

Coupled-cluster theory for three-body Hamiltonians

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We derive coupled-cluster equations for three-body Hamiltonians. The equations for the one- and two-body cluster amplitudes are presented in a factorized form that leads to an efficient numerical implementation. We employ low-momentum two- and three-nucleon interactions and calculate the binding energy of ${}^4\text{He}$. The results show that the main contribution of the three-nucleon interaction stems from its density-dependent zero-, one-, and two-body terms that result from the normal ordering of the Hamiltonian in coupled-cluster theory. The residual three-body terms that remain after normal ordering can be neglected.

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I. INTRODUCTION

One of the central challenges in nuclear theory is to understand and predict the structure of nucleonic matter based on microscopic nucleon-nucleon (NN) and many-nucleon interactions. In recent years, there has been significant progress in exact calculations of ground and excited states of light nuclei based on various high-precision interactions fitted to NN data [1–7]. These results clearly show that three-nucleon forces (3NFs) contribute significantly: Without 3NFs, the binding energies depend strongly on the NN potential used, which can be traced to scheme and model dependences in any theory restricted to NN interactions. The existence of 3NFs is not surprising, because nucleons are not point particles. There are always virtual excitations (high-momentum nucleons and Δ -isobars) or internal degrees of freedom (quarks and gluons) that have been “integrated out.” This directly leads to three- and many-nucleon interactions. The modern understanding is that nuclear interactions are effective interactions and depend on the resolution scale given by a cutoff Λ of the effective theory (see, for example, Ref. [8]). Exact calculations are cutoff independent up to the effects of omitted higher-order interactions, and therefore varying the resolution scale is a powerful tool to analyze the predictive power of theoretical calculations.

The study of 3NFs in systems beyond the lightest nuclei is an important goal. This requires a flexible technique to solve the many-body problem, including NN and 3N interactions. Coupled-cluster theory is a promising tool for this endeavor. This method originated in nuclear physics [9,10] and is today mostly propelled through its importance in quantum chemistry [11,12]. For reviews, we refer the reader to Refs. [13–17]. After the seminal work by the Bochum group [13], Heisenberg and Mihaila employed coupled-cluster theory for structure calculations of ${}^{16}\text{O}$ based on realistic NN potentials [18]. For their calculation of the charge form factor of ${}^{16}\text{O}$, they also included selected contributions from 3NFs that could be

cast into the form of density-dependent NN interactions [19]. Another recent approach employed *ab initio* coupled-cluster theory for structure calculations in closed-shell nuclei ${}^4\text{He}$ and ${}^{16}\text{O}$ [20–22], in open-shell nuclei as the neighbors of ${}^{16}\text{O}$ [23], and weakly bound and unbound helium isotopes [24]. These calculations were limited to NN interactions. It is the purpose of this article to develop coupled-cluster theory for three-body Hamiltonians. This extension of the coupled-cluster method might also find applications in condensed-matter theory and quantum chemistry [25–27].

A current frontier in nuclear structure theory is to determine consistent 3NFs corresponding to the different NN interactions and with predictive power when extrapolated to the extremes of isospin and to moderate densities. Several theoretical approaches are currently being used. The Tucson-Melbourne 3NF already employed symmetries of quantum chromodynamics (QCD) in its construction [28,29]. Existing phenomenological 3NFs include the Fujita-Miyazawa force based on 2π exchange with an intermediate Δ -isobar [30] and also various shorter-ranged 3NFs [31–34]. This approach has led to a very successful description of light nuclei (for a review see Ref. [2]).

Another approach is to systematically construct NN and higher-order interactions within the framework of chiral effective field theory (EFT) [35–38]. Chiral interactions are expanded in powers of a typical momentum of nucleons in nuclei and in powers of the pion mass, both generically called Q , over the EFT breakdown scale $\Lambda_\chi \sim 1\text{ GeV}$ ($\approx 5\text{ fm}^{-1}$). The EFT power counting naturally explains the hierarchy of NN , 3N, and higher-body interactions [39], which enter only in subleading orders and make calculations for complex nuclei based on NN and 3N interactions meaningful. At this point, the leading 3NF has been implemented [40–42].

Nuclear interactions require regularization and renormalization to be meaningful, and with EFT, usually a momentum cutoff scheme is used [37,38,43,44]. The description of the

NN data is not completely independent of the cutoff, but the cutoff variation decreases with increasing order, because the cutoff dependences can be absorbed by higher-order contact interactions. In this way, chiral EFT implements the renormalization group (RG) running of the interactions up to higher-order terms. This is reflected in the leading chiral 3NF, which includes a 3N contact term to be fitted to 3N data. In principle, nuclear structure calculations based directly on chiral interactions are feasible [7,45–47]. However, the typical cutoffs ($\Lambda \approx 2.5\text{--}3\text{ fm}^{-1}$) are somewhat too large to use chiral interactions without resummations or prediagonalizations in many-body calculations.

Renormalization group methods can be used to evolve nuclear interactions to lower momenta, which leads to improved convergence in few- and many-body calculations [48–51]. The resulting low-momentum interactions, known generically as $V_{\text{low } k}$, have variable momentum cutoffs and are approximately independent of the starting NN interaction for cutoffs $\Lambda \lesssim 2\text{ fm}^{-1}$ [48,49]. The RG evolution preserves the long-range parts and, starting from chiral EFT interactions, generates all higher-order contact operators needed to reproduce low-energy NN observables. Moreover, with increasing orders in EFT, the resulting $V_{\text{low } k}$ interactions are very similar to low-momentum interactions obtained from conventional potentials [49,52]. Because chiral EFT represents the most general low-momentum expansion of nuclear forces, the above observations motivate combining low-momentum interactions with 3NFs from chiral EFT [50]. We will follow this approach and employ $V_{\text{low } k}$ with the corresponding low-momentum 3NF adjusted to the binding energies of ^3H and ^4He [50]. In this way, we can expect to define approximately consistent NN and 3N interactions. In this first study, we focus on the development of the coupled-cluster method to include 3NFs and will present results for only one cutoff. A study of the cutoff variation and associated uncertainties due to higher-order many-body interactions will be left to future work.

This article is organized as follows. In Sec. II we present coupled-cluster theory and its extension to three-body Hamiltonians. In Sec. III, we solve the resulting coupled-cluster amplitude equations at the singles and doubles (CCSD) level for ^4He based on low-momentum NN and 3N interactions. Our findings are especially promising as they show that the binding energy of ^4He can be calculated based on those parts of the 3NF that can be viewed as density-dependent zero-, one-, and two-body forces. We summarize our results in Sec. IV.

II. COUPLED-CLUSTER EQUATIONS FOR THREE-NUCLEON FORCES

This is the main technical section of this article. In the first subsection, we briefly recapitulate coupled-cluster theory and the reformulation of the three-body Hamiltonian in normal-ordered form. The second subsection deals with the diagrammatic derivation and the factorization of the coupled-cluster equations due to the residual three-body force that remains after normal ordering of the Hamiltonian.

A. Coupled-cluster theory

We consider a pure three-body Hamiltonian. Coupled-cluster theory for one- and two-body Hamiltonians is a mature field, and we refer the reader to the reviews [13–17]. The three-body Hamiltonian is written as

$$\hat{H}_3 = \frac{1}{36} \sum_{pqrstu} \langle pqr || stu \rangle \hat{a}_p^\dagger \hat{a}_q^\dagger \hat{a}_r^\dagger \hat{a}_u \hat{a}_t \hat{a}_s. \quad (1)$$

Here, $\langle pqr || stu \rangle$ denotes the antisymmetrized three-body matrix elements, whereas \hat{a}_p^\dagger and \hat{a}_p create and annihilate a fermion in the single-particle orbital p , respectively.

In coupled-cluster theory, the Fermi vacuum is a single-particle product state $|\phi\rangle = \prod_{i=1}^A \hat{a}_i^\dagger |0\rangle$, where the A lowest-energy orbitals are occupied. In a first step, we cast the three-body interaction into a normal-ordered form with respect to this vacuum. In what follows it is assumed that the indices i, j, k, l, m label the occupied orbitals of $|\phi\rangle$, whereas a, b, c, d, e refer to the unoccupied orbitals of $|\phi\rangle$. The former indices run over the number $n_o \equiv A$ of occupied orbitals, whereas the latter run over the remaining number n_u of unoccupied orbitals. Typically, one has $n_u \gg n_o$. Indices referring to all orbitals are denoted as p, q, r, s, t, u ; see, for example, Eq. (1). The normal-ordered Hamiltonian is thus

$$\begin{aligned} \hat{H}_3 = & \frac{1}{6} \sum_{ijk} \langle ijk || ijk \rangle + \frac{1}{2} \sum_{ijpq} \langle ijp || ijq \rangle \{\hat{a}_p^\dagger \hat{a}_q\} \\ & + \frac{1}{4} \sum_{ipqrs} \langle ipq || irs \rangle \{\hat{a}_p^\dagger \hat{a}_q^\dagger \hat{a}_r \hat{a}_s\} + \hat{h}_3, \end{aligned} \quad (2)$$

where \hat{h}_3 denotes the residual three-body Hamiltonian

$$\hat{h}_3 \equiv \frac{1}{36} \sum_{pqrstu} \langle pqr || stu \rangle \{\hat{a}_p^\dagger \hat{a}_q^\dagger \hat{a}_r^\dagger \hat{a}_u \hat{a}_t \hat{a}_s\}. \quad (3)$$

Here, we used the $\{\dots\}$ to denote normal ordering. Writing the three-body Hamiltonian in normal-ordered form, it is clear that the Hamiltonian separates into a zero-, one-, two-, and a three-body term. The first three sums in Eq. (2) are the vacuum expectation value, and the “density-dependent” one- and two-body terms, respectively. Their treatment is standard in coupled-cluster theory, as they simply modify the normal-ordered two-body Hamiltonian from the NN interaction. In this section, we focus on the residual three-body operator of Eq. (3).

In coupled-cluster theory, the correlated state $|\psi\rangle$ is given by a correlation operator $\exp(\hat{T})$ that acts onto a single-particle product state $|\phi\rangle = \prod_{i=1}^A \hat{a}_i^\dagger |0\rangle$ of the A -body system by

$$|\psi\rangle = e^{\hat{T}} |\phi\rangle. \quad (4)$$

The cluster operator,

$$\hat{T} = \hat{T}_1 + \hat{T}_2 + \dots + \hat{T}_A, \quad (5)$$

consists of a one-body cluster operator

$$\hat{T}_1 = \sum_{ia} t_i^a \hat{a}_a^\dagger \hat{a}_i, \quad (6)$$

a two-body cluster operator

$$\hat{T}_2 = \frac{1}{4} \sum_{ijab} t_{ij}^{ab} \hat{a}_a^\dagger \hat{a}_b^\dagger \hat{a}_j \hat{a}_i, \quad (7)$$

and so forth. Note that the two-body cluster amplitudes $t_{ij}^{ab} = -t_{ij}^{ba} = -t_{ji}^{ab} = t_{ji}^{ba}$ are fully antisymmetric. Clearly, \hat{T}_1 and \hat{T}_2 induce 1p-1h and 2p-2h excitations, respectively. In what follows, we will limit the expansion, Eq. (5), of the cluster operator to the one-body cluster, Eq. (6), and the two-body cluster, Eq. (7), respectively. This approximation is referred to as CCSD (“coupled-cluster theory with single and double excitations”). CCSD is a powerful approximation and a compromise between accuracy on the one hand and computational effort on the other hand. One inserts the ansatz Eq. (4) into the Schrödinger equation, multiplies with $\exp(-\hat{T})$ from the left, and obtains the following set of equations:

$$E = \langle \phi | \bar{H} | \phi \rangle, \quad (8)$$

$$0 = \langle \phi_i^a | \bar{H} | \phi \rangle, \quad (9)$$

$$0 = \langle \phi_{ij}^{ab} | \bar{H} | \phi \rangle. \quad (10)$$

Here $|\phi_{i_1 \dots i_n}^{a_1 \dots a_n}\rangle = \hat{a}_{a_n}^\dagger \dots \hat{a}_{a_1}^\dagger \hat{a}_{i_1} \dots \hat{a}_{i_n} |\phi\rangle$ is a np - nh excitation of the product state $|\phi\rangle$, and

$$\bar{H} = \exp(-\hat{T}) \hat{H} \exp(\hat{T}) \quad (11)$$

is the similarity-transformed Hamiltonian. This Hamiltonian is a sum of one-, two-, and three-body Hamiltonians, i.e., $\hat{H} = \hat{H}_1 + \hat{H}_2 + \hat{H}_3$. The treatment of this Hamiltonian within coupled-cluster theory is well known, except for the residual three-body term, Eq. (3), of the normal-ordered three-body Hamiltonian \hat{H}_3 .

The CCSD Eqs. (9) and (10) determine the one-particle and two-particle cluster amplitudes t_i^a and t_{ij}^{ab} , respectively. These amplitudes can then be inserted into the first of the CCSD equations, Eq. (8), to determine the ground-state energy. Note that the similarity-transformed Hamiltonian of Eq. (11) is not Hermitian, and CCSD is not a variational approach. However, the similarity-transformed Hamiltonian can be evaluated *exactly* for any truncation of the cluster operator. In what follows, we will compute the corrections to the energy, Eq. (8), and to the CCSD Eqs. (9) and (10) that arise due to the residual three-body Hamiltonian.

B. Derivation of coupled-cluster equations

In this subsection, we derive the contribution of the residual three-body term \hat{H}_3 from Eq. (3) to the energy, Eq. (8), and the cluster amplitudes, Eqs. (9) and (10). The final results are given in Eq. (12) for the energy and in Eqs. (15) and (16) for the cluster amplitudes, respectively. The contributions to the energy and the coupled-cluster amplitudes can be derived, for instance, by following the approach of Ref. [16]. The resulting expressions contain 2, 15, and 51 terms for the energy and the cluster amplitudes. Many of the individual terms consist of subterms of similar structure. The theoretical derivation presented in this subsection exploits this structure and leads to an efficient numerical implementation.



FIG. 1. (Color online) Energy contributions of the residual three-body Hamiltonian, Eq. (3), in the CCSD approximation.

It is most convenient to evaluate the matrix elements in Eq. (8) of the similarity-transformed Hamiltonian, Eq. (11), in a diagrammatic form (see, for example, Refs. [16,53]). Diagrams are a useful book keeping device to keep track of the (considerable) number of possible Wick contractions. We refer the reader to the literature for a more detailed description.

In a first step, we determine the correction to the CCSD energy, Eq. (8), that is due to the residual three-body Hamiltonian (3). The matrix element $\langle \phi | \bar{H}_3 | \phi \rangle$ is a sum of all topologically different diagrams, where the Hamiltonian, Eq. (3), is fully contracted by the cluster operators. The two diagrams that enter this expression are presented in Fig. 1. In these diagrams, the thick horizontal bar represents the residual three-body Hamiltonian, Eq. (3); the thin horizontal bars denote the one-body and two-body cluster operators, respectively. Particle and hole lines are also shown. If uncontracted, the former have an arrow that points upward, whereas the latter have an arrow that points downward. The corresponding algebraic expression for the energy correction is

$$e_3 = \langle \phi | (\hat{H} \hat{T}_1 \hat{T}_2)_c | \phi \rangle + \langle \phi | \left(\frac{1}{6} \hat{H} \hat{T}_1^3 \right)_c | \phi \rangle, \quad (12)$$

$$= \frac{1}{4} \sum_{klm cde} \langle klm || cde \rangle t_k^c t_l^{de} + \frac{1}{6} \sum_{klm cde} \langle klm || cde \rangle t_k^c t_l^d t_m^e.$$

This is the energy correction due to the residual three-body Hamiltonian. Note that the computational effort of the energy, Eq. (12), scales as $O(n_u^3 n_o^3)$. Note also that the summation over $\sum_{kc} t_k^c$ is common to both diagrams and might therefore be factored out. This is the basic idea behind the factorization [53] and will be presented in detail below.

Let us consider the matrix element $\langle \phi_i^a | \bar{H} | \phi \rangle$ appearing in the CCSD Eq. (9). It is given by the sum over all topologically different diagrams where one particle line and one hole line are not contracted. Figure 2 shows the 15 diagrams that enter this matrix element.

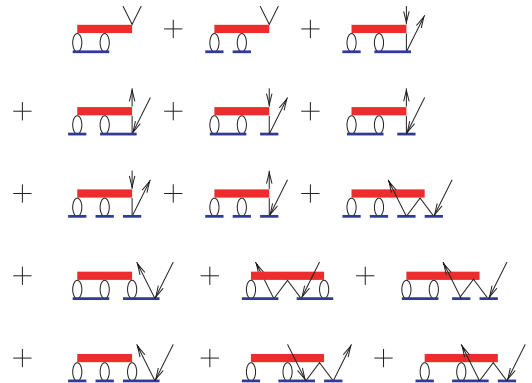


FIG. 2. (Color online) Contributions of the three-body Hamiltonian, Eq. (3), to the \hat{T}_1 cluster equation in the CCSD approximation.

TABLE I. Memory multiplication factor m and computational cost factor c for forming intermediates through contraction with the k -body cluster operator $T_k(p^n h^m)$ via n particle lines and m hole lines.

	$T_2(p^2 h^2)$	$T_1(ph)$	$T_2(p^2 h)$	$T_1(p)$	$T_2(p^2)$	$T_2(ph^2)$	$T_2(ph)$	$T_2(p)$	$T_2(h^2)$	$T_1(h)$	$T_2(h)$
m	$(n_u n_o)^{-2}$	$(n_u n_o)^{-1}$	n_u^{-2}	n_o/n_u	$(n_o/n_u)^2$	n_o^{-2}	1	n_o^2	$(n_u/n_o)^2$	n_u/n_o	$n_u^2 c$
1	1	n_o	n_o	n_o^2	n_u	$n_u n_o$	$n_u n_o^2$	n_u^2	n_u	$n_u^2 n_o$	

The translation of these diagrams into algebraic expressions is straightforward, but there are two reasons to not present them here. First, the naive numerical implementation of the resulting expressions would be inefficient. To see this, consider, for instance, the third diagram in Fig. 2. The corresponding algebraic expression is

$$-\frac{1}{2} \sum_{ckdlm} \langle klm || cdi \rangle t_k^c t_{lm}^{da}, \quad (13)$$

and its naive evaluation costs $O(n_u^3 n_o^4)$ operations. However, performing the summations involving the \hat{T}_1 cluster operator first yields the intermediate

$$I_{di}^{lm} \equiv \frac{1}{2} \sum_{ck} \langle klm || cdi \rangle t_k^c, \quad (14)$$

which only costs $O(n_u^2 n_o^4)$ operations and requires $O(n_u n_o^3)$ in memory. The subsequent contraction of this intermediate with the remaining \hat{T}_2 cluster operator costs only $O(n_u^2 n_o^3)$ operations. Clearly, the memory cost of the intermediate is overcompensated by the reduction of computational cycles. Note also that the intermediate Eq. (14) enters in the evaluation of the seventh diagram depicted in Fig. 2. The second reason is that the complexity of the involved diagrams increases rapidly. The number of diagrams increases from 2 to 15 to 51, when going from Eq. (8) to Eq. (9) to Eq. (10), respectively. The construction of each individual diagram and its inspection regarding the construction and use of intermediates then becomes cumbersome, and a more systematic approach is called for. Similar comments apply when improving the coupled-cluster wave function through the inclusion of three-body or four-body cluster amplitudes. One therefore considers a factorization of the coupled-cluster equations [53]. This approach yields a very compact form of the coupled-cluster equations and is particularly useful for the numerical implementation [54–56]. So far, factorized coupled-cluster equations have been derived in a two-step procedure. The first step consists of constructing all topologically different coupled-cluster diagrams. In a second step, these diagrams are analyzed and repeatedly decomposed into simpler intermediates that undergo single contractions. Here, we proceed differently and present a *direct* diagrammatic derivation of the factorized coupled-cluster equations. Our derivation avoids the explicit construction of all individual coupled-cluster diagrams.

We have to decide in which order multiple contractions of the Hamiltonian with the cluster operators should be performed. Let $T_k(p^n h^m)$ denote the contraction of the Hamiltonian with the k -body cluster operator \hat{T}_k via n particle lines and m hole lines. The contraction of $T_k(p^n h^m)$ with an object

of i particle lines and j hole lines costs $cn_u^i n_o^j$ computational operations and results in an object of size $mn_u^i n_o^j$. Here, c and m denote the computational cost and memory multiplier, respectively, and one finds

$$c = n_u^{k-n} n_o^{k-m},$$

$$m = n_u^{k-2n} n_o^{k-2m}.$$

Based on this analysis, we find that the cost of two subsequent contractions labeled T_A and T_B , respectively, is proportional to $c(T_A) + m(T_A)c(T_B)$ when contraction T_A is first and proportional to $c(T_B) + m(T_B)c(T_A)$ when contraction T_B is first. Table I shows (from left to right) the optimal order in which subsequent contractions should be performed, under the condition that $1 \ll n_o \ll n_u$ and $n_o^2 < n_u$. We also listed the relative computational cost c , and the memory multiplier m . For the first seven entries, the order is easily understood. These contractions do not yield an increase of the size of the contracted object (because $m \leq 1$), and the order is therefore determined by the computational cost. The remaining four contractions increase the size of the contracted object (because $m > 1$), and it is usually most efficient to perform the computationally more expensive contraction before performing the second contraction on an object with increased size.

Let us now turn to the diagrammatic factorization of the coupled-cluster equations. Considering Eq. (9) and Fig. 2, we have to construct all coupled-cluster diagrams with one incoming hole line and one outgoing particle line. These diagrams should be constructed from simpler diagrams, adding one contraction at each step. For the residual three-body Hamiltonian, the root is clearly given by the diagram 1 in Fig. 3. This diagram is the only one that has three incoming particle lines and three outgoing hole lines. Diagrams 2 and 3 have one outgoing particle line and one incoming hole line, respectively, and also have a total of six outgoing and incoming lines. They are sums of two diagrams. The first is the residual three-body Hamiltonian with this appropriate particle and hole lines, whereas the second diagram is a contraction of diagram 1 with a \hat{T}_1 cluster operator.

In what follows, we adopt the following convention. For a diagram that is the sum of diagrams, we label the first, second, third, etc., term of the sum by a, b, and c, respectively. For example, diagram 2 is the sum of diagram 2a and diagram 2b. These labels are not printed in Fig. 3. Diagram 4 is the sum of four diagrams, namely the corresponding residual three-body Hamiltonian and three contractions of previously generated diagrams. Note that of the two diagrams of diagram 2, only diagram 2a enters. This is due to the order specified in Table I. Diagram 2b is a contraction of diagram 1 with $T_1(p)$, and this contraction cannot be followed by the contraction $T_1(p)$. It is

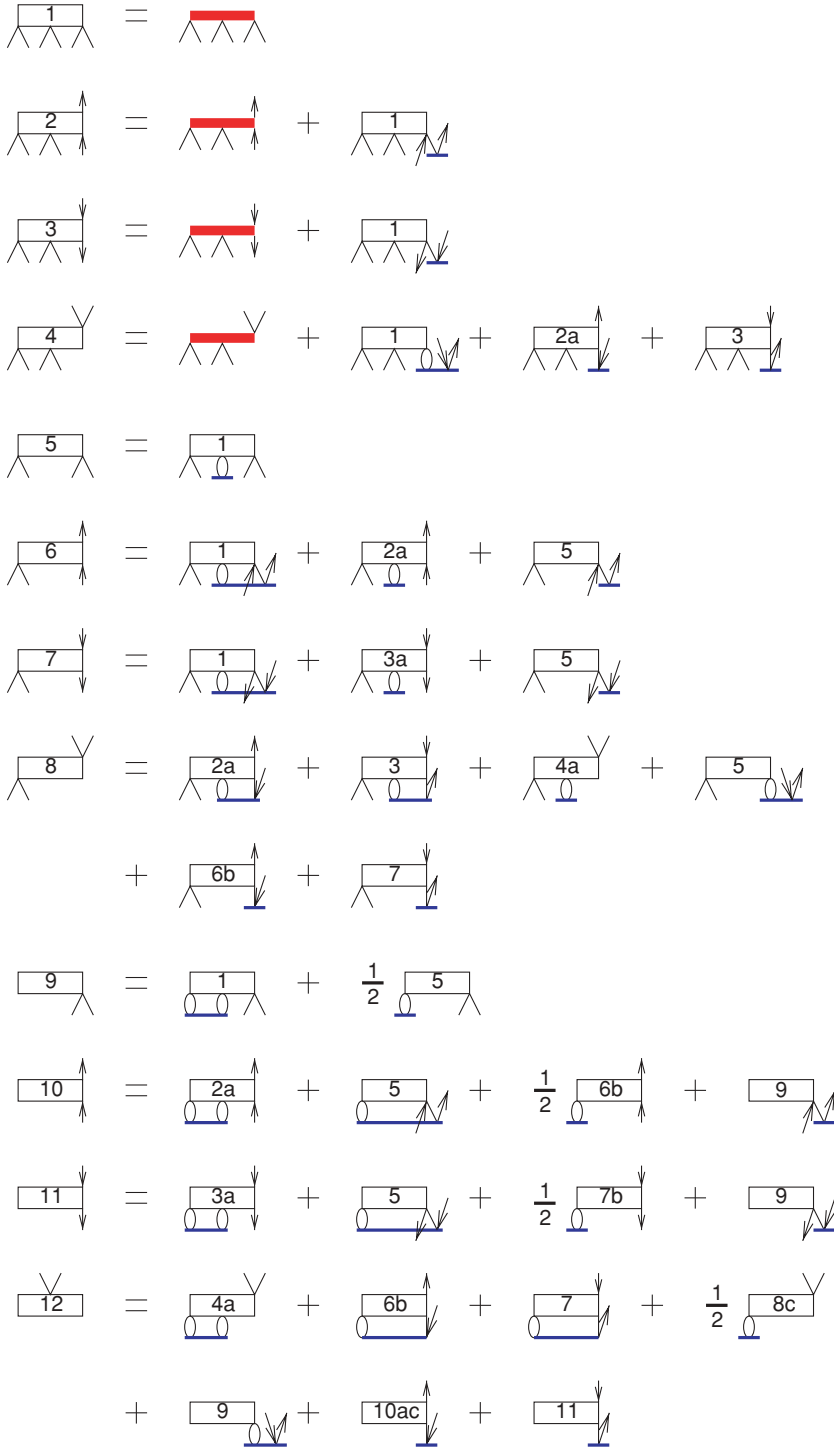


FIG. 3. (Color online) Matrix elements of the residual three-body part of the similarity-transformed Hamiltonian, Eq. (11), that enter the amplitude Eq. (9).

now clear how to proceed. Diagrams 5 to 8 have a total of four incoming and outgoing lines, whereas diagrams 9 to 12 have a total of two incoming and outgoing lines, respectively. Note the peculiar factor $1/2$ in front of diagram 9b. This diagram is a $T_1(ph)$ contraction of diagram 5b, which itself is also a $T_1(ph)$ contraction. The factor $1/2$ will be needed for the translation into algebraic expressions. Similar comments apply to diagrams 10c, 11c, and 12d. Note that diagram 12 consists of all diagrams with one incoming hole line and one outgoing

particle line. The recursive expansion of the corresponding right-hand side yields indeed all diagrams depicted in Fig. 2, and therefore factors the CCSD Eq. (9).

Let us translate Fig. 2 into algebraic expressions. We use the convention that the intermediate $I_{p_1, p_2, \dots}^{q_1, q_2, \dots}$ has incoming lines p_1, p_2, \dots and outgoing lines q_1, q_2, \dots , respectively. In what follows, the intermediate $I(v)$ corresponds to diagram v of Fig. 3. We restrict ourselves to those intermediates that are needed for the construction of diagram 12 and denote

the diagrams 1, 2a, 3a, and 4a directly in terms of the corresponding three-body matrix elements. The result is

$$\begin{aligned}
 I(5)_{ce}^{km} &= \frac{1}{4} \sum_{dl} \langle klm || cde \rangle t_l^d, \\
 I(6b)_{ce}^{ka} &= \frac{1}{2} \sum_{dl} \langle kla || cde \rangle t_l^d, \\
 I(7b)_{ci}^{km} &= \frac{1}{2} \sum_{dl} \langle klm || cdi \rangle t_l^d, \\
 I(7)_{ci}^{km} &= \frac{1}{4} \sum_{lde} \langle klm || cde \rangle t_{li}^{de} + I(7b)_{ci}^{km} \\
 &\quad + 2 \sum_e I(5)_{ce}^{km} t_i^e, \\
 I(8c)_{ci}^{ka} &= \sum_{dl} \langle kla || cdi \rangle t_l^d, \\
 I(9)_c^k &= \frac{1}{4} \sum_{delm} \langle klm || cde \rangle t_{lm}^{de} + 2 \sum_{em} I(5)_{ce}^{km} t_m^e,
 \end{aligned}$$

and

$$\begin{aligned}
 I(10ac)_e^a &= \frac{1}{4} \sum_{ckdl} \langle kla || cde \rangle t_{kl}^{cd} + \sum_{kc} I(6b)_{ce}^{ka} t_k^c, \\
 I(11)_i^m &= \frac{1}{4} \sum_{ckdl} \langle klm || cdi \rangle t_{kl}^{cd} + 2 \sum_{kce} I(5)_{ce}^{km} t_{ki}^{ce} \\
 &\quad + \sum_{kc} I(7b)_{ci}^{km} t_k^c + \sum_c I(9)_c^m t_i^c.
 \end{aligned}$$

In terms of these intermediates, the correction to the right-hand side of the CCSD Eq. (9) reads

$$\begin{aligned}
 &\frac{1}{4} \sum_{ckdl} \langle kla || cdi \rangle t_{kl}^{cd} + \sum_{cke} I(6b)_{ce}^{ka} t_{ki}^{ce} - \sum_{ckm} I(7)_{ci}^{km} t_{km}^{ca} \\
 &\quad + \frac{1}{2} \sum_{kc} I(8c)_{ci}^{ka} t_k^c + \sum_{ck} I(9)_c^k t_{ki}^{ca} + \sum_e I(10ac)_e^a t_i^e \\
 &\quad - \sum_m I(11)_i^m t_m^a. \tag{15}
 \end{aligned}$$

These terms have to be added to the well-known CCSD equation for the T_1 -cluster amplitudes based on two-body Hamiltonians.

For the factorization of the CCSD Eq. (10), one has to construct all diagrams with two incoming hole lines and two outgoing particle lines. Their number is 51, and we directly derive them in a factorized form. Figure 4 shows the factorized diagrams that are needed for the recursive construction of the CCSD Eq. (10). The corresponding algebraic expressions for the intermediates are

$$\begin{aligned}
 I(3b)_{cdi}^{klm} &= \frac{1}{12} \sum_e \langle klm || cde \rangle t_i^e, \\
 I(3)_{cdi}^{klm} &= \frac{1}{12} \langle klm || cdi \rangle + I(3b)_{cdi}^{klm}, \\
 I(4c)_{cdi}^{kla} &= \frac{1}{4} \sum_e \langle kla || cde \rangle t_i^e,
 \end{aligned}$$

$$I(7c)_{ci}^{km} = 2 \sum_e I(5)_{ce}^{km} t_i^e,$$

$$I(7ab)_{ci}^{km} = I(7)_{ci}^{km} - I(7c)_{ci}^{km},$$

$$I(8ac)_{ci}^{ka} = \frac{1}{2} \sum_{lde} \langle kla || cde \rangle t_{li}^{de} + I(8c)_{ci}^{ka},$$

$$I(8b)_{ci}^{ka} = -6 \sum_{lmd} I(3)_{cdi}^{klm} t_{lm}^{da},$$

$$I(8d)_{ci}^{ka} = 4 \sum_{me} I(5)_{ce}^{km} t_{mi}^{ea},$$

$$I(8e)_{ci}^{ka} = 2 \sum_e I(6b)_{ce}^{ka} t_i^e,$$

$$I(8abce)_{ci}^{ka} = I(8ac)_{ci}^{ka} + I(8b)_{ci}^{ka} + I(8e)_{ci}^{ka},$$

$$I(10abc)_e^a = I(10ac)_e^a - 2 \sum_{kmc} I(5)_{ce}^{km} t_{km}^{ca},$$

$$\begin{aligned}
 I(14)_{idj}^{klm} &= \frac{1}{12} \langle klm || idj \rangle + \frac{1}{24} \sum_{ce} \langle klm || cde \rangle t_{ij}^{ce} \\
 &\quad + \sum_c \left[\frac{1}{6} \langle klm || cdj \rangle + I(3b)_{cdj}^{klm} \right] t_i^c,
 \end{aligned}$$

and

$$I(15)_{ce}^{ab} = \frac{1}{2} \sum_{ld} \langle alb || cde \rangle t_l^d,$$

$$\begin{aligned}
 I(16)_{ij}^{km} &= \frac{1}{4} \sum_{cdl} \langle klm || cdj \rangle t_{li}^{dc} + \frac{1}{2} \sum_{ce} I(5)_{ce}^{km} t_{ij}^{ce} \\
 &\quad + \sum_c \left[I(7ab)_{cj}^{km} + \frac{1}{2} I(7c)_{cj}^{km} \right] t_i^c \\
 &\quad + \frac{1}{4} \sum_{ld} \langle klm || idj \rangle t_l^d,
 \end{aligned}$$

$$\begin{aligned}
 I(18)_{idj}^{klb} &= \frac{1}{4} \langle klb || idj \rangle + \frac{1}{8} \sum_{ce} \langle klb || cde \rangle t_{ij}^{ce} \\
 &\quad + \sum_c \left[\frac{1}{2} \langle klb || cdj \rangle + I(4c)_{cdj}^{klb} \right] t_i^c,
 \end{aligned}$$

$$I(19ac)_{cj}^{ab} = \frac{1}{2} \sum_{lde} \langle alb || cde \rangle t_{ij}^{de} + \sum_{ld} \langle alb || cdj \rangle t_l^d,$$

$$I(19b)_{cj}^{ab} = 2 \sum_e I(15)_{ce}^{ab} t_j^e,$$

$$\begin{aligned}
 I(20a - g)_{ij}^{kb} &= \frac{1}{2} \sum_{cld} \langle klb || cdj \rangle t_{il}^{cd} + \frac{1}{2} \sum_{ec} I(6b)_{ce}^{kb} t_{ij}^{ce} \\
 &\quad + 2 \sum_{mc} I(7)_{ci}^{mk} t_{mj}^{cb} + \frac{1}{2} \sum_{ld} \langle klb || idj \rangle t_l^d \\
 &\quad + \sum_c \left[I(8ac)_{cj}^{kb} + I(8e)_{cj}^{kb} \right] t_i^c \\
 &\quad + \frac{1}{2} \sum_c I(9)_c^k t_{ij}^{cb} - 3 \sum_{ldm} I(14)_{idj}^{klm} t_{lm}^{db},
 \end{aligned}$$

$$I(20h)_{ij}^{kb} = -2 \sum_m I(16)_{ij}^{km} t_m^b.$$

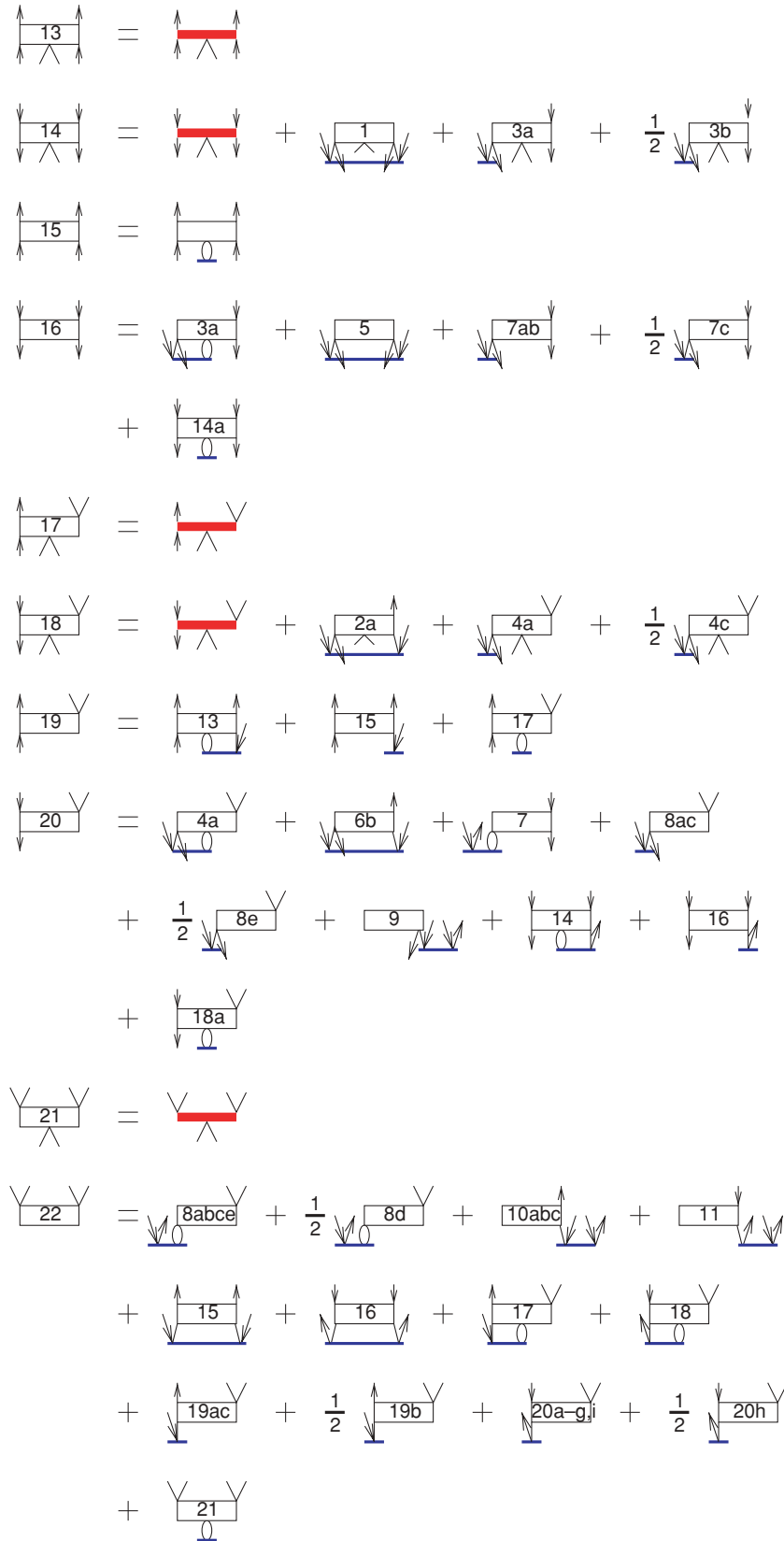


FIG. 4. (Color online) Additional diagrams needed for the contributions of the residual three-body Hamiltonian, Eq. (3), to the CCSD Eq. (10).

The residual three-body Hamiltonian, Eq. (3), thus leads to the following correction, to be added to the right-hand side of the CCSD Eq. (10),

$$\begin{aligned}
& \sum_{ld} \langle alb || idj \rangle t_l^d + \frac{1}{2} P(ij) \sum_{lcd} \langle alb || cdj \rangle t_{li}^{dc} \\
& + P(ab) P(ij) \sum_{kc} \left[I(8abce)_{ci}^{ka} + \frac{1}{2} I(8d)_{ci}^{ka} \right] t_{kj}^{cb} \\
& + P(ab) \sum_e I(10abc)_e^a t_{ij}^{eb} - P(ij) \sum_m I(11)_i^m t_{mj}^{ab} \\
& + \sum_{ce} I(15)_{ce}^{ab} t_{ij}^{ce} + P(ij) \sum_{km} I(16)_{ij}^{km} t_{km}^{ab} \\
& - P(ab) P(ij) \sum_{kld} I(18)_{idj}^{klb} t_{lk}^{da} + P(ij) \\
& \times \sum_c \left[I(19ac)_{cj}^{ab} + \frac{1}{2} I(19b)_{cj}^{ab} \right] t_i^c - P(ab) P(ij) \\
& \times \sum_k \left[I(20a - g, i)_{ij}^{kb} + \frac{1}{2} I(20h)_{ij}^{kb} \right] t_k^a. \quad (16)
\end{aligned}$$

Here the permutation $P(ab)$ implies $P(ab)I_{ab} = I_{ab} - I_{ba}$. Again, these have to be added to the right-hand side of the CCSD Eq. (10). Equation (12) for the energy correction, and the expressions of Eqs. (15) and (16) for the cluster amplitudes are the main technical results of this article. On expansion of diagram 22 in Fig. 4, one gets indeed all 51 diagrams that enter the coupled-cluster Eq. (10). The numerical implementation of the terms in Eqs. (15) and (16) is straightforward. It is interesting to analyze the resulting computational costs. The most expensive intermediates $I(18)$ and $I(19ac)$ cost $n_u^4 n_o^4$ and $n_u^5 n_o^2$ computational cycles, respectively. The most memory-expensive object is the interaction $\langle alb || cde \rangle$, which enters the construction of intermediates $I(15)$ and $I(19ac)$ and requires the storage of $n_u^5 n_o$ real numbers. In our largest calculations for ${}^4\text{He}$, we have $n_o = 4$ and $n_u = 220$.

III. APPLICATION TO ${}^4\text{He}$

We present the first *ab initio* coupled-cluster calculations including 3NFs. Our results are based on low-momentum NN [49] and 3N [50] interactions,

$$H = T + V_{\text{low}k}(\Lambda) + V_{3N}(\Lambda). \quad (17)$$

In this exploratory study we use a sharp cutoff $\Lambda = 1.9 \text{ fm}^{-1}$, and $V_{\text{low}k}$ is derived from the Argonne v_{18} potential [57]. The corresponding 3N interaction is based on the leading chiral 3NF and has been fitted to the ${}^3\text{H}$ and ${}^4\text{He}$ binding energies in Ref. [50]. This 3NF consists of a long-range 2π -exchange part, determined by the low-energy coefficients c_1 , c_3 , and c_4 , an intermediate-range 1π -exchange (“ D -term”) and a short-range contact interaction (“ E -term”) [40,42]. The 3NF operators are multiplied by regulating functions of the incoming and outgoing Jacobi momenta, $f_R(p, q) = \exp\{-(p^2 + \frac{3q^2}{4})/\Lambda^2\}$, with the same cutoff value Λ as in $V_{\text{low}k}$. For additional details on the 3NF, we refer the reader

to Refs. [50,51]. These low-momentum interactions can be directly employed within coupled-cluster theory.

In this application of coupled-cluster theory we restrict ourselves to a proof-of-principle calculation. For simplicity, we have therefore considered only the 3NF channel with total isospin $T = 1/2$, total angular momentum $J = 1/2$, and positive parity. This partial wave is the dominant contribution to the binding energies of light nuclei: For ${}^4\text{He}$, the corresponding Faddeev-Yakubovsky result is $E = -28.20(5) \text{ MeV}$, which differs only by 100 keV from the exact energy including all partial waves, $E = -28.30(5) \text{ MeV}$ [50].

The coupled-cluster calculations are performed in a harmonic oscillator (HO) basis, with basis parameters given by the oscillator spacing $\hbar\omega$ and the number N of oscillator shells. We will present our results as a function of these parameters. The matrix elements of the 3NF are calculated first in relative HO states given by the expansion in 3N partial waves. For the transformation from relative HO states to the single-particle (“ m -scheme”) basis, we essentially follow the Appendix of Ref. [7]. However, we can also start from a 3NF in the non-antisymmetrized HO basis, which simplifies the transformations given in Eqs. (B.9), (B.11), and (B.12) of Ref. [7], and explicitly antisymmetrize the m -scheme matrix elements in the last step of our transformation. This approach does not require coefficients of fractional parentage. We verified that our matrix elements agree with those obtained from a transformation based on antisymmetrized relative HO matrix elements. We further have checked that the transformation preserves the unit matrix and that it yields identical matrix elements for $\sigma_1 \cdot \sigma_2$ and $\sigma_2 \cdot \sigma_3$ in antisymmetrized states.

In Fig. 5, our CCSD results for the binding energy of ${}^4\text{He}$ are shown with and without 3NFs as a function of the oscillator spacing and with increasing model space size. The 3NF contribution is repulsive, in agreement with the Faddeev-Yakubovsky calculation [50], and correspondingly, the minima in $\hbar\omega$ are shifted to smaller oscillator spacings. We observe a slow convergence at the last few 100-keV level, which is

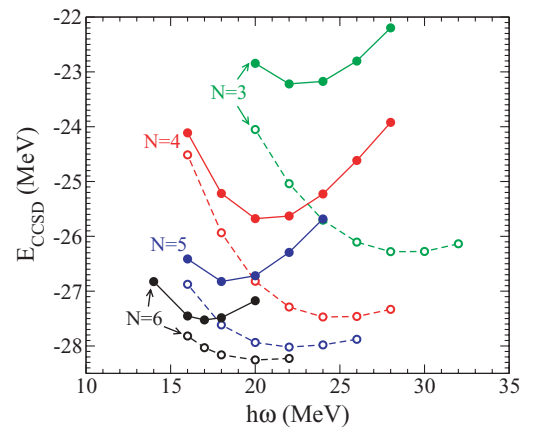


FIG. 5. (Color online) CCSD results for the binding energy of ${}^4\text{He}$ as a function of the oscillator spacing and for model spaces consisting of $N = 3$ to $N = 6$ oscillator shells. The CCSD calculations are based on low-momentum NN and 3N interactions, where the full and dashed lines, respectively, denote the energy obtained with and without 3NFs.

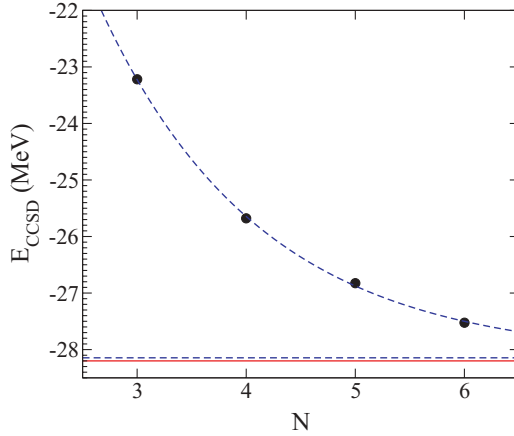


FIG. 6. (Color online) (Data points) CCSD results (taken at the $\hbar\omega$ minima) for the binding energy of ^4He with 3NFs as a function of the number of oscillator shells. (Dashed lines) Exponential fit to the data and asymptote of the fit. (Full line) Exact result.

due to the sharp cutoff in $V_{\text{low } k}$. This might be improved by using low-momentum interactions with smooth cutoffs [58]. Using the minima of the CCSD results with 3NFs, we make an exponential fit of the form $E(N) = E_\infty + a \exp(-bN)$ to the data points. The result is shown in Fig. 6. The extrapolated infinite model space value is $E_\infty = -28.09$ MeV, which is very close to the exact result $E = -28.20(5)$ MeV.

It is interesting to analyze the different contributions ΔE to the binding energy E . The individual contributions are given in Fig. 7 for a model space of $N = 4$ oscillator shells and $\hbar\omega = 20$ MeV. The main contribution stems from the low-momentum NN interaction. The contributions from 3NFs account only for about 10% of the total binding energy. This

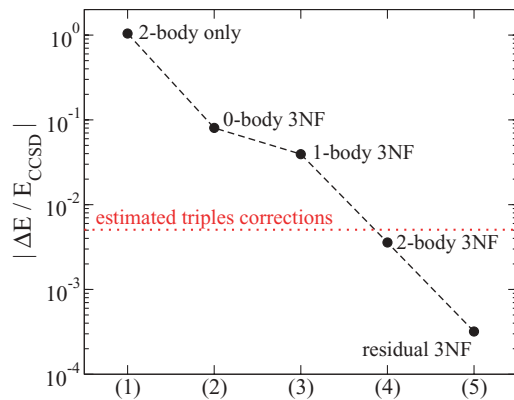


FIG. 7. (Color online) Relative contributions $|\Delta E/E|$ to the binding energy of ^4He at the CCSD level. The different points denote the contributions from (1) low-momentum NN interactions, (2) the vacuum expectation value of the 3NF, (3) the normal-ordered one-body Hamiltonian due to the 3NF, (4) the normal-ordered two-body Hamiltonian due to the 3NF, and (5) the residual 3NFs. The dotted line estimates the corrections due to omitted three-particle/three-hole clusters.

is consistent with the chiral EFT power-counting estimate $\langle V_{3N} \rangle \sim (Q/\Lambda_\chi)^3 \langle V_{\text{low } k} \rangle \approx 0.1 \langle V_{\text{low } k} \rangle$ [50] (see also Table I in Ref. [52]). The second, third, and fourth largest contribution are due to the first, second, and third term on the right-hand side of Eq. (2). These are the density-dependent zero-, one-, and two-body terms, which resulted from the normal ordering of the three-body Hamiltonian in coupled-cluster theory. The contributions from the residual three-body Hamiltonian, Eq. (3), are very small and are represented by the last point in Fig. 7. Recall that the residual 3NF contributes to the energy directly through Eq. (12) and indirectly through a modification of the cluster amplitudes via Eqs. (15) and (16). Apparently, both contributions are very small. In addition and independent of the result that low-momentum 3N interactions are perturbative for cutoffs $\Lambda \lesssim 2 \text{ fm}^{-1}$ [50], we find here that the contributions of 3NFs decrease rapidly with increasing rank of the normal-ordered terms.

The small contribution from the residual three-body Hamiltonian is the most important result of our study. It suggests that one can neglect the residual terms of the 3NF when computing binding energies of light nuclei. This is not unexpected and has been anticipated in several earlier studies. Mihaila and Heisenberg [19] computed the charge form factor for ^{16}O within coupled-cluster theory and found a very good agreement with experimental data by considering only the density-dependent one- and two-body parts of 3NFs. Similarly, Navrátil and Ormand [59] observed in no-core shell-model calculations that density-dependent two-body terms are the most significant contributions of effective three-body forces. Our finding also support Zuker's [60] idea that monopole corrections to valence-shell interactions are due to the density-dependent terms of 3NFs. Note finally that the modeling of three-body interactions in terms of density-dependent two-body Hamiltonians has a long history, see, e.g., Ref. [61]. Note that all these examples and the present study employ sufficiently “soft” or “effective” interactions. We expect that the smallness of residual 3NFs is a property of such interactions. We will study the cutoff dependence of this finding in future work. Finally, the smallness of residual 3NFs is also encouraging for future improved nuclear matter calculations, which currently include low-momentum 3NFs through density-dependent NN interactions [51].

The smallness of the residual three-body terms is also for coupled-cluster calculations a most welcome result. This is attractive for two reasons. First, the inclusion of the residual three-nucleon Hamiltonian, as described in subsection II B, is computationally expensive. It exceeds the cost of a CCSD calculation for two-body Hamiltonians by a factor of order $O(n_u) + O(n_o^2)$ and is therefore significant for a large number of unoccupied orbitals and/or large number of nucleons. Second, the omission of the residual three-body Hamiltonian will allow us to treat 3NFs within the standard coupled-cluster theory developed for two-body Hamiltonians (after normal ordering). As a result, we can take the CCSD calculations one step further and include perturbative corrections of three-particle/three-hole clusters [62].

Let us neglect the residual 3NF terms of Eq. (3) and perform CCSD(T) calculations for the binding energy of ^4He . The approximate inclusion of three-particle/three-hole clusters

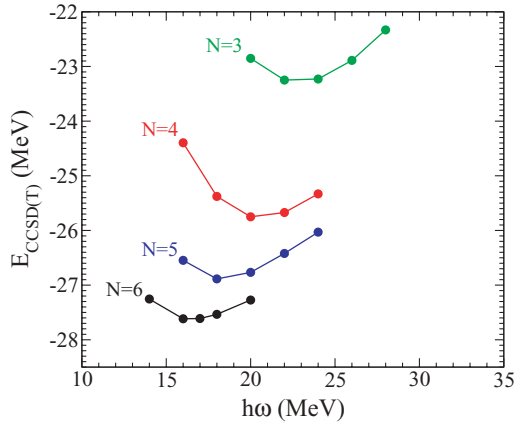


FIG. 8. (Color online) CCSD(T) results for the binding energy of ${}^4\text{He}$ as a function of the oscillator spacing and for model spaces consisting of $N = 3$ to $N = 6$ oscillator shells. The contributions from 3NFs are limited to the density-dependent zero-, one-, and two-body terms and exclude its residual three-body terms.

improves the accuracy of our calculations. Our results are shown in Fig. 8. The comparison with the CCSD full lines in Fig. 5 shows that the triples corrections add about 100–200 keV of additional binding energy (at the minimum) for fixed number N of oscillator shells and somewhat weaken the $\hbar\omega$ dependence.

An exponential extrapolation of the (approximate) CCSD(T) minima to an infinite model space is shown in Fig. 9 and yields $E_\infty = -28.24$ MeV. This is in excellent agreement with the exact Faddeev-Yakubovsky result $E = -28.20(5)$ MeV. In our largest model space at the minimum $\hbar\omega = 17$ MeV, the ground-state expectation values for the center-of-mass Hamiltonian is $\langle H_{\text{cm}} \rangle \approx 20$ keV, whereas the expectation value for the angular momentum is zero for a closed-shell nucleus by construction. These results are very

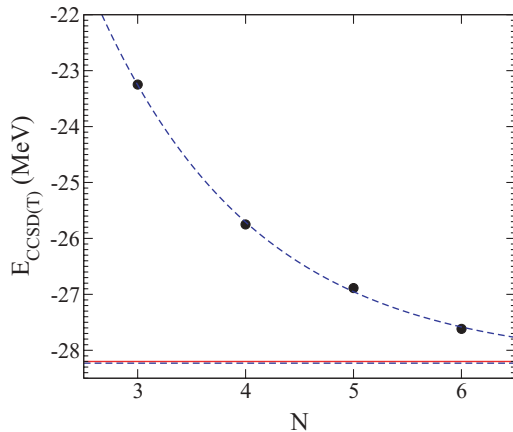


FIG. 9. (Color online) (Data points) CCSD(T) results (taken at the $\hbar\omega$ minima) for the binding energy of ${}^4\text{He}$ with 3NFs as a function of the number of oscillator shells. (Dashed lines) Exponential fit to the data and asymptote of the fit. (Full line) Exact result.

good, and it remains to be seen whether a more sophisticated treatment of triples excitations [63] would lead to further improvements. We expect that the expectation value of the center-of-mass Hamiltonian decreases with increasing size of the model space.

Finally, we also show the size of the CCSD(T) corrections in Fig. 7 as the horizontal dotted line. Clearly, these contributions are more important than the contributions from the residual 3NF terms, and this observation fully justifies the omission of the latter.

IV. SUMMARY

We have developed coupled-cluster theory for three-body Hamiltonians in the two-particle/two-hole cluster approximation (CCSD). We derived the corresponding coupled-cluster equations directly in a factorized form and thereby avoided the explicit construction and analysis of a considerable number of diagrams that enter these equations. The resulting formulas were used for a very efficient numerical implementation.

We have performed *ab initio* coupled-cluster calculations based on low-momentum NN and $3N$ interactions for the binding energy of ${}^4\text{He}$ and compared to the exact Faddeev-Yakubovsky result. The 3NF contributions to the zero-, one-, and two-body terms of the normal-ordered Hamiltonian are dominant. The contributions from residual 3NFs are smaller than the corrections due to three-particle/three-hole cluster excitations and can therefore be safely neglected. Future work will include all $3N$ partial waves and studies of the cutoff dependence and of the convergence properties using low-momentum interactions with smooth cutoffs. Our findings tremendously simplify the computational cost of coupled-cluster theory with 3NFs. This opens the avenue to explore 3NFs in medium-mass nuclei and to investigate questions related to modern nuclear interactions.

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- [1] H. Kamada, A. Nogga, W. Glöckle, E. Hiyama, M. Kamimura, K. Varga, Y. Suzuki, M. Viviani, A. Kievsky, S. Rosati, J. Carlson, S. C. Pieper, R. B. Wiringa, P. Navrátil, B. R. Barrett, N. Barnea, W. Leidemann, and G. Orlandini, *Phys. Rev. C* **64**, 044001 (2001).
- [2] S. C. Pieper and R. B. Wiringa, *Annu. Rev. Nucl. Part. Sci.* **51**, 53 (2001).
- [3] R. B. Wiringa and S. C. Pieper, *Phys. Rev. Lett.* **89**, 182501 (2002).
- [4] A. Nogga, H. Kamada, W. Glöckle, and B. R. Barrett, *Phys. Rev. C* **65**, 054003 (2002).
- [5] P. Navrátil, J. P. Vary, and B. R. Barrett, *Phys. Rev. C* **62**, 054311 (2000).
- [6] P. Navrátil and W. E. Ormand, *Phys. Rev. C* **68**, 034305 (2003).
- [7] A. Nogga, P. Navrátil, B. R. Barrett, and J. P. Vary, *Phys. Rev. C* **73**, 064002 (2006).
- [8] G. P. Lepage, lectures delivered at the 9th Jorge Andre Swieca Summer School: Particles and Fields, São Paulo, Brazil, February, 1997, nucl-th/9706029.
- [9] F. Coester, *Nucl. Phys.* **7**, 421 (1958).
- [10] F. Coester and H. Kümmel, *Nucl. Phys.* **17**, 477 (1960).
- [11] J. Cizek, *J. Chem. Phys.* **45**, 4256 (1966).
- [12] J. Cizek, *Adv. Chem. Phys.* **14**, 35 (1969).
- [13] H. Kümmel, K. H. Lührmann, and J. G. Zabolitzky, *Phys. Rep.* **36**, 1 (1978).
- [14] R. J. Bartlett, *J. Phys. Chem.* **93**, 1697 (1989).
- [15] J. Paldus and X. Li, *Adv. Chem. Phys.* **110**, 1 (1999).
- [16] T. D. Crawford and H. F. Schaefer III, *Rev. Comput. Chem.* **14**, 33 (2000).
- [17] P. Piecuch, K. Kowalski, I. S. O. Pimienta, P.-D. Fan, M. Lodriguito, M. J. McGuire, S. A. Kucharski, T. Kus, and M. Musiał, *Theor. Chem. Acc.* **112**, 349 (2004).
- [18] J. H. Heisenberg and B. Mihaila, *Phys. Rev. C* **59**, 1440 (1999).
- [19] B. Mihaila and J. H. Heisenberg, *Phys. Rev. C* **61**, 054309 (2000); *Phys. Rev. Lett.* **84**, 1403 (2000).
- [20] D. J. Dean and M. Hjorth-Jensen, *Phys. Rev. C* **69**, 054320 (2004).
- [21] K. Kowalski, D. J. Dean, M. Hjorth-Jensen, T. Papenbrock, and P. Piecuch, *Phys. Rev. Lett.* **92**, 132501 (2004).
- [22] M. Włoch, D. J. Dean, J. R. Gour, M. Hjorth-Jensen, K. Kowalski, T. Papenbrock, and P. Piecuch, *Phys. Rev. Lett.* **94**, 212501 (2005).
- [23] J. R. Gour, P. Piecuch, M. Hjorth-Jensen, M. Włoch, and D. J. Dean, *Phys. Rev. C* **74**, 024310 (2006).
- [24] G. Hagen, D. J. Dean, M. Hjorth-Jensen, and T. Papenbrock, nucl-th/0610072.
- [25] F. Wegner, *Ann. Phys. (Leipzig)* **3**, 77 (1994).
- [26] S. D. Glazek and K. G. Wilson, *Phys. Rev. D* **49**, 4214 (1994).
- [27] S. R. White, *J. Chem. Phys.* **117**, 7472 (2002).
- [28] S. A. Coon, M. D. Scadron, P. C. McNamee, B. R. Barrett, D. W. E. Blatt, and B. H. J. McKellar, *Nucl. Phys.* **A317**, 242 (1979).
- [29] S. A. Coon and M. T. Pena, *Phys. Rev. C* **48**, 2559 (1993).
- [30] J. Fujita and H. Miyazawa, *Prog. Theor. Phys.* **17**, 360 (1957).
- [31] M. R. Robilotta and H. T. Coelho, *Nucl. Phys.* **A460**, 645 (1986).
- [32] J. A. Eden and M. F. Gari, *Phys. Rev. C* **53**, 1510 (1996).
- [33] B. S. Pudliner, V. R. Pandharipande, J. Carlson, S. C. Pieper, and R. B. Wiringa, *Phys. Rev. C* **56**, 1720 (1997).
- [34] S. C. Pieper, V. R. Pandharipande, R. B. Wiringa, and J. Carlson, *Phys. Rev. C* **64**, 014001 (2001).
- [35] P. F. Bedaque and U. van Kolck, *Annu. Rev. Nucl. Part. Sci.* **52**, 339 (2002).
- [36] E. Epelbaum, *Prog. Part. Nucl. Phys.* **57**, 654 (2006).
- [37] D. R. Entem and R. Machleidt, *Phys. Rev. C* **68**, 041001(R) (2003).
- [38] E. Epelbaum, W. Glöckle, and U. G. Meißner, *Nucl. Phys.* **A747**, 362 (2005).
- [39] S. Weinberg, *Phys. Lett.* **B251**, 288 (1990); *Nucl. Phys.* **B363**, 3 (1991).
- [40] U. van Kolck, *Phys. Rev. C* **49**, 2932 (1994).
- [41] J. L. Friar, D. Hüber, and U. van Kolck, *Phys. Rev. C* **59**, 53 (1999).
- [42] E. Epelbaum, A. Nogga, W. Glöckle, H. Kamada, U. G. Meißner, and H. Witala, *Phys. Rev. C* **66**, 064001 (2002).
- [43] C. Ordonez, L. Ray, and U. van Kolck, *Phys. Rev. C* **53**, 2086 (1996).
- [44] E. Epelbaum, W. Glöckle, and U. G. Meißner, *Nucl. Phys.* **A671**, 295 (2000).
- [45] P. Navrátil and E. Caurier, *Phys. Rev. C* **69**, 014311 (2004).
- [46] C. Forssen, P. Navrátil, W. E. Ormand, and E. Caurier, *Phys. Rev. C* **71**, 044312 (2005).
- [47] P. Navrátil, V. G. Gueorguiev, J. P. Vary, W. E. Ormand, and A. Nogga, nucl-th/0701038.
- [48] S. K. Bogner, T. T. S. Kuo, A. Schwenk, D. R. Entem, and R. Machleidt, *Phys. Lett.* **B576**, 265 (2003).
- [49] S. K. Bogner, T. T. S. Kuo, and A. Schwenk, *Phys. Rep.* **386**, 1 (2003).
- [50] A. Nogga, S. K. Bogner, and A. Schwenk, *Phys. Rev. C* **70**, 061002(R) (2004).
- [51] S. K. Bogner, A. Schwenk, R. J. Furnstahl, and A. Nogga, *Nucl. Phys.* **A763**, 59 (2005).
- [52] A. Schwenk, *J. Phys. G* **31**, S1273 (2005).
- [53] S. A. Kucharski and R. J. Bartlett, *Adv. Quantum Chem.* **18**, 281 (1986).
- [54] S. A. Kucharski and R. J. Bartlett, *Theor. Chim. Acta* **80**, 387 (1991).
- [55] S. A. Kucharski and R. J. Bartlett, *J. Chem. Phys.* **97**, 4282 (1992).
- [56] P. Piecuch, S. A. Kucharski, K. Kowalski, and M. Musiał, *Comput. Phys. Commun.* **149**, 71 (2002).
- [57] R. B. Wiringa, V. G. J. Stoks, and R. Schiavilla, *Phys. Rev. C* **51**, 38 (1995).
- [58] S. K. Bogner, R. J. Furnstahl, S. Ramanan, and A. Schwenk, *Nucl. Phys.* **A784**, 79 (2007).
- [59] P. Navrátil and W. E. Ormand, *Phys. Rev. Lett.* **88**, 152502 (2002).
- [60] A. P. Zuker, *Phys. Rev. Lett.* **90**, 042502 (2003).
- [61] J. W. Negele, *Rev. Mod. Phys.* **54**, 913 (1982).
- [62] M. J. O. Deegan and P. J. Knowles, *Chem. Phys. Lett.* **227**, 321 (1994).
- [63] P. Piecuch and M. Włoch, *J. Chem. Phys.* **123**, 224105 (2005).