In The Name of God



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45-784: Advanced Orbital Mechanics

CH#4: Orbit Representation Methods

- We need six quantities to define the translational state of a satellite in space. These quantities may take on many equivalent forms:
- Cartesian (inertial) coordinates

$$\mathbf{X} = \begin{bmatrix} \mathbf{r}^T & \mathbf{v}^T \end{bmatrix}^T = \begin{bmatrix} x & y & z & \dot{x} & \dot{y} & \dot{z} \end{bmatrix}^T$$

True classical element set

$$\mathbf{X} = \begin{bmatrix} a & e & i & \Omega & \omega & t \end{bmatrix}^T$$

Singular at e = 0 or 1, i = 0 or π

Modified classical element set

$$\mathbf{X} = \begin{bmatrix} a & e & i & \Omega & \omega & \theta \end{bmatrix}^T$$

Flight elements (ADBARV elements)

Spherical elements	Geographic elements	Flight elements
α	h	α
δ	λ	δ
$lpha_{_{\scriptscriptstyle \mathcal{V}}}$	$oldsymbol{\phi}_{gc}$	$oldsymbol{eta}$
$\delta_{_{v}}$	ν	\boldsymbol{A}
r	γ	r
ν	\boldsymbol{A}	v

Singular at poles

$$\sin(\alpha) = \frac{y}{\sqrt{x^2 + y^2}}, \quad \sin(\delta) = \frac{z}{r}, \quad r = \sqrt{x^2 + y^2 + z^2}, \quad v = \sqrt{\dot{x}^2 + \dot{y}^2 + \dot{z}^2}$$

$$[\mathbf{v}]^{SEZ} = T^{SEZ-GCRF} [\mathbf{v}]^{GCRF} = T^{SEZ-GCRF} \begin{bmatrix} \dot{x} \\ \dot{y} \\ \dot{z} \end{bmatrix}$$

$$T^{SEZ-GCRF} = R_2 (90 - \delta) R_3 (\alpha)$$

$$[\mathbf{v}]^{SEZ} = \begin{bmatrix} v_s \\ v_E \\ v_z \end{bmatrix} = v \begin{bmatrix} -\cos A \cos \gamma \\ \sin A \cos \gamma \\ \sin \gamma \end{bmatrix}$$

$$\sin \gamma = \frac{v_z}{v}, \qquad \sin A = \frac{v_E}{\sqrt{v_s^2 + v_E^2}}$$

$$\sin \phi_{gc} = \frac{r_z}{r}, \qquad \tan \lambda = \frac{r_y}{r_x}$$

$$\text{where } [\mathbf{r}]^{ITRF} = T^{ITRF-GCRF} [\mathbf{r}]^{GCRF} \rightarrow \begin{bmatrix} r_x \\ r_y \\ r \end{bmatrix} = T^{ITRF-GCRF} \begin{bmatrix} x \\ y \\ z \end{bmatrix}$$

Nonsingular (e < 1) Euler parameters $\mathbf{X} = \begin{bmatrix} a & e & q_1 & q_2 & q_3 & M \end{bmatrix}^T$ $T^{GCRF-PCS} = R_3(-\Omega)R_1(-i)R_3(-\omega)$ $|q_1^2 - q_2^2 - q_3^2 + q_4^2| 2(q_1q_2 - q_3q_4) 2(q_1q_3 + q_2q_4)$ $= \begin{vmatrix} 2(q_1q_2 + q_3q_4) & -q_1^2 + q_2^2 - q_3^2 + q_4^2 & 2(q_2q_3 - q_1q_4) \end{vmatrix}$ $\left| 2(q_1q_3 - q_2q_4) \quad 2(q_2q_3 + q_1q_4) \quad -q_1^2 - q_2^2 + q_3^2 + q_4^2 \right|$ where $\mathbf{q} = \begin{bmatrix} q_1 & q_2 & q_3 & q_4 \end{bmatrix}^T$, $\mathbf{q}^T \mathbf{q} = 1$ $q_4 = \frac{1}{2}\sqrt{1 + t_{11} + t_{22} + t_{33}} = \cos\frac{i}{2}\cos\frac{\Omega + \omega}{2}$ Undefined for $e \ge 1$. $q_1 = \frac{t_{32} - t_{23}}{4q_4} = \sin\frac{i}{2}\cos\frac{\Omega - \omega}{2}$ $q_2 = \frac{t_{13} - t_{31}}{4q_A} = \sin\frac{i}{2}\sin\frac{\Omega - \omega}{2}, \qquad q_3 = \frac{t_{21} - t_{12}}{4a_A} = \cos\frac{i}{2}\sin\frac{\Omega + \omega}{2}$

Quaternions

- We How to compute unit quaternion from a given DCM?
 - 1. Form the symmetric matrix

$$\mathbf{K} = \frac{1}{3} \begin{bmatrix} t_{11} - t_{22} - t_{33} & t_{21} + t_{12} & t_{31} + t_{13} & t_{23} - t_{32} \\ t_{21} + t_{12} & -t_{11} + t_{22} - t_{33} & t_{32} + t_{23} & t_{31} - t_{13} \\ t_{31} + t_{13} & t_{32} + t_{23} & -t_{11} - t_{22} + t_{33} & t_{12} - t_{21} \\ t_{23} - t_{32} & t_{31} - t_{13} & t_{12} - t_{21} & t_{11} + t_{22} + t_{33} \end{bmatrix}$$

2. Solve the eigenvalue problem $\mathbf{Ke} = \lambda \mathbf{e}$ for the largest eigenvalue (λ_{max}). The corresponding eigenvector is the quaternion.

Nonsingular (e < 1) quaternion element set: $\mathbf{X} = [\mathbf{q}^T \quad e_x \quad e_y]^T$

$$q_{1} = p^{\frac{1}{4}} \sin \frac{i}{2} \cos \frac{\Omega - \omega - M}{2}$$

$$q_{2} = p^{\frac{1}{4}} \sin \frac{i}{2} \sin \frac{\Omega - \omega - M}{2}$$

$$q_{3} = p^{\frac{1}{4}} \cos \frac{i}{2} \sin \frac{\Omega + \omega + M}{2}$$

$$q_{4} = p^{\frac{1}{4}} \cos \frac{i}{2} \cos \frac{\Omega + \omega + M}{2}$$

$$e_{x} = e \cos M$$

$$e_{y} = e \sin M$$

Equinoctial elements: $\mathbf{X} = \begin{bmatrix} a & h & k & p & q & \lambda \end{bmatrix}^T$

$$a = a$$

$$h = e \sin(\omega + f_r \Omega)$$

$$k = e\cos(\omega + f_r\Omega)$$

$$p = \tan^{f_r} (i/2) \sin \Omega = \frac{\sin i \sin \Omega}{1 + \cos^{f_r} i}$$

$$q = \tan^{f_r} (i/2) \cos \Omega = \frac{\sin i \cos \Omega}{1 + \cos^{f_r} i}$$

$$\lambda = M + \omega + \Omega$$
: mean longitude

 $f_r = \begin{cases} +1 & : \text{Prograde orbits} \\ -1 & : \text{Retrograde orbits} \end{cases}$

Undefined for $e \ge 1$.

Modified Equinoctial elements: $\mathbf{X} = [p \ f \ g \ h \ k \ L]^T$

$$p = a(1 - e^2)$$

$$h = \tan^{f_r} (i/2) \cos(\Omega)$$

Nonsingular for all eccentricities and inclinations

$$f = e\cos(\omega + f_r\Omega)$$

$$f = e\cos(\omega + f_r\Omega)$$
 $k = \tan^{f_r}(i/2)\sin(\Omega)$

$$g = e \sin(\omega + f_r \Omega)$$
 $L = \omega + f_r \Omega + \theta$

$$L = \omega + f_r \Omega + \theta$$

Classical OE from MEOE

$$e = \sqrt{f^2 + g^2}$$

$$a = \frac{p}{1 - e^2}$$
Assuming $f_r = 1$: $\tan(i/2) = \sqrt{h^2 + k^2}$

$$\tan \Omega = \frac{k}{h}$$

$$\omega = \tan^{-1}(g/f) - \Omega$$

$$\theta = L - (\Omega + \omega)$$

 \bigcirc Delaunay elements: $\mathbf{X} = [M \ \Omega \ \omega \ L \ h \ H]^T$

$$L = \sqrt{\mu a}$$
 Singular at $e = 0$ or 1 , $i = 0$ or π $H = \sqrt{\mu p} \cos i$

Solution Poincaré elements: $\mathbf{X} = [\Lambda \quad \xi \quad \eta \quad u \quad \upsilon \quad \lambda]^T$

$$\Lambda = \sqrt{\mu a}$$

$$\xi = e \sin(\Omega + \omega) \sqrt{2\Lambda/(1 + \sqrt{1 - e^2})}$$

$$\eta = e\cos(\Omega + \omega)\sqrt{2\Lambda/(1 + \sqrt{1 - e^2})}$$

$$u = \sin i \sin \Omega \sqrt{2\Lambda \sqrt{1 - e^2}} / (1 + \cos i)$$

$$\upsilon = \sin i \cos \Omega \sqrt{2\Lambda \sqrt{1 - e^2}} / (1 + \cos i)$$

$$\lambda = M + \Omega + \omega$$

Nonsingular for small eccentricities and inclinations

Undefined for $e \ge 1$.

Solution Alternate form of Poincaré elements: $\mathbf{X} = \begin{bmatrix} L & g_p & h_p & G & H & \lambda \end{bmatrix}^T$

$$L = \sqrt{\mu a}$$

$$g_p = \cos(\Omega + \omega)\sqrt{2L(1 - \sqrt{1 - e^2})}$$

$$h_p = \cos(\Omega)\sqrt{2L\sqrt{1 - e^2}(1 - \cos i)}$$

$$G = \sin(\Omega + \omega)\sqrt{2L(1 - \sqrt{1 - e^2})}$$

$$H = \sin(\Omega)\sqrt{2L\sqrt{1 - e^2}(1 - \cos i)}$$

$$\lambda = M + \Omega + \omega$$

Nonsingular for small eccentricities and inclinations

Undefined for $e \ge 1$.

- Two-line Element Set (TLE) is a data format encoding a list of orbital elements of an Earth-orbiting object for a given point in time.
- A TLE set consists of two 69-character lines of data which can be used together with SGP4/SDP4 orbital model to determine the position and velocity of the associated satellite.

ISS (ZARYA)

1 25544U 98067A 08264.51782528 -.00002182 00000-0 -11606-4 0 2927 2 25544 51.6416 247.4627 0006703 130.5360 325.0288 15.72125391563537

Line 0 is a twenty-four character name.

Uline 1

1 25544U 98067A 08264.51782528 -.00002182 00000-0 -11606-4 0 2927

Field	Columns	Content	Example
1	01-01	Line number of Element Data	1
2	03-07	Satellite number	25544
3	08-08	Classification (U=Unclassified)	U
4	10-11	International Designator (Last two digits of launch year)	98
5	12–14	International Designator (Launch number of the year)	067
6	15–17	International Designator (piece of the launch)	A
7	19–20	Epoch Year (last two digits of year)	08
8	21-32	Epoch (day of the year and fractional portion of the day)	264.51782528
9	34-43	First Time Derivative of the Mean Motion divided by two $(\dot{n}/2)$	00002182
10	45–52	Second Time Derivative of Mean Motion divided by six (decimal point assumed) $(\ddot{n}/6)$	00000-0
11	54-61	BSTAR drag term (decimal point assumed)	-11606-4
12	63-63	The number 0 (originally this should have been "Ephemeris type")	0
13	65–68	Element set number. Incremented when a new TLE is generated for this object	292
14	69–69	Checksum (modulo 10) (Letters, blanks, periods, plus signs = 0; minus signs = 1)	7

$$B^* = \frac{1}{2} \frac{c_D A}{m} \rho_o R_{\oplus} \longrightarrow BC = \frac{1}{12.741621B^*} \frac{\text{kg}}{\text{m}^2}$$

Une 2

2 25544 51.6416 247.4627 0006703 130.5360 325.0288 15.72125391563537

Field	Columns	Content	Example
1	01–01	Line number of Element Data	2
2	03–07	Satellite number	25544
3	09–16	Inclination (degrees)	51.6416
4	18–25	Right ascension of the ascending node (degrees)	247.4627
5	27–33	Eccentricity (decimal point assumed)	0006703
6	35–42	Argument of perigee (degrees)	130.5360
7	44–51	Mean Anomaly (degrees)	325.0288
8	53–63	Mean Motion (revolutions per day)	15.72125391
9	64–68	Revolution number at epoch (revolutions)	56353
10	69–69	Checksum (modulo 10)	7