Planetary waves: a numerical study of the wave equation

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December 11, 2018

Abstract

In this article, we study Rossby waves using different both implicit and explicit schemes for solving the wave equation numerically. We consider both the advantages and disadvantages of both methods, as well as their efficiency and accuracy. From this analysis, we find that

1 Introduction

We encounter wave phenomena everywhere in the natural sciences. From quantum mechanics to oceanography, we find that be it the motion of a particle or the ocean, we require knowledge of wave-like behaviour to solve the problem. In quantum mechanics, a particle's wave function is described by a complex-valued diffusion equation, the Schrödinger equation, while in oceanography, we can describe ocean waves using the wave equation,

$$\frac{\partial^2 u}{\partial x^2} = \frac{\partial^2 u}{\partial t^2}, \tag{1}$$

where x and t denote the spatial and temporal coordinates, respectively. The wave equation will be the topic of this paper, in particular, we will model Rossby waves, first identified by Rossby (1939). These are inertial, planetary waves in the Earth's atmosphere and ocean which motions contribute to extreme weather (Mann et al., 2017), might drive the El-Ninõ southern oscillation (ENSO) (Bosc and Delcroix, 2008), and is also produced by ENSO, see Battisti (1989). While we hope the reader appreciate the wide range of phenomena related to these waves, our article is a numerical study of the waves isolated from other processes. We therefore begin by describing fundamental theory of waves and partial differential equations in the Theory section, present our algorithm and the technicalities relating to its implementation. In the Results section, we present our data as figures, before discussing their implications in the Discussion section. Concluding our paper, we present our final thoughts on the topic of simulating Rossby waves.

2 Theory

2.1 Partial differential equations

Partial differential equations can often be separated into a set of coupled ordinary differential equations which are easier to solve.

$\mathbf{n}$	$\mathbf{t_g/t_s}$	$\mathbf{t_{LU}/t_s}$
10	2.08	3.70
10^2	1.89	$1.00 \cdot 10^2$
10^3	1.48	$1.05 \cdot 10^4$
10^4	1.43	$1.18 \cdot 10^6$
10^5	1.39	-
10^6	1.41	-
10^7	1.39	-

Table 1: Ratio between CPU time for the general algorithm ($\mathbf{t_g}$), the special algorithm ($\mathbf{t_g}$) and the LU decomposition algorithm ($\mathbf{t_{LU}}$) for different matrix sizes ($\mathbf{n}$). The LU decomposition crashed for $\mathbf{n}$ greater than 10^4.

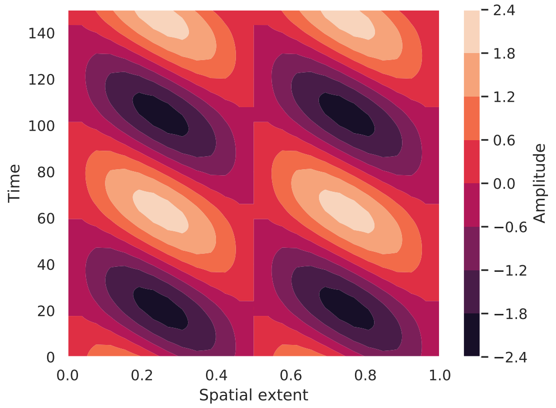

Figure 1: Hov Müller diagram of a bounded Rossby wave with a initial sine wave using a explicit scheme.

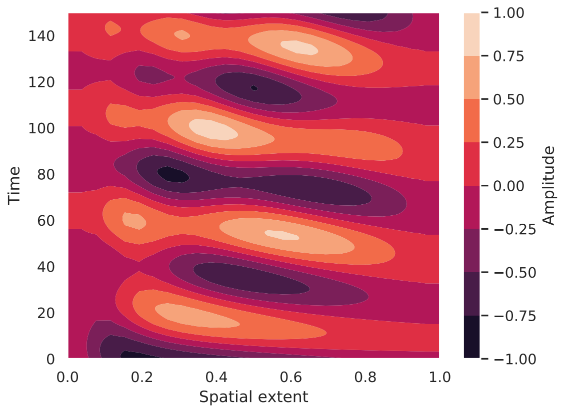

Figure 2: Hov Müller diagram of a bounded Rossby wave initially a gaussian using a explicit scheme.

2.2 Wave theory

2.3 Implementation

3 Results

4 Discussion

5 Conclusion

References

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A

$$A = \begin{bmatrix} b_1 & c_1 & 0 & \dots & \dots & 0 \\ a_1 & b_2 & c_2 & 0 & \dots & 0 \\ 0 & a_2 & b_3 & c_3 & \dots & 0 \\ \vdots & \ddots & \ddots & \ddots & \ddots & \vdots \\ 0 & \dots & \ddots & a_{n-2} & b_{n-1} & c_{n-1} \\ 0 & \dots & \dots & 0 & a_{n-1} & b_n \end{bmatrix},$$