

X WSS, H LTI \leftarrow Even fun
then $S_x = \mathcal{F}\{R_x\} \longleftrightarrow$ non neg. real

$$X \rightarrow [H] \rightarrow Y$$

then X, Y jointly WSS

$$S_y = |H|^2 S_x, M_y = H(0) M_X$$

Def. X, Y jointly WSS, we define

$$S_{XY}(w) = \mathcal{F}\{R_{XY}\}$$

"Cross-Spectral Density"

Gaussian Processes

Def random variables X_1, \dots, X_n are jointly Gaussian if every linear comb. of X_i 's is Gaussian

⊕ Special Case: X_1, \dots, X_n independent, identical Gaussian

Def a random process $X(t)$ is a Gaussian process if $\forall n \in \mathbb{N}, \forall (t_1, \dots, t_n) \in \mathbb{R}^n$

$\{X(t_1), X(t_2), \dots, X(t_n)\}$ are jointly Gaussian.

⊕ Special Case indep. gaussian at all $t \in \mathbb{R}$
identical

Def Jointly Gaussian Processes $X(t), Y(t)$ j.g. if $\forall n, m \in \mathbb{N}$
 $\forall (t_1, \dots, t_n, \tau_1, \dots, \tau_m) \in \mathbb{R}^{n+m}, \{X(t_1), \dots, Y(\tau_m)\}$ j.g.
as r.v.s

$\forall (t_1, \dots, t_n, \tau_1, \dots, \tau_m) \in \mathbb{R}^{n+m}$, $\{X(t_i), \dots, Y(\tau_m)\}$ i.g.
as r.v.s

1) If $X(t)$ Gaussian, H LTI

$$X \xrightarrow{H} Y$$

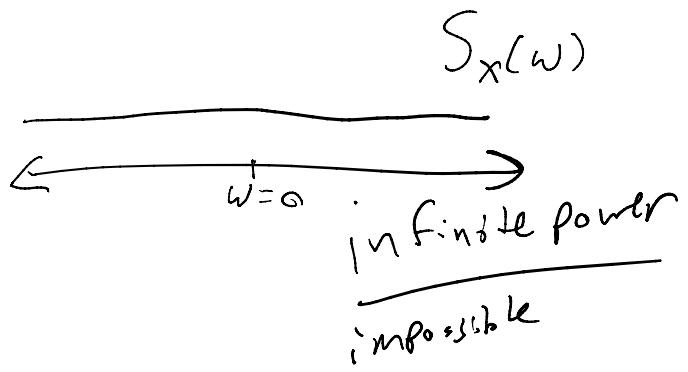
$X(t), Y(t)$ are jointly gaussian

2) For j.g. processes

Uncorrelated = independent

Def. A process is white if it has flat PSD

i.e. $S_x(\omega)$ is constant freq



$$\mathcal{P}_x = \infty$$

usually, when we say "white" in practice, we mean
flat PSD over the band of interest

if $S_x(\omega) = C \forall \omega$, then

$$\mathcal{F}^{-1}\{C\} = R_x(\tau) = C \delta(\tau)$$

qntcorr. of a white process is S , i.e. no two time indices are correlated with each other

Ex. $X(t)$ gaussian process, uncorr=indp.

So; indp. identically gaussian at each $t \Rightarrow$ white

{most noise in this class is modeled
" " " " " }

↑ noise in this class is modeled
this way

A W G N

d h a o
z i u i
i s s
+ s e
v :
e a
r n

Signal $m(t)$

model for AWGN is

$$r(t) = m(t) + n(t)$$

↑ r.p. white
gaussian

usually means indep. identical gaussian r.v.

Independent $N(0, \sigma^2)$ at all time

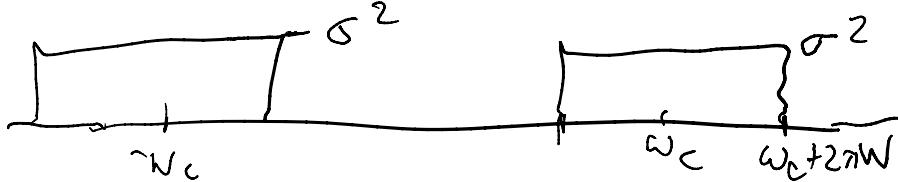
$$R(\tau) = 0 \vee \tau \neq 0$$

$$R(0) = E[X^2] = \sigma^2$$

$$\mathcal{F}\{R(\tau)\} = \mathcal{F}\{\sigma^2 \delta(t-t_0)\} = \boxed{\sigma^2 = S_X(\omega)}$$

We will have a receiver which will filter all incoming signals to BW of interest

So filtered noise looks like

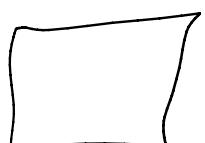


PSD of the
filtered
noise

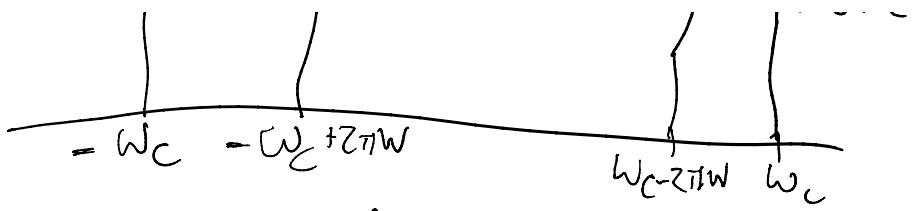
by convention we say

$$\sigma^2 = N_0/2 \quad \text{← noise power}$$

S_N



$\boxed{N_0/2}$



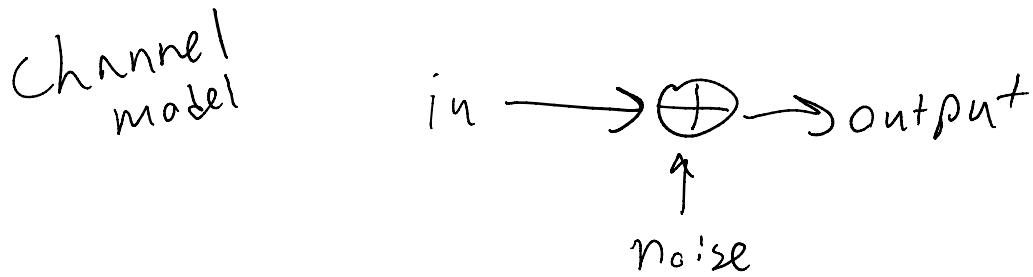
$$P_N = \frac{1}{2\pi} \int S_N(\omega) d\omega = \frac{N_0}{2} \frac{(2\pi W)^2}{2\pi} = N_0 W$$

power in SSB noise is $N_0 W = 2\sigma^2 W$

power in DSB noise is $2N_0 W = 4\sigma^2 W$

Noise on Baseband Signals

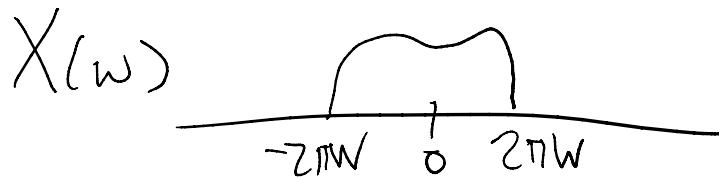
Model noise as white $S_N(\omega) = N_0/2$



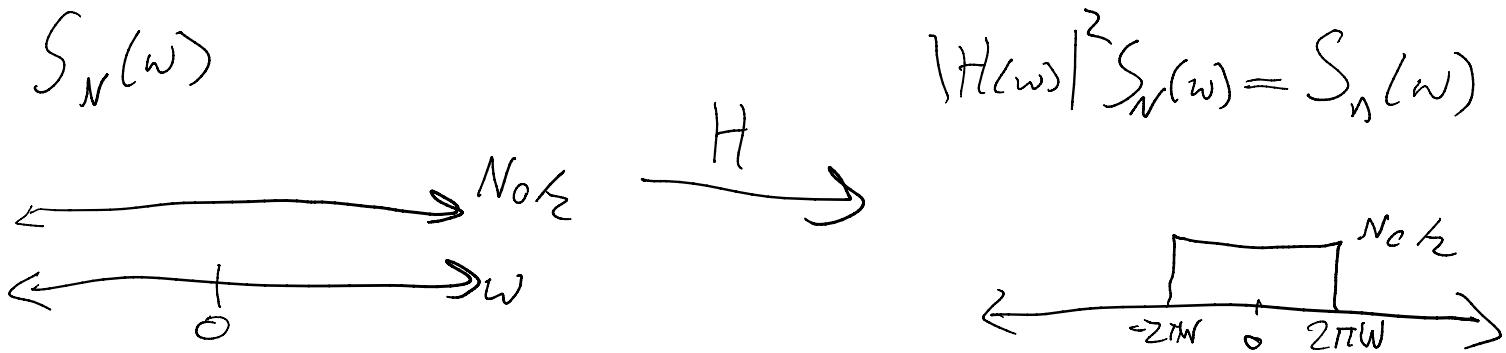
AWN model



Send bb signal



at the receiver, we apply BPF $H(\omega) = \begin{cases} 1, |\omega| \leq 2\pi W \\ 0, \text{else} \end{cases}$



$X(\omega)$ $\xrightarrow{H} X(\omega)$ b/c $X(\omega) = 0$ for $(|\omega| \geq 2\pi W)$

$$X(w) \xrightarrow{\quad} X(w) \text{ b/c } \begin{array}{l} X(w)=0 \text{ TO} \\ (|w| \geq 2\pi W) \end{array}$$

$P_N = \infty$ white noise has infinite power

$$P_n = \frac{1}{2\pi} \int_{-2\pi W}^{2\pi W} \frac{N_0}{2} dw = \frac{4\pi W}{2\pi} \frac{N_0}{2} = N_0 W \leftarrow \begin{array}{l} \text{power is} \\ \text{baseband} \\ \text{f. filtered noise} \end{array}$$

$\left(\text{if AWGN, } N_0 = 2^{-2W} \right)$

Obvious but important: additive $\rightarrow P_n$ is indep. of signal power

... Sort of...

P_n is a func of W , so $W \nearrow$ (\therefore is a prop. of the signal)
 $P_n \nearrow$

first example of the power-bw relationship
 (tradeoff)

Signal-to-Noise ratio (SNR) is the ratio
 of received signal power over received noise power
 (usually reported in dB)

in our case $SNR_{bb} = \frac{P_r}{N_0 W}$ \leftarrow power of received message signal

in our case

$$SNR_{bb} = \frac{P_r}{N_c W}$$

(*) $SNR_{dB} = 10 \log_{10} \left(\frac{P_r}{P_N} \right)$

10 b/c it's power

Ex. AWGN, variance 5×10^{-12}

affects bb signal w/ bw 10 kHz

I transmit w/ 100 kW of power

but the channel attenuates by a factor of 10^{-10}

$$P_R = (100 \text{ kW}) (10^{-10}) = 10^{-5} \text{ W}$$

$$P_N = N_0 W = (10 \text{ kHz}) (\sigma^2) = (10^4 \text{ Hz}) (10^{-11}) \\ = 10^{-7} \text{ W}$$

(linear scale)

$$\rightarrow SNR = \frac{P_R}{P_N} = \frac{10^{-5}}{10^{-7}} = 100$$

log scale $\rightarrow SNR_{dB} = 10 \log_{10}(100) = 20 \text{ dB}$

Noise in AM

Tuesday, September 22, 2020 6:56 PM

DSB-SC

We transmit $u(t) = m(t) \cos \omega_c t$

$|U(\omega)|$



Model: received = signal + AWGN

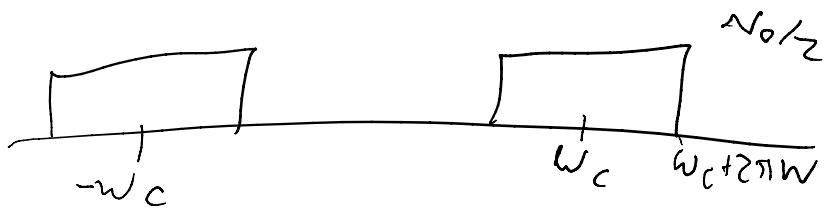
$$r(t) = u(t) + n(t)$$

\curvearrowleft filtered white noise

$$\omega_c - 2\pi W \leq |\omega| \leq \omega_c + 2\pi W$$

Sidebands - talk about filtered white noise that is at bb

$S_N(\omega)$



just like w/ deterministic signals, we can write a bb rep. of this noise:

$$y(t) = I_n(t) \cos \omega_c t - Q_n(t) \sin \omega_c t$$

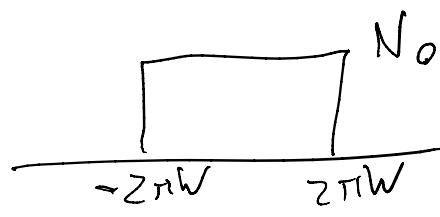
$\curvearrowleft \quad \curvearrowright$
random IQ processes

If it can be shown that if X is white, Gaussian

1) I_n, Q_n are zero-mean, baseband, jointly WSS
and jointly gaussian

$$2) P_Y = P_{I_n} = P_{Q_n} = \frac{1}{2\pi} \int S_Y(\omega) d\omega$$

3) I_n and Q_n have same PSD

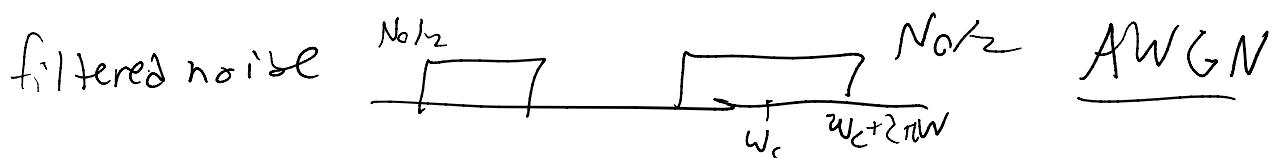


$$\begin{aligned} S_o P_{an} &= P_{In} \\ &= P_Y = 2N_0 W \end{aligned}$$

4) If ω_c is an axis of symmetry (i.e. in DSB)
then I_n, Q_n are indep.



$$u(t) = A_c m(t) \cos \omega_c t$$



$$\text{Can write } r(t) = u(t) + h(t)$$

$$= u(t) + n_I(t) \cos \omega_c t + n_Q(t) \sin \omega_c t$$

Then we demodulate $r(t)$ by product w/ $\cos(\omega_c t + \varphi)$

Then baseband filter $\cos a \cos b = \frac{1}{2} \cos(a+b) + \frac{1}{2} \cos(a-b)$

$$\cos a \sin b = \frac{1}{2} (\sin(a+b) - \sin(a-b))$$

1/11 . . . 1 . . . 1 . . . 1 - 1 . . . 1 . . . 1

$$\cos a \sin b = \frac{1}{2} (\sin(a+b) - \sin(a-b))$$

$$r(t) \cos \omega_c t = A_c m(t) \cos \omega_c t \cos(\omega_c t + \varphi) + n_I(t) \cos \omega_c t \cos(\omega_c t + \varphi) \\ - n_Q(t) \sin \omega_c t \cos(\omega_c t + \varphi)$$

↓ BOF filter plus simplify

$$y(t) = \underbrace{\frac{1}{2} A_c m(t) \cos \varphi}_{\text{What we had w/ no noise}} + \underbrace{\frac{1}{2} (n_I(t) \cos \varphi + n_Q(t) \sin \varphi)}_{+ \text{bb noise}}$$

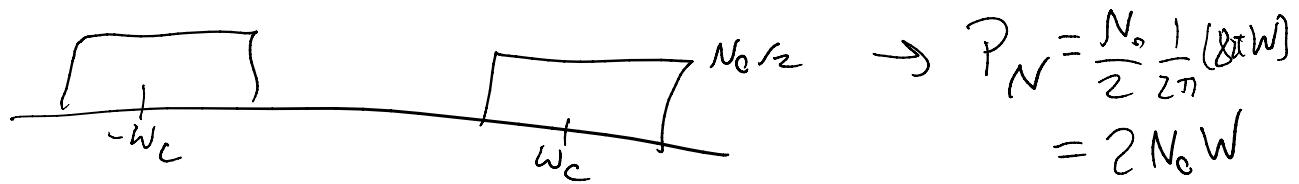
↑ this is stopping pt for non-coherent + assume coherent, then $\varphi = 0$

$$y(t) = \underbrace{\frac{1}{2} A_c m(t)}_{\text{What we had w/ no noise for coh.}} + \underbrace{\frac{1}{2} n_I(t)}_{P_{n_I} = P_N \text{ by point (2)}}$$

So power in signal (as before) is $P_o = \frac{A_c^2}{4} P_M$

Power in noise is $\frac{1}{4} P_N = P_{n_o}$

$$S_N(\omega)$$



So power in noise $P_{n_o} = \frac{N_o W}{\pi} = \sigma^2 W$

$$S_o \text{ power in noise } P_{n_o} = \frac{N_o W}{2} = \sigma^2 W$$

$$\begin{aligned} SNR &= \frac{\frac{A_c^2/4 P_m}{N_o W/2}}{\boxed{\frac{A_c^2}{2 N_o W} P_m}} \\ &\stackrel{DSB-SC}{\text{coherent}} = \frac{A_c^2}{4 \sigma^2 W} P_m \end{aligned}$$

$$SNR_{bb} \text{ was } \frac{P_R}{N_o W}, \quad P_R = \frac{1}{T} \int_0^T (A_c \cos \omega_c t m(t))^2 dt$$

$$\boxed{SNR_{bb} = \frac{P_R}{N_o W} = \frac{A_c^2 P_m}{2 N_o W} = SNR_{PSB-SC}}$$

coherent + DSB-SC
has same SNR as BB

SSB $r(t) = u(t) + n(t)$ AWGN

$$u(t) = A_c m(t) \cos \omega_c t \quad \begin{matrix} \swarrow A_c \hat{m}(t) \sin \omega_c t \\ \nwarrow LSSB \end{matrix}$$

$$r(t) = (A_c m(t) + n_I(t)) \cos \omega_c t - (\pm A_c \hat{m}(t) + n_Q(t)) \sin \omega_c t$$

↓ demod + BB filter, assume coherent

$$y(t) = \frac{A_c}{2} m(t) + \frac{1}{2} n_i(t) \quad (\text{for UDSB or LSSB})$$

$$= P_o + P_n = \frac{A_c^2}{4} P_M + \frac{1}{4} P_n \rightarrow$$

$$SNR_{SSB} = \frac{\frac{1}{4} A_c^2 P_M}{\frac{1}{4} N_e W} = \boxed{\frac{A_c^2 P_M}{N_e W}}$$

$$P_n = \frac{1}{2\pi} \frac{N_o}{2} 4\pi W = N_e W$$

Seems better than DSB

but we transmitted $m(t)$ and $\hat{m}(t)$ at same power
so we receive 2 times signal per

so if $P_R^{DSB} = \frac{A_c^2}{2} P_M$, then $P_R^{SSB} = A_c^2 P_M$
 so

$$\boxed{SNR_{SSB} = \frac{P_R}{N_e W} = SNR_{DSB} = SNR_{DSB-SC}}$$

P_R is more directly related to transmitted power ← more cost

whereas P_{Signal} in SNR computation is demod signal power

SSB - transmit 2* power, half BW, same SNR

Power/BW tradeoff

Conventional

Tuesday, September 22, 2020 7:43 PM

$$u(t) = A_c(1 + a_m(t)) \cos \omega_c t$$

$$r(t) = (A_c(1 + a_m(t)) + n_I(t)) \cos \omega_c t - n_Q(t) \sin \omega_c t$$

before if I $\otimes \cos \omega_c t$,

I get rid of $n_Q(t)$ by $\cos t \sin$

here, demand is rectifier + LPF \rightarrow doesn't
get rid of n_Q

Using conventional demand \rightarrow quadrature component noise
 \rightarrow Worse than DSB or SSB

What if I have conventional AM on 2 demod line regular?

= Like DSB AM, with a pilot tone



$$y_1(t) = \underbrace{\frac{1}{2} A_c(1 + a_m(t))}_{\text{DC}} + \frac{1}{2} n_I(t)$$

\downarrow use a DC-block (capacitor)

$$y_2(t) = \frac{a_m}{2} m(t) + \frac{1}{2} n_I(t)$$

$$\overline{P}_N = \overline{P}_N^{\text{DSB}} = \frac{Z N_c W}{4},$$

$$\overline{P}_R = \frac{A_c^2}{2} (a^2 P_M + 1)$$

$$\overline{P}_o = \frac{a^2 A_c^2 P_M}{4}$$

$$P_o = \frac{\alpha^2 A_c^2 P_M}{4}$$

$$\text{So } SNR_{\text{DSB-conv}} = \frac{\alpha^2 A_c^2 P_M}{2 N_0 W} = \frac{\alpha^2 P_M}{1 + \alpha^2 P_M} \cdot \frac{\frac{A_c^2}{2} (\alpha^2 P_M + 1)}{N_0 W}$$

$$\gamma = \frac{\alpha^2 P_M}{1 + \alpha^2 P_M} < 1$$

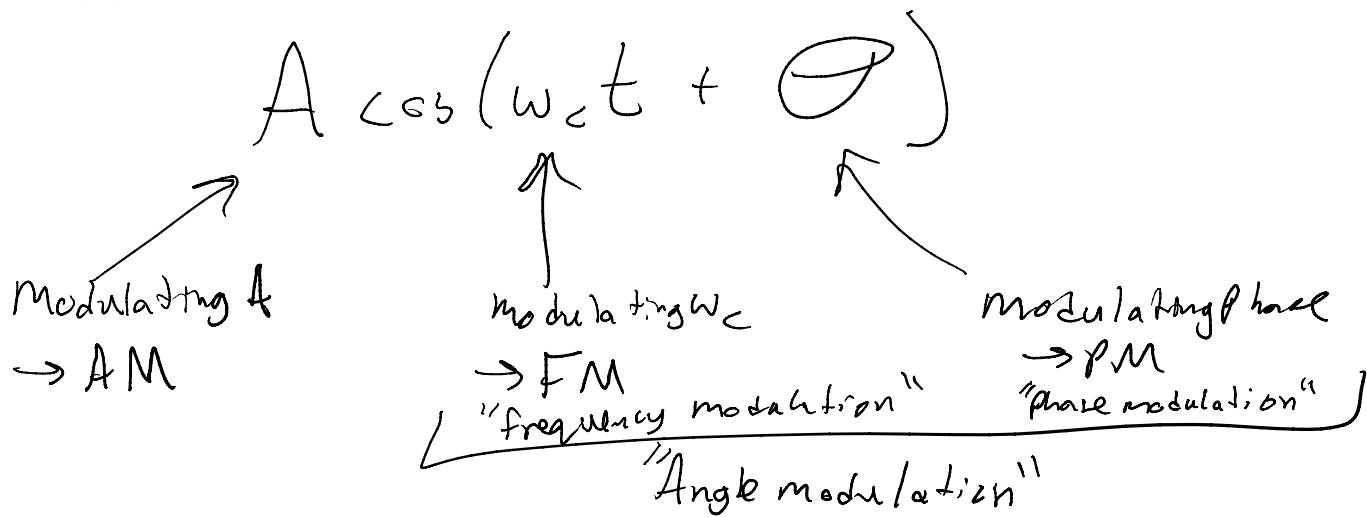
$$= \gamma \frac{P_R}{N_0 W} = \gamma SNR_{\text{bb}}$$

So With a pilot tone, SNR decreased by γ
waste power in pilot tone

Unless carrier component \rightarrow large part of P_R

FM/PM

Tuesday, September 22, 2020 7:55 PM



- Big Problem
- not intuitively all
 - nonlinear - modulating the argument of a transcendental function

We have to rely on math more than intuition
hard

Generally: write any angle-modulated signal as

$$u(t) = A_c \cos(w_c t + \varphi(t))$$

how will this affect frequency content of the signal?

One simple example let $\varphi(t) = \omega_0 t + \Theta$ — general linear phase

$$u(t) = A_c \cos((w_c + \omega_0)t + \Theta) \quad \text{increased freq}$$

but still narrowband

what if $\varphi(t) = \sin \omega_0 t$? — We have no tools to ...

What if $\varphi(t) = \sin \omega_0 t$? — We have no tools to deal with this yet



Def. The instantaneous frequency of an angle-modulated signal $u(t) = A_c \cos(\omega_c t + \varphi(t))$ is

$$f_i(t) = f_c + \frac{1}{2\pi} \frac{d}{dt} \varphi(t)$$

Or $\omega_i(t) = \omega_c + \frac{d}{dt} \varphi(t)$ in radian freq.



In above example, $\varphi(t) = \omega_0 t + \theta$, $\varphi'(t) = \omega_0$

$$\omega_i = \omega_c + \omega_0 = \text{freq. of mod. signal}$$

"More-or-less" freq. of signal at time t

in PM

Tuesday, September 22, 2020 8:10 PM

We let $\varphi(t) \approx k_p m(t)$ \rightarrow phase \propto message

in FM

We let the instantaneous freq. deviation is prop-
to the message $f_i - f_c \propto m$

$$f_i(t) - f_c = k_f m(t) = \frac{1}{2\pi} \frac{d}{dt} \varphi(t)$$

in FM $\varphi \propto \int m(t) dt$

Specifically

$$\varphi(t) = 2\pi k_f \int_{-\infty}^t m(\tau) d\tau$$

FM might better be called "instantaneous freq. mod"

How do we demod?
we want to find $\varphi(t)$ somehow

Ex. Let $m(t) = \alpha \cos \omega_0 t$

$$\text{PM: } \varphi(t) = \alpha K_p \cos \omega_0 t$$

$$\text{FM: } \varphi(t) = 2\pi \alpha K_f \int_{-\infty}^t \cos \omega_0 \tau d\tau$$

$$= \frac{2\pi \alpha K_f}{\omega_0} \sin \omega_0 t$$

$$u_{PM}(t) = A_c \cos(\omega_c t + \alpha K_p \cos \omega_0 t)$$

$$u_{FM}(t) = A_c \cos(\omega_c t + \frac{2\pi \alpha K_f}{\omega_0} \sin \omega_0 t)$$

$$\text{def } \tilde{\beta}_p = \alpha K_p, \quad \tilde{\beta}_f = \frac{\alpha K_f}{f_0}$$

$$u_{PM}(t) = A_c \cos(\omega_c t + \tilde{\beta}_p \cos \omega_0 t)$$

$$u_{FM}(t) = A_c \cos(\omega_c t + \tilde{\beta}_f \sin \omega_0 t)$$

$\text{PM: Phase changes by at most } \tilde{\beta}_p$
More generally: $K_p \max(|m(t)|) = \Delta \varphi_{\max}$

$$= \Delta \varphi_{\max} = \pi - \arccos(m(t))$$

$$FM: \frac{k_f \max(|m(\epsilon)|)}{W} = \Delta\varphi_{\max}, W \text{ BW of } M(\omega)$$

Def modulation index

$$PM: \beta_p = k_p \max(|m(t)|) = \Delta\varphi_{\max}$$

$$FM: \beta_f = \frac{k_f \max(|m(t)|)}{W} = \frac{\Delta f_{\max}}{W}$$

Special Case $\varphi(t) \ll 1 \forall t$ (low mod. index)

"low index" or "Narrowband" FM/PM

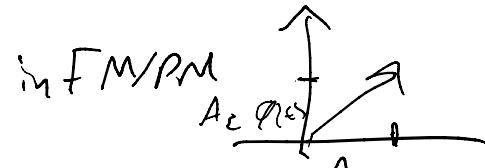
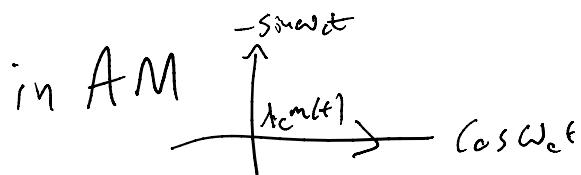
$$u(t) = A_c \cos(\omega_c t + \varphi(t))$$

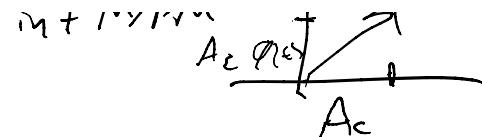
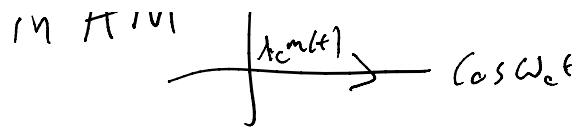
$$= A_c \cos(\omega_c t) \cos(\varphi(t)) - A_c \sin(\omega_c t) \sin(\varphi(t))$$

$$\approx A_c \cos \omega_c t - A_c \varphi(t) \sin \omega_c t$$

$\cos x \approx 1$ $\sin x \approx x$
for small x

$$I(t) = A_c, Q(t) = A_c \varphi(t)$$

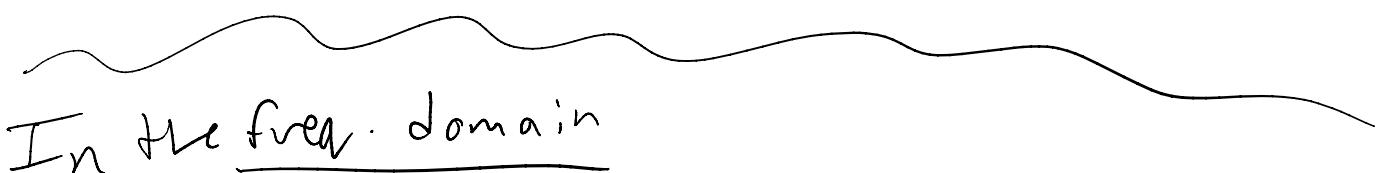
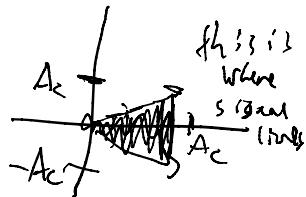




$$\text{narrowband} \rightarrow I(t) = A_c \quad Q(t) = A_c \cos \varphi(t)$$

$$\text{More generally} \rightarrow I(t) = A_c \cos \varphi(t), \quad Q(t) = A_c \sin \beta(t)$$

(Range $[-A_c, A_c]$)



$$\text{Special case } m(t) = a \cos \omega_0 t$$

$$u_{FM}(t) = A_c \cos(\omega_c t + \beta_f \sin \omega_0 t)$$

$$= \operatorname{Re} (A_c e^{j\omega_c t} e^{j\beta_f \sin \omega_0 t})$$

\leftarrow periodic with period

$$T = \frac{2\pi}{\omega_0}$$

So we can rep. as a Fourier Series

$$e^{j\beta_f \sin \omega_0 t} = \sum_{n=-\infty}^{\infty} c_n e^{j\omega_0 n t}$$

def. of
F.S.

$$c_n = \frac{1}{T} \int_0^T e^{j\beta_f \sin \omega_0 t} e^{-jn\omega_0 t} dt$$

$$, \pi^2 \pi : (\beta \sin u - nu)$$

$$= \frac{1}{2\pi} \int_0^{2\pi} e^{j(\beta \sin u - nu)} du , \quad u = w_0 +$$

Evaluate this integral \times it's impossible

This is an indexed set of functions of β with a name!

Def. The n^{th} order Bessel Function of the First Kind
for integer n $J_n : \mathbb{R} \rightarrow \mathbb{R}$ is def as

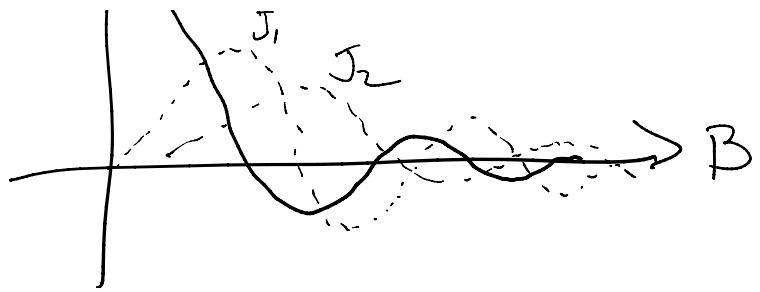
$$J_n(\beta) = \frac{1}{2\pi} \int_0^{2\pi} e^{j(\beta \sin u - nu)} du$$

Shockingly useful

We have a lot of info about Bessel Funcs

- 1) $J_{-n} = (-1)^n J_n$, so $|J_n(\beta)| = |J_{-n}(\beta)|$
- 2) $\max_{\beta > 0} (J_n(\beta)) > \max_{\beta > 0} (J_{n+1}(\beta)) \forall n$
- 3) I will draw them for you Poorly





$$4) J_n(\beta) = \sum_{k=0}^{\infty} \frac{(-1)^k (\beta/2)^{n+2k}}{k! (n+k)!}$$

for small β , $J_n(\beta) \approx \frac{\beta^n}{2^n n!}$ so as $n \rightarrow \infty$
 $J_n(\beta) \downarrow$ for small β

For higher β , you need more modes to estimate power

for $\beta \leq 8$ can get 98% power est. with < 10 terms of sum

$$\begin{aligned} u(t) &= \operatorname{Re} \left(A_c \sum_{n=-\infty}^{\infty} J_n(\beta) e^{j\omega_0 nt} e^{j\omega_c t} \right) \\ &= A_c \sum_{n=-\infty}^{\infty} J_n(\beta) \cos((\omega_c + n\omega_0)t) \end{aligned}$$

uhoh...

I transmitted one tone and my FM signal has
infinite BW

freq. of signal was mod. \rightarrow nonlinear op

but $J_n \downarrow$ as $n \rightarrow \infty$, it decays in freq.

So angle mod. signals are infinite BW

but approximately Finite BW

So we can lose a little info in high freq. components
and filter before TX and after RX

$$4X'(t) + X(t-\tau)$$

$$X(t) \rightarrow [H_1] \rightarrow X(t-\tau)$$

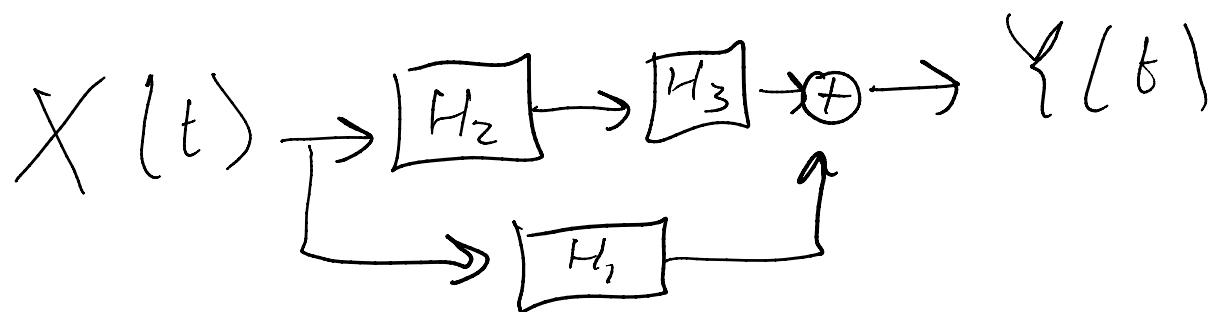
$$H_1 = e^{j\omega\tau}$$

$$X(t) \rightarrow [H_2] \rightarrow 4X(t)$$

$$H_2 = 4$$

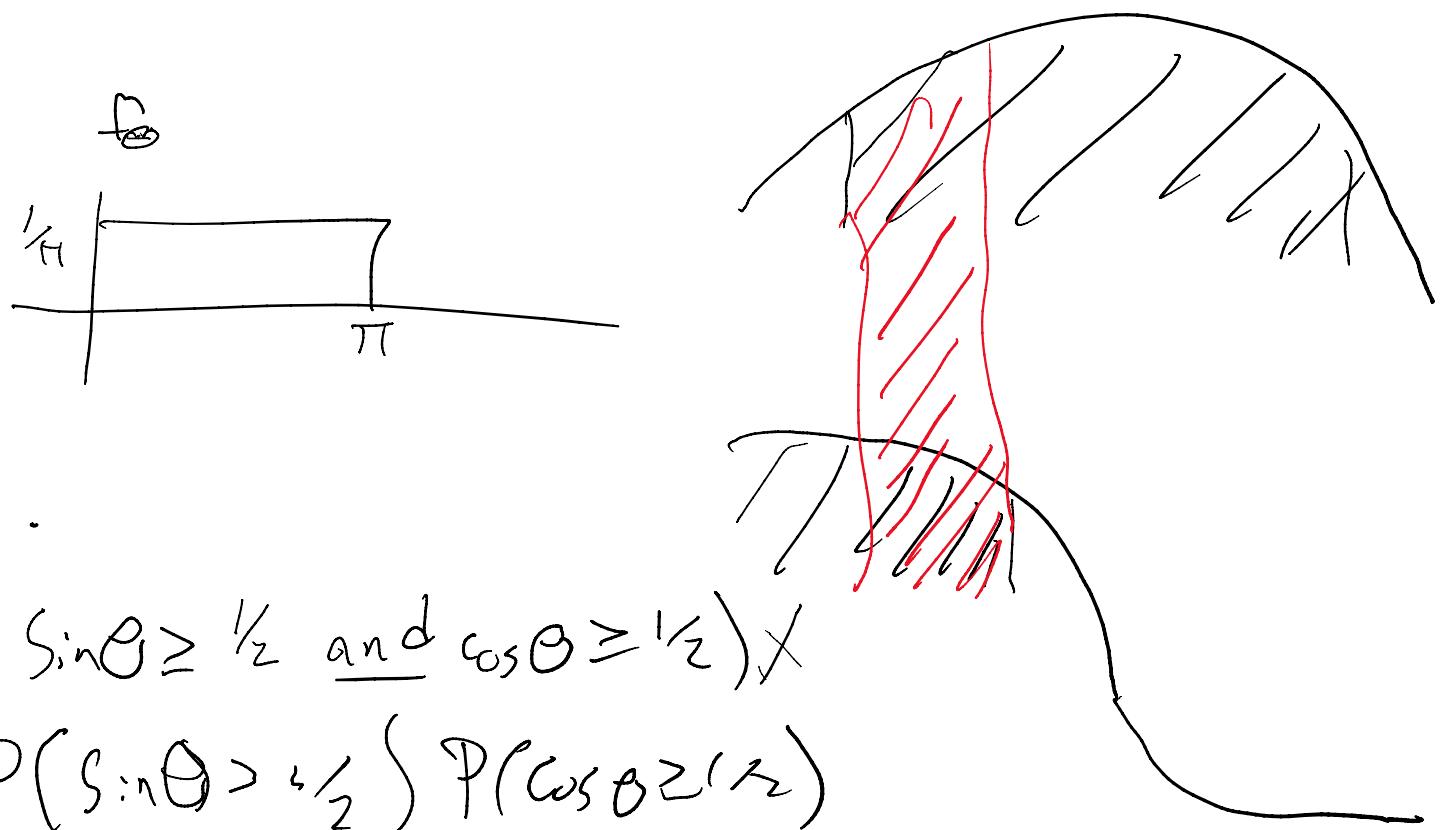
$$X(t) \rightarrow [H_3] \rightarrow X'(-t)$$

$$H_3 = j\omega$$



$$X(t) \rightarrow [H] \rightarrow Y(t)$$

$$\begin{aligned} H &= H_2 H_3 + H_1, \\ &= 4j\omega + e^{j\omega T} \end{aligned}$$



$$R_x = E[\tilde{x}]$$

Sample autocorr = Sample mean(\tilde{x})

\downarrow as $N \uparrow$

$$E[\tilde{x}]$$