Home Work #1

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March 16, 2023

1 Question 1

 $h = 200_{km} \rightarrow r = R_e + h = 6378.137 + 200 = 6578.1137, \quad \mu = 3.986 \times 10^{14}_{m^3/s^2} = 3.986 \times 10^5_{km^3/s^2}$ The orbit is circular.

1.1 Part a

$$T = 2\pi \sqrt{\frac{r^3}{\mu}} = 2\pi \sqrt{\frac{6578.1137^3}{3.986 \times 10^5}} = 5309.62_{\text{sec}}$$
 (1)

$$T\omega = 2\pi \to \omega = \frac{2\pi}{T} = \frac{2\pi}{5309.62} = 0.00118_{rad/sec}$$
 (2)

1.2 Part b

$$v = \sqrt{\frac{\mu}{r}} = \sqrt{\frac{3.986 \times 10^5}{6578.1137}} = 7.78_{km/\text{sec}}$$
 (3)

The new velocity is calculated as:

$$v_{new} = v + 0.5_{km.\,\text{sec}} = 8.28_{km.\,\text{sec}} \tag{4}$$

Assume the new orbit is circular just changed altitude and has a new velocity.

$$v_{new} = \sqrt{\frac{\mu}{r_{new}}} \to r_{new} = \frac{\mu}{v_{new}^2} = 5808_{km}$$
 (5)

$$T = 2\pi \sqrt{\frac{r_{new}^3}{\mu}} = 2\pi \sqrt{\frac{5808^3}{3.986 \times 10^5}} = 4405.08_{\text{sec}}$$
 (6)

$$T_{new}\omega_{new} = 2\pi \to \omega_{new} = \frac{2\pi}{T_{new}} = \frac{2\pi}{4405.08} = 0.00143_{rad/sec}$$
 (7)

 r_{new} is smaller than the earth's radius.

2 Question 2

Assume \mathbf{r}_0 and \mathbf{v}_0 as:

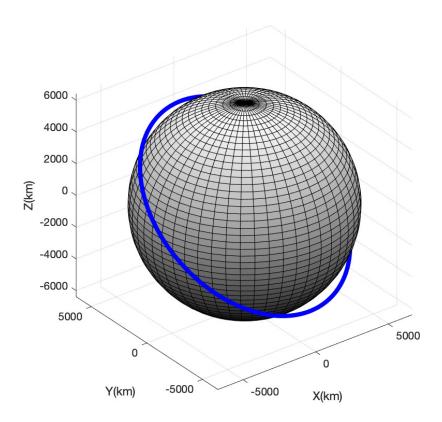
$${m r}_0 = \begin{bmatrix} 1600 & 5310 & 3800 \end{bmatrix}_{km}^{\mathrm{T}}, \quad {m v}_0 = \begin{bmatrix} -7.350; & 0.4600 & 2.470 \end{bmatrix}_{km/\,\mathrm{sec}}^{\mathrm{T}}$$

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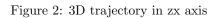
2.1 Part a

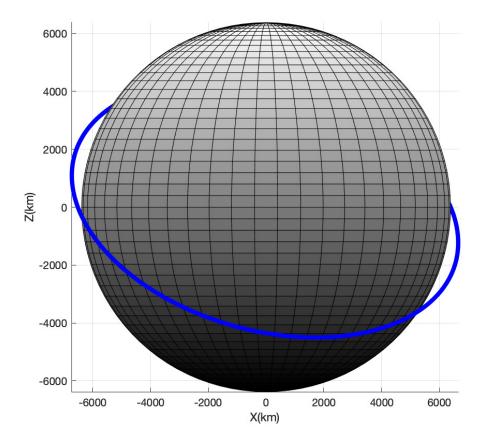
the n-body problem was solved with MATLAB with the n-body function in the question directory. The results are illustrated below.

Figure 1: 3D trajectory



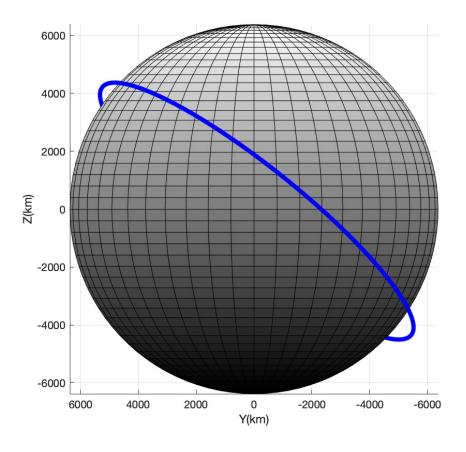
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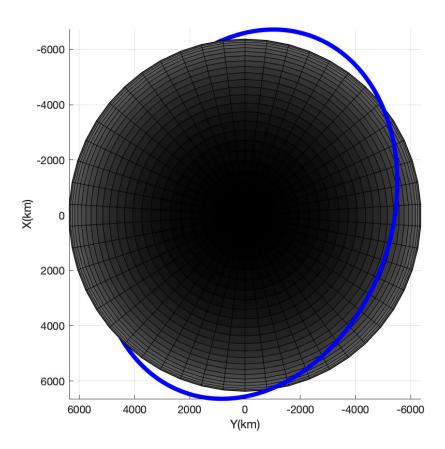


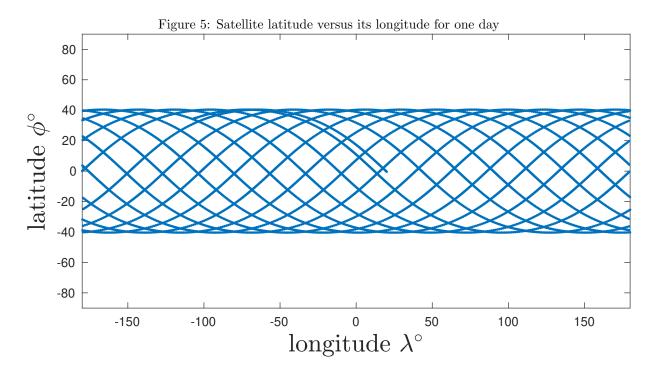
Figure 4: 3D trajectory in xy axis

2.2 Part b

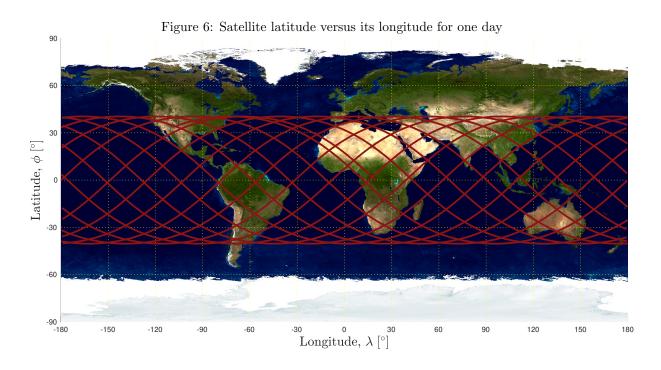
Using the below transfer matrix to transfer from ECI coordinate to the ECEF coordinate.

$$\boldsymbol{T}^{ECCF-ECI} = \begin{bmatrix} \cos(\omega_E t) & -\sin(\omega_E t) & 0\\ \sin(\omega_E t) & \cos(\omega_E t) & 0\\ 0 & 0 & 1 \end{bmatrix}$$
$$\boldsymbol{\phi} = \arccos(\frac{\boldsymbol{r}(3)}{r})$$
$$\boldsymbol{\lambda} = \begin{cases} \arctan(\frac{\boldsymbol{r}(1)}{r_{xy}}), & \boldsymbol{r}(2) > 0\\ 2\pi - \arctan(\frac{\boldsymbol{r}(1)}{r_{xy}}), & \boldsymbol{r}(2) \le 0 \end{cases}$$

Ali BaniAsad 401209244 2.2 Part b



Below the figure drawn provided by tamaskis, please click here to see the source code. Please use the mentioned library to run code or skip the part on earth fig.

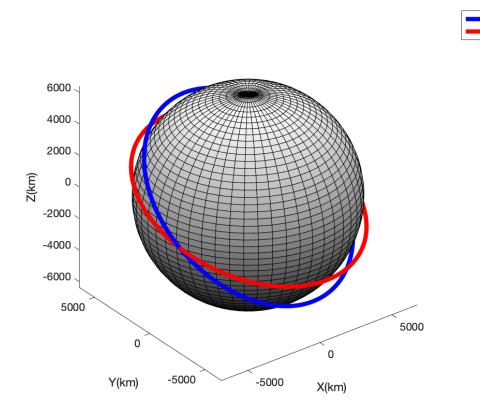


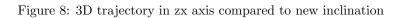
2.3 Part c (Bonus)

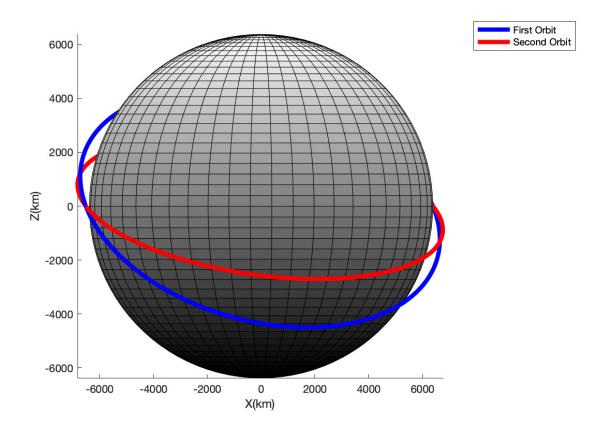
In this section find the orbital elements, then, change the inclination to 0.4_{rad} (using oe2ecf and vec2orbElem functions). Then, find \mathbf{r}_0 and \mathbf{v}_0 by new orbital elements. The results are illustrated below.

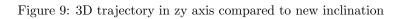
Figure 7: 3D trajectory compared to new inclination

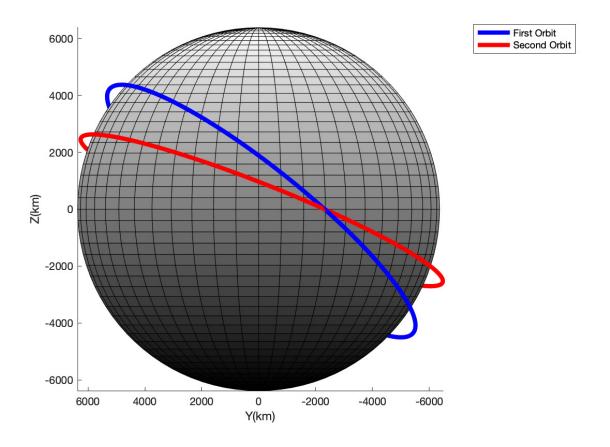
First Orbit Second Orbit

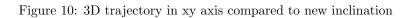


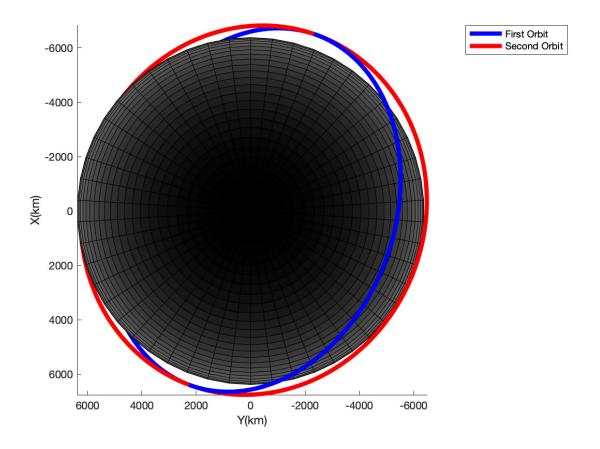


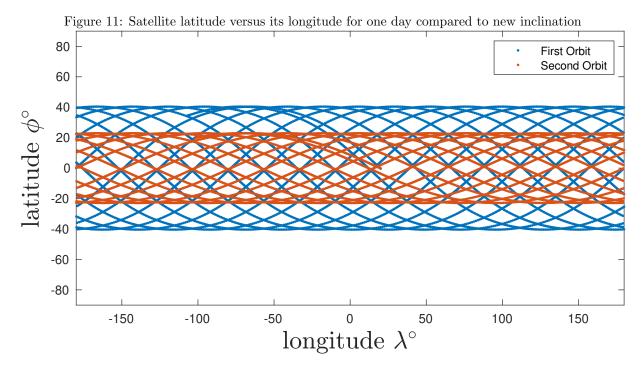












In general, changing the inclination of a satellite's orbit will cause its ground track to move to different latitudes and longitudes on the Earth's surface. In this example when the inclination changed from 0.7_{rad} to 0.4_{rad} , the range of longitude decreased.

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