Geodesica Alrededor de una Particula con Masa en el Espacio de Minkowski

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Clase de Metodos Geometricos





Partiendo de la metrica de Minkowski en esfericas:

$$ds^2_{\text{Minkowski}} = -dt^2 + dr^2 + r^2 d\Omega^2$$

Necesitamos una metrical que mantenga la forma del angulo solido, así que podemos multiplicarlo por una funcion radial:

$$ds^{2} = -e^{2\alpha(r)}dt^{2} + e^{2\beta(r)}dr^{2} + e^{2\gamma(r)}r^{2}d\Omega^{2}$$

Aplicamos la siguiente transformacion:

$$\bar{r} = e^{\gamma(r)}r$$
 $d\bar{r} = \left(1 + r\frac{d\gamma}{dr}\right)e^{\gamma}dr$

La metrica nos queda:

$$ds^{2} = -e^{2\alpha(r)}dt^{2} + \left(1 + r\frac{d\gamma}{dr}\right)^{-2}e^{2\beta(r) - 2\gamma(r)}d\bar{r}^{2} + \bar{r}^{2}d\Omega^{2}$$



$$ds^2 = -e^{2\alpha(r)}dt^2 + \left(1 + r\frac{d\gamma}{dr}\right)^{-2}e^{2\beta(r)-2\gamma(r)}d\bar{r}^2 + \bar{r}^2d\Omega^2$$

Haciendo ar r o r obtenemos que: $\left(1+rrac{d\gamma}{dr}
ight)^{-2}e^{2eta(r)-2\gamma(r)} o e^{2eta}$

Por lo tanto podemos escribir la metrica como:

$$ds^{2} = -e^{2\alpha(r)}dt^{2} + e^{2\beta(r)}dr^{2} + r^{2}d\Omega^{2}$$





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Calculamos los simbolos de Christoffel:

$$\begin{split} \Gamma^t_{tr} &= \partial_r \alpha & \Gamma^r_{tt} &= e^{2(\alpha - \beta)} \partial_r \alpha & \Gamma^r_{rr} &= \partial_r \beta \\ \Gamma^\theta_{r\theta} &= \frac{1}{r} & \Gamma^r_{\theta\theta} &= -re^{-2\beta} & \Gamma^\phi_{r\phi} &= \frac{1}{r} \\ \Gamma^r_{\phi\phi} &= -re^{-2\beta} \sin^2 \theta & \Gamma^\theta_{\phi\phi} &= -\sin \theta \cos \theta & \Gamma^\phi_{\theta\phi} &= \frac{\cos \theta}{\sin \theta} \end{split}$$





Con los simbolos de Christoffel podemos obtener el tensor de Riemann:

$$\begin{split} R^t_{rtr} &= \partial_r \alpha \partial_r \beta - \partial_r^2 \alpha - (\partial_r \alpha)^2 \\ R^t_{\theta t \theta} &= -r e^{-2\beta} \partial_r \alpha \\ R^t_{\phi t \phi} &= -r e^{-2\beta} \sin^2 \theta \partial_r \alpha \\ R^r_{\theta r \theta} &= r e^{-2\beta} \partial_r \beta \\ R^r_{\phi r \phi} &= r e^{-2\beta} \sin^2 \theta \partial_r \beta \\ R^\theta_{\phi \theta \phi} &= \left(1 - e^{-2\beta}\right) \sin^2 \theta \end{split}$$





Tomando la contraccion obtenemos el tensor de Ricci:

$$R_{tt} = e^{2(\alpha - \theta)} \left[\partial_r^2 \alpha + (\partial_r \alpha)^2 - \partial_r \alpha \partial_r \beta + \frac{2}{r} \partial_r \alpha \right]$$

$$R_{rr} = -\partial_r^2 \alpha - (\partial_r \alpha)^2 + \partial_r \alpha \partial_r \beta + \frac{2}{r} \partial_r \beta$$

$$R_{\theta\theta} = e^{-2\beta} \left[r \left(\partial_r \beta - \partial_r \alpha \right) - 1 \right] + 1$$

$$R_{\phi\phi} = \sin^2 \theta \ R_{\theta\phi}$$





Reemplazamos el tensor de Ricci en la ecuacion de campo de Einstein en el vacio para hallar α y β :

$$R_{\mu\nu}-\frac{1}{2}Rg_{\mu\nu}=0$$

Obteniendo:

$$e^{2(\beta-\alpha)}e^{2(\alpha-\beta)}\left[\partial_r^2\alpha + (\partial_r\alpha)^2 - \partial_r\alpha\partial_r\beta + \frac{2}{r}\partial_r\alpha\right] \\ - \partial_r^2\alpha - (\partial_r\alpha)^2 + \partial_r\alpha\partial_r\beta + \frac{2}{r}\partial_r\beta = 0$$
$$\Longrightarrow \frac{2}{r}\partial_r\alpha + \frac{2}{r}\partial_r\beta = 0$$

Para que se cumpla esta ecuación necesitamos que $\alpha = -\beta + c$



$$\alpha = -\beta$$

Reemplazando en $R_{\theta\theta} = 0$:

$$e^{2\alpha}(2r\partial_r\alpha+1)=1\Longrightarrow\partial_r(re^{2\alpha})=1$$

Resolviendo esta ecuacion obtenemos:

$$e^{2\alpha} = 1 - \frac{R_s}{r}$$

De esta forma obtenemos que podemos escribir la metrica como:

$$ds^2 = -\left(1 - \frac{R_s}{r}\right)dt^2 + \left(1 - \frac{R_s}{r}\right)^{-1}dr^2 + r^2d\Omega^2$$





EXPLICAR POR QUÉ EL RADIO DE Schwarzschild ES $R_S = 2GM$





Geodesicas de Schwarzschild

$$ds^{2} = -\left(1 - \frac{2GM}{r}\right)dt^{2} + \left(1 - \frac{2GM}{r}\right)^{-1}dr^{2} + r^{2}d\Omega^{2}$$

Calculamos los simbolos de Christoffel

```
m = var('m')
M = Manifold(4, 'R^4', start_index=1)
c spher.<t,r,th,ph> = M.chart(r't:(0,+oo) r:(0,+oo) th:(0,pi):\theta ph:(0,2*pi):\phi')
q = M.metric('q')
g[1,1], g[2,2], g[3,3], g[4,4] = (-1)*(1-2*m/r), (1-2*m/r)^{(-1)}, r^2, r^2*\sin(th)^2
print(q.christoffel_symbols_display(chart=c_spher))
BL.<t,r,th,ph> = M.chart(r"t r:(0,+oo) th:(0,pi):\theta ph:(0,2*pi):\phi")
M.default frame() is BL.frame()
xi = BL.frame()[0]
xi_form = xi.down(q)
print(xi_form.display())
```

Geodesicas de Schwarzschild

Simbolo de Christoffel para la metrica:

$$\begin{array}{lll} \Gamma^t_{\ tr} & = & -\frac{M}{2\,Mr-r^2} \\ \Gamma^r_{\ tt} & = & -\frac{2\,M^2-Mr}{r^3} \\ \Gamma^r_{\ rr} & = & \frac{M}{2\,Mr-r^2} \\ \Gamma^r_{\ \theta\theta} & = & 2\,M-r \\ \Gamma^r_{\ \theta\phi} & = & (2\,M-r)\sin(\theta)^2 \\ \Gamma^\theta_{\ r\theta} & = & \frac{1}{r} \\ \Gamma^\theta_{\ \phi\phi} & = & -\cos(\theta)\sin(\theta) \\ \Gamma^\phi_{\ r\phi} & = & \frac{1}{r} \\ \Gamma^\phi_{\ \theta\phi} & = & \frac{\cos(\theta)}{\sin(\theta)} \end{array}$$





Geodesicas de Schwarzschild

Reemplazando los simbolos de Christoffel en la ecuacion de la geodesica:

$$\ddot{t} + \frac{2M}{r(r-2M)}\dot{r}\dot{t} = 0$$

$$\ddot{r} + \frac{M}{r^3}(r-2M)\dot{t}^2 - \frac{M\dot{r}}{r(r-2M)} - (r-2M)\dot{\phi}^2 = 0$$

$$\ddot{\phi} + \frac{2}{r}\dot{\phi}\dot{r} = 0$$



