GRAVITY RELATED DERIVATIONS

$\begin{array}{c} Stem \ Astronomy \ Club \\ Academic \ committee \\ Academic \ 23 \\ I^{A}T_{E}X \end{array}$

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1 Symbols meaning

F is force

m is mass

a is acceleration

v is velocity

t is one cycle time

r is the distance

G is gravitational constant

2 Kepler's Law

1- Newton's Second Law

$$F = ma (1)$$

Because the gravitational force is a centripetal one, then it will have a centripetal acceleration:

$$F = m\frac{v^2}{r} \tag{2}$$

since the motion of the body is circular, the covered distance in one complete cycle will be a circumference of a circle

$$F = m \frac{\left(\frac{2\pi r}{t}\right)^2}{r} \tag{3}$$

From Newton's law of gravity

$$F = G\frac{Mm}{r^2} \tag{4}$$

from equation 3 and 4 we get:

$$\begin{split} \frac{m4\pi^2 r}{t^2} &= G \frac{Mm}{r^2} \\ \frac{4\pi^2 r}{t^2} &= G \frac{M}{r^2} \\ \frac{4\pi^2 r}{t^2} &= G \frac{M}{r^2} \\ 4\pi^2 r^3 &= GMt^2 \\ \frac{4\pi^2 r^3}{GM} &= t^2 \end{split} \tag{5}$$

by substituting r with the semi-major a, and t with the orbital period P

$$P^2 = \frac{4\pi^2}{GM}a^3\tag{6}$$

(Kepler's law)

3 Orbital Velocity

from eq. 2 and eq. 4

$$m\frac{v^2}{r} = G\frac{Mm}{r^2}$$

$$v^2 = \frac{GM}{r}$$

$$v = \sqrt{\frac{GM}{r}}$$
(7)

4 Escape Velocity

the potential energy due to gravity equals:

$$E = G\frac{Mm}{r} \tag{8}$$

the Kinetic energy of the body equals:

$$E = \frac{1}{2}mv^2 \tag{9}$$

for the body to have an escape velocity (least velocity required to leave its orbit), the potential energy should equal the kinetic energy. from equations 8 and 9 we get:

$$G\frac{Mm}{r} = \frac{1}{2}mv^{2}$$

$$\frac{2GM}{r} = v^{2}$$

$$v_{esc} = \sqrt{\frac{2GM}{r}}$$
(10)

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