

Local and global models

- We can make models of light-matter interaction in two ways
 - Locally: we treat each object in a scene separately from any other object
 - Globally: we treat all objects together, and model the interactions between objects
- we'll develop a simple local illumination model in detail
- we'll look at global models in COMP37111 next year

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Illumination models

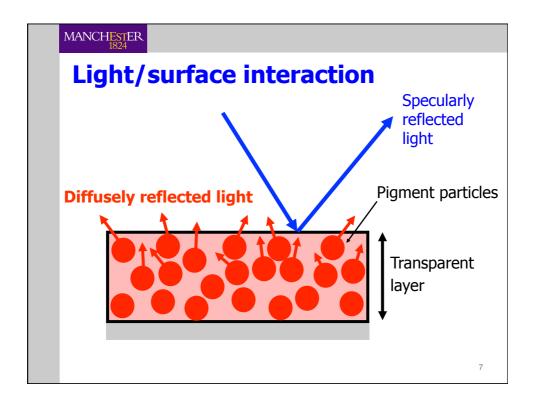
- The interaction of light and matter is an extremely complex process
- In computer graphics we try to model this process. In other words, we approximate it
- Computer pictures are digital, with finite precision. We can only ever approximate.

Local illumination: elements

- We'll develop a model step-by-step, to include the following:
 - Ambient illumination
 - Diffuse reflection
 - Positional light source
 - Specular reflection
 - Colour of lights and surfaces

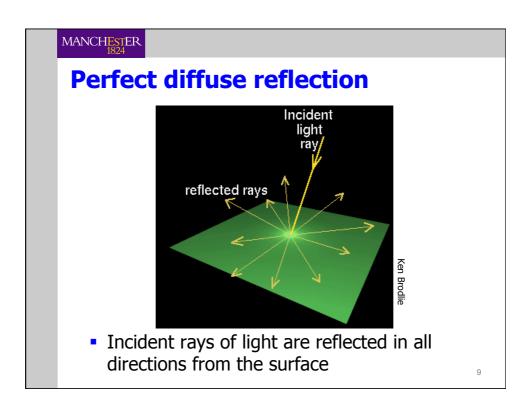
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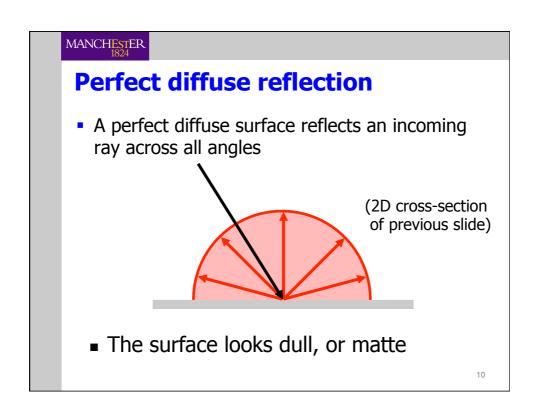
Reflectivity • There are three kinds of reflection: • Perfect diffuse reflection • Perfect specular reflection • Imperfect specular reflection



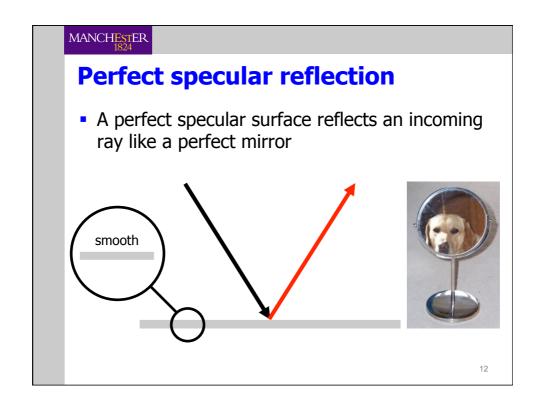
Diffuse and specular reflection

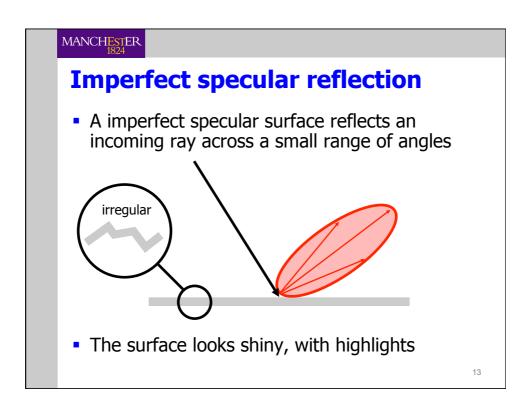
- Diffuse reflection is absorption and uniform reradiation
 - Some wavelengths are absorbed, some are reflected
 - a green object looks green because it only reflects green
- Specular reflection is reflection at the air/ surface interface
 - To a first approximation, the colour of the specular reflection is that of the light source

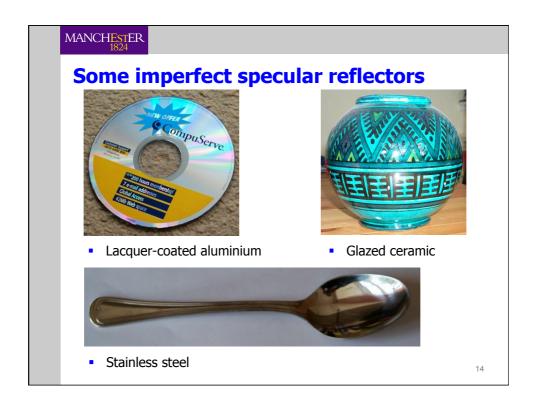






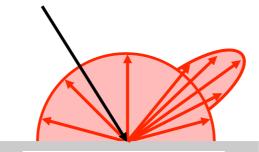






Diffuse/specular surfaces

 Most surfaces exhibit a combination of diffuse and specular reflection



 We can model these effects with varying degrees of realism

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Illumination sources

- We begin our development of the local illumination model by considering diffuse reflection, and sources of illumination:
 - Ambient illumination
 - Point illumination source at infinity (directional illumination)
 - Point illumination source in the scene

Ambient illumination

- Consider an environment with a light source
- Multiple reflections cause a general level of illumination in the scene



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Ambient illumination

- lacksquare If intensity of ambient light is I_a
- Amount of light diffusely reflected from a surface is

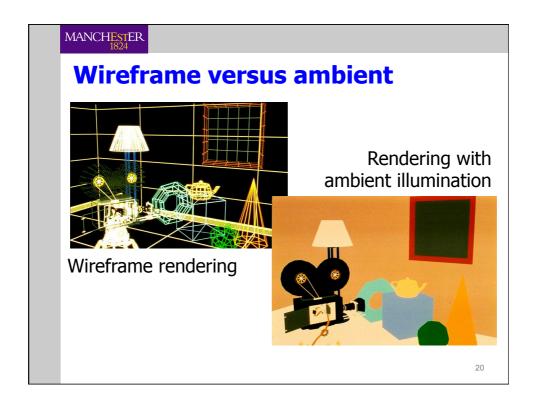
$$I_{\text{ambient}} = k_a I_a$$

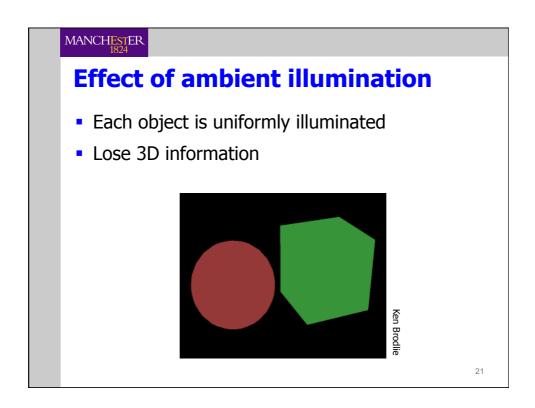
Where k_a is the ambient reflection coefficient of the surface

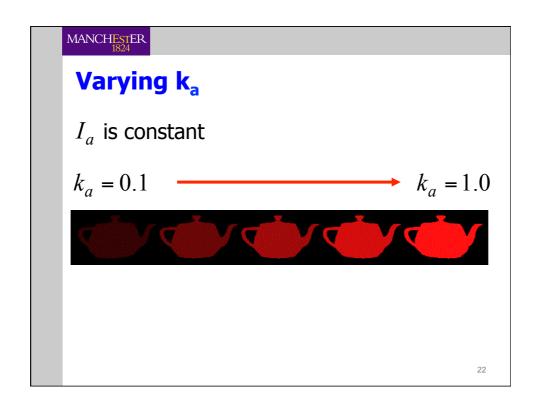
$$0 \le k_a \le 1$$

Local illumination model v1

- We now have the first term in the model we're developing
- *I* = ambient
- $I = k_a I_a$







True ambient lighting

Note: the

 $I = k_a I_a$

term is a gross simplification of true ambient illumination, which is **not** constant in a scene



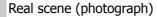
Advanced Interfaces Group

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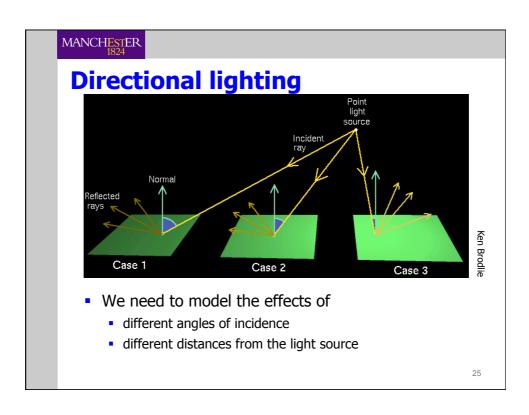
True ambient lighting

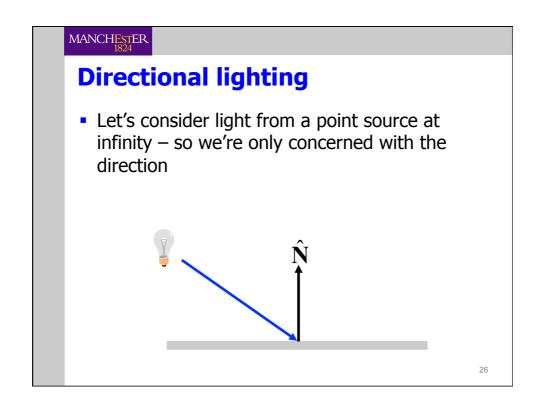




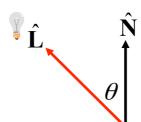


Scene modelled and rendered with accurate global illumination model





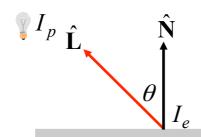
Describing surface orientation



- $\ \ \, \hat{N}$ is surface normal
- ${}^{\color{red} \bullet}$ \hat{L} is direction of light source
- ullet heta is angle of incidence
- Note the vectors are normalised

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Diffuse reflection: Lambert's Law



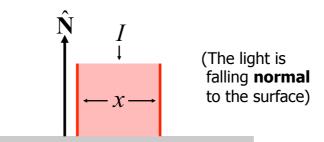
- Light source of intensity \boldsymbol{I}_p Effective intensity received is \boldsymbol{I}_e
- Lambert's Law: $I_e = I_p \cos \theta$



Johann Heinrich Lambert (1728-1777)

Lambert's Law derived

• Consider light of intensity I and cross-sectional width x falling on a surface:



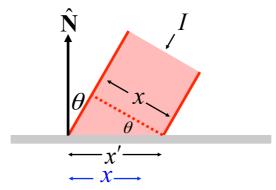
 So width x on surface receives all of intensity I

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Lambert's Law derived

- Now consider the light inclined at heta
- $x' = x/\cos\theta$, so $x = x'\cos\theta$
- So width x receives intensity $I\cos\theta$



Diffuse reflectivity

- We express how good a diffuse reflector a surface is using k_d
- k_d is the **diffuse reflection coefficient** of the surface, $0 \le k_d \le 1$
- Amount of diffusely reflected light is
 - $I_{\text{diffuse}} = I_p k_d \cos \theta$

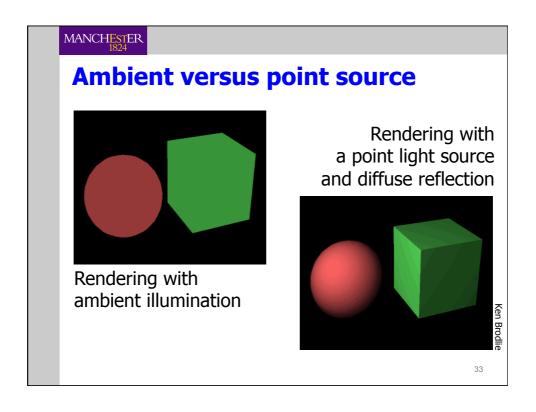
 $I_{\text{diffuse}} = I_p k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$

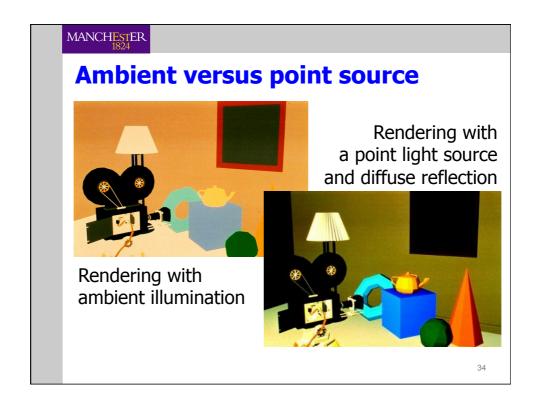
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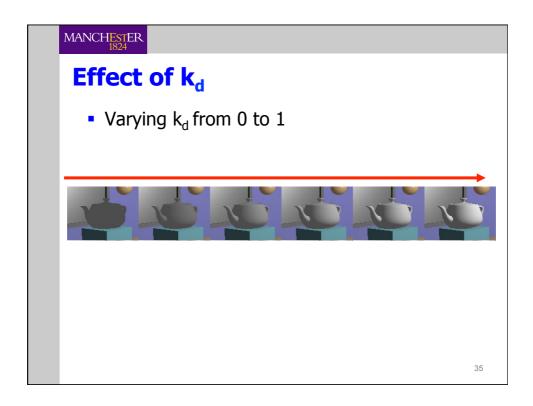
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Local illumination model v2

- I = ambient + diffuse
- $I = k_a I_a + I_p k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$







Light source distance

- Physically, light intensity falls off with the square of distance travelled
- $\ ^{\bullet}$ After travelling d , original intensity I_{p} is now I_{e}

$$I_e = \frac{I_p}{4\pi d^2}$$

Light source distance

- While this is physically correct, it doesn't always work well for computer graphics
- We only have a limited number of pixel intensities, and often the d^2 term changes too rapidly, so instead we use:

$$I_e = \frac{I_p}{k_c + k_l d + k_q d^2}$$
 We can choose k_c, k_l, k_q for the best results

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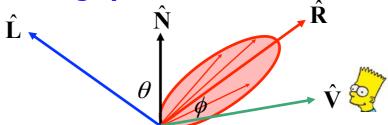
Local illumination model v3

• I = ambient + distance (diffuse)

$$I = k_a I_a + \frac{I_p}{d'} k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}})$$

• Where $d' = k_c + k_l d + k_q d^2$

Modelling specular reflection

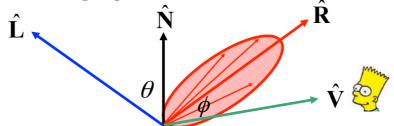


- $\hat{\mathbf{R}}$ is a vector giving the direction of maximum specular reflection.
- $\hat{\mathbf{R}}$ makes an angle of heta with $\hat{\mathbf{N}}$, as does $\hat{\mathbf{L}}$
 - (for a perfect mirror, angle of incidence == angle of reflection)
- $f \hat{V}$ is a vector pointing to the observer's position

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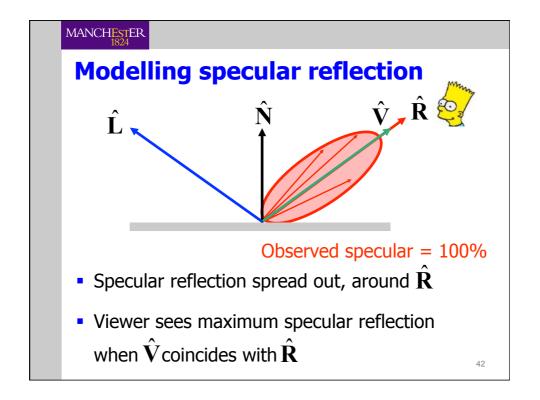
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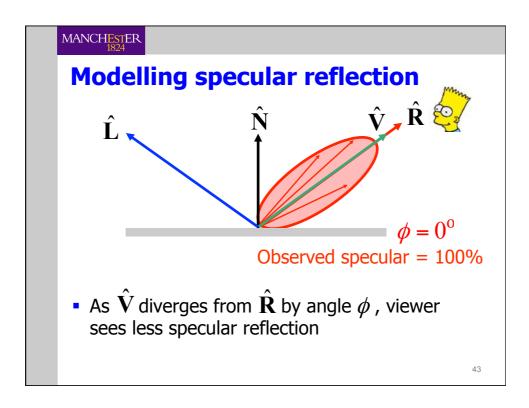
Modelling specular reflection

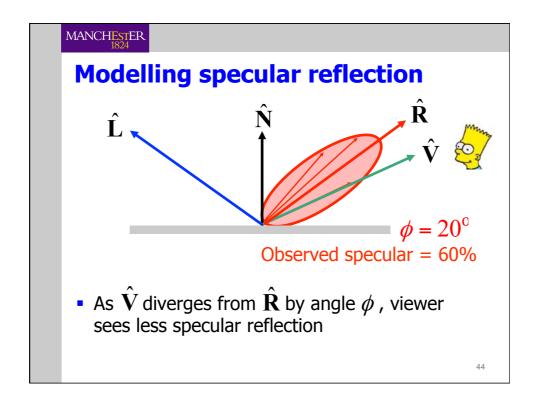


- Experiment shows specular reflection varies with:
 - Angle ϕ between $\hat{f R}$ and $\hat{f V}$
 - ullet Incident angle heta and light wavelength λ
 - So we seek a function $I_{\text{specular}} = S(\phi, \theta, \lambda)$

Manchester Modelling specular reflection • What is this function $S(\phi, \theta, \lambda)$? • We'll look at the effects of ϕ, θ, λ in turn • First, ϕ $\hat{\mathbf{L}}$





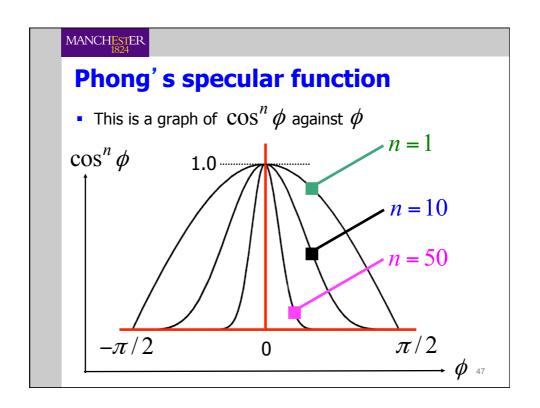


Modelling specular reflection $\hat{L} \qquad \hat{N} \qquad \hat{R} \qquad \hat{V} \qquad \hat{\phi} = 40^{\circ}$ Observed specular = 5% $\hat{V} \qquad \hat{V} \qquad \hat$

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Modelling specular reflection

- Variation of observed specular = $F(\phi)$
- But what is the function *F* ?
- Bui-Tuong Phong (1942-1975) proposed using the function $\cos^n \phi$



Phong's specular function

So we now have

•
$$I_{\text{specular}} = I_p \cos^n \phi$$

• Which we can rewrite using vectors as

$$I_{\text{specular}} = I_p (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n$$

• Normally we use $1 \le n \le 200$









tichard Lobb

n = 10

n = 20

n = 80

n = 200

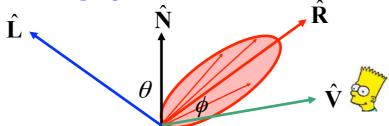
Incident angle and wavelength

- Recall $I_{\text{specular}} = S(\phi, \theta, \lambda)$
- We've accounted for ϕ
- Now we look at the effects of θ and λ

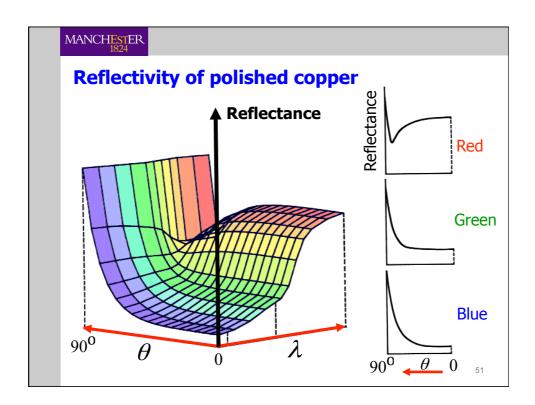
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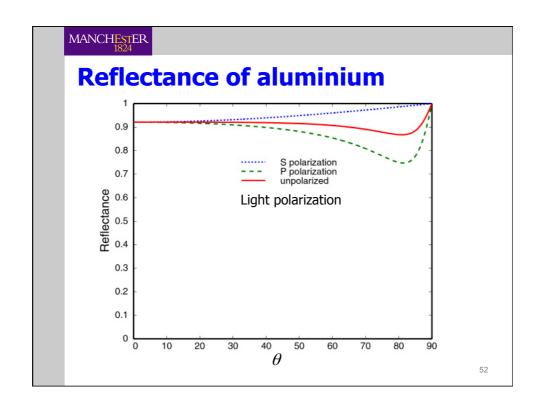
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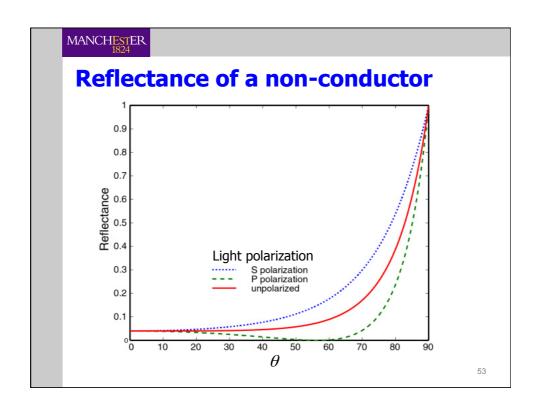
Modelling specular reflection



- Experiment shows specular reflection varies with:
 - Angle ϕ between $\hat{\mathbf{R}}$ and $\hat{\mathbf{V}}$
 - ullet Incident angle heta and light wavelength λ
 - So we seek a function $I_{\text{specular}} = S(\phi, \theta, \lambda)$







Incident angle and wavelength

- This complex variation is expressed by the Fresnel equation (Augustin-Jean Fresnel, 1788-1827)
 - ➤ A founder of the wave theory of light.
 - Developed theory of diffraction of light.
 - Obtained circularly polarised light
 - Developed the use of compound lenses instead of mirrors for lighthouses.

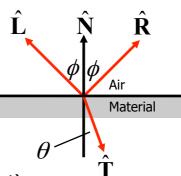


Incident angle and wavelength

 This complex variation is expressed by the Fresnel equation (Augustin-Jean Fresnel, 1788-1827)

$$F = \frac{1}{2} \left[\frac{\sin^2(\phi - \theta)}{\sin^2(\phi + \theta)} + \frac{\tan^2(\phi - \theta)}{\tan^2(\phi + \theta)} \right]$$

- *F* is the fraction of light reflected
- $\sin \theta = \sin \phi / \mu$
- μ is the refractive index of the material (λ dependent)



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Incident angle and wavelength

- In practice, we usually approximate F with a single constant $k_{\rm s}$
- k_s is the **specular reflection coefficient** of the surface, $0 \le k_s \le 1$
- $I_{\text{specular}} = I_p \mathbf{k}_s (\mathbf{R} \cdot \mathbf{V})^n$
- But, we sacrifice accuracy for efficiency

Local illumination model v4

I = ambient + distance (diffuse + specular)

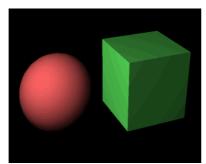
$$I = k_a I_a + \frac{I_p}{d'} \left[k_d (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}}) + k_s (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n \right]$$

• Where $d' = k_c + k_l d + k_q d^2$

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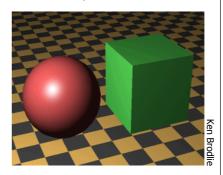
MANCHESTER 1824

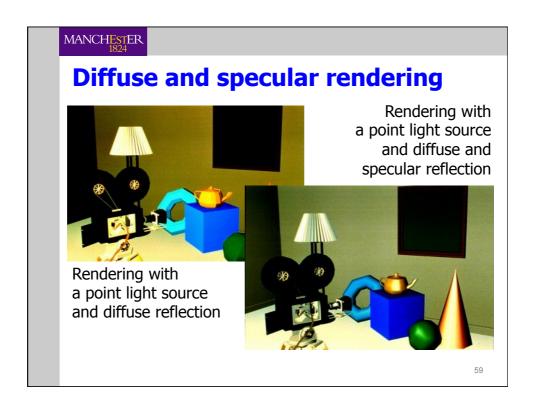
Diffuse and specular rendering

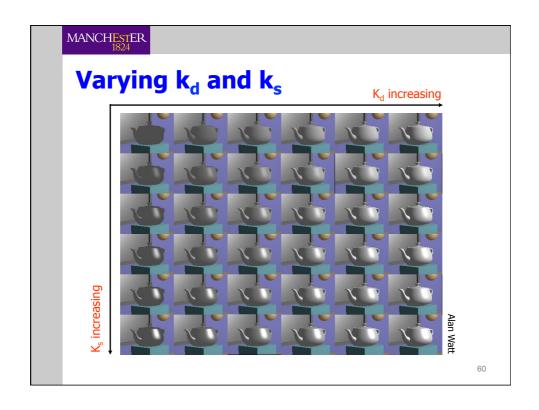


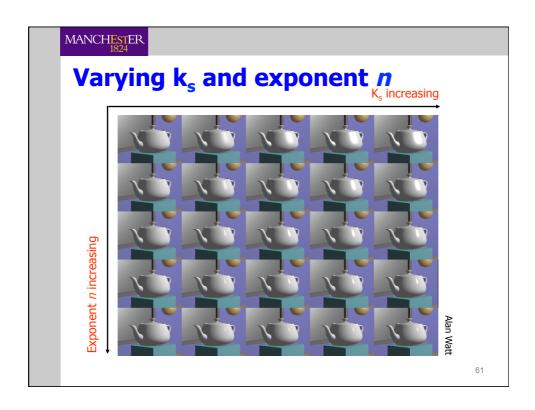
Rendering with a point light source and diffuse reflection

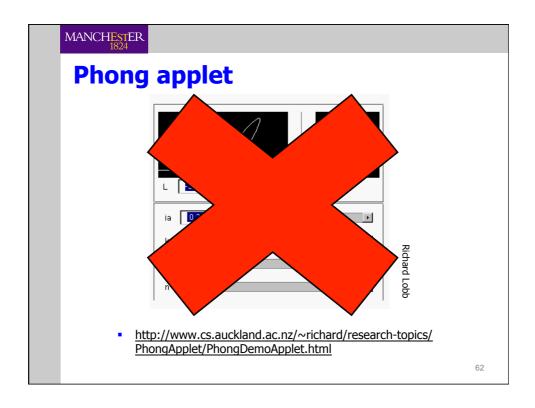
Rendering with a point light source and diffuse and specular reflection



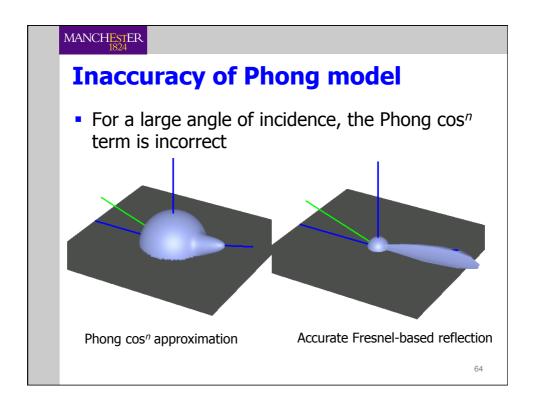












Incorporating colour

- So far, we've only considered light intensity, not colour
- It's easy to incorporate we express light colour as a triple of RGB intensities:

•
$$I_{pR}, I_{pG}, I_{pB}$$

- And correspondingly we express surface colour using
 - k_{aR}, k_{aG}, k_{aB}
 - k_{dR}, k_{dG}, k_{dB}

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Local illumination model v5

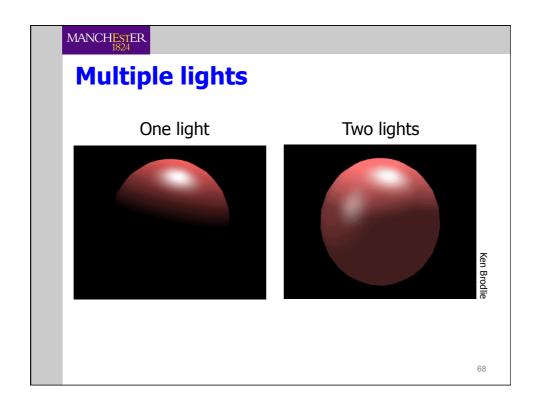
- We now have a separate expression for each colour component. For example, for red:
- I_R = ambient_{Red} + distance (diffuse_{Red} + specular)

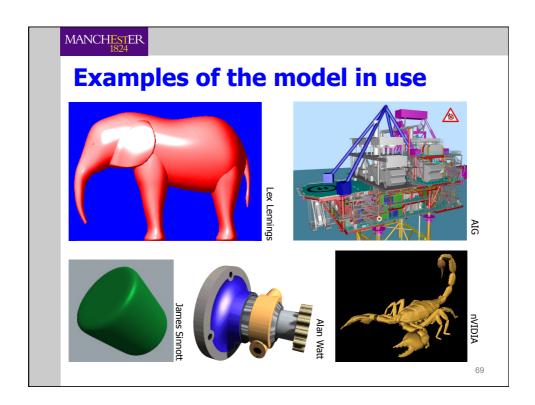
$$I_R = k_{aR} I_{aR} + \frac{I_{pR}}{d'} \left[k_{dR} (\hat{\mathbf{N}} \cdot \hat{\mathbf{L}}) + k_s (\hat{\mathbf{R}} \cdot \hat{\mathbf{V}})^n \right]$$

Multiple lights

- Finally, what if we have more than one light?
- Easy, compute illumination separately for each and sum. For M lights:

•
$$I = ambient + \sum_{i=1}^{M} (diffuse_i + specular_i)$$





The "standard" model

- The local illumination model we've developed is the "standard" model in use today
- Implemented in OpenGL
- Implemented in hardware on consumer 3D graphics cards