = [\lambda^+] + [\lambda^-]$

So that

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}^+}{\partial x} + \frac{\partial \mathbf{F}^-}{\partial x} = 0$$

Are the positive and negative eigenvalues of A



Computational Fluid Dynamics Flux Splitting

For nonlinear equation the splitting is not unique, different matrices can have the same eigenvalues

The history of flux splitting is interesting in that complex splitting were discovered first and the simpler one later

The simplest one is that Zha-Bilgen

Other examples are due to van Leer and Steger-Warming



Computational Fluid Dynamics The Euler Equations

Zha-Bilgen flux splitting

$$\mathbf{F}^+ = \max(u, 0) \begin{bmatrix} \rho \\ \rho u \\ \rho e \end{bmatrix} + \begin{bmatrix} 0 \\ p^+ \\ (pu)^+ \end{bmatrix}$$

$$\mathbf{F}^- = \min(u, 0) \begin{bmatrix} \rho \\ \rho u \\ \rho e \end{bmatrix} + \begin{bmatrix} 0 \\ p^- \\ (pu)^- \end{bmatrix},$$

$$p^+ = p \begin{cases} 0, & M \leq -1 \\ \frac{1}{2}(1+M), & -1 < M < 1, \\ 1, & M \geq 1 \end{cases}, \quad p^- = p \begin{cases} 1, & M \leq -1 \\ \frac{1}{2}(1-M), & -1 < M < 1, \\ 0, & M \geq 1 \end{cases},$$

$$(pu)^+ = p \begin{cases} 0, & M \leq -1 \\ \frac{1}{2}(u+c), & -1 < M < 1 \\ u, & M \geq 1 \end{cases}, \quad (pu)^- = p \begin{cases} u, & M \leq -1 \\ \frac{1}{2}(u-c), & -1 < M < 1 \\ 0, & M \geq 1 \end{cases},$$

Separate the advective fluxes and linearly interpolate the pressure terms between $M=1$ and $M=-1$.



Computational Fluid Dynamics The Euler Equations

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}^+}{\partial x} + \frac{\partial \mathbf{F}^-}{\partial x} = 0,$$

$$\mathbf{f}_j^{n+1} = \mathbf{f}_j^n - \frac{\Delta t}{\Delta x} \left(\mathbf{F}^+(\mathbf{f}_j^n) - \mathbf{F}^+(\mathbf{f}_{j-1}^n) + \mathbf{F}^-(\mathbf{f}_{j+1}^n) - \mathbf{F}^-(\mathbf{f}_j^n) \right)$$

$$\mathbf{f}_j^{n+1} = \mathbf{f}_j^n - \frac{\Delta t}{\Delta x} \left(\hat{\mathbf{F}}_{j+1/2}^n - \hat{\mathbf{F}}_{j-1/2}^n \right)$$

$$\hat{\mathbf{F}}_{j+1/2}^n = \mathbf{F}^+(\mathbf{f}_j^n) + \mathbf{F}^-(\mathbf{f}_{j+1}^n).$$

Using that $\max(f, 0) = 0.5*(f+|f|)$ and $\min(f, 0) = 0.5*(f-|f|)$

$$\begin{aligned} \hat{\mathbf{F}}_{j+1/2}^n &= \mathbf{F}(\mathbf{f}_{j+1}^n) + \mathbf{F}(\mathbf{f}_j^n) - \frac{1}{2} \left(|u_{j+1}^n| \begin{bmatrix} \rho_{j+1}^n \\ \rho_{j+1}^n u_{j+1}^n \\ \rho_{j+1}^n e_j^n \end{bmatrix} - |u_j^n| \begin{bmatrix} \rho_j^n \\ \rho_j^n u_j^n \\ \rho_j^n e_j^n \end{bmatrix} \right) \\ &\quad - \frac{1}{2} \left(\begin{bmatrix} 0 \\ p_{j+1}^n M_{j+1}^n \\ p_{j+1}^n e_j^n \end{bmatrix} - \begin{bmatrix} 0 \\ p_j^n M_j^n \\ p_j^n e_j^n \end{bmatrix} \right) \end{aligned}$$



Computational Fluid Dynamics Summary

```
%upwind-Zha-Bilgen flux splitting
nx=40;dx=1;rho=0.05;u=1.0;tme=0;gg=1.4;
hx=dx*1.0;for i=1:nx,u(i)=hx*(tme-1);end;
p_left=100000,p_right=10000,r_left=1,r_right=0.125,u_left=0;
r=zeros(1,nx);u=zeros(1,nx);E=zeros(1,nx);
p=zeros(1,nx);p_left=zeros(1,nx);p_right=zeros(1,nx);
F1=zeros(1,nx);F2=zeros(1,nx);F3=zeros(1,nx);

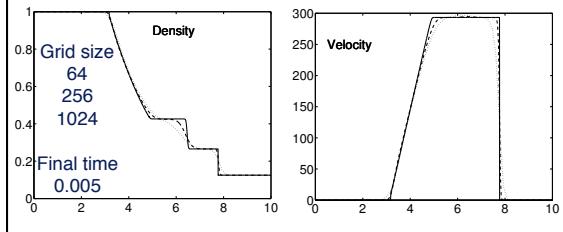
for i=1:nx, r(i)=r_left;right(r(i))=0.0;E(i)=p_left;right(gg-1);end
for i=1:nx/2, r(i)=r_left;right(gg-1);end
cmax=sqrt(max(gg*p_left*r_left,right(gg*p_left)*right(r_left)));
dt=0.45*cmax;maxstep=tme/dt;
if m(j)>1, F2(j)=0;F3(j)=0;end
for i=1:nx-1
    % Find fluxes
    F1(i)=0.5*(u(i)+u(i+1))*(E(i)-0.5*(u(i)*u(i)*r(i)));end
    for i=1:nx,c(i)=sqrt((gg*p(i)*p(i))/r(i));end
    for i=1:nx,m(i)=u(i)/c(i);end
    for i=1:nx-1
        if m(i)>1, F2(i)=0;F3(i)=0;end
        if m(i)<-1, F2(i)=sqrt((gg*p(i)*p(i))/r(i));F3(i)=0;end
        if m(i)<1, F3(i)=sqrt((gg*p(i)*p(i))/r(i));F2(i)=0;end
    end
    if m(j)>1, F2(j)=0;F3(j)=0;end
    if m(j)<-1, F3(j)=sqrt((gg*p(j)*p(j))/r(j));F2(j)=0;end
    if m(j)<1, F3(j)=sqrt((gg*p(j)*p(j))/r(j));F2(j)=0;end
    time=tme+d*step
    end
end
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Computational Fluid Dynamics The Euler Equations

Shock tube problem of G.A. Sod, JCP 27:1, 1978

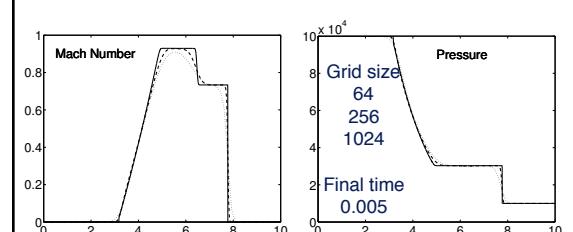
$$\begin{aligned} p_L &= 10^5; \quad \rho_L = 1.0; \quad u_L = 0 \\ p_R &= 10^4; \quad \rho_R = 0.125; \quad u_R = 0 \end{aligned}$$



Computational Fluid Dynamics The Euler Equations

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Computational Fluid Dynamics First order upwind

Another example: one-dimensional Euler equation using the van Leer vector flux splitting

Van Leer

$$\mathbf{F}^+ = \frac{\rho}{4c}(u+c)^2 \begin{bmatrix} 1 \\ \frac{(\gamma-1)u+2c}{\gamma} \\ \frac{[(\gamma-1)u+2c]^2}{2(\gamma^2-1)} \end{bmatrix} \quad \mathbf{F}^- = -\frac{\rho}{4c}(u-c)^2 \begin{bmatrix} 1 \\ \frac{(\gamma-1)u-2c}{\gamma} \\ \frac{[2c-(\gamma-1)u]^2}{2(\gamma^2-1)} \end{bmatrix}$$



Computational Fluid Dynamics First order upwind

van Leer vector flux splitting

$$\begin{bmatrix} \rho u \\ \rho u^2 + p \\ \rho u(E + \frac{p}{\rho}) \end{bmatrix} = \frac{\rho}{4c}(u+c)^2 \begin{bmatrix} 1 \\ \frac{(\gamma-1)u+2c}{\gamma} \\ \frac{[(\gamma-1)u+2c]^2}{2(\gamma^2-1)} \end{bmatrix} - \frac{\rho}{4c}(u-c)^2 \begin{bmatrix} 1 \\ \frac{(\gamma-1)u-2c}{\gamma} \\ \frac{[2c-(\gamma-1)u]^2}{2(\gamma^2-1)} \end{bmatrix}$$

For example, the mass flux:

$$\begin{aligned} \frac{\rho}{4c}(u+c)^2 - \frac{\rho}{4c}(u-c)^2 \\ = \frac{\rho}{4c}(u^2 + 2uc + c^2 - u^2 + 2uc - c^2) = \frac{\rho}{4c}4uc = \rho u \end{aligned}$$



Computational Fluid Dynamics First order upwind

For 1D flow the fluxes are

$$\mathbf{F}^\pm = \pm \frac{\rho}{4c}(u \pm c)^2 \begin{bmatrix} 1 \\ \frac{(\gamma-1)u \pm 2c}{\gamma} \\ \frac{[2c \pm (\gamma-1)u]^2}{2(\gamma^2-1)} \end{bmatrix} = \pm \frac{\rho c}{4}(M \pm 1)^2 \begin{bmatrix} 1 \\ \frac{2c(\gamma-1)}{2} \left(M \pm 1 \right) \\ \frac{2c^2}{\gamma^2-1} \left(1 \pm \frac{\gamma-1}{2} M \right)^2 \end{bmatrix}$$



Computational Fluid Dynamics First order upwind

```

for istep=1:maxstep
    for i=1:nx,c(i)=sqrt(gg*(gg-1)*(rE(i)-0.5*(ru(i)^2/r(i)))/r(i));end
    for i=1:nx,u(i)=ru(i)/r(i);end; for i=1:nx,m(i)=u(i)/c(i);end

    for i=2:nx-1
        %upwind
        rr(i)=r(i)-(dr/dh)*...
        (0.25*r(i-1)*c(i)*(m(i-1)+1)^2) - (0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*...
        (-0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2) - (0.25*r(i)*c(i)*(m(i-1)+1)^2);
        run(i)=ru(i)-(dt/h)*...
        (0.25*r(i)*c(i)*(m(i)+1)^2)*((1+0.5*(gg-1)*m(i))*2*c(i)/gg) - ...
        (0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1+0.5*(gg-1)*m(i-1))*2*c(i-1)/gg) + ...
        (-0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1-0.5*(gg-1)*m(i+1))*2*c(i+1)/gg) - ...
        (0.25*r(i)*c(i)*(m(i)+1)^2)*((1-0.5*(gg-1)*m(i))*2*c(i)/gg);

        rEn(i)=rE(i)-(dt/h)*...
        (0.25*r(i)*c(i)*(m(i)+1)^2)*((1+0.5*(gg-1)*m(i))^2 * 2*c(i)^2 / (gg^2-1)) - ...
        (0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1+0.5*(gg-1)*m(i-1))^2 * 2*c(i-1)^2 / (gg^2-1)) + ...
        (-0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1-0.5*(gg-1)*m(i+1))^2 * 2*c(i+1)^2 / (gg^2-1)) - ...
        (0.25*r(i)*c(i)*(m(i)+1)^2)*((1-0.5*(gg-1)*m(i))^2 * 2*c(i)^2 / (gg^2-1));
        end
    end
    rE(i)=rE(i)-(dt/h)*...
    (0.25*r(i)*c(i)*(m(i)+1)^2)*((1+0.5*(gg-1)*m(i))^2 * 2*c(i)^2 / (gg^2-1)) - ...
    (0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1+0.5*(gg-1)*m(i-1))^2 * 2*c(i-1)^2 / (gg^2-1)) + ...
    (-0.25*r(i-1)*c(i-1)*(m(i-1)+1)^2)*((1-0.5*(gg-1)*m(i+1))^2 * 2*c(i+1)^2 / (gg^2-1)) - ...
    (0.25*r(i)*c(i)*(m(i)+1)^2)*((1-0.5*(gg-1)*m(i))^2 * 2*c(i)^2 / (gg^2-1));
    end

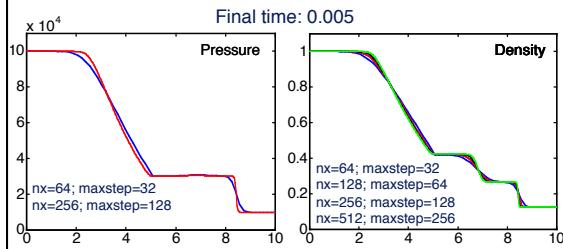
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Computational Fluid Dynamics First order upwind

Effect of resolution

Shocktube problem of G.A. Sod, JCP 27:1, 1978 $p_L = 10^5$; $\rho_L = 1.0$; $u_L = 0$
 $p_R = 10^4$; $\rho_R = 0.125$; $u_R = 0$



Computational Fluid Dynamics First order upwind

For 2D flow the van Leer fluxes are

$$\mathbf{F}^\pm = \pm \frac{\rho}{4c}(u \pm c)^2 \begin{bmatrix} \frac{1}{\gamma} \\ \frac{(\gamma-1)u \pm 2c}{\gamma} \\ \frac{v^2 + [(\gamma-1)u \pm 2c]^2}{2(\gamma^2-1)} \end{bmatrix}; \quad \mathbf{G}^\pm = \pm \frac{\rho}{4c}(v \pm c)^2 \begin{bmatrix} \frac{1}{\gamma} \\ \frac{(\gamma-1)v \pm 2c}{\gamma} \\ \frac{u^2 + [(\gamma-1)v \pm 2c]^2}{2(\gamma^2-1)} \end{bmatrix}$$

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}^+}{\partial x} + \frac{\partial \mathbf{F}^-}{\partial x} + \frac{\partial \mathbf{G}^+}{\partial y} + \frac{\partial \mathbf{G}^-}{\partial y} = 0$$



Computational Fluid Dynamics First order upwind

Several other splitting schemes are possible, such as:

Steger-Warming

$$\mathbf{F}^+ = \frac{\rho}{2\gamma} \begin{bmatrix} (2\gamma-1)u + c \\ 2(\gamma-1)u^2 + (u+c)^2 \\ (\gamma-1)u^3 + \frac{1}{2}(u+c)^3 + \frac{3-\gamma}{2(\gamma-1)}(u+c)c^2 \end{bmatrix}$$

$$\mathbf{F}^- = \frac{1}{2}\frac{\rho}{\gamma}(u-c) \begin{bmatrix} 1 \\ u-c \\ \frac{1}{2}(u-c)^2 + \frac{1}{2}c^2 \left(\frac{3-\gamma}{\gamma-1} \right) \end{bmatrix}$$



Computational Fluid Dynamics First order upwind

Rewrite the flux terms in terms of Mach number:

$$\mathbf{F}^+ = \begin{pmatrix} \frac{\rho}{4c}(u+c)^2 & \begin{pmatrix} 1 \\ \frac{(\gamma-1)u+2c}{\gamma} \\ \frac{[(\gamma-1)u+2c]^2}{2(\gamma^2-1)} \end{pmatrix} \end{pmatrix} = \begin{pmatrix} \frac{\rho c}{4} \left(\frac{u}{c} + 1 \right)^2 & \begin{pmatrix} 1 \\ \frac{2c}{\gamma} \left(1 + \frac{\gamma-1}{2} \frac{u}{c} \right) \\ \frac{2c^2}{\gamma^2-1} \left(1 + \frac{\gamma-1}{2} \frac{u}{c} \right)^2 \end{pmatrix} \end{pmatrix}$$

$$= \begin{pmatrix} \frac{\rho c}{4} (M+1)^2 & \begin{pmatrix} 1 \\ \frac{2c}{\gamma} \left(1 + \frac{\gamma-1}{2} M \right) \\ \frac{2c^2}{\gamma^2-1} \left(1 + \frac{\gamma-1}{2} M \right)^2 \end{pmatrix} \end{pmatrix}$$



Computational Fluid Dynamics First order upwind

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}^+}{\partial x} + \frac{\partial \mathbf{F}^-}{\partial x} = 0$$

$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho u \\ \rho E \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} 1 \\ \frac{\rho c}{4} (M+1)^2 \\ \frac{2c}{\gamma} \left(1 + \frac{\gamma-1}{2} M \right) \\ \frac{2c^2}{\gamma^2-1} \left(1 + \frac{\gamma-1}{2} M \right)^2 \end{pmatrix} + \frac{\partial}{\partial x} \begin{pmatrix} 1 \\ -\frac{\rho c}{4} (M-1)^2 \\ \frac{2c}{\gamma} \left(-1 + \frac{\gamma-1}{2} M \right) \\ \frac{2c^2}{\gamma^2-1} \left(-1 + \frac{\gamma-1}{2} M \right)^2 \end{pmatrix} = 0$$



Computational Fluid Dynamics First order upwind

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}^+}{\partial x} + \frac{\partial \mathbf{F}^-}{\partial x} = 0$$

Solve by

$$\mathbf{f}_j^{n+1} = \mathbf{f}_j^n - \left(\frac{\Delta t}{h} \right) \left(\mathbf{F}_j^+ - \mathbf{F}_{j-1}^+ \right)^n - \left(\frac{\Delta t}{h} \right) \left(\mathbf{F}_{j+1}^- - \mathbf{F}_j^- \right)^n$$

Where

$$\mathbf{F}^+ = \frac{\rho c}{4} (M+1)^2 \begin{pmatrix} 1 \\ \frac{2c}{\gamma} \left(1 + \frac{\gamma-1}{2} M \right) \\ \frac{2c^2}{\gamma^2-1} \left(1 + \frac{\gamma-1}{2} M \right)^2 \end{pmatrix}; \quad \mathbf{F}^- = -\frac{\rho c}{4} (M-1)^2 \begin{pmatrix} 1 \\ \frac{2c}{\gamma} \left(-1 + \frac{\gamma-1}{2} M \right) \\ \frac{2c^2}{\gamma^2-1} \left(-1 + \frac{\gamma-1}{2} M \right)^2 \end{pmatrix}$$



Computational Fluid Dynamics <http://www.nd.edu/~gtryggva/CFD-Course/>

WENO-3



Computational Fluid Dynamics

Third order WENO

$$\frac{\partial}{\partial t} \mathbf{f} + \frac{\partial}{\partial x} \mathbf{F} = 0$$

The semi-discrete equation is

$$\frac{d\mathbf{f}_j}{dt} + \frac{1}{\Delta x} (\mathbf{F}_{j+1/2} - \mathbf{F}_{j-1/2}) = 0$$

The fluxes are the weighted sum

$$\mathbf{F}_{j+1/2} = \omega_1 \mathbf{F}_{j+1/2}^{(1)} + \omega_2 \mathbf{F}_{j+1/2}^{(2)}$$

where

$$\mathbf{F}_{j+1/2}^{(1)} = -\frac{1}{2} \mathbf{F}_{j-1} + \frac{3}{2} \mathbf{F}_j$$

2nd order upwind

$$\mathbf{F}_{j+1/2}^{(2)} = \frac{1}{2} \mathbf{F}_j + \frac{1}{2} \mathbf{F}_{j+1}$$

2nd order central

The weights are

$$\tilde{\omega}_l = \frac{\gamma_l}{(\varepsilon + \beta_l)^2}$$

$$\omega_m = \sum_{l=1}^2 \tilde{\omega}_l$$

$$\beta_1 = (f_j - f_{j-1})^2$$

$$\beta_2 = (f_{j+1} - f_j)^2$$

$$\gamma_1 = \frac{1}{3}; \quad \gamma_2 = \frac{2}{3}$$

$$\varepsilon = 10^{-6}$$

The logo for Computational Fluid Dynamics (CFD) features the letters 'CFD' in a large, bold, black font. The letter 'C' has a blue gradient fill, 'F' has an orange gradient fill, and 'D' has a green gradient fill. A small blue square is positioned at the bottom right corner of the logo.

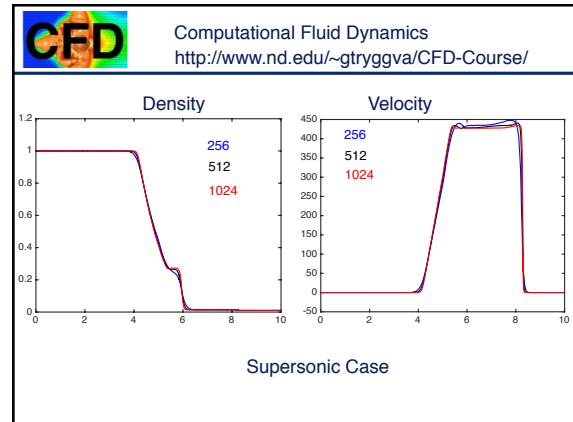
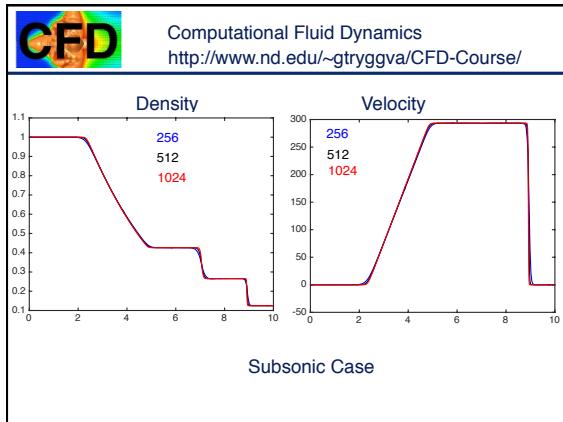
Computational Fluid Dynamics

$$\frac{df_j}{dt} + \frac{1}{\Delta x} (\mathbf{F}_{j+1/2} - \mathbf{F}_{j-1/2}) = 0$$

The time integration is done by a third order Runge-Kutta

$$\begin{aligned}f_j^{(1)} &= f_j^n + \Delta t L(f_j^n, t^n) \\f_j^{(2)} &= \frac{3}{4} f_j^n + \frac{1}{4} f_j^{(1)} + \frac{1}{4} \Delta t L(f_j^{(1)}, t^n + \Delta t) \\f_j^{n+1} &= \frac{1}{3} f_j^n + \frac{2}{3} f_j^{(2)} + \frac{2}{3} \Delta t L(f_j^{(2)}, t^n + \frac{1}{2} \Delta t)\end{aligned}$$

where $L(f, t) = -\frac{\partial F}{\partial x}$



Computational Fluid Dynamics

The Euler Equations

In general we solve:

$$\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{F}}{\partial x} = 0$$

by

$$\mathbf{f}_j^{n+1} = \mathbf{f}_j^n - \frac{\Delta t}{h} \left(\mathbf{F}_{j+1/2}^n - \mathbf{F}_{j-1/2}^n \right)$$

Where the fluxes are found in a variety of ways

$$\mathbf{F}_{j+1/2} = F \left(\left(f^L \right)_{j+1/2}, \left(f^R \right)_{j+1/2} \right)$$

In principle we can solve this problem, the Riemann problem, exactly by assuming constant states and then integrate the fluxes over the time step.

	Computational Fluid Dynamics The Euler Equations
Predictor-corrector method: Limiting the variables	
Predictor step	$f_j^{n+1/2} = f_j^n - \frac{\Delta t}{2h} (F_{j+1/2}^n - F_{j-1/2}^n)$
Variables	$f_{j+1/2}^L = f_j^{n+1/2} + \frac{1}{2} \Psi^L (f_j^{n+1/2} - f_{j-1}^{n+1/2})$ $f_{j+1/2}^R = f_{j+1}^{n+1/2} - \frac{1}{2} \Psi^R (f_{j+1}^{n+1/2} - f_j^{n+1/2})$
Find:	$F_{j+1/2}^{n+1/2} = F \left((f_L)_{j+1/2}^{n+1/2}, (f_R)_{j+1/2}^{n+1/2} \right)$
Final step	$f_j^{n+1} = f_j^n - \frac{\Delta t}{h} (F_{j+1/2}^{n+1/2} - F_{j-1/2}^{n+1/2})$



Computational Fluid Dynamics

The Riemann problems for a linear hyperbolic system



Computational Fluid Dynamics The Euler Equations

The Riemann problem for a system of linear equations with constant coefficients. We start with:

$$\frac{\partial \mathbf{f}}{\partial t} + \mathbf{A} \frac{\partial \mathbf{f}}{\partial x} = 0$$

The eigenvalues and eigenvectors of \mathbf{A} are given by

$$\mathbf{A}\mathbf{e}_p = \lambda_p \mathbf{e}_p$$

If we normalize the eigenvectors to unity $|\mathbf{e}_p| = 1$ then the eigenvectors form an orthonormal basis and we can write

$$\mathbf{f} = \sum_{p=1}^3 w_p \mathbf{e}_p \quad w_p = \mathbf{f} \cdot \mathbf{e}_p$$



Computational Fluid Dynamics The Euler Equations

Substituting for \mathbf{f} in the original equation, we get

$$\sum_{p=1}^3 \frac{\partial w_p}{\partial t} \mathbf{e}_p + \sum_{p=1}^3 \frac{\partial w_p}{\partial x} \mathbf{A} \mathbf{e}_p = 0$$

Or, using the definition of the eigenvalue and the fact that the eigenvectors are orthonormal, we have three sets of independent equations

$$\frac{\partial w_p}{\partial t} + \lambda_p \frac{\partial w_p}{\partial x} = 0 \quad \mathbf{A}\mathbf{e}_p = \lambda_p \mathbf{e}_p$$

Whose solutions are given by

$$w_p(x, t) = w_{p0}(x - \lambda_p t)$$

Representing a simple translation of the initial conditions

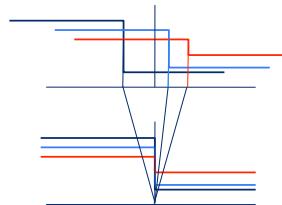


Computational Fluid Dynamics The Euler Equations

For discontinuous initial conditions, where

$$\mathbf{w} = \mathbf{w}^L, \quad x < 0 \quad \& \quad \mathbf{w} = \mathbf{w}^R, \quad x > 0; \quad \mathbf{w} = (w_1, w_2, w_3)$$

The different components of the solution remain piecewise constant but the discontinuity moves left or right, depending on the sign of the eigenvalues.



Computational Fluid Dynamics The Euler Equations

Since the solution is piecewise constant the fluxes are easily found. In particular, for each of the linear equations the fluxes are:

$$\begin{aligned} w_p \lambda_p &= \begin{cases} w_p^L \lambda_p & \text{if } \lambda_p > 0 \\ w_p^R \lambda_p & \text{if } \lambda_p < 0 \end{cases} = \frac{1}{2} \left\{ (\lambda_p + |\lambda_p|) w_p^L + (\lambda_p - |\lambda_p|) w_p^R \right\} \\ &= \frac{1}{2} \left\{ \lambda_p (w_p^L + w_p^R) - |\lambda_p| (w_p^R - w_p^L) \right\} \end{aligned}$$

Or, for the system

$$\mathbf{F}_{j+1/2} = \sum_p \frac{1}{2} \left\{ \lambda_p (w_p^L + w_p^R) - |\lambda_p| (w_p^R - w_p^L) \right\} \mathbf{e}_p$$



Computational Fluid Dynamics The Euler Equations

For the Euler equations the situation is more complex.
Start with the primitive form

$$\frac{\partial \mathbf{w}}{\partial t} + \mathbf{C} \frac{\partial \mathbf{w}}{\partial x} = 0$$

$$\mathbf{C} = \begin{pmatrix} u & \rho & 0 \\ 0 & u & 1/\rho \\ 0 & \rho c^2 & u \end{pmatrix} \quad \mathbf{w} = \begin{pmatrix} \rho \\ u \\ p \end{pmatrix}$$

We can find the eigenvectors in standard ways and use them to make the matrix diagonal



Computational Fluid Dynamics The Euler Equations

The eigenvector of

$$\mathbf{C} = \begin{pmatrix} u & \rho & 0 \\ 0 & u & 1/\rho \\ 0 & \rho c^2 & u \end{pmatrix}$$

are (using arbitrary scaling)

Right eigenvectors

$$\mathbf{Cr}_p = \lambda_p \mathbf{r}_p$$

$$\mathbf{r}_1 = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}; \quad \mathbf{r}_2 = \begin{pmatrix} \frac{\rho}{2c} \\ \frac{1}{2} \\ \frac{\rho c}{2} \end{pmatrix}; \quad \mathbf{r}_3 = \begin{pmatrix} -\frac{\rho}{2c} \\ \frac{1}{2} \\ -\frac{\rho c}{2} \end{pmatrix}$$

Left eigenvectors

$$\mathbf{C}^T \mathbf{l}_p = \lambda_p \mathbf{l}_p$$

$$\mathbf{l}_1 = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}; \quad \mathbf{l}_2 = \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}; \quad \mathbf{l}_3 = \begin{pmatrix} -\frac{1}{c^2} \\ 0 \\ 1 \end{pmatrix}$$



Computational Fluid Dynamics The Euler Equations

Form the matrices

$$Q_C = \begin{pmatrix} 1 & \frac{\rho}{2c} & -\frac{\rho}{2c} \\ 0 & \frac{1}{2} & \frac{1}{2} \\ 0 & \frac{\rho c}{2} & -\frac{\rho c}{2} \end{pmatrix} \quad Q_C^{-1} = \begin{pmatrix} 1 & 0 & -\frac{1}{c^2} \\ 0 & 1 & \frac{1}{\rho c} \\ 0 & 1 & -\frac{1}{\rho c} \end{pmatrix} \quad \mathbf{w} = \begin{pmatrix} \rho \\ u \\ p \end{pmatrix}$$

Then

$$Q_C^{-1} \mathbf{C} Q_C = \begin{pmatrix} u & 0 & 0 \\ 0 & u+c & 0 \\ 0 & 0 & u-c \end{pmatrix} \quad \mathbf{C} = \begin{pmatrix} u & \rho & 0 \\ 0 & u & 1/\rho \\ 0 & \rho c^2 & u \end{pmatrix}$$

$$\frac{\partial \mathbf{w}}{\partial t} + \mathbf{C} \frac{\partial \mathbf{w}}{\partial x} = 0$$



Computational Fluid Dynamics The Euler Equations

$$\text{Multiply } \frac{\partial \mathbf{w}}{\partial t} + \mathbf{C} \frac{\partial \mathbf{w}}{\partial x} = 0$$

To get the characteristic form

$$Q_C^{-1} \frac{\partial \mathbf{w}}{\partial t} + Q_C^{-1} \mathbf{C} \frac{\partial \mathbf{w}}{\partial x} = 0$$

Which can be written as

or

$$\frac{\partial \mathbf{v}}{\partial t} + \Lambda \frac{\partial \mathbf{v}}{\partial x} = 0 \quad \frac{\partial v_0}{\partial t} + u \frac{\partial v_0}{\partial x} = 0$$

$$\Lambda = Q_C^{-1} \mathbf{C} Q_C = \begin{pmatrix} u & 0 & 0 \\ 0 & u+c & 0 \\ 0 & 0 & u-c \end{pmatrix} \quad \frac{\partial v_+}{\partial t} + (u+c) \frac{\partial v_+}{\partial x} = 0$$

$$\frac{\partial v_-}{\partial t} + (u-c) \frac{\partial v_-}{\partial x} = 0$$



Computational Fluid Dynamics The Euler Equations

Although the Riemann problem can be solved for the general case when the fluids are not stationary, the solution is expensive (involving solving a nonlinear equation for the pressure ratio across the shock) and the only information that we need from the solution is the flux across the cell boundary.

Therefore, usually we use approximate Riemann solvers. Different approximations are possible but generally the result is a complex process. Here we will only outline it briefly