## Gamma Cross Sections

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### Gamma Cross Sections

- Analyzing the Data
- Issues encountered in Analysis
- Resolution

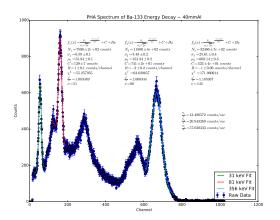
The general theory of each step will be interspersed throughout.

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### Gaussian Fits to Data



#### Notes:

- Notice the fit function linear background
- $ilde{\chi}^2$  gives confidence in the form of the fit as well as the fit itself

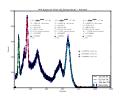
### Gaussian Fits to Data

#### Fit Function:

$$f(x) = \frac{N}{\sigma\sqrt{2\pi}}e^{\frac{-(x-\mu)^2}{2\sigma^2}} + Ax + B$$

- Linear background is not a rate in time; it's a rate over the channels
- ▶ ⇒ Background measures:
  - susceptibility (to noise) by channel A
  - overall susceptibility (to noise) B
- We assume that this noise is random and is not caused by our apparatus
  - It comes from another source(s)
- Why aren't they delta functions?
  - energies are one value, width comes from detector

### Gaussian Fits to Data

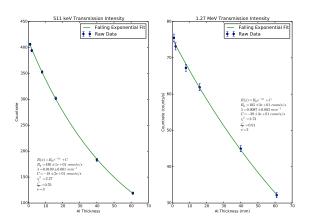


### Then what are these other peaks?

- Compton Edge and Backscatter
  - Peak energies are known ignore these other features<sup>†</sup>

 $<sup>^\</sup>dagger$  We should never ignore features, but once we verify that we aren't interested in them we can effectively ignore them.

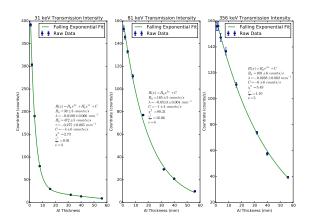
# Fitting to Find Countrates



Good fits, good agreement with the theory (for the fit at least)



# Fitting to Find Countrates



Return to the Ba countrate fits later on



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# How do we stack up?

So the Na fit is good, are the values any good?

$$\mu = \frac{\mu}{\rho} \times \rho = 0.223 \text{cm}^{-1}$$

[1]

Uh oh...

$$\lambda_{exp} \pm \sigma_{\lambda} = 0.19 \pm 0.01 \, cm^{-1}$$
 $\sigma_{\lambda} = .01$ 

 $+3\sigma \implies$  we either made a discovery or we have a confidence of

$$100 - 99.74 = 0.26\%$$
.

[3]

# How do we stack up?

All is not lost, however! We should test the next measurement.

$$\mu = \frac{\mu}{\rho} \times \rho = 0.143 cm^{-1}$$

Hmm, not looking good.

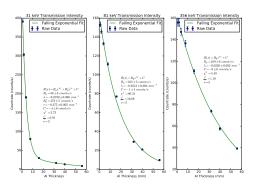
$$\lambda_{exp} \pm \sigma_{\lambda} = 0.09 \pm 0.03 \, cm^{-1}$$
 $\sigma_{\lambda} = .03$ 

 $-1.67\sigma \implies$  confidence of

$$100 - 90.5 = 9.5\%$$

[3] which just tells us we're using student-grade lab equipment in a room of people doing the same experiment.

### Ba Silliness



$$R(x) = R_0 e^{\lambda x} + R_0' e^{\tau x} + C$$
 31 keV

$$R(x) = R_0 e^{\lambda x} + R_0' e^{\tau x} + C$$

 $R_0e^{\lambda x} + C$  is the normal expression.

- Does not account for photoelectric effect dominance at low energies (like 31 keV)!
- $ightharpoonup \lambda = \mathsf{Compton} \ \mathsf{effect} \ \mathsf{linear} \ \mathsf{attenuation} \ \mathsf{coefficient}$
- au T = Photoelectric effect linear attenuation coefficient
- $\lambda + \tau = 2.94 \text{ cm}^{-1}$

## Ba Silliness

So, what's the literature value?

$$\mu = 2.98 \text{ cm}^{-1}$$

$$(\lambda + \tau) \pm \sigma_{\lambda + \tau} = 2.94 \pm 0.11 \text{ cm}^{-1}$$

$$\implies 0.36\sigma$$

Which gives us a confidence value of

$$100 - 28.12 = 71.88\%$$

in our measurement!

Note that the extra significant figures were added to give a sense of how close the value is, the measurement is properly reported as  $2.9\pm0.11\,cm^{-1}$ 

# And the other energies?

Well those are easy.

$\mu$	$\lambda \pm \sigma_{\lambda}$	$\sigma$
$0.533  cm^{-1}$	$0.51 \pm .04~cm^{-1}$	$0.17\sigma$
$0.260  cm^{-1}$	$.23 \pm .02  cm^{-1}$	$0.67\sigma$

values of lambda are not right on, just best I could find

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other stuff here???

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### References



Physics.NIST.gov - Table of XRay Mass Attenuation Coefficients http://physics.nist.gov/PhysRefData/XrayMassCoef/ElemTab/z13.html



University of Chicago, PHYS 211 Lab Manual - P211 Wiki



An Introduction to Error Analysis - John Taylor