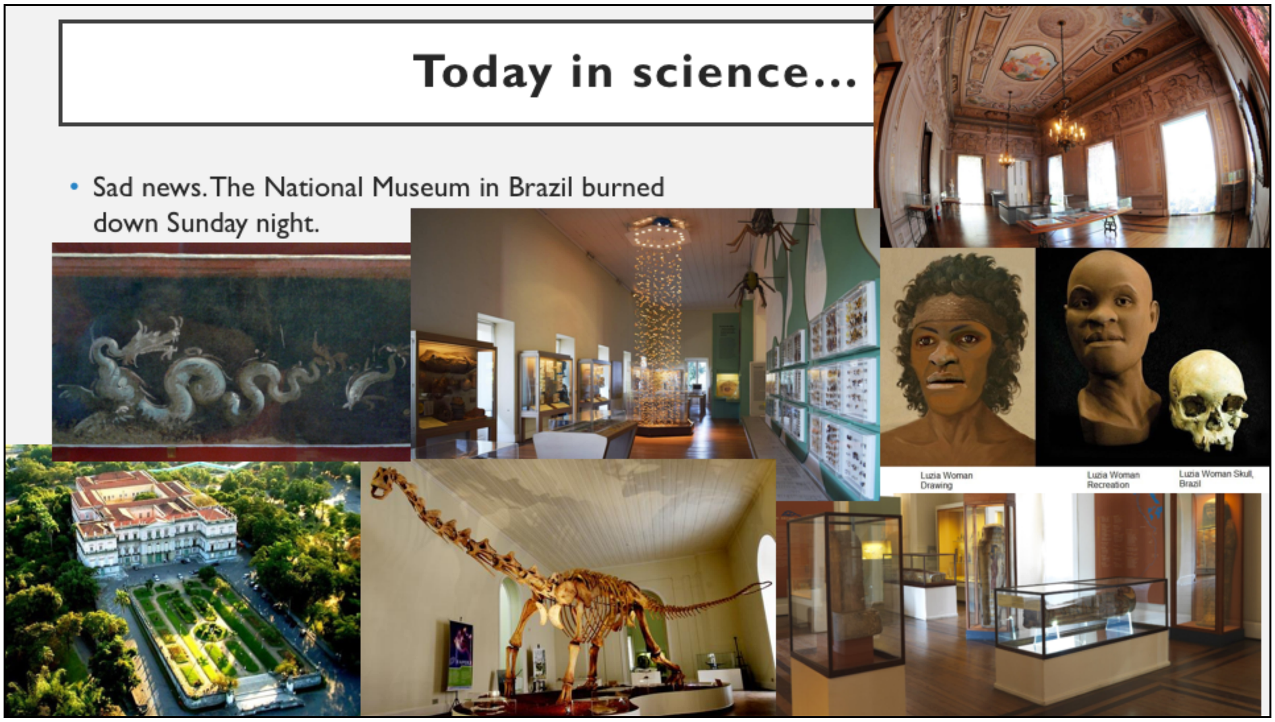


Incalculable loss.

They had just recently secured funding to install fire prevention systems, after years of their budget being reduced again and again.

Brief syllabus recap: on homeworks & labs

- The method to the assignments

AST 235. Astronomy: Stars, Galaxies, and Cosmology - Fall 2018

Tentative Lecture Schedule, week by week (plus helpful dates to remember)

Date	Chapter	Topic	Homework
Aug 14		Introduction to the Course	Tutorial assigned*
Aug 16	1.1-1.4	Modern View of the Universe	Ch 1 assigned
Aug 21	2.1-2.4	What We See in the Sky	Tutorial due**
Aug 23	2.1-2.4	Half class in the UNCG Planetarium: Earth's Orbit and its Effects	Ch 1 due** Ch 2 assigned, Lab 1 assigned
Aug 28	3.1-3.3	History of Astronomy, The Scientific Method	
Aug 30	3.4, 4.1-4.2	Basic Laws of Physics, I.	Lab 1 due, Ch 2 due 9/1 Ch 3 assigned
Sep 4	4.3-4.4	Basic Laws of Physics, II.	Ch 4 assigned
Sep 6	5.1-5.2	Properties of Light	Ch 3 due Lab 2 assigned
Sep 11	5.3-5.4	Properties of Matter	Ch 4 due Ch 5 assigned

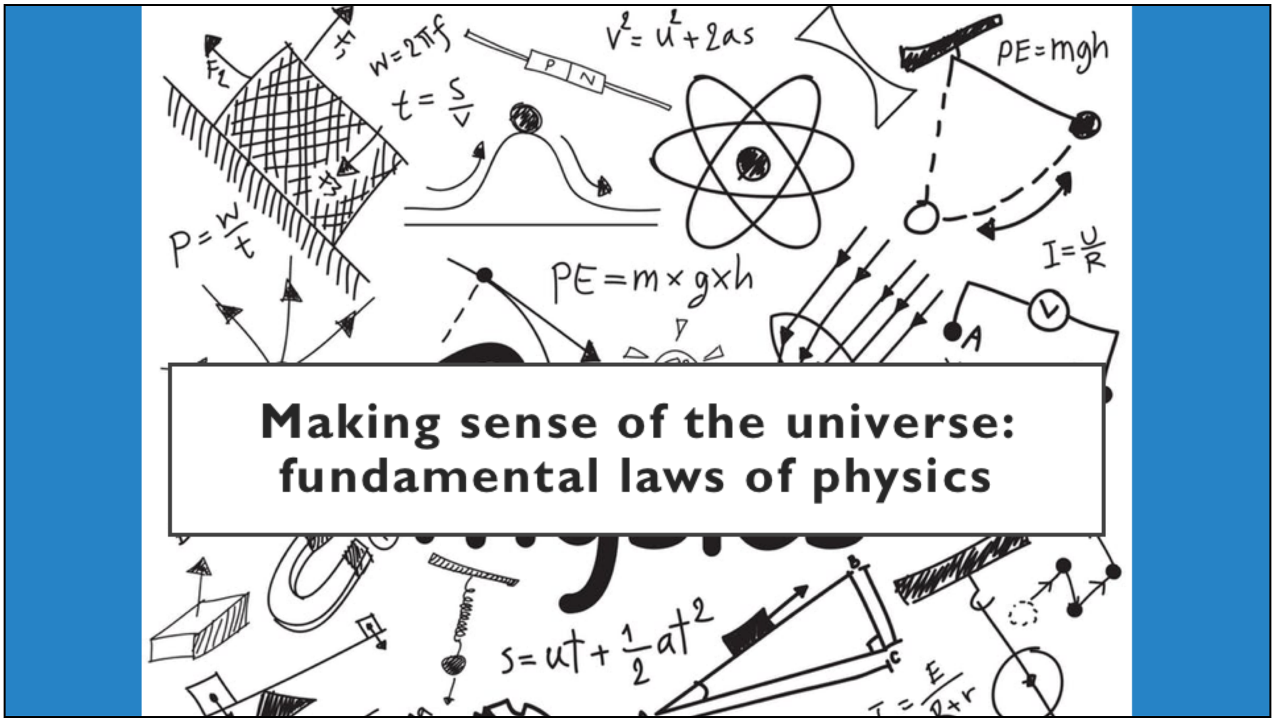
Brief syllabus recap: on homeworks & labs

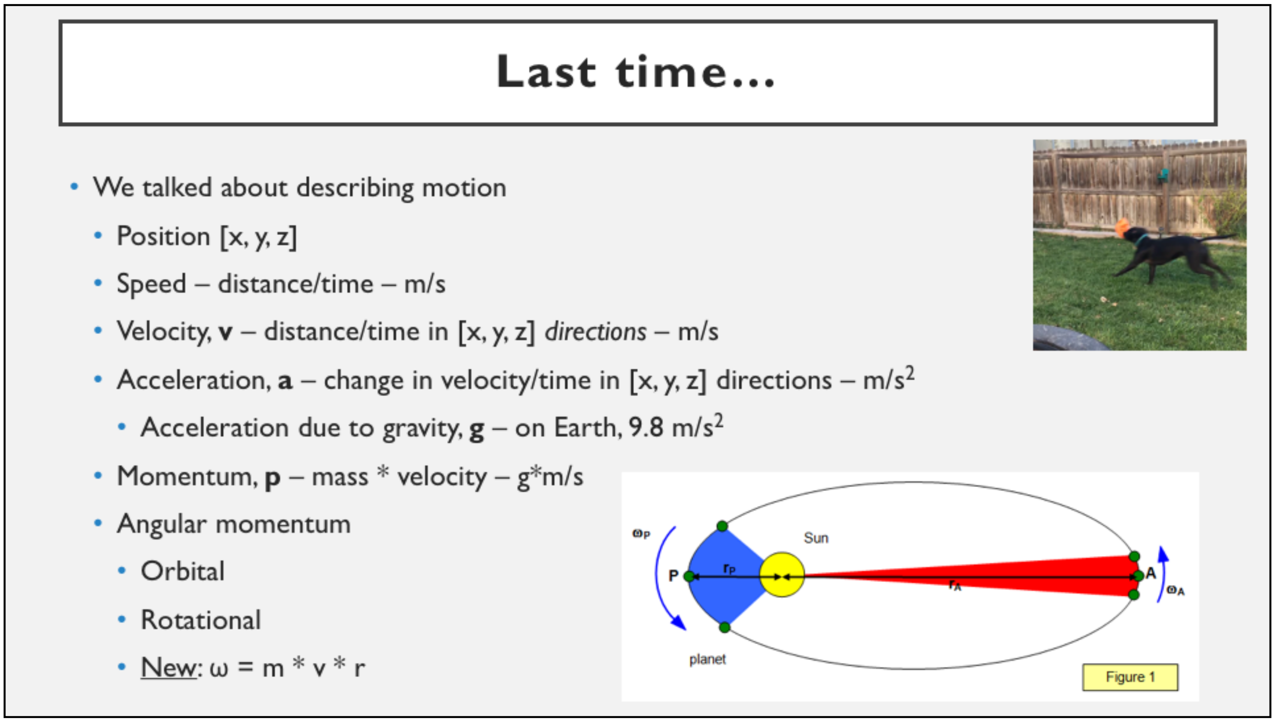
- Deep breaths! Lots of extra credit options. And,
- I've added midterm exam review the days before exams.

Course Grading

Activity	% of final grade	Extra credit	% possible
Online homework	20	Extra online homework	10
In-class quizzes	15	Visit to Three College Observatory	5
Online labs	10	*Total*	*15*
UNCG planetarium visit	5	Bonus project, can replace one homework or lab	(ask Prof. A. for details if interested)
Midterm exam 1	15		
Midterm exam 2	15		
Final exam	20		

Sep 18	14.1-14.2	The Sun: Internal Structure and Energy Source	Ch 5 due
Sep 20		Midterm Review, Ch. 1-6, 14.1-14.2	Ch 6 due, Lab 3 due
Sep 25	1-6, 14.1, 14.2	Mid-Term Exam 1	


Making sense of the universe:
fundamental laws of physics
W=2πf
v²=u²+2as
PE=mgh
t=s/v
P=W/t
PE=m×g×h
I=U/R
s=ut+½at²



ω is the greek letter omega (lower case). It’s used to denote angular momentum as distinct from linear momentum, p.

Last time…

- Force: acts to change an object's momentum
  - Net force: total sum of all forces working on an object
- Torque changes an object's *angular* momentum
- Weight vs Mass: a force vs an inherent property
- Free-fall and weightlessness: absence of force acting against downward weight

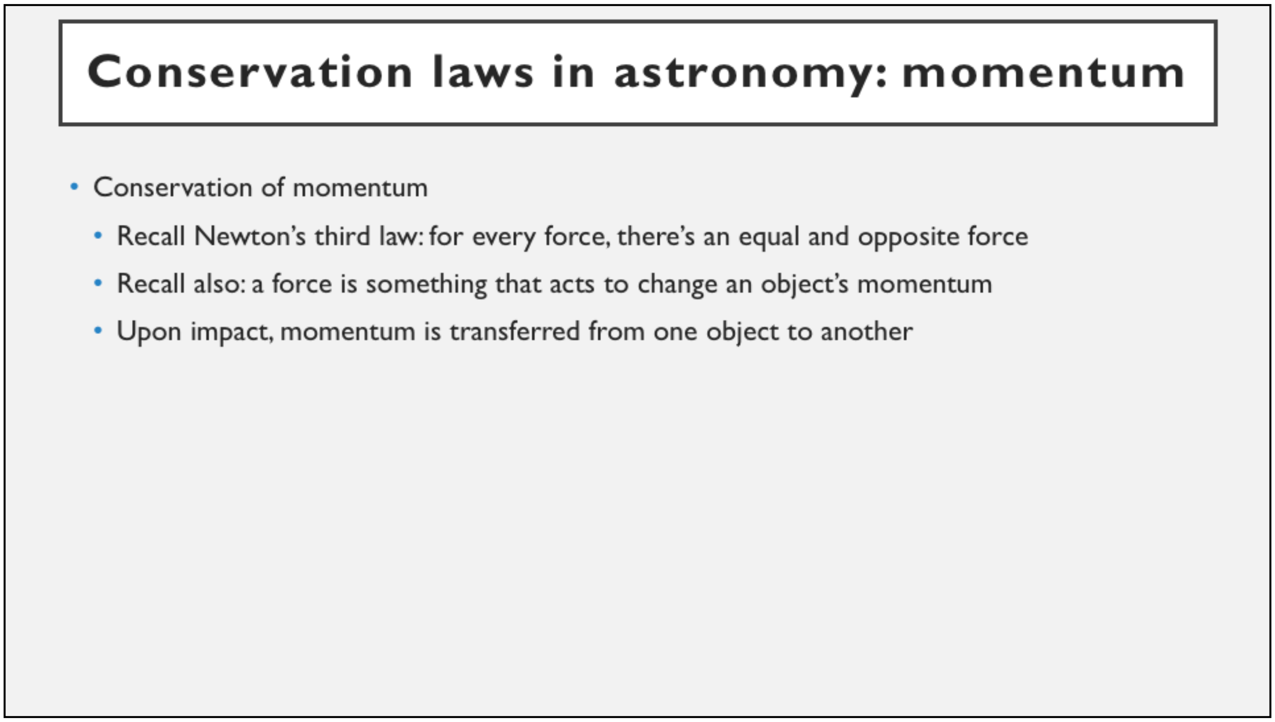
Newton's laws

1. An object at rest will remain at rest unless acted upon by a force
2. F=ma
3. For any force, there is an equal and opposite force

http://www.cakeburger.com/comic/cat-physics/

Evolution of ideas

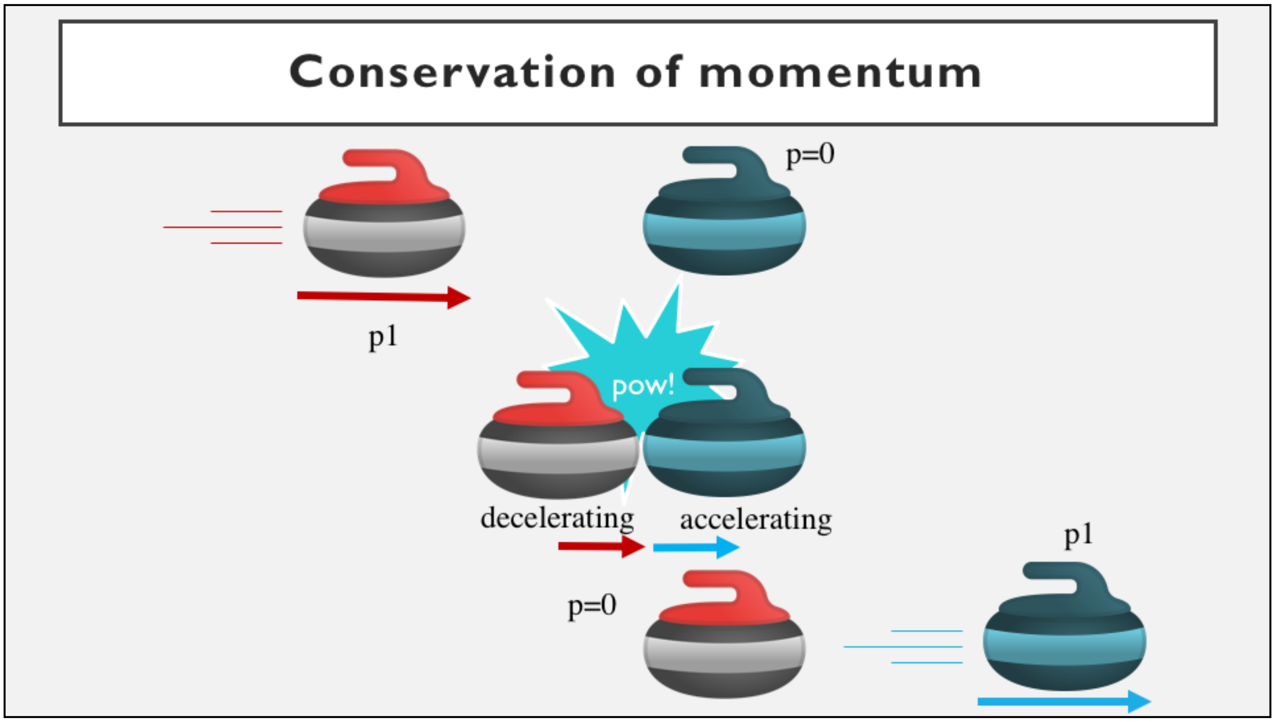
- Early observers collected the data later mathematicians and astrophysicists worked with later
- Kepler came up with the framework of how:
  - How do the planets move in the sky?
- Newton came up with the *why*
  - Newton's laws, in combination with what we observe, give rise to what we call *conservation laws*
  - When a quantity is conserved, it means that over an entire system, the **net** amount of it stays the same



Something we observe here on Earth: conservation of linear momentum; that is to say, momentum acting in a straight path

Conservation of momentum

The book gives you an example of playing pool- curling is another good example, when the stones impact each other



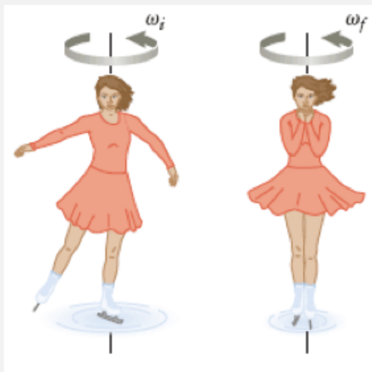
Let’s break down one of those impacts…

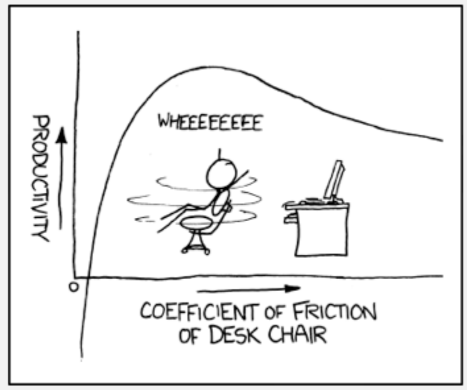
Conservation of momentum

The book gives you an example of playing pool- curling is another good example, when the stones impact each other

Conservation laws in astronomy: angular momentum

- Conservation of angular momentum (recall: $\omega = m * v * r$)
  - Rotational



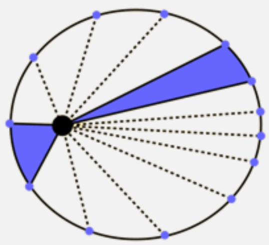


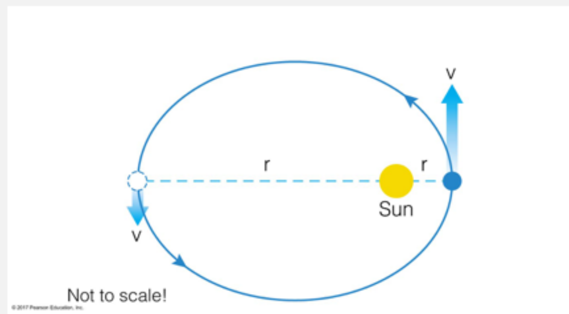
https://xkcd.com/815/

As figure skater's radius, r, decreases, to conserve angular momentum, her rotational velocity must increase

Conservation laws in astronomy: angular momentum

- Conservation of angular momentum (recall: ω = m * v * r)
  - Rotational
  - Orbital
- Newton's 1st law: if no force is acting on the orbiting planet, it will keep on orbiting
- Kepler saw orbital speeds changing, deduced elliptical orbits
  - ➢its angular momentum is conserved





We see this in the observations Kepler's laws were built from

Kepler's second law: the path of the planet's orbit sweeps out equal areas for equal time

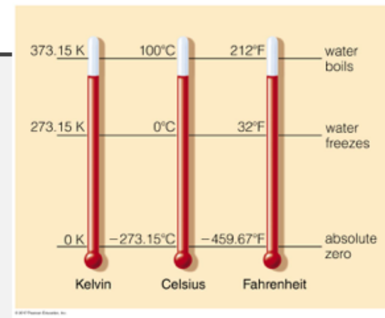
Energy

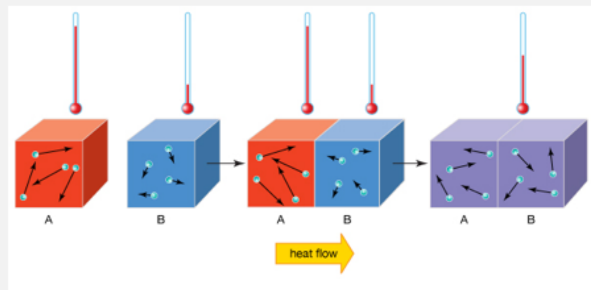
- Energy is a property that enables something to do work
- When a force acts to move an object a certain distance, it has done work
  - Work = force * distance
- For astronomy, important types of energy:
  - Kinetic
  - Radiative
  - Potential
- Affiliated conservation law: energy is not created or destroyed

We'll talk a lot about radiative energy in chapter 5

Kinetic energy

- The energy of a moving object: ½ mv^2
- Important form of kinetic energy for us: thermal energy
  - Kinetic energy of atoms or molecules in a substance
- Temperature: measures the *average* kinetic energy
- Temperature tells us which way energy will move when two systems come into contact
- Rate of energy flow depends on density
- Absolute zero (Kelvin): thermal energy is 0

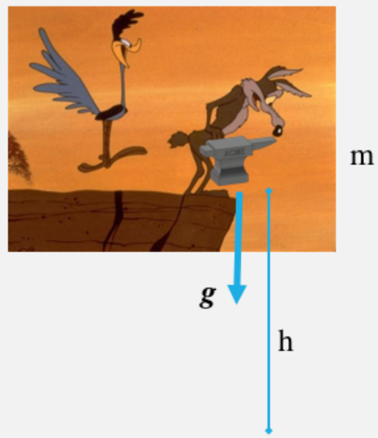




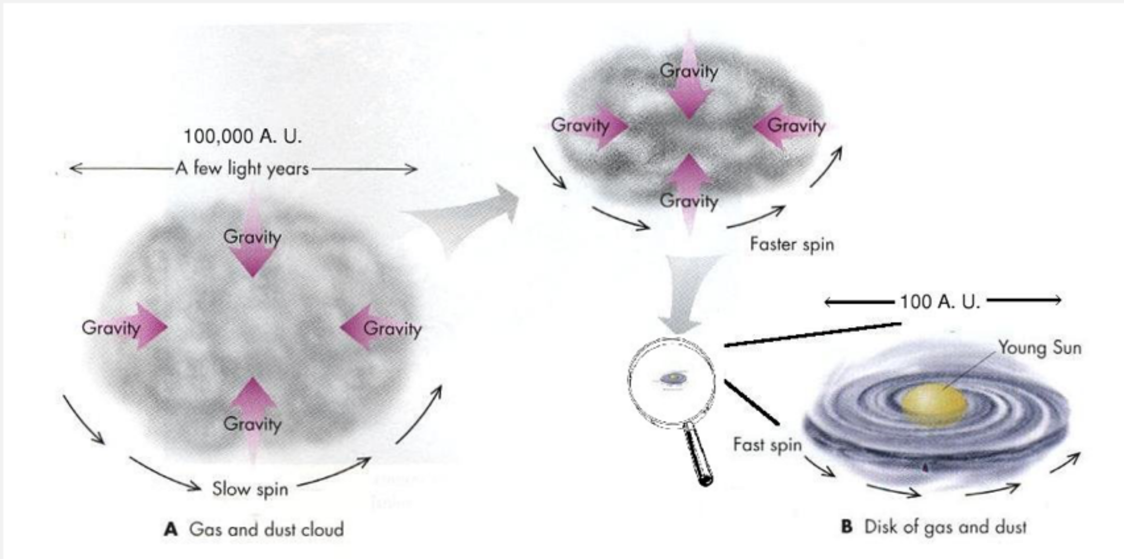
The book gives a good example of a boiling pot of water vs an oven. The oven may be hotter, meaning the particles have higher average kinetic energy, but there are fewer of them than say, in a pot of boiling water.

Potential energy

- Potential energy is the capacity of something to do work
  - Example: if Wile E. Coyote is holding an anvil over a cliff, about to drop it, its potential energy depends on
    - The mass of the anvil
    - The acceleration due to gravity
    - The height of the cliff
      - PE = m*g*h



Gravitational potential energy



Gravitational potential energy decreases the smaller the protostellar cloud gets

Thermal energy increases because density is increasing; the star forms in the middle, getting heated up from conversion of potential to kinetic energy

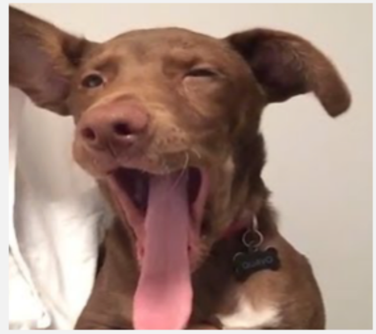
Due to conservation of angular momentum, we think material speeds up as it collapses

http://faculty.virginia.edu/skrutskie/ASTR1210/notes/ssrecip.html

Conservation of energy

- Energy is not created or destroyed, it simply changes form

- Example: you raise your hand to stifle a yawn
  - Your muscles exerted energy to perform the mechanical action of raising your hand
  - The energy your muscles used came from what you ate for breakfast
  - What you ate for breakfast got its energy from sunlight, one way or another
  - The sun is radiating energy in the form of light that came from fusing H atoms at its core, converting mass into energy
  - H atoms came from the big bang, which was a conversion of energy into mass



Universal law of gravitation

- Newton, not content to stop at laws of motion, calculus, optics, etc… carried on to develop a framework for understanding gravity

1. Every mass attracts every other mass through a force called gravity
2. the strength of the force of gravity is *directly proportional* to the product of those two masses
3. the strength of the force of gravity decreases as the square of the distance between the two masses' centers

$F_g = G*M_1*M_2 / d^2$

Note the big G- this is a universal constant, not to be confused with the little g for acceleration due to gravity on earth

Universal law of gravitation

- $F_g = G*M_1*M_2 / d^2$
  - Say $M_1 = M_2$; $F_g = GM_2^2/d^2$

- If $M_1 = M_2$, but d is now doubled…
  - $F_g = GM_2^2/(2d)^2 = GM_2^2/4d^2$

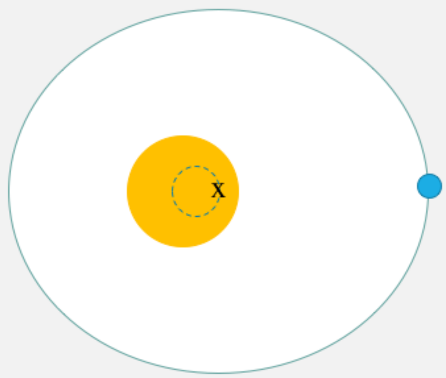
- Make $M_1 = 0.5\ M_2$; $F_g = GM_2^2/2d^2$

- Strength of the force of gravity is
  - Directly proportional to the product of the masses- they're in the numerator, and ^1 power
  - Inversely proportional to the square of the separation of the objects (d is in the denominator, and ^2)

Examples of what would happen to the force of gravity between two objects if their masses or separation changed

Kepler's laws and Newton's laws

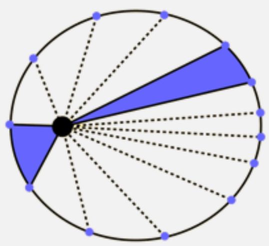
- Kepler found the how, Newton found the why
  1. Planets travel in elliptical orbits around the Sun
     - This is a natural consequence of Newton's law of gravitation: both bodies are orbiting the center of mass of the Sun+planet system

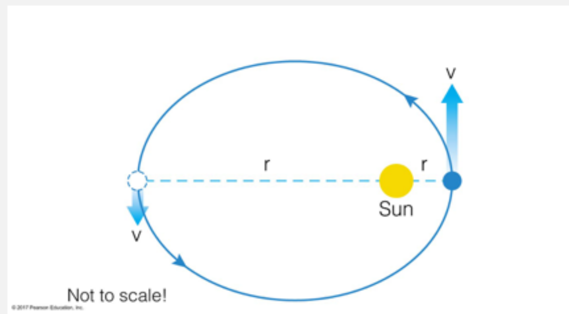


Exaggerated for effect…

Kepler's laws and Newton's laws

- Kepler found the how, Newton found the why
  1. Planets travel in elliptical orbits around the Sun
  2. Planets cover equal areas in equal time along their orbits
     - Conservation of angular momentum



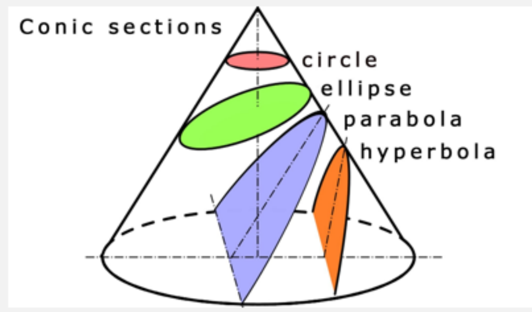


Kepler's laws and Newton's laws

- Kepler found the how, Newton found the why
  1. Planets travel in elliptical orbits around the Sun
  2. Planets cover equal areas in equal time along their orbits
  3. Planets' orbital periods and semimajor axis radii are related as $p^2 = a^3$
     - Planets experience less force due to gravity farther away from the Sun, and they move more slowly ($F_g \sim 1/d^2$)

Beyond Kepler's laws

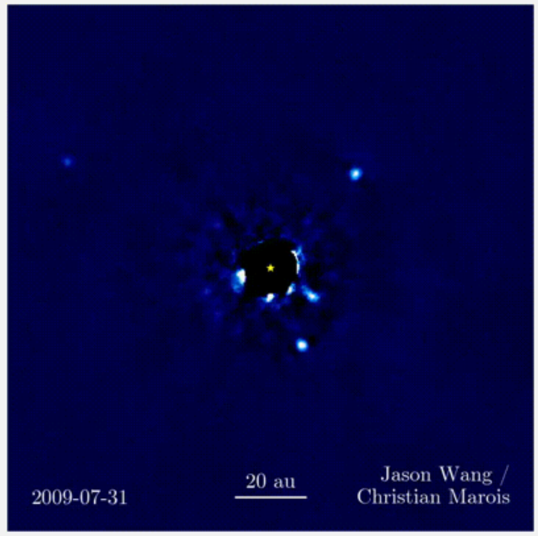
- Planets aren't the only bodies that travel in elliptical orbits
  - So do moons
  - And asteroids
- Ellipses aren't the only kinds of orbits
  - Unbound parabolic orbits
  - Unbound hyperbolic orbits
- Knowing orbital parameters, we can now calculate masses of orbiting bodies using the universal law of gravitation



This is the cone Neil deGrasse Tyson was talking about in the video last week

Beyond Kepler’s laws

- Newton applied law of gravitation <u>universally</u>



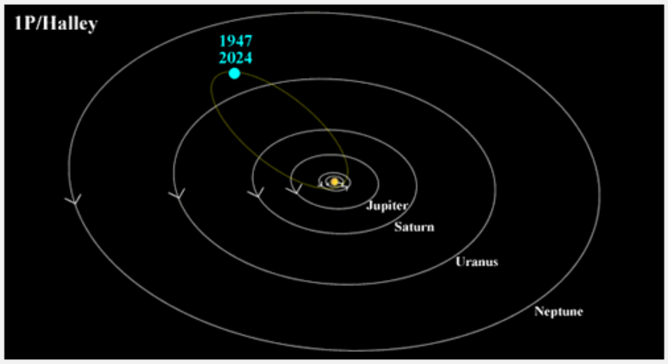
https://www.forbes.com/sites/startswithabang/2017/01/30/watch-four-gas-giants-in-orbit-around-another-star-for-the-first-time/

Orbital energy & gravitational encounters

- A body's total orbital energy is the sum of its kinetic energy (energy of motion, related to speed) and gravitational potential energy (dependent upon distance)
- These components vary throughout the planet's orbit, but the total energy is conserved

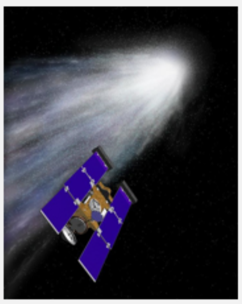
- A close pass between a planet and a body (a comet, a satellite, another planet) can exchange orbital energy
- The change in energy is almost negligible for the larger body

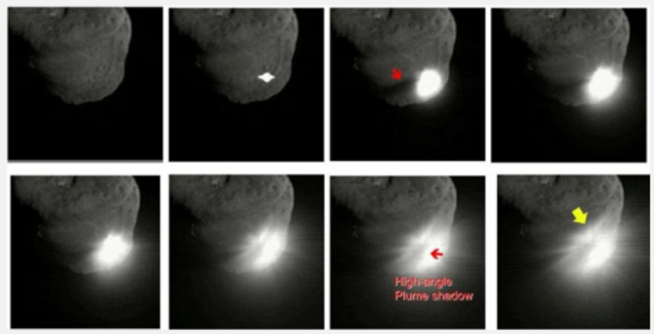
Orbital energy & gravitational encounters



Comet shenanigans

- Stardust mission
  - Flew into coma of comet Wild 2 to collect material comet was outgassing
- Deep Impact
  - Dropped a refrigerator-sized impactor onto comet Tempel 1 to see what would come out





https://arstechnica.com/science/2013/09/deep-impact-probe-down-and-out-ice-back-for-maybe-more/

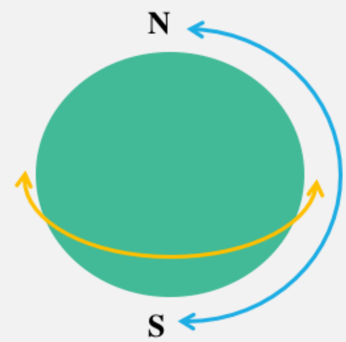
Atmospheric drag & escape velocity

- Back to orbital energy considerations…
- Orbital energy will be conserved, unless a force acts to change it
  - Atmospheric drag can reduce orbital energy, causing satellites' trajectories to change
    - This could be bad: satellites we launch need to overcome/get far enough away to be safe
    - This could be good/interesting: planets in the outer solar system may have captured a lot of moons this way

- As you increase its orbital energy, body will get farther and farther away
  - At some velocity, body can entirely leave its bound orbit: *escape* velocity

Tidal effects

- The Earth's rotation makes it an oblate spheroid shape
  - Centrifugal forces push outward, most at the equator
    - From N->S poles, Earth's diameter is 12,713 km
    - Around Equator, Earth's diameter is 12,756 km!
- Gravitational forces between the Earth and the Moon also add to Earth's bulge (~1m in water, ~30cm in rock) as do gravitational forces between the Earth and the Sun (about half as much as the Moon's effect)



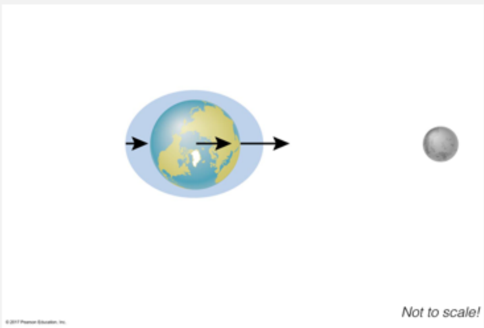
http://www.slate.com/blogs/bad_astronomy/2008/09/08/ten_things_you_dont_know_about_the_earth.html

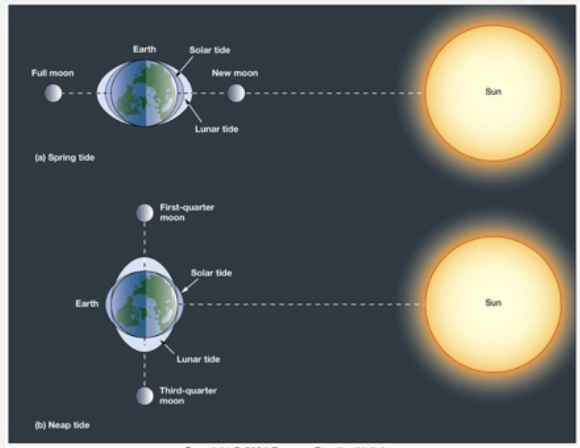
Also

http://www.slate.com/blogs/bad_astronomy/2006/08/30/when_i_say_centrifugal_i_mean_centrifugal.html

Tidal effects

- Amount of bulge depends on distance (it's a gravitational effect!)

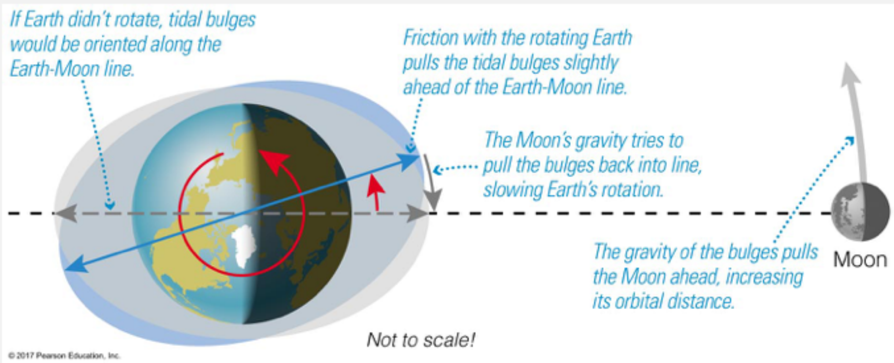




- Sun and Moon can work together or against each other, depending on where Earth, and Moon are in orbits

Tidal friction

- Gravitational force between Earth and Moon has raised a bulge on Earth
- Earth is rotating, so a tug-of-war begins as bulge starts to outpace the Moon



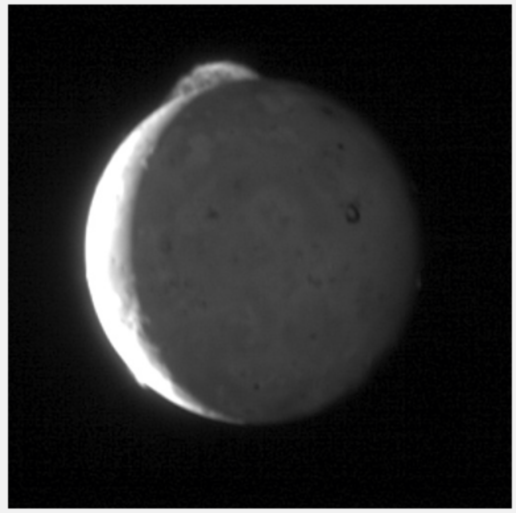
The moon's synchronous rotation

- As the Moon is creating tidal effects on Earth, so the Earth has also done to the Moon
  - Moon gains orbital angular momentum from the Earth as it slows earth down
  - Moon's distance from Earth starts to grow (conservation of angular momentum)
- The Moon used to be a lot closer to us than it is now
- It also likely rotated faster
- 2014: Scientists were able to measure a "lunar body tide," a 20 inch bulge in the Moon's surface due to Earth's gravitational pull

https://www.space.com/26246-lunar-tide-seen-from-space.html

Tidal effects on other worlds

- Synchronous rotation of moons is common
- Tidal bulging also common on other moons
  - the more massive the planet, the stronger the force
  - Jupiter's moon Io is volcanically active from being squeezed by Jupiter, Ganymede, and Europa (numerous tidal bulges!)



Video from New Horizons spacecraft while flying by

Why do all objects fall at the same rate?

- We can understand this using two laws Newton discovered.
- $F_g = \frac{G\, M_{Earth}\, M_{ball}}{R_{Earth}^2}$
- $F = m_{ball} a$

- Set them equal to each other…

$$\frac{G\, M_{Earth}\, \cancel{M_{ball}}}{R_{Earth}^2} = \cancel{m_{ball}}\, a$$

$$\ldots = g$$