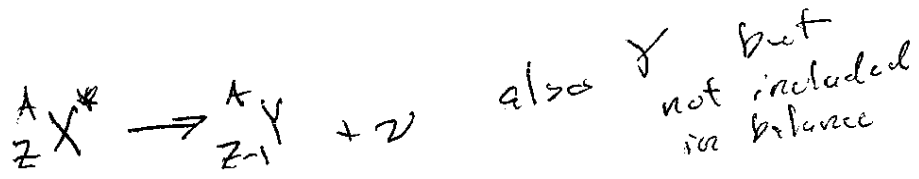


Internal Conversion vs Electron Capture

EC \rightarrow orbital electron captured by nucleus
 has to accelerate, so γ ray emitted
 γ ray has energy depending on
 Q value of EC



IC \rightarrow excited nucleus, transfers energy
 to orbital electron w/

$$E_c = E_{ex} - E_b$$

\nwarrow binding energy
 in shell

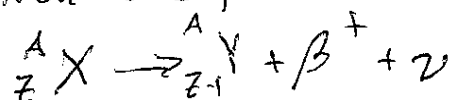
practically monoenergetic b.c. shell energy

Decay Situations

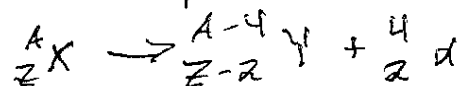
given

 ${}^A_Z X_N$ w/ $M({}^A_Z)$ is ~~nuclear mass~~ $M({}^A_Z)$ is neutral atomic mass $m_e \rightarrow$ mass e^- 931.5 $\frac{\text{MeV}}{\text{amu}}$

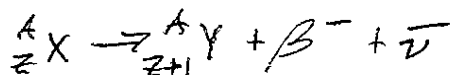
positron decay



alpha decay



Beta Decay



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QF studying Mfrcs Alex Hagen

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Given

$$\frac{dE}{dx} = c \left(\frac{Z}{v} \right)^2 \quad (\text{Bethe-Bloch})$$

for α, d, p, α

if mass ion increased

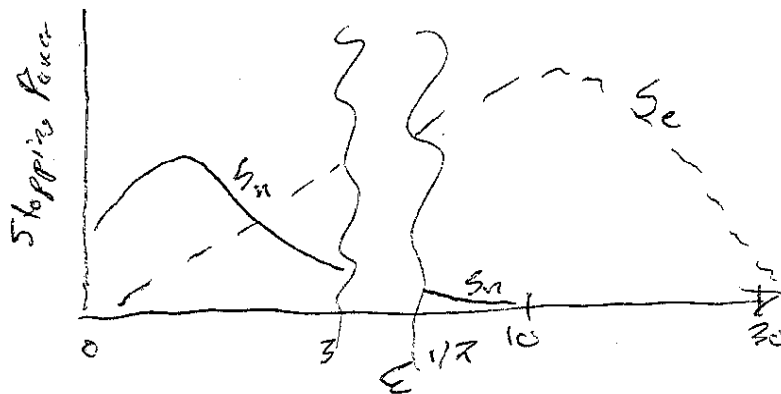
→ more less penetrative

~~charge~~
charge increased

→ less penetrative

look at (Bethe-Bloch eqn)

$$-\frac{dE}{dx} \Big|_c = \frac{4\pi Z_1^2 e^4 n_e}{m_e v^2} \ln \frac{v_{max}}{v_{min}}$$



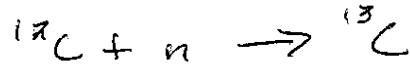
$$-\frac{dE}{dx} = \frac{4\pi Z_1^2 e^4}{m_e v^2} N \beta$$

$$\beta \equiv Z \left[\ln \frac{2m_e v^2}{I} - \ln \left(1 - \frac{v^2}{c^2} \right) - \frac{v^2}{c^2} \right]$$

Binding Energy Problem

Calculate binding energy of last neutron
in ^{13}C

so ^{13}C is equal to



use Q-value eqn

~~$$Q = \left(\frac{BE(^{13}\text{C})}{13} \right) - \left(\frac{BE(^{12}\text{C})}{12} + \frac{BE(n)}{1} \right) \times 931.5$$~~

$$Q = \left([M(^{12}\text{C}) + M(n)] - [M(^{13}\text{C})] \right) \times 931.5 \text{ MeV}$$

$$Q = (12 + 1.00866 - 13.00335) \times 931.5 \text{ MeV}$$

$$Q = \underline{\underline{4.95 \text{ MeV}}}$$

Calculate binding energy per nucleon in
 ^{13}C

$$BE = \frac{1300325}{13} \times 931.5 \text{ MeV} =$$

$$BE_{\text{nucleon}} = \frac{[M_n + 6M_p - M_{^{13}\text{C}}] \times 931.5 \text{ MeV}}{13}$$

Particle wavelengths (relativistic)

mass bearing

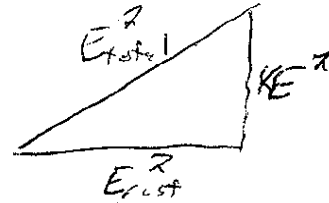
$$h = 4.135 \times 10^{-15} \text{ eV}\cdot\text{s}$$

$$E = \frac{hc}{\lambda}$$

$$\lambda = \frac{hc}{\sqrt{E_{\text{total}}^2 - E_{\text{rest}}^2}}$$

$$E_{\text{total}} = KE + E_{\text{rest}}$$

$$p = \frac{1}{c} \sqrt{T(T + 2m_0 c^2)}$$



zero rest mass

$$\lambda = \frac{hc}{E}$$

Beam of 100 keV γ is Compton scattered.
 $\theta = 30^\circ$ calc E' , E_{recoil} , θ_{recoil} , T_{max}

solve for E'

$$E' = \frac{EE_e}{E(1 - \cos\theta) + E_e}$$

solve for E_{recoil}

also rest mass cancel

assume at rest

$$E_e + E^\# \Rightarrow E_r + E'$$

$$E_r = E' - E$$

$$p_e + p_\gamma = p_r + p_{e'}$$

x dir $\rightarrow \frac{E}{c} = \frac{E' \cos(30)}{c} + \cos(\theta) \sqrt{T_r(T_r + 2m_e c^2)}$

y dir $\rightarrow 0 = \frac{E'}{c} \sin(30) + \sin(\theta) \sqrt{T_r(T_r + 2m_e c^2)}$

solve for θ

T_{max} occurs when the γ recoils at $\theta = 180^\circ$,

so $\theta_e = 0^\circ \rightarrow$

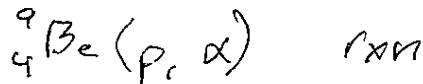
$$E' = \frac{EE_e}{E(1 - \cos\theta) + E_e}$$

assume at rest

$$E_e + E = E_r + E'$$

$$E_r = E' - E$$

Consider

 T_p, T_α given α emerges at angle of 90° . Find T_R

Energy is conserved

$$T_{\text{Be}} + m_{\text{Be}} c^2 + T_p + m_p c^2 = T_R + m_R c^2 + T_\alpha + m_\alpha c^2$$

$$T_R = T_\alpha - T_p + m_p c^2 + m_\alpha c^2 - m_R c^2 - m_{\text{Be}} c^2$$

momentum is conserved

$$P_p = P_R + P_\alpha$$

$$\frac{1}{c} \sqrt{T_\alpha(T_\alpha + 2m_\alpha c^2)} = \frac{1}{c} \sqrt{T_R(T_R + 2m_R c^2)} + \frac{1}{c} \sqrt{T_\alpha(T_\alpha + 2m_\alpha c^2)}$$

$$\sqrt{T_\alpha(T_\alpha + 2m_\alpha c^2)} = \sqrt{T_R(T_R + 2m_R c^2)} + \sqrt{T_\alpha(T_\alpha + 2m_\alpha c^2)}$$

solve

1/22/14

QE studying M11

Alex Hecy

1/1

$$\Delta t = 24 \text{ hr}$$

$$N_0 = 1683 \text{ counts/10 min}$$

$$N_1 = 1683 \text{ counts/20 min}$$

$$N_b = 50 \text{ cpm}$$

$$\frac{N_0}{N_1} = \exp(-\lambda \Delta t)$$

$$\lambda = \frac{1}{\Delta t} \ln\left(\frac{N_1}{N_0}\right)$$

$$t_{1/2} = +\ln(2) \Delta t \frac{\ln\left(\frac{N_1}{N_0}\right)}{\ln(N_1) - \ln(N_0)}$$

$$\sigma_{t_{1/2}} = \sqrt{\left(\frac{\partial t_{1/2}}{\partial \Delta t}\right)^2 \sigma_{\Delta t}^2 + \left(\frac{\partial t_{1/2}}{\partial N_1}\right)^2 \sigma_{N_1}^2 + \left(\frac{\partial t_{1/2}}{\partial N_0}\right)^2 \sigma_{N_0}^2}$$

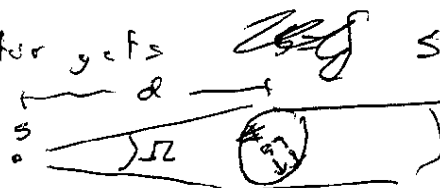
careful w/ derivatives

Stereoscopy problem

given efficiency ϵ_{ip} , counts N (full energy)

isotropic, no attenuation

so detector gets ~~where~~ $S = N \frac{4\pi}{\epsilon_{ip} \Omega}$

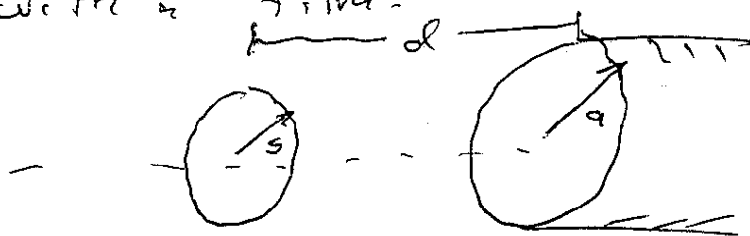


then angle is $\Omega = \int_A \frac{\cos \alpha}{r^2} dA$ ↙ angle between source dir and normal of det.

for geo above $\Omega = 2\pi \left(1 - \frac{d}{\sqrt{d^2 + a^2}}\right)$

if $d \gg a \Rightarrow \Omega = \frac{A}{d^2} = \frac{\pi a^2}{d^2}$

Or with a film:



$$\Omega = \frac{4\pi a}{s} \int_0^\infty \frac{\exp(-dk) J_1(sk) J_1(ak)}{k} dk$$

$$\Omega \approx 2\pi \left[1 - \frac{1}{(1+\beta)^{1/2}} - \frac{3}{8} \frac{\alpha\beta}{(1+\beta)^{3/2}} + \alpha^2 [F1] - \alpha^3 [F2] \right]$$

$$F1 = \frac{5}{16} \frac{\beta}{(1+\beta)^{3/2}} - \frac{35}{64} \frac{\beta^2}{(1+\beta)^{5/2}}$$

$$F2 = \frac{35}{128} \frac{\beta}{(1+\beta)^{3/2}} - \frac{315}{256} \frac{\beta^2}{(1+\beta)^{5/2}} + \frac{1155}{1024} \frac{\beta^3}{(1+\beta)^{7/2}}$$

$$\alpha = \left(\frac{s}{d}\right)^2 \quad \beta = \left(\frac{a}{d}\right)^2$$

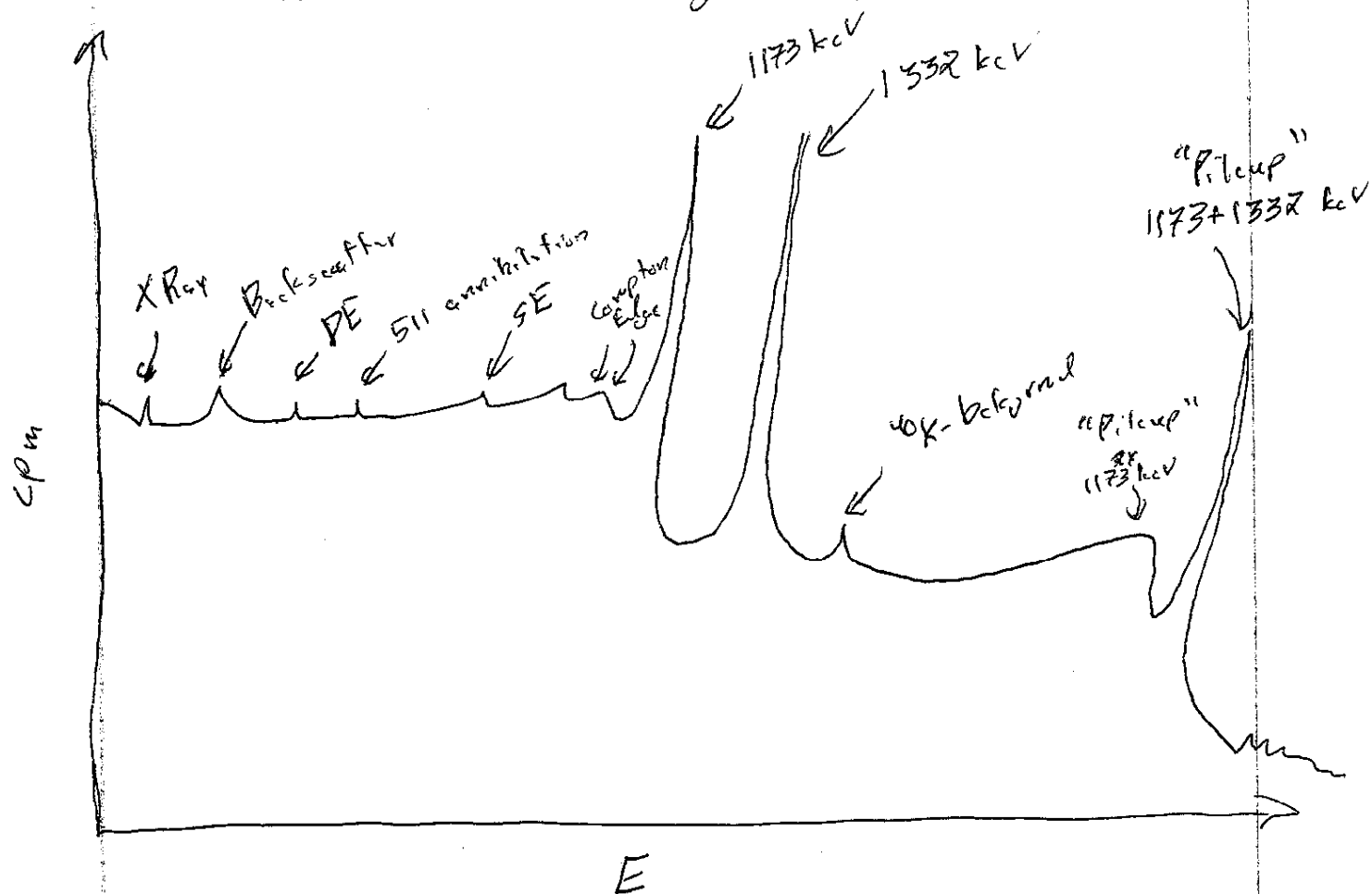
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QE studying Mtl

Aser Ityom

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Peaks in a Pulse height spectrum



E

Compton Continuum \rightarrow backscatter and C-EdgeDouble Escape \rightarrow Pair Production both gammas escape (1.022 MeV less than peak)Single Escape \rightarrow Pair Production, only one gamma escapes (.511 MeV less than peak)annihilation \rightarrow in β^+ decay, always annihilates and makes two 0.511 MeV photons.Backscatter \rightarrow 0.2 - 0.75 MeV, anything above $\approx 110^\circ$ is essentially same energy so we get peak.X-Ray \rightarrow electrons falling shells

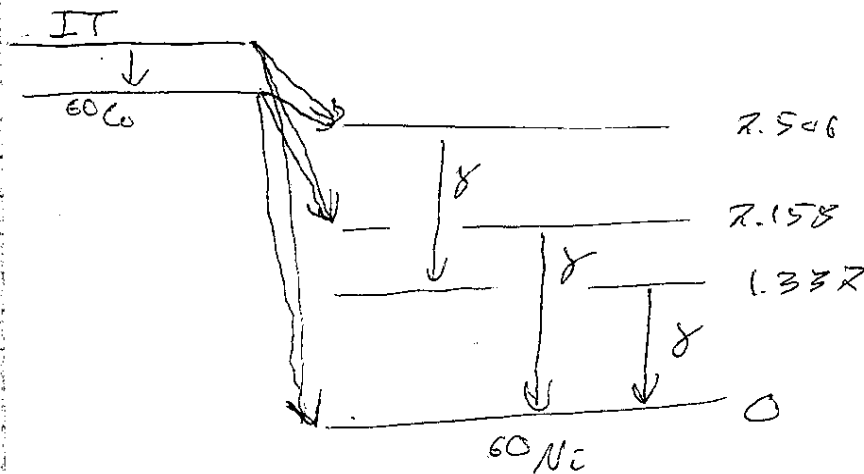
5) Radioactive Decay, decay rate, activity

Reference - Lamarsh (S. p. 18-26)

n poor - $Z > N$ - β^+ decay - $X \xrightarrow{\beta^+} X-1 + \nu$

n rich - $Z < N$ - β^- decay - $X \xrightarrow{\beta^-} X+1 + \bar{\nu}$
competes with e capture

n rich - heavy, very unstable - α decay - $X \xrightarrow{\alpha} X-2 + \alpha$



Decay rates

$$\lambda = \frac{\ln(2)}{T_{1/2}}$$

$$N(t) = N(0)e^{-\lambda t}$$

Single decay:

$$d(t) = d_0 e^{-\lambda t}$$

Binary system:
with production

$$\frac{dn}{dt} = -\lambda n + R$$

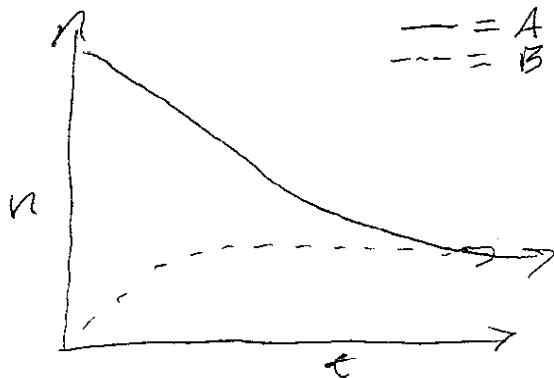
$$n = n_0 e^{-\lambda t} + \frac{R}{\lambda} (1 - e^{-\lambda t})$$

Binary system:

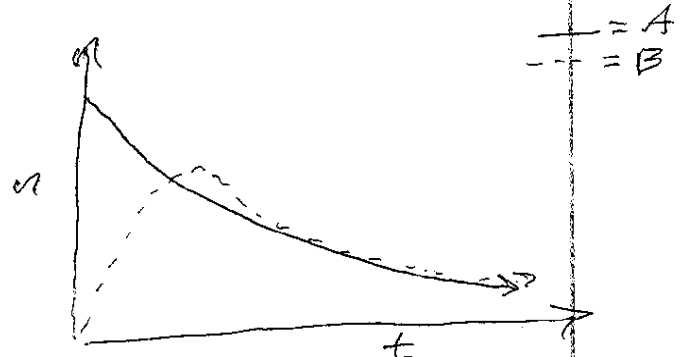
$$\frac{dn_B}{dt} = -\lambda_B n_B + \lambda_A n_A(t)$$

$$n_B = n_{B0} e^{-\lambda_B t} + \frac{\lambda_A n_{A0}}{\lambda_B - \lambda_A} (e^{-\lambda_A t} - e^{-\lambda_B t})$$

$$t^* = \frac{T_P T_d}{\ln(2)(T_P - T_d)} \ln(T_P/T_d)$$



secular
 $\lambda_B > \lambda_A$



transient
 $\lambda_B < \lambda_A$

3) Binding Energy for atoms and nuclei, mass defect

Reference - Lamarsh (2, p. 11)

mass \rightarrow energy by $E_{\text{rest}} = m_0 c^2$

then remember that KE is difference between total and rest mass energy

$$E = mc^2 - m_0 c^2 = m_0 c^2 \left[\frac{1}{\sqrt{1-\beta^2}} - 1 \right]$$

leads to $\frac{1}{2}mv^2$ for $v \ll c$

Binding Energy:
remember

$$Q = [(M_a + M_b) - (M_c + M_d)] \overbrace{931 \text{ MeV/amu}}^{c^2}$$

so the mass defect is difference between constituents and the actual atom

$$D = ZM_p + NM_n - M_A \quad (Z, N \text{ are coefficients, } A \text{ is product})$$

so binding energy

$$BE(A) = Z_a \underbrace{M(^1\text{H})}_{\text{in MeV}} + N_a \underbrace{(M_n)}_{\text{in MeV}} - \underbrace{M_a}_{\text{in MeV}}$$

LAST

neutron is the last bound neutron

$$E_s = [M_n + \underbrace{M(A-1, Z)}_{\text{nuclei}} - \underbrace{M(A, Z)}_{\text{nuclei}}] 931 \text{ MeV/amu}$$

Possible problems:

- Calculate binding energy of the last neutron in ^{13}C . (Lamarsh, p. 33)

2) Nuclear structure and Nuclear Radii Estimate

Reference - Lemarsh (I), p 11, 33-36

$$R = 1.25 \text{ fm} \times A^{1/3}$$

 \Rightarrow indicates uniform density

for all atoms w/ several shells, atomic radius is close to $2 \times 10^{-10} \text{ m}$

Reference - Lemarsh (I) p. 60
 $r = 1.2 \times 10^{-15} \times A^{1/3}$

Nuclear Models

shell - Pauli exclusion principle
 there are $2j+1$ possible states for
 substates w/ total momentum j .
 • orbits. Helps solve binding
 energy problems
 • magic numbers correspond to
 closed shells in both M and P

Liquid
Drop

$$M = NM_n + ZM_p - \alpha A + \beta A^{2/3} + \gamma Z^2/A^{1/3} + \xi(A - 2Z)/A + \delta$$

δ : pairing term
 the bond between $n+n$ or $p+p$
 is stronger than $n+p$

ξ : symmetry term, stronger for
 Z and N the same

γ : coulombic term \rightarrow from potential energy

β : surface tension ($R^2 \propto A^{2/3}$)

α : total volumetric correction, total
 energy of one bond

Possible problems:

probably qualitative or a small
 part of other problems