

IMSRG with flowing 3N operators

some first explorations

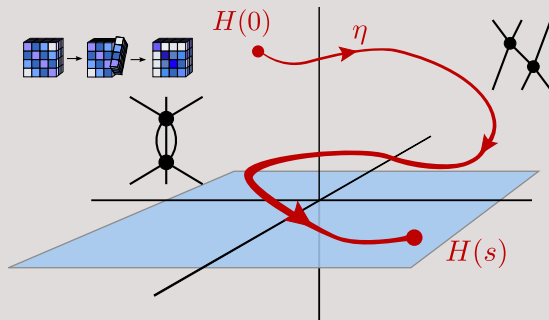
Ragnar Stroberg

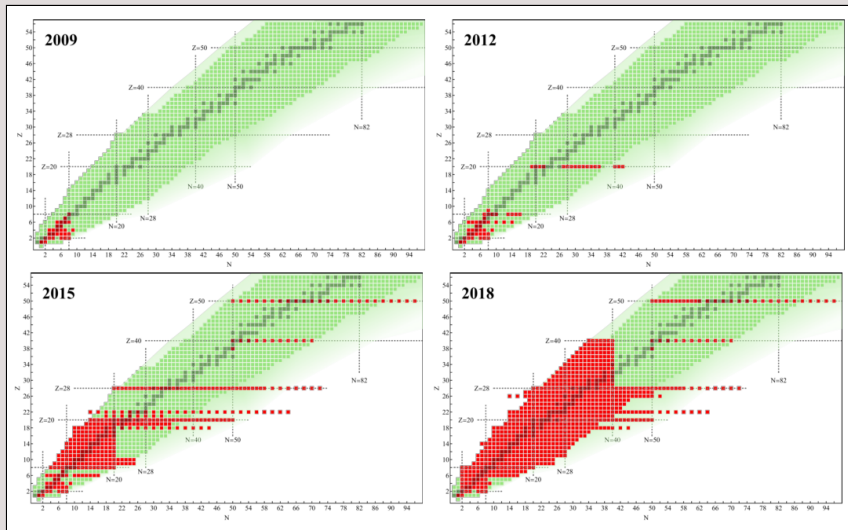
University of Washington

Progress in Ab initio Techniques for Nuclear Physics

TRIUMF Vancouver, BC

March 5, 2020





Hebeler 2020; figure credit Heiko Hergert

$$H(s) = U(s)HU^\dagger(s)$$

$$\frac{dU}{ds} = \eta(s)U(s)$$

$$\frac{dH}{ds} = [\eta(s), H(s)]$$

Flow equations:

$$H(s) = H^d(s) + H^{od}(s)$$

$$H^{od}(s) \rightarrow 0 \text{ as } s \rightarrow \infty$$

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$$H^{od}(s) \rightarrow 0 \text{ as } s \rightarrow \infty$$

$$H = E_0 + \sum_{ij} f_{ij} \{a_i^\dagger a_j\} + \frac{1}{4} \sum_{ijkl} \Gamma_{ijkl} \{a_i^\dagger a_j^\dagger a_l a_k\} \\ + \frac{1}{36} \sum_{ijklmn} W_{ijklmn} \{a_i^\dagger a_j^\dagger a_k^\dagger a_n a_m a_l\}$$

$$\text{Flow equations: } \frac{dE_0}{ds} = [\eta(s), H(s)]_{0b}$$

$$\frac{df}{ds} = [\eta(s), H(s)]_{1b}$$

$$\frac{d\Gamma}{ds} = [\eta(s), H(s)]_{2b}$$

$$\frac{dW}{ds} = [\eta(s), H(s)]_{3b}$$

$$H(s) = U(s)H U^\dagger(s)$$

$$\frac{dU}{ds} = \eta(s)U(s)$$

$$\frac{dH}{ds} = [\eta(s), H(s)]$$

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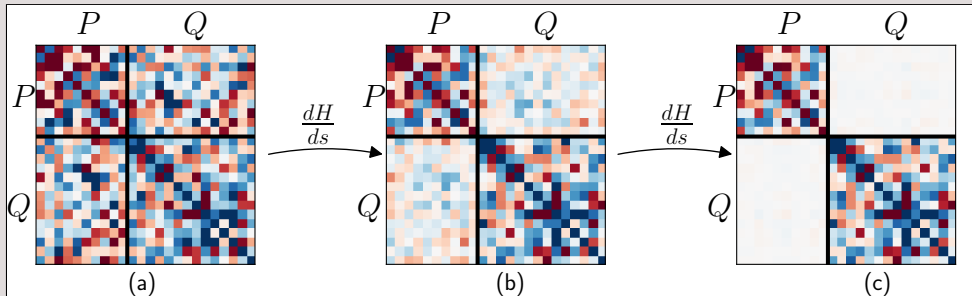
Flow equations: $\frac{dE_0}{ds} = [\eta(s), H(s)]_{0b}$

$$\frac{df}{ds} = [\eta(s), H(s)]_{1b}$$

$$\frac{d\Gamma}{ds} = [\eta(s), H(s)]_{2b}$$

$$\frac{dW}{ds} = [\eta(s), H(s)]_{3b}$$

IMSRG(2)



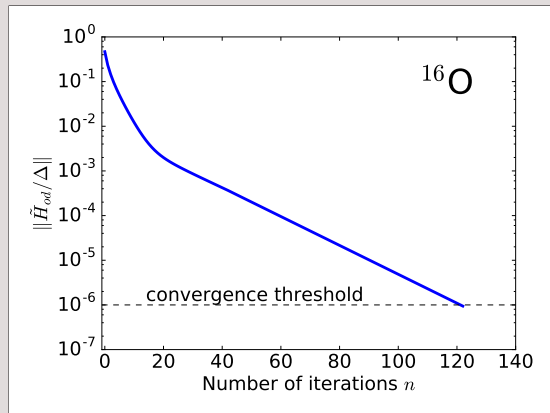
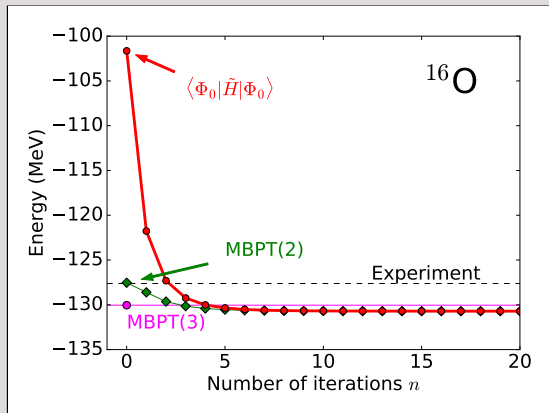
As $s \rightarrow \infty$, P space is decoupled from Q space.

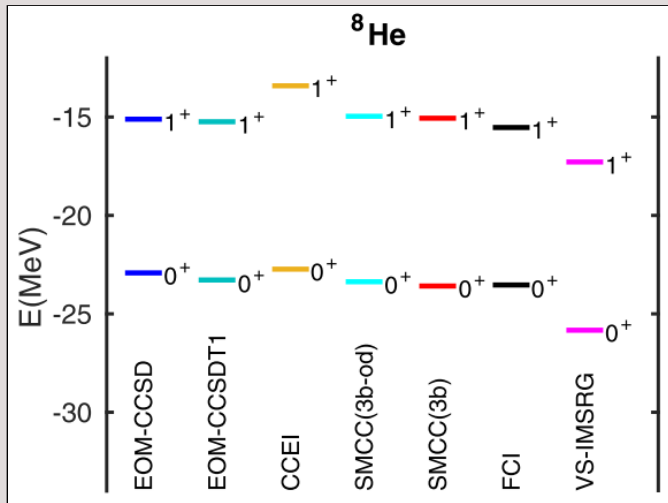
Output is a shell model effective interaction
→ diagonalize with NuShellX, Antoine, Kshell etc.

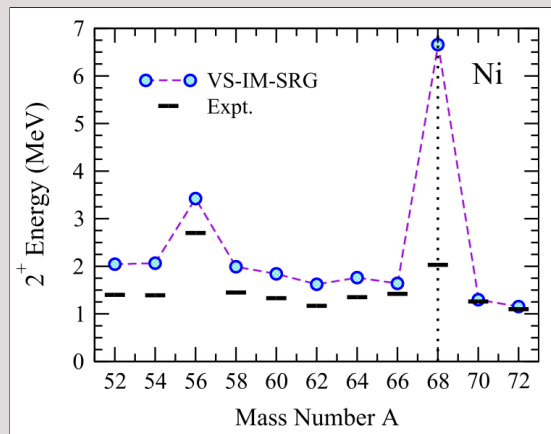
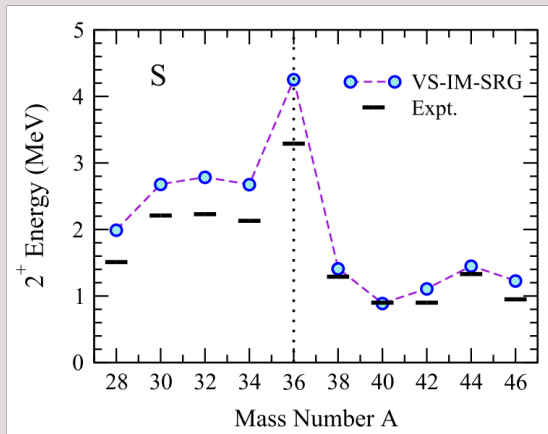
$$U(s) = e^{\Omega(s)} \quad , \text{ Magnus operator } \Omega$$

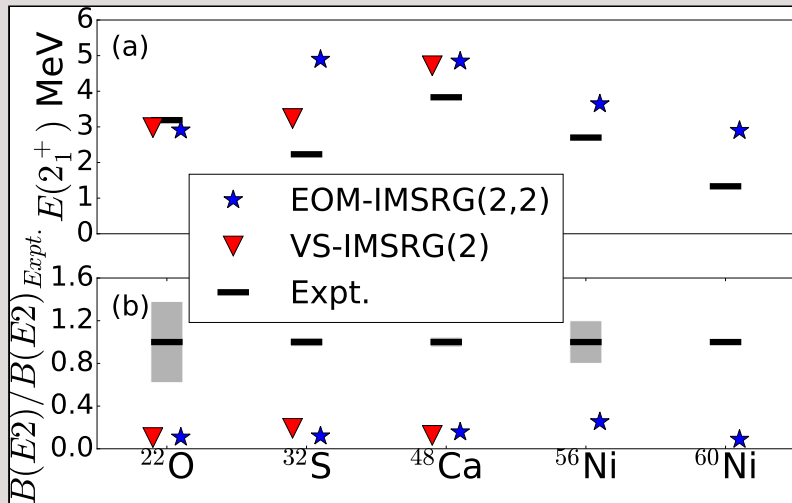
$$\begin{aligned} H(s) &= e^{\Omega(s)} H e^{-\Omega(s)} \\ &= H + [\Omega, H] = \frac{1}{2}[\Omega, [\Omega, H]] + \dots \\ &= \sum_{k=0}^{\infty} \frac{1}{k!} [\Omega, H]^k \end{aligned}$$

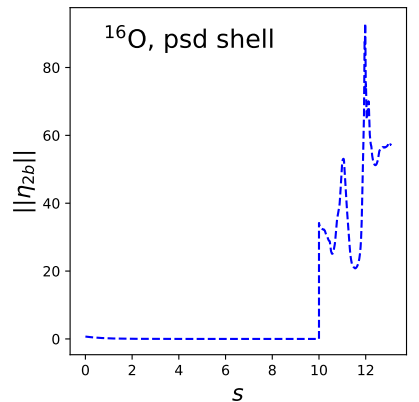
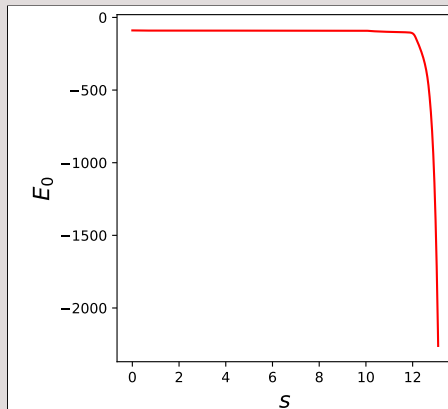
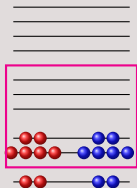
$$\text{Flow eq. for } \Omega: \frac{d\Omega}{ds} = \sum_{k=0}^{\infty} \frac{B_k}{k!} [\Omega, \eta]^k$$

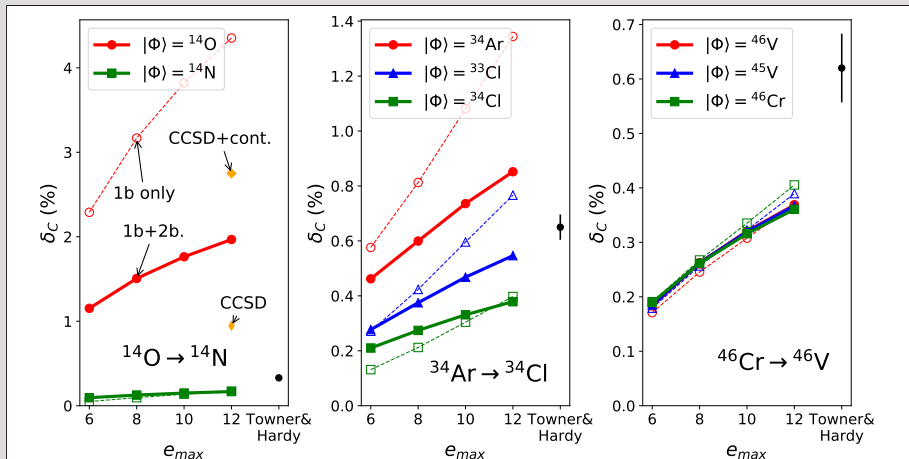




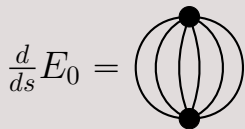






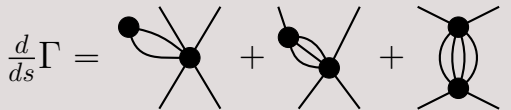


Gaute Hagen, priv. comm.



$[3, 3] \rightarrow 0$

$\sim n^6$



$[1, 3] \rightarrow 2$

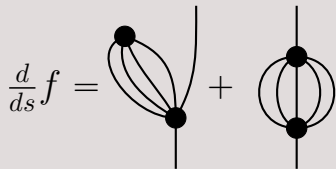
$\sim n^6$

$[2, 3] \rightarrow 2$

$\sim n^7$

$[3, 3] \rightarrow 2$

$\sim n^8$

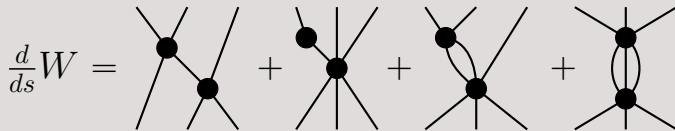


$[2, 3] \rightarrow 1$

$\sim n^6$

$[3, 3] \rightarrow 1$

$\sim n^7$



$[2, 2] \rightarrow 3$

$\sim n^7$

$[1, 3] \rightarrow 3$

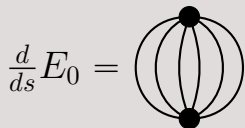
$\sim n^7$

$[2, 3] \rightarrow 3$

$\sim n^8$

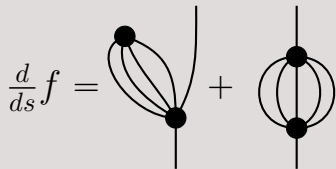
$[3, 3] \rightarrow 3$

$\sim n^9$



$$[3, 3] \rightarrow 0$$

$$\sim n^6$$

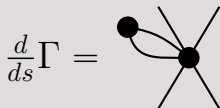


$$[2, 3] \rightarrow 1$$

$$\sim n^6$$

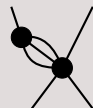
$$[3, 3] \rightarrow 1$$

$$\sim n^7$$



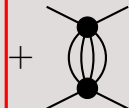
$$[1, 3] \rightarrow 2$$

$$\sim n^6$$



$$[2, 3] \rightarrow 2$$

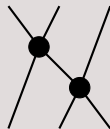
$$\sim n^7$$



$$[3, 3] \rightarrow 2$$

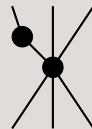
$$\sim n^8$$

$$\frac{d}{ds}W =$$



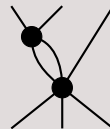
$$[2, 2] \rightarrow 3$$

$$\sim n^7$$



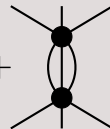
$$[1, 3] \rightarrow 3$$

$$\sim n^7$$



$$[2, 3] \rightarrow 3$$

$$\sim n^8$$



$$[3, 3] \rightarrow 3$$

$$\sim n^9$$

$$\frac{d}{ds}E_0 =$$

$$[3,3] \rightarrow 0$$

$$\sim n^6$$

$$\frac{d}{ds}\Gamma =$$

$$[2,3] \rightarrow 2$$

$$\sim n^7$$

$$[3,3] \rightarrow 2$$

$$\sim n^8$$

$$\frac{d}{ds}f =$$

$$[2,3] \rightarrow 1$$

$$\sim n^6$$

$$[3,3] \rightarrow 1$$

$$\sim n^7$$

Error, keeping only n^7 terms:

- $E_0 \sim 5\text{th order}$
- $f \sim 4\text{th order}$
- $\Gamma \sim 4\text{th order}$
- $W \sim 3\text{rd order}$

$$\frac{d}{ds}W =$$

$$[2,2] \rightarrow 3$$

$$\sim n^7$$

$$[1,3] \rightarrow 3$$

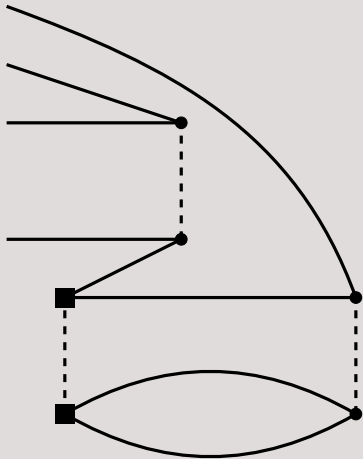
$$\sim n^7$$

$$[2,3] \rightarrow 3$$

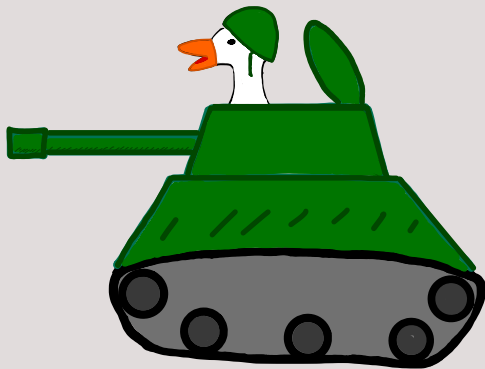
$$\sim n^8$$

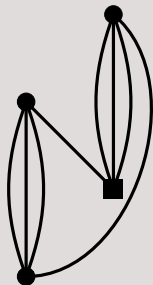
$$[3,3] \rightarrow 3$$

$$\sim n^9$$

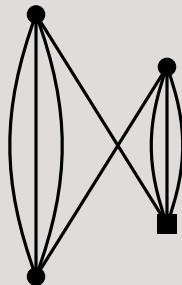


$$[\Omega, [\Omega, H]_{3b}]_{2b}$$



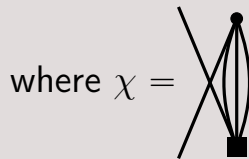


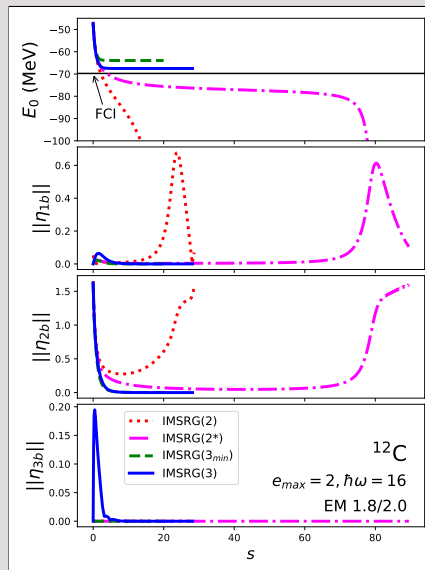
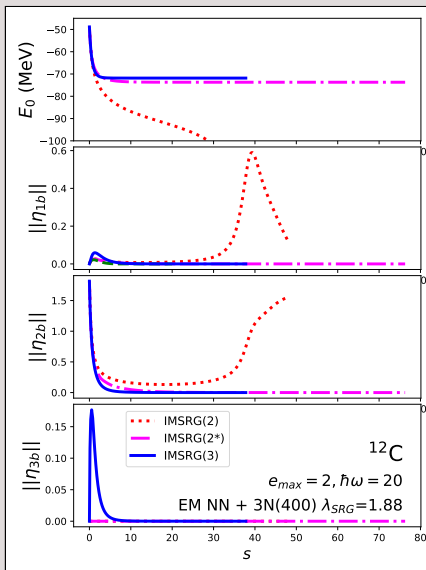
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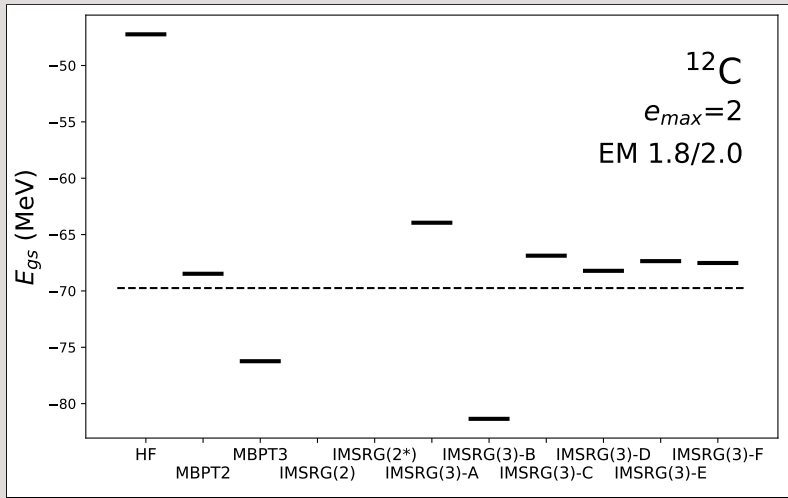


Included

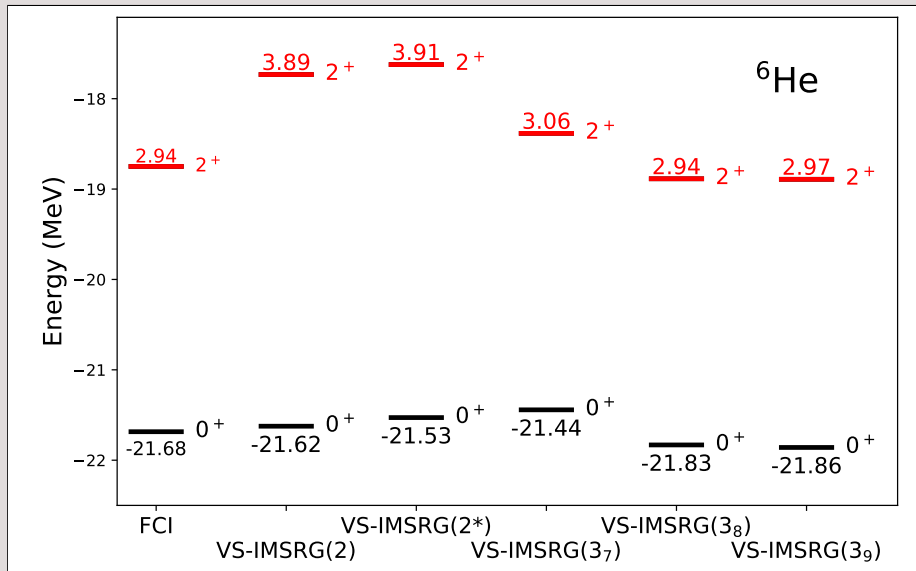
Remedy: $[\Omega, [\Omega, H]] \rightarrow [\Omega, [\Omega, H]] + [\Omega, \chi]$
 [IMSRG(2*) approx]

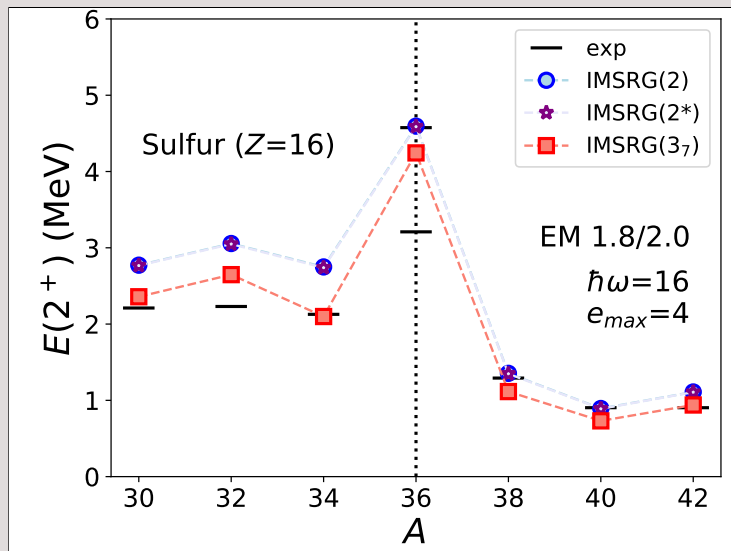






- A: $[2, 2] \rightarrow 3 + [2, 3] \rightarrow 2, \eta_{2b}, \Omega_{2b}$
- B: $A + \eta_{3b}, \Omega_{3b}$
- C: $B + [1, 3] \rightarrow 3$
- D: $C + [3, 3] \rightarrow 0$
- E: $D + [2, 3] \rightarrow 1$
- F: $E + [1, 3] \rightarrow 2$



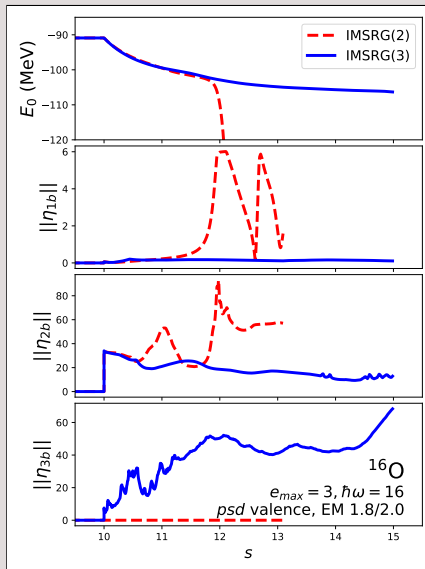
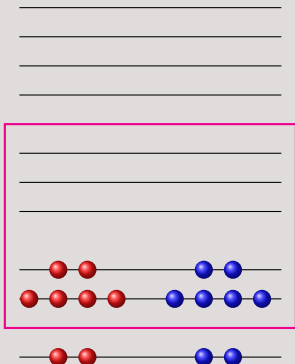


IMSRG(37) calculations made with cut

$$\tilde{e}_a + \tilde{e}_b + \tilde{e}_c \leq \tilde{E}_{max} = 2,$$

$$\tilde{e}_a \equiv 2n_a + \ell_a - \epsilon_F$$

Multi-shell valence spaces and the intruder problem



Summary

- 3-body contributions can help stabilize IMSRG flow (e.g. ^{12}C), and improve excitation energies
- Multi-shell decoupling leads to rapid growth of 3-body terms and loss of cluster hierarchy. Keeping 3 body terms is no help.
- Main effects appear to come from n^7 scaling terms
- To do:
 - Optimize, factorize nested commutator terms to n^6
 - How best to use IMSRG(3) to assess uncertainty?

Thanks to Titus Morris, Takayuki Miyagi, Julien Ripoche, Alex Tichai, and Roland Wirth for helpful discussions.

Additional figures

- Hebeler, Kai (2020). "Three-Nucleon Forces: Implementation and Applications to Atomic Nuclei and Dense Matter". In: arXiv: 2002.09548. URL: <http://arxiv.org/abs/2002.09548>.
- Parzuchowski, N. M. et al. (2017). "Ab initio electromagnetic observables with the in-medium similarity renormalization group". In: *Phys. Rev. C* 96.3, p. 034324. ISSN: 2469-9985. DOI: 10.1103/PhysRevC.96.034324. arXiv: 1705.05511. URL: <https://link.aps.org/doi/10.1103/PhysRevC.96.034324>.
- Simonis, J. et al. (2017). "Saturation with chiral interactions and consequences for finite nuclei". In: *Phys. Rev. C* 96.1, p. 014303. ISSN: 2469-9985. DOI: 10.1103/PhysRevC.96.014303. arXiv: 1704.02915. URL: <http://arxiv.org/abs/1704.02915><http://link.aps.org/doi/10.1103/PhysRevC.96.014303>.
- Stroberg, S. R. et al. (2019). "Non-Empirical Interactions for the Nuclear Shell Model: An Update". In: *Annu. Rev. Nucl. Part. Sci.* 69.1, pp. 307–362. ISSN: 0163-8998. DOI: 10.1146/annurev-nucl-101917-021120. arXiv: 1902.06154. URL: <http://arxiv.org/abs/1902.06154>.
- Sun, Z. H. et al. (2018). "Shell-model coupled-cluster method for open-shell nuclei". In: *Phys. Rev. C* 98.5, p. 054320. ISSN: 2469-9985. DOI: 10.1103/PhysRevC.98.054320. arXiv: 1806.07405. URL: <http://arxiv.org/abs/1806.07405><https://link.aps.org/doi/10.1103/PhysRevC.98.054320>.