



ZTF-FCT
Zientzia eta Teknologia Fakultatea
Facultad de Ciencia y Tecnología



CONSTANT MEAN CURVATURE INVARIANT SURFACES IN \mathbb{L}^3 AND A BLASCHKE'S VARIATIONAL PROBLEM

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in Minkowski Space

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OBJECTIVE 1 (EXTENDED BLASCHKE'S PROBLEM)

Completely **solve** an extended Blaschke's Variational problem in the Minkowski n-space, \mathbb{L}^n .

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- Get a **geodesic foliation** of S .
- Leaves are **critical for the Blaschke's problem**.
- Finally, study **isometric deformations**.

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5. Planar Critical Curves

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- Take into account that $\kappa = \mu$ would be a global minimum if we were considering $L^1([0, L])$ as the space of curves.
- Observe that the case $\mu = 0$ was studied by Blaschke in the Euclidean 3-Space, [2].

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A critical point of Θ must lie in a 3-dimensional totally geodesic submanifold of \mathbb{L}^n .

Thus, we are interested in studying critical curves in the Minkowski 3-Space, \mathbb{L}^3 .

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The Euler-Lagrange equations for the curvature energy functional $\Theta(\gamma) = \int_{\gamma} \sqrt{\kappa - \mu}$, in \mathbb{L}^3 can be written as

$$\frac{d^2}{ds^2} \left(\frac{\varepsilon_2}{\sqrt{\kappa - \mu}} \right) + \frac{1}{\sqrt{\kappa - \mu}} (\varepsilon_1 \kappa^2 - \varepsilon_3 \tau^2) = 2\varepsilon_1 \kappa \sqrt{\kappa - \mu},$$

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Under suitable boundary conditions, solutions of these equations are critical curves for our energy functional.

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- Thus, γ must be a Frenet helix.
- Moreover, substituting this in the first Euler-Lagrange equation we get the relation

$$\kappa_o = \mu + \sqrt{\mu^2 - \varepsilon_1 \varepsilon_3 \tau_o^2}.$$

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$$c = -\varepsilon_1 \varepsilon_2 - \varepsilon_2 \varepsilon_3 e^2,$$

and $\Delta = 4ac - b^2 = -16d^2 - 16\varepsilon_1 \mu d + 4\varepsilon_1 \varepsilon_3 \mu^2 e^2$, where d, e are real constants (constants of integration).

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Thus, the following cases are not possible:

1. $\Delta \geq 0$ and $c < 0$,
2. $a \leq 0$, $2d = -\varepsilon_1 \mu$ and $e^2 = -\varepsilon_1 \varepsilon_3$.

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1. If $\Delta \neq 0$ and $a \neq 0$

$$\kappa(s) = \frac{2a + \mu(b + \sqrt{|\Delta|}f(2\mu s))}{-b + \sqrt{|\Delta|}f(2\mu s)},$$

where, $f(x) = \sinh x$, if $\Delta > 0$ and $a > 0$; $f(x) = \cosh x$, if $\Delta < 0$ and $a > 0$; and $f(x) = \sin x$, if $\Delta < 0$ and $a < 0$.

SOLUTIONS WITH NON-CONSTANT CURVATURE

2. If $\Delta = 0$ and $a > 0$

$$k(s) = \frac{\mu + (2a - b\mu) \exp 2\mu s}{1 - b \exp 2\mu s}.$$

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$$k(s) = \frac{1}{2\sqrt{cs}}.$$

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The **curvature** and **torsion** (together with the causal characters ε ; of the Frenet frame) determine a **unique curve** up to rigid motions.

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- When $\mu = 0$, our critical curves are **Lancret curves**, that is, curves making a constant angle with a fixed direction.
- On the other hand, if $\mu \neq 0$, our critical curves are **Bertrand curves**.

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DELAUNAY CURVES ([1], [3])

Critical curves of $\Theta(\gamma) = \int_{\gamma} \sqrt{\kappa - \mu} ds$ in \mathbb{R}^2 are precisely the Delaunay curves, that is, the roulettes of foci of conics (lines, circles, catenaries, nodaries and undularies).

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If γ lies in a Lorentzian plane, $\gamma \subset \mathbb{L}^2$, we can prove

HANO-NOMIZU CURVES ([1], [5])

The locus of the origin when a part of a spacelike quadratic curve is rolled along a spacelike line is a spacelike critical curve for $\Theta(\gamma) = \int_{\gamma} \sqrt{\kappa - \mu} ds$ in \mathbb{L}^2 .

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ASSOCIATED KILLING VECTOR FIELD ALONG γ ([3])

The vector field $\mathcal{I} = \frac{1}{2\sqrt{\kappa-\mu}} B$ is a Killing vector field along γ , if and only if, γ verifies the Euler-Lagrange equations (γ may have a restriction on its length).

EVOLUTION OF CRITICAL CURVES

Since \mathbb{L}^3 is complete,

1. Consider the one-parameter group of isometries determined by the flow of

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FOLIATIONS OF INVARIANT SURFACES

For the converse, assume that S is a non-degenerate \mathcal{G}_ξ -invariant surface of \mathbb{L}^3 , i.e., for any $x \in S$ and $\Phi_t \in \mathcal{G}_\xi$ we have $\Phi_t(S) = S$.

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GEODESIC FOLIATION BY CRITICAL CURVES ([1])

A ξ -invariant CMC surface S of \mathbb{L}^3 admits a local geodesic parametrization where the leaves provide a geodesic foliation by critical curves of the extended Blaschke's problem with $\mu = -\varepsilon_1 \varepsilon_2 H$.

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Thus, every ξ -invariant CMC surface is

- A ruled surface or,
- It is generated by evolving a critical curve of $\Theta(\gamma) = \int_\gamma \sqrt{\kappa + \varepsilon_1\varepsilon_2 H} ds$ under the flow of the Killing vector field ξ .

ISOMETRIC DEFORMATIONS

1. Deformations by Isometric Surfaces

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2. Orbits of Deformations

DEFORMATIONS BY ISOMETRIC SURFACES (1)

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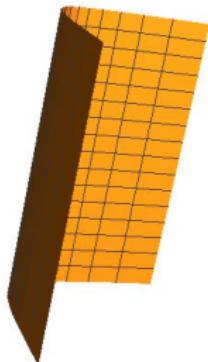
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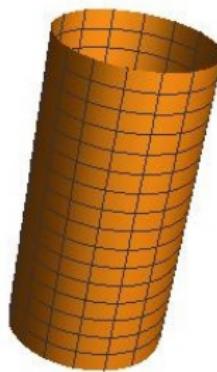
Suppose that the critical curve γ of $\Theta(\gamma) = \int_{\gamma} \sqrt{\kappa - \mu} ds$ has constant curvature κ_o , then we know that γ is a **Frenet helix** with

$$\begin{aligned}\kappa &= \kappa_o = \mu + \sqrt{\mu^2 - \varepsilon_2 \varepsilon_3 \tau_o^2}, \\ \tau &= \tau_o,\end{aligned}$$

and these helices can only generate **congruence surfaces** to the following ones (depending on the causal character of ε_i)



(E) $\varepsilon_1 \varepsilon_2 = -1$



(F) $\varepsilon_3 = -1$

DEFORMATIONS BY ISOMETRIC SURFACES (2)

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4. If $\Delta = a = 0$, (as $d = 0$, there is **no biparametric family**)
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ISOMETRIC DEFORMATIONS ([1])

For each real constant μ , let $\{S_\gamma\}_{d,e}$ be the family of ξ -invariant surfaces shaped on a critical curve γ of

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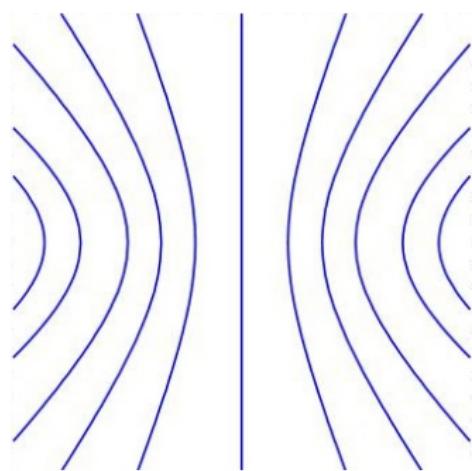
$\Theta(\gamma) = \int_\gamma \sqrt{\kappa - \mu} ds$. Under the **relations** above (except for case 2), the family $\{S_\gamma\}_d$ is generated by **isometric surfaces** with the same constant mean curvature $H = -\varepsilon_1 \varepsilon_2 \mu$.

DEFORMATIONS OF RIEMANNIAN CMC SURFACES

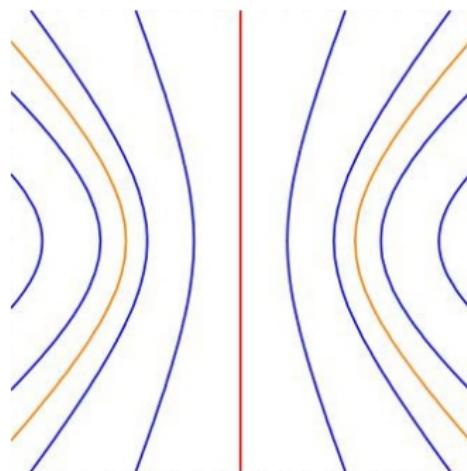
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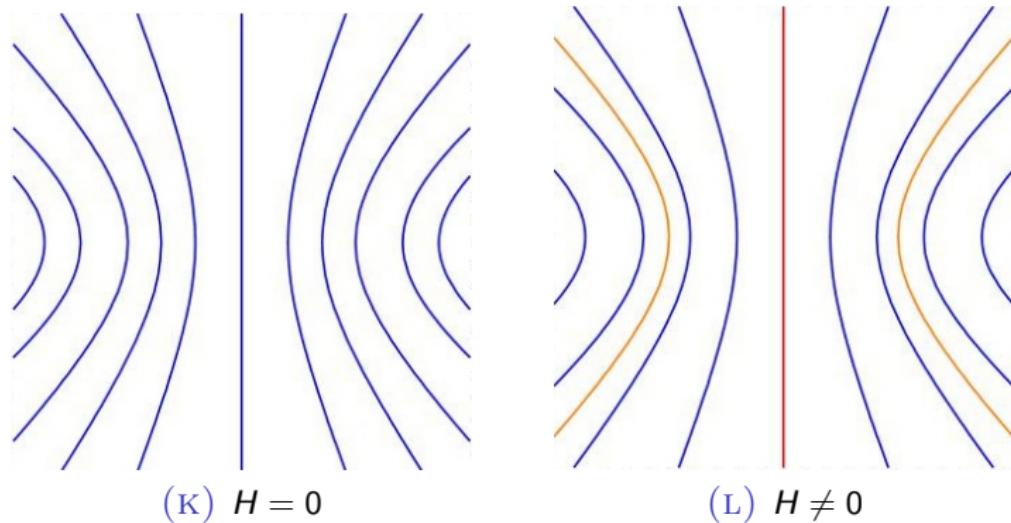
(I) $H = 0$



(J) $H \neq 0$

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ROTATIONAL SPACELIKE SURFACES ([1])

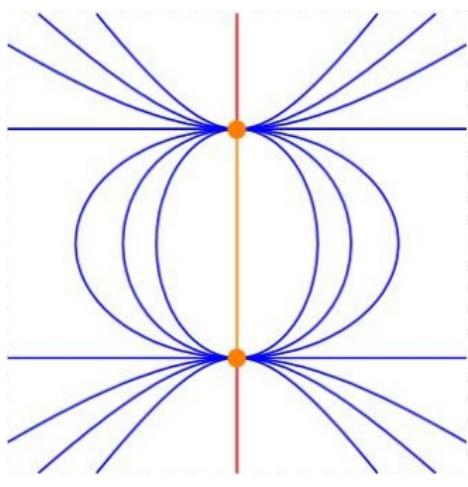
Any spacelike surface of CMC can be isometrically deformed into a spacelike rotational CMC surface, except for the yellow cases.

DEFORMATIONS OF LORENTZIAN SURFACES

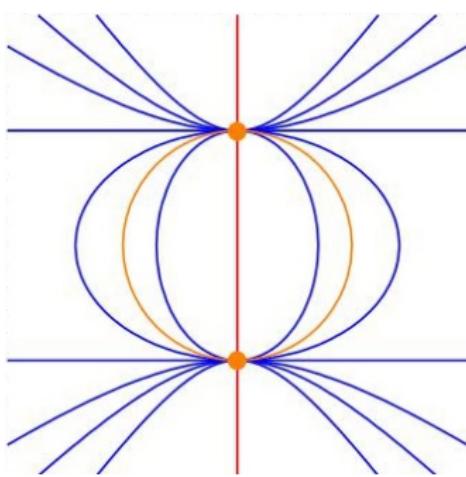
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DEFORMATIONS OF LORENTZIAN SURFACES

Let S_γ be a Lorentzian surface of \mathbb{L}^3 with timelike profile curve γ . Then, we have the following orbits of the isometric deformations



(o) $H = 0$



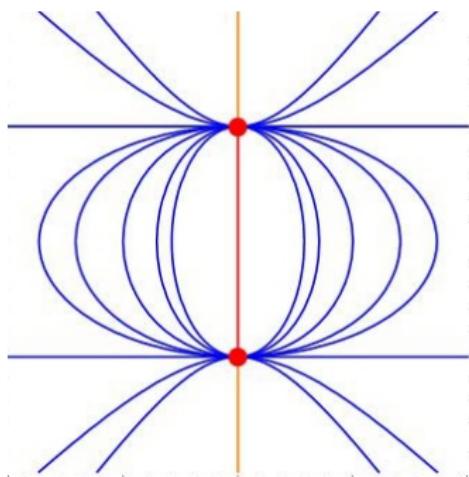
(P) $H \neq 0$

DEFORMATIONS OF LORENTZIAN SURFACES

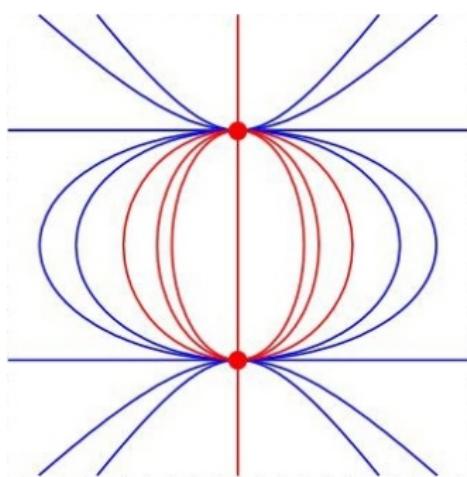
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DEFORMATIONS OF LORENTZIAN SURFACES

Now, if γ is a **spacelike profile curve** of a **Lorentzian surface S_γ** of \mathbb{L}^3 , the **isometric deformations** appear in the following diagrams



(s) $H = 0$

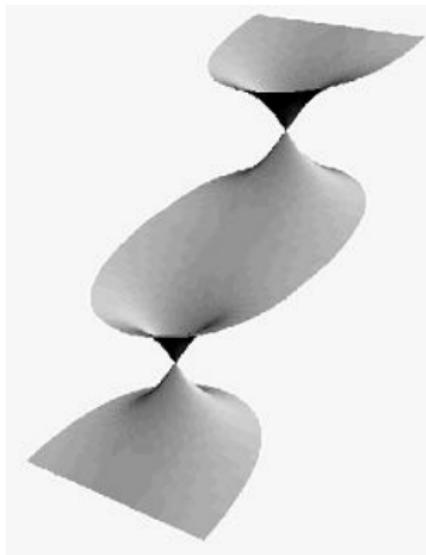


(t) $H \neq 0$

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THE END



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