A Spacetime Finite Elements Method to Solve the Dirac Equation

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Abstract

In this work, a fully implicit numerical approach based on space-time finite element method is presented to solve the Dirac equation in 1 (space) + 1 (time), 2 + 1, and 3 + 1 dimensions. We utilize PETSc/Tao library to implement our linear system and for using Krylov subspace based solvers such as GMRES. We demonstrate our method by analyzing several different cases including plane wave solution, Zitterbewegung, and Klein paradox. Parallel performance of this implementation is also presented.

Keywords: Finite Elements, Dirac Equation, PETSc

1. Introduction

The Dirac equation governs spin 1/2 particles, known as fermions. It has been applied and studied extensively to many fields of physics and chemistry including relativistic heavy ion collisions, heavy ion spectroscopy, laser-matter interaction, and condensed matter physics [1, 2]. Although the Dirac equation

has been utilized profusely, solving this equation is still a challenging problem. Due to the complicated nature of the Dirac equation, only highly symmetric systems can be studied analytically, i.e. the more realistic case should be based on an approximation-based method such as semi-classical theory [3] and numerical calculations. However, the usual time scales of the fermion dynamics is often much smaller than the time scales of interesting phenomena. Thus, obtaining the numerical solutions is extremely difficult, and it is hard to maintain computational efficiency. Furthermore, certain numerical schemes, such as naive symmetric spatial differencing, are often encountered with fermion doubling problem [4, 5, 6]. Previously, there are several different numerical approaches to explore the Dirac equation. In [7, 8, 9, 10, 11], some variations of the operator splitting method along with a spectral scheme are presented. Also, the finite element scheme [4, 12] and finite difference scheme [13, 14, 15] have been exploited.

In this work, we apply spacetime finite element method (FEM) to solve the Dirac equation in 1 (space) + 1 (time), 2 + 1, 3 + 1 dimensions. This work applies spacetime FEM in a similar way as in previous works [16, 17, 18]. The spacetime FEM has advantages for numerical simulations. The method explored in this work is a fully-implicit method. Instead of time integration, spatial and temporal dimensions are considered simultaneously. Thus, approximated solutions can avoid accumulated time integration errors. Furthermore, this approach can use time varying computational domain, higher order approaches, and unstructured meshes. Moderate size of simulation based on the spacetime FEM requires solving linear systems with millions of unknowns. Therefore, the major challenge of the method is the significant memory overhead requirement that entire spacetime problem needs to fit in memory all at once.

The rest of this paper is organized as follows: in Section 2, the numerical approach is provided, including space-time finite element discretization and solvers, in Section 3, implementation description of our work is presented; Section 4 presents results with different example cases while Section 5 contains the conclusions and future works

2. Numerical Approaches

2.1. Weak Formulation of the Dirac Equation

The standard form of the Dirac equation in a gauge-free case is:

$$(i\hbar\gamma^{\mu}\partial_{\mu} - m)\Psi(x^{\mu}) = 0 \tag{1}$$

where $x^{\mu} \in \mathbb{R}^4$, a four dimensional Minkowski space. Traditionally $\mu = 0$ corresponds to time and $\mu = 1, 2, 3$ corresponds to the three spatial dimensions. In this work, we adapt the natural unit system such as $c = \hbar = 1$.

And, the weak form of the Dirac boundary value problem is to find Ψ_h such that:

$$\int_{\Omega} \overline{\Phi}_h(i\hbar\gamma^\mu \partial_\mu - m) \Psi_h d\Omega = 0$$

$$\Psi_h(x^i) = \Psi_0(x^i) \text{ for } x^i \in \partial\Omega$$
(2)

for every Φ_h where:

$$\Phi_h = \begin{bmatrix} \Phi_1 n_1(x^\mu) & \Phi_2 n_2(x^\mu) & \dots & \Phi_N n_N(x^\mu) \end{bmatrix}^T
\Psi_h = \begin{bmatrix} \Psi_1 n_1(x^\mu) & \Psi_2 n_2(x^\mu) & \dots & \Psi_N n_N(x^\mu) \end{bmatrix}^T
N \in \mathbb{N}, \Phi_i, \Psi_i \in \{ \begin{bmatrix} \phi_1 & \phi_2 & \phi_3 & \phi_4 \end{bmatrix}^T \phi_{1-4} \in \mathbb{C} \}
n(x^\mu) : \mathbb{R}^4 \to \mathbb{R}$$

where \mathbb{R}^4 is the 4D Minkowski space, and $\overline{\Phi}_h = \Phi^* \gamma_0$. Note that both Φ_h and Ψ_h share the same set of interpolation functions $n(x^{\mu})$. Substituting for the definition of Φ_h and Ψ_h allows us to move to the matrix form and simplify the expression.

$$\int_{\Omega} \left[\overline{\Phi}_1 n_1(x^{\mu}) \quad \overline{\Phi}_2 n_2(x^{\mu}) \quad \dots \quad \overline{\Phi}_N n_N(x^{\mu}) \right] (i\hbar \gamma^{\mu} \partial_{\mu} - m) \begin{bmatrix} \Psi_1 n_1(x^{\mu}) \\ \Psi_2 n_2(x^{\mu}) \\ \dots \\ \Psi_N n_N(x^{\mu}) \end{bmatrix} d\Omega = 0 \quad (3)$$

The function vectors Φ_h and Ψ_h can be simplified by using the subscripts i and j to refer to the columns and rows of the resulting matrix. Thus, we can rewrite

this as follows

$$\left[\int_{\Omega} \overline{\Phi}_{j} n_{j}(x^{\mu}) (i\hbar \gamma^{\mu} \partial_{\mu} - m) \Psi_{i} n_{i}(x^{\mu}) d\Omega \right]_{ij} = 0 \text{ where } 1 \leq i, j \leq N$$
 (4)

All of the complex constants in Φ_h can then be factored out to the front of the matrix

$$\overline{\Phi} \left[\int_{\Omega} n_j(x^{\mu}) (i\hbar \gamma^{\mu} \partial_{\mu} - m) n_i(x^{\mu}) d\Omega \right]_{ij} \Psi_i = 0 \text{ where } \overline{\Phi} = \left[\overline{\Phi}_1 \quad \overline{\Phi}_2 \quad \dots \quad \overline{\Phi}_N \right]$$
 (5)

and eliminated by multiplying both sides by their inverse

$$\frac{\Phi}{\|\Phi\|^2} \cdot \overline{\Phi} \left[\int_{\Omega} n_j(x^{\mu}) (i\hbar \gamma^{\mu} \partial_{\mu} - m) n_i(x^{\mu}) d\Omega \right]_{ij} \Psi_i = \frac{\Phi}{\|\Phi\|^2} \cdot 0$$

$$\Rightarrow \left[\int_{\Omega} n_j(x^{\mu}) (i\hbar \gamma^{\mu} \partial_{\mu} - m) n_i(x^{\mu}) d\Omega \right]_{ij} \Psi_i = 0$$
(6)

We will refer to Eqn. 6 as the simplified weak form written as

$$D_w \Psi = 0 \tag{7}$$

Here D_w is a $N \times N$ block matrix with each entry a 4×4 matrix giving a total dimension of $4N \times 4N$. We begin by supposing that the test functions are chosen such that the boundary value at t=0 defines the first $K \in \mathbb{N}$ spinor values of Ψ_h uniquely, such that $\Psi_0(\mathbf{x}^\mu) = \begin{bmatrix} \Psi_1 n_1(x^\mu) & \Psi_2 n_2(x^\mu) & \dots & \Psi_K n_K(x^\mu) \end{bmatrix}$. This allows us to partition $D_w\Psi$ into known and unknown components as follows.

$$\begin{bmatrix} \begin{bmatrix} d_{1,1} & \cdots & d_{1,K} \\ \vdots & D_{w11} & \vdots \\ d_{K,1} & \cdots & d_{K,K} \end{bmatrix} & \begin{bmatrix} d_{1,K+1} & \cdots & d_{1,N} \\ \vdots & D_{w12} & \vdots \\ d_{K,K+1} & \cdots & d_{K,N} \end{bmatrix} \\ \begin{bmatrix} d_{K+1,1} & \cdots & d_{K+1,K} \\ \vdots & D_{w21} & \vdots \\ d_{N,1} & \cdots & d_{N,K} \end{bmatrix} & \begin{bmatrix} d_{K+1,K+1} & \cdots & d_{K+1,N} \\ \vdots & D_{w22} & \vdots \\ d_{N,K+1} & \cdots & d_{N,N} \end{bmatrix} \end{bmatrix} \times \begin{bmatrix} \begin{bmatrix} \Psi_1 \\ \vdots \\ \Psi_K \end{bmatrix} \\ \begin{bmatrix} \Psi_{K+1} \\ \vdots \\ \Psi_N \end{bmatrix} \end{bmatrix} = 0 \quad (8)$$

Using block matrix multiplication, this can be rewritten as the sum of four

smaller matrix operation as follows.

$$\begin{bmatrix} d_{1,1} & \cdots & d_{1,K} \\ \vdots & D_{w11} & \vdots \\ d_{K,1} & \cdots & d_{K,K} \end{bmatrix} \begin{bmatrix} \Psi_1 \\ \vdots \\ \Psi_K \end{bmatrix} + \begin{bmatrix} d_{1,K+1} & \cdots & d_{1,N} \\ \vdots & D_{w12} & \vdots \\ d_{K,K+1} & \cdots & d_{K,N} \end{bmatrix} \begin{bmatrix} \Psi_{K+1} \\ \vdots \\ \Psi_N \end{bmatrix} + \dots$$

$$\begin{bmatrix} d_{K+1,1} & \cdots & d_{K+1,K} \\ \vdots & D_{w21} & \vdots \\ d_{N,1} & \cdots & d_{N,K} \end{bmatrix} \begin{bmatrix} \Psi_1 \\ \vdots \\ \Psi_K \end{bmatrix} + \begin{bmatrix} d_{K+1,K+1} & \cdots & d_{K+1,N} \\ \vdots & D_{w22} & \vdots \\ d_{N,K+1} & \cdots & d_{N,N} \end{bmatrix} \begin{bmatrix} \Psi_{K+1} \\ \vdots \\ \Psi_N \end{bmatrix} = 0 \quad (9)$$

From the definition of the matrix, the first line corresponding to spinor values $\Psi_{1\to K}$ is equal to zero. Thus, we may remove it and rewrite the equation as follow

$$\begin{bmatrix} d_{K+1,K+1} & \cdots & d_{K+1,N} \\ \vdots & D_{w22} & \vdots \\ d_{N,K+1} & \cdots & d_{N,N} \end{bmatrix} \begin{bmatrix} \Psi_{K+1} \\ \vdots \\ \Psi_{N} \end{bmatrix} = - \begin{bmatrix} d_{K+1,1} & \cdots & d_{K+1,K} \\ \vdots & D_{w21} & \vdots \\ d_{N,1} & \cdots & d_{N,K} \end{bmatrix} \begin{bmatrix} \Psi_{1} \\ \vdots \\ \Psi_{K} \end{bmatrix}$$
(10)

Therefore, the equation can be written as using only the matrix subscript as

$$D_{w22}\Psi_{K+1\to N} = -D_{w21}\Psi_{1\to K} \tag{11}$$

We also can express Eqn. 11 in terms of the boundary value problems

$$\left[\int_{\Omega} n_{j}(x^{\mu})(i\hbar\gamma^{\mu}\partial_{\mu} - m)n_{i}(x^{\mu})d\Omega\right]_{ij}\Psi_{i} = -\left[\int_{\Omega} n_{k}(x^{\mu})(i\hbar\gamma^{\mu}\partial_{\mu} - m)n_{i}(x^{\mu})d\Omega\right]_{ik}\Psi_{k} (12)$$
where $K + 1 \leq i, j \leq N$ and $1 \leq k \leq K$

Equation 12 is the weak formulation of the Dirac equation. Each n_j is interpolate functions for finite element space.

2.2. Space-Time Finite Element Discretization

A space-time FEM using continuous approximation functions in both space and time is used to solve the system. The space-time FEM divides the domain into a finite number of subsets, which are called elements, and confines the function space of the weak form to test functions that are non-zero on only a finite number of elements in the domain Ω . Since the test functions have no value outside of the local element, we note that Eqn. 12 is only non-zero when n_i and n_j belong to the same finite element. Thus, we compute Eqn. 12 as an integral over the basis function (n_i) of each element, which results in the element stiffness matrix, and then sum each element stiffness matrices to assemble the stiffness matrix for entire domain Ω .

In this viewpoint, the FEM establishes an algebraic relationship between nodes. If the elements are chosen in a specific grid pattern, this algebraic relationship may become equivalent to a finite difference stencil. The discretization of the Dirac equation in this paper is an extension of previous works presented in [16, 17, 18].

There are several different choices for interpolation functions n_i . For example, we use Lagrange tensor elements for the finite element spaces. In 1+1 case, the Lagrangian interpolation polynomials are:

$$n_1(x,t) = (h-x)(h-t)/h^2$$
, (13)

$$n_2(x,t) = x(h-t)/h$$
, (14)

$$n_3(x,t) = (h-x)t/h$$
, (15)

$$n_4(x,t) = xt/h^2$$
 (16)

where h is size of element

For 2+1 and 3+1 cases, we examine the interpolation polynomials with respect to each coordinates x, y, z, t. Different choice of interpolation function will provide different numerical efficiencies. A detailed study of interpolation function for 1+1 case is presented in [19]. Detailed element stiffness matrix calculations are provided in Appendix A.

Due to the large size of the unsymmetrical system of the discretized matrix, iterative methods based on Krylov subspace (KSP) methods such as the generalized minimal residual method (GMRES) [20] are used.

3. Implementation Description

In this section, we describe our implementation details. We utilize PETSc (Portable Extensible Toolkit for Scientific Computation) to build and solve our system. PETSc, developed at Argonne National Laboratory, is at its core a highly efficient library for parallel linear algebra. It was designed for use in C, C++, Fortran, and Python. PETSc handles the parallel distribution of matrices and vectors, in addition to a variety of linear and nonlinear solvers that ease the use of parallel computing.

Since the Dirac equation contains complex number, there are several important steps to import PETSc correctly. Here, we describe details on configuration and installation of PETSc

To configure PETSc on a Linux or Mac OS to use our code, the following is required. First after downloading PETSc from [21], there are two environmental variables that need to be defined. PETSC_DIR is used to point to the directory where PETSc is located, for example \$HOME\user\petsc\,, while PETSC_ARCH is the build name. PETSC_ARCH does not need to be named anything specific but is commonly used to identify how PETSc was configured. Assuming Linux terminal an example of these would be,

```
export PETSC_DIR=$HOME\user\petsc\
export PETSC_ARCH=complex_petsc
```

Note that we highly recommended you to declare PETSC_DIR and PETSC_ARCH into your bash files such as .bashrc or .bash_profile otherwise PETSC_DIR and PETSC_ARCH will need to be redefined every time you reopen the terminal. This can be done by adding above export commands into your bash files

Before PETSc can be configured, first make sure that you have compatible C, C++, Fortran, and python compliers installed; then run and follow the given instructions.

```
--with-cxx=<your C++ compiler>
--with-fc=<your Fortran compiler>
--with-scalar-type=complex
```

After successful configuration, PETSc introduce next step to install automatically. Or, you can simply type make all check to install PETSc.

3.1. PETSc configure options explained

fblaslapack option installs both BLAS and LAPACK which are both numerical linear algebra libraries written in Fortran. If already installed this option can be disregarded, however if PETSc configure cannot find it in PATH run --with-blaslapack-dir=<location of BLAS/LAPACK>.

mpich is an option for MPI communication, if a MPI is already installed configure without this command, however if it is not found a warning will appear during the configure process and the option -{with-mpi-dir=<location of MPI>.

- < Compiler> is the name of whatever compiler you have installed for that language.
- --with-scalar-type=complex must be called for our code to work as default PETSc assumes real number system.

More detailed information about the configuration and installation can be found in PETSc official website [21].

3.2. Dirac software Algorithm

Algorithm 1 Main Algorithm

- 1: Given global desired mesh construct global size of vectors
- 2: Define initial conditions of ψ
- 3: Create Dirac D_w matrix
- 4: Separate D_w into D_{w22} and D_{w21} \triangleright (11)

 \triangleright (7)

- 5: Create A matrix using D_{w22} and b vector using D_{w21} \triangleright Construct Ax = b
- 6: Solve the linear system by utilizing KSP

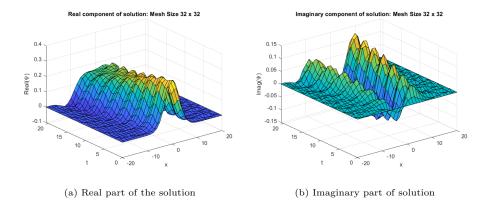


Figure 1: Gaussian plane wave solution for 1+1 case. We use 32×32 meshes on the domain $-20 \le x \le 20$ and $0 \le t \le 20$ with a initial wave packet centered at x=0

Algorithm 1 shows our main algorithm to solve the Dirac equation using spacetime FEM. We also describe parameters in this code in Appendix C

4. Results

In this section, we provide several test cases to demonstrate validations and functionalities. All tests are performed in LANL supercomputer Badger and SDState Roaring Thunder. Badger is an 4-SU cluster running RHEL Linux v.7.7; it has dual socket 2.1 GHz 18 core Intel Broadwell E5 2695v4 processor with 45MB of cache and 128GB of RAM on each node. Roaring Thunder consists of 56 compute nodes, 5 large memory nodes, 4 NVIDIA GPU nodes (V100/P100), and a 1.5 PB high-performance GPFS parallel file system.

4.1. Gaussian Plane Wave

As a first step, we examine Gaussian plane wave solutions. Detailed derivation for plane wave solution is presented in Appendix B.

Fig. 1 shows an example plane wave solution in 1+1 case. We use 32×32 meshes on the domain $-20 \le x \le 20$ and $0 \le t \le 20$ with a initial wave packet centered at x=0.

Mesh Size $(Nt \times Nx)$	Solver	Initial Amplitude	Relative Errors (%)
32×32	GMRES	1	1.1332
64×64	GMRES	1	0.2863
128×64	GMRES	1	0.1127
32×32	GMRES	0.5	1.2012
64×64	GMRES	0.5	0.0897
128×64	GMRES	0.5	0.0503
32×32	BiCGSTAB	1	1.3119
64×64	BiCGSTAB	1	0.3245
128×64	BiCGSTAB	1	0.1574
32×32	BiCGSTAB	0.5	1.2878
64×64	BiCGSTAB	0.5	0.2919
128×64	BiCGSTAB	0.5	0.1073

Table 1: 1+1 case

Since we can obtain the analytic expression for Gaussian wave packet, we compare our numerical results with analytic solution to examine our toolkit for all different dimensional cases.

For these testing, we varied mesh sizes, solvers, and initial amplitude by introducing additional scalar value to initial data with different dimensional cases with respect to each dimensions on the domain $0 \le x \le 1$ and $0 \le t \le 1$ with an initial wave packet centered at x = 0.5. This particular choice of domain is small enough to resolve all wave motions with different mesh sizes. Our desired tolerance is 10^{-7} for all tests and we compute relative error by $|(\Psi_{\text{numeric}} - \Psi_{\text{exact}})/\Psi_{\text{exact}}| \times 100$.

Tables 1, 2, 3 show the Gaussian wave studies for 1+1, 2+1, and 3+1 respectively. We choose two different KSP solvers, GMRES and BiCGSTAB. As we increase number of meshes, relative errors decrease for all cases. We observe that both GMRES and BiCGSTAB agree with analytic solution well for all cases. Furthermore varying initial amplitude value didn't change solution quality too.

${\text{Mesh Size } (Nt \times Nx \times Ny)}$	Solver	Initial Amplitude	Relative Errors (%)
$24 \times 24 \times 24$	GMRES	1	2.7147
$32 \times 32 \times 32$	GMRES	1	1.1526
$64 \times 32 \times 32$	GMRES	1	0.5724
$24 \times 24 \times 24$	GMRES	0.5	2.5958
$32 \times 32 \times 32$	GMRES	0.5	1.0984
$64 \times 32 \times 32$	GMRES	0.5	0.5322
$24 \times 24 \times 24$	BiCGSTAB	1	2.7981
$32 \times 32 \times 32$	BiCGSTAB	1	1.1673
$64 \times 32 \times 32$	BiCGSTAB	1	0.5802
$24 \times 24 \times 24$	BiCGSTAB	0.5	2.7093
$32 \times 32 \times 32$	BiCGSTAB	0.5	1.1424
$64 \times 32 \times 32$	BiCGSTAB	0.5	0.5427

Table 2: 2+1 case

${\text{Mesh Size } (Nt \times Nx \times Ny \times Nz)}$	Solver	Initial Amplitude	Relative Errors (%)
$12 \times 12 \times 12 \times 12$	GMRES	1	5.3212
$24\times24\times24\times24$	GMRES	1	0.3426
$32\times32\times32\times32$	GMRES	1	0.1052
$12\times12\times12\times12$	GMRES	0.5	5.1447
$24\times24\times24\times24$	GMRES	0.5	0.3168
$32\times32\times32\times32$	GMRES	0.5	0.1136
$12\times12\times12\times12$	BiCGSTAB	1	5.2618
$24 \times 24 \times 24 \times 24$	BiCGSTAB	1	0.3157
$32\times32\times32\times32$	BiCGSTAB	1	0.1039
$12\times12\times12\times12$	BiCGSTAB	0.5	5.5192
$24\times24\times24\times24$	BiCGSTAB	0.5	0.3321
$32 \times 32 \times 32 \times 32$	BiCGSTAB	0.5	0.1022

Table 3: 3+1 case

For 3+1 case, there is comparably larger error for smallest mesh size because mesh size for each direction might be small to resolve the system well. However, as we increase the mesh size, relative error decreases significantly. Note the matrix size for the smallest 3+1 case is 82944×82944 which is reasonably large matrix system. In general, all cases show that our results agree well with analytic case.

The plane waves propagate vacuum so energy should be conserved. We monitor energy during the evolution to check energy loss. Using this energy loss, we also measure order of convergence of our scheme.

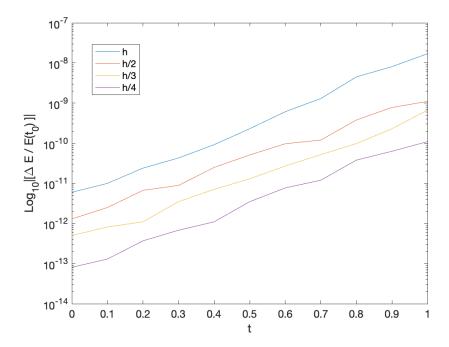


Figure 2: Energy loss plot for plane wave based on different resolutions. The mesh h is a space resolution which can be obtained by dividing physical domain size by the number of collocation point. Here, h is defined by 1/100 = 0.01 which is chosen arbitrary. The energy loss $\Delta E(t) \equiv E(t_f) - E(t_0)$ is plotted until the initial data reaches to end of our time domain $t_f = 1$. As resolution increases, energy loss decreases. Using this data, the self convergence test is evaluated $(||\Delta E_{h/4} - \Delta E_{h/2}||_2)/(||\Delta E_{h/2} - \Delta E_h||_2) = 4.091$ which indicates the order of self convergence is second order

Fig. 2 shows the energy loss plots for plane wave solution. The energy loss is defined by $\Delta E(t) \equiv E(t_f) - E(t_0)$ where t_0 is initial time and t_f is final time i.e. end of our time domain. Ideally, the energy loss $|\Delta E(t)|$ should be zero. As shown in Fig. 2, energy loss decreases as resolution increase. Using this, a self convergence test is performed by evaluating $(||\Delta E_{h/4} - \Delta E_{h/2}||_2)/(||\Delta E_{h/2} - \Delta E_h||_2)$. This was taken where the same initial conditions were used but with different resolutions. The self convergence test value is 4.091 which indicates second order convergence.

4.2. Zitterbewegung

Zitterbewegung ('trembling motion' in German) [22] is a phenomena exclusive to relativistic quantum mechanics describing the oscillatory behavior of a wave packet which is intrinsic to the Dirac equation. This is quite different from the smooth relaxation of a packet predicted by the Schrödinger equation.

There are many different discussions on Zitterbewegung and its interpretation. In the Newton-Wigner theory [23], the position operator leaves the positive and negative energy sub-spaces invariant and thus it does not display the Zitterbewegung behavior. In the Foldy-Wouthuysen representation [24], this operator is identical to the standard position operator. Regardless of these interpretation, we will focus on Zitterbewegung in position, x(t) as a test case for our implementation.

Using Heisenberg's picture, we can obtain analytic expression on position operator

$$i\dot{x} = [H_0, x],\tag{17}$$

where H_0 is free-field Dirac Hamiltonian. In this section, we keep c to track physical constant.

Solving these equations will provide operator solutions such that

$$x(t) = x + c^{2} p_{x} H_{0}^{-1} t + \frac{ic}{2} (cp_{x} + H_{0}^{-2} - \alpha_{x} H_{0}^{-1}) (1 - e^{-2iH_{0}t}).$$
 (18)

Note that the last term Eqn. 18 shows oscillatory behavior.

We are interested to obtain expectation value to compare this analytic expression with our numerical simulation.

To display this behavior using our method, the initial condition was set to a Gaussian wave packet centered at x=0.

$$\psi(x,0) = \left(\frac{1}{32\pi}\right)^{1/4} e^{-\frac{x^2}{16}} \begin{pmatrix} 1\\1 \end{pmatrix} \tag{19}$$

The spatial expectation value $\langle x(t) \rangle$ was calculated using the standard position operator. The following integrals were evaluated numerically using trapezoidal sums.

$$\langle x(t) \rangle = \int_{D} x ||\psi(x,t)||_{2}^{2} dx.$$
 (20)

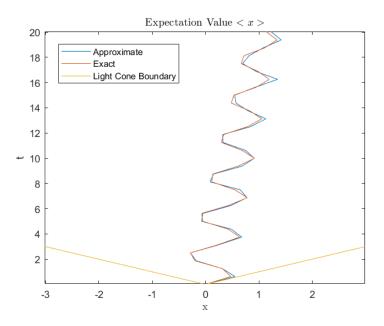


Figure 3: Zitterbewegung behavior).

Figure 3 shows the numerical and exact spatial expectation value, $\langle x(t) \rangle$ as well as the light cone boundary. The light cone boundary represents the trajectory of a particle at the speed of light. Any results outside this boundary

would be non-physical, as it would imply part of the wave packet is traveling faster than the speed of light. As it is presented in Fig. 3, the numerical results seem to compare well to the analytic solution and lie inside the light cone boundary as expected.

4.3. Klein Paradox

The Klein paradox [25] is a well-known example for which the single particle interpretation of the Dirac equation can lead to some paradoxical predictions. In non-relativistic quantum mechanics, the wave function in the region where potential, V, is non-zero is decaying exponentially if the energy, E, of the incident wave is lower than the V. Thus, most of the wave function is reflected and transmission coefficient is negligible. However, in the relativistic case, a new phenomenon appears when $E < V - mc^2$. In this regime, a plane wave solution can exist, resulting in a non-negligible transmission coefficient even if E < V. Many previous works have been reported to "resolve" the Klein paradox in the context of the Dirac see picture and second quantization [26, 27] In this subsection, we explore this phenomena in 1+1 case.

The Dirac equation for a particle in the presence of an external scalar potential field is

$$\begin{pmatrix} i\partial_t + V - m & i\partial_x \\ -i\partial_x & -i\partial_t - V - m \end{pmatrix} \begin{pmatrix} \psi^1(x) \\ \psi^2(x) \end{pmatrix} = 0.$$
 (21)

Gaussian wave packet is chose for initial data such that

$$\psi(x,t=0) = \left(\frac{1}{32\pi}\right)^{1/4} e^{-100(x-0.6)^2} e^{35ix} \begin{pmatrix} 1\\1 \end{pmatrix}.$$
 (22)

This is a wave packet centered at x = 0.6 with an average momentum of 35. Also, we consider a smooth potential to avoid numerical issues related with discontinuous function. This potential, called Sommerfeld potential is given by

$$V(x) = \frac{V_0}{2} [1 + \tanh((x - x_0)/L)]. \tag{23}$$

Here, V_0 and x_0 are the magnitude and central location of the potential step, respectively, and L corresponds to the steepness of the transition to the step.

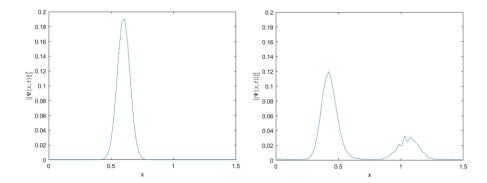


Figure 4: Example wave density $||\psi(x,t)||_2^2$ before (left) and after (right) the interaction of the initial wave packet with the potential barrier. Potential barrier is located at $x_0 = 0.8$.

Fig. 4 shows the wave distribution $||\psi(x,t)||_2^2$ before and after scattering at the Sommerfeld potential. In this figure, the parameters for this potential were set to $V_0 = 36.1$, L = 0.001, and $x_0 = 0.8$,

The potential (Eqn. 23) is already studied by [28, 29]. The exact formula for the transmission coefficient for a Sommerfeld potential is given by

$$T = \frac{\sinh(\pi p L) \sinh(\pi p' L)}{\sinh(\pi (V_0 + p + p') \frac{L}{2}) \sinh(\pi (V_0 - p - p') \frac{L}{2})},$$
(24)

where $p = \sqrt{(E_p - V_0)^2 - 1}$, $p' = -\sqrt{E_p^2 - 1}$, and $E_p = \sqrt{p_0^2 + 1}$. Using this, reflection coefficients can be obtained by R = 1 - T.

To compare this result with an analytic expression, the T and R were calculated such that

$$T = \frac{\int_{\Omega^{+}} ||\psi(x, t_{f})||_{2}^{2} dx}{\int_{\Omega} ||\psi(x, t_{f})||_{2}^{2} dx}, \quad R = \frac{\int_{\Omega^{-}} ||\psi(x, t_{f})||_{2}^{2} dx}{\int_{\Omega} ||\psi(x, t_{f})||_{2}^{2} dx}, \tag{25}$$

where Ω is whole spatial domain, Ω^+ is transmitted domain, and Ω^- is reflected domain with t_f being the final time value, after the wave has scattered at the potential. These integrals were evaluated numerically using trapezoidal sums.

Fig. 5 shows scattering coefficients as a function of the steepness of the transition to the step, L. For this test, we use 256×256 mesh size and after 10,000 iterations of the GMRES. Different L were tested to compare with analytic values of transmission and reflection coefficients. All relative errors are less than

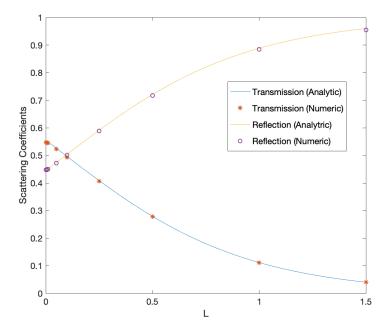


Figure 5: Scattering coefficients as function of the steepness of the transition to the step. We vary different values of L = [0.001, 0.005, 0.01, 0.05, 0.1, 0.25, 0.5, 1.0, 1.5] to compare with analytic expression (Eqn 24).

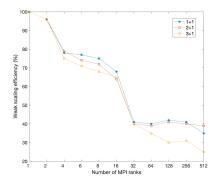
0.5% which shows our numerical results agree well with analytic expression.

4.4. Performance Tests

We perform both weak and strong scaling tests to measure parallel efficiency of our implementation. on the LANL supercomputing cluster, Badger.

The plane wave with different dimensions was used for both tests. For jobs with number of MPI ranks from 1 to 32, we used a single node, and for larger jobs, we use multiple nodes with 32 rank per nodes. For strong scaling tests, we use meshes of 256×256 , $64 \times 64 \times 64$, and $32 \times 32 \times 32 \times 32$ elements for 1+1, and 2+1, and 3+1 respectively. For weak scaling we use 256, 192, and 128 elements per rank for 1+1, and 2+1, and 3+1 respectively. All tests are evolved with 100 GMRES iterations. Then we compute strong scaling efficiency, $t_1/t_N' \times 100\%$ where t_1 is the

amount of time to complete a task with 1 processing element, t_N is the amount of time to complete the same task with N processing elements, and t'_N is the amount of the time to complete N of the same task with N processing elements.



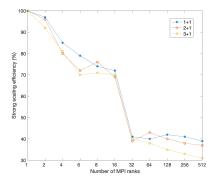


Figure 6: Weak (left) and strong (right) scaling efficiency on Badger clusters. Strong scaling tests used meshes of 512×512 , $128 \times 128 \times 128$, and $64 \times 64 \times 64 \times 64$ elements for 1+1, and 2+1, and 3+1 respectively. Weak scaling used 256, 192, and 128 elements per rank for 1+1, and 2+1, and 3+1 respectively. The plane wave functions are evolved with 100 GMRES iterations.

Fig. 6 shows strong scaling and weak scaling efficiencies for 1+1, and 2+1, and 3+1 cases. Note that efficiencies degrade between 32 and 64 ranks due to the inter-node data transfer, but then remain relatively flat. Overall all cases show comparable scaling. Having a better treatment to partitioning matrix system for higher numbers of MPI ranks will provide better scaling efficiencies.

5. Conclusion and Future Works

In this work, we present a spacetime FEM to solve the Dirac equation. We demonstrate several different application examples, Gaussian plane waves, Zittwerbewegung, and Klein paradox to validate our method and implementation. All of these tests show good agreement with analytic cases. Further, we explore our parallel efficiency. Although certain limitations need to overcome such as inter-node data transfer, all cases show comparable scaling.

As its nearest goal, this method will solve more realistic problems including an inhomogeneous Dirac equation. Further, complicated and realistic problems will require more computational costs thus exploring a proper preconditioner for the system to solve the problem is required. In [16, 18], they applied time-additive Schwarz method as a time decomposition method. Since the nature of Dirac equation is complex, it will be interesting subject to apply the time decomposition idea for this problem and monitor how linear solver such as GMRES performance will be changed. This might provide better performance.

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Appendix A. Matrix Element Calculations

We consider 1+1 case in Eqn. 12 for simplicity. Then equation becomes

$$D_{ij} = \int_{\Omega} n_j(x,t) \left(i\hbar \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \partial_t + i\hbar \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \partial_x - m\mathbb{I} \right) n_i(x,t) dx dt$$

Note that D_{ij} is written as a double integral of three separate terms, which can be separated as

$$D_{ij} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} i\hbar D_{ij}^t + i\hbar \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} D_{ij}^x - m\mathbb{I}D_{ij}^0,$$

where the individual terms D_{ij}^t , D_{ij}^x , and D_{ij}^0 are given by

$$D_{ij}^{t} = \int_{0}^{h} \int_{0}^{h} n_{j}(x,t)\partial_{t}n_{i}(x,t)dxdt$$

$$D_{ij}^{x} = \int_{0}^{h} \int_{0}^{h} n_{j}(x,t)\partial_{x}n_{i}(x,t)dxdt$$

$$D_{ij}^{0} = \int_{0}^{h} \int_{0}^{h} n_{j}(x,t)n_{i}(x,t)dxdt.$$

Using Eqns. 13 - 16, we can compute all of these integrations. As results, each D_{ij}^x , D_{ij}^y , D_{ij}^z are 4×4 matrices. We can extend this idea to higher dimensional cases. We increase spatial dimension into n_i , add extra spatial derivatives for D_{ij} , and increase dimension of gamma matrices for each higher dimension respectively. Thus, we have four 8×8 matrices and five 16×16 matrices for 2+1 and 3+1 respectively. We use MAXIMA [30] to perform integration. Our Maxima script can be found in https://gitlab.com/resundermann/dirac/-/tree/master/tools/integration

Appendix B. Solutions for Plane Gaussian Waves

The first case of the Dirac equation examined in this study is the free-field form, meaning that the particle described by the Dirac equation has no external fields or forces interacting with it. The exact solution to the free-field Dirac equation is relatively simple given the initial state $\psi(x,0)$, making it an appropriate candidate for comparison to a numerical solution. Suppose the initial state of the free particle is a wave packet given by

$$\psi(x,0) = \left(\frac{1}{32\pi}\right)^{1/4} e^{-\frac{x^2}{16} + i\frac{3x}{4}} \begin{pmatrix} 1\\1 \end{pmatrix}.$$

A wave packet is a superposition of waves with a range of momentum which is often used to describe a particle in quantum mechanics. Taking a Fourier transform of the initial state yields the wavefunction as a function of momentum rather than space, $\phi(p,t=0)$. Therefore, the formula for the wavefunction $\phi(p,t=0)$ is

$$\phi(p, t = 0) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \psi(x, 0) e^{-ipx} dx$$
$$= \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \left(\frac{1}{32\pi}\right)^{1/4} e^{-\frac{x^2}{16} + i\frac{3x}{4}} \begin{pmatrix} 1\\1 \end{pmatrix} e^{-ipx} dx.$$

The above integral can be evaluated analytically and is shown in Appendix A. The result is the following wavefunction,

$$\phi(p, t = 0) = \left(\frac{8}{\pi}\right)^{1/4} e^{-4(p-\frac{3}{4})^2} \begin{pmatrix} 1\\1 \end{pmatrix}.$$

Now, this momentum-dependent wavefunction can be written as the linear combination of vectors u_{pos} , u_{neg} so that the proper time-dependency can be assigned to each component [6]. This form of the wavefunction is

$$\phi(p,0) = \phi^{+}(p)u_{pos}(p) + \phi^{-}(p)u_{neg}(p).$$

Here, the coefficients ϕ^{\pm} are the scalar products between the initial wavefunction and the vectors u_{pos} and u_{min} , given by

$$\phi^{\pm} = \langle u_{pos,neg}(p), \phi(p) \rangle_{2},$$

$$u_{pos}(p) = \frac{1}{\sqrt{2}} \begin{pmatrix} (1+1/\sqrt{k^{2}+1})^{1/2} \\ sgn(p)(1-1/\sqrt{k^{2}+1})^{1/2} \end{pmatrix},$$

and

$$u_{neg}(p) = \frac{1}{\sqrt{2}} \begin{pmatrix} -sgn(p)(1-1/\sqrt{k^2+1})^{1/2} \\ (1+1/\sqrt{k^2+1})^{1/2} \end{pmatrix}.$$

Multiplying the upper and lower components by their respective time dependency, e^{-iEt} and e^{iEt} where $E = \sqrt{p^2 + m^2}$ (relativistic energy), yields the time dependent wavefunction,

$$\phi(p,t) = \phi^+(p)u_{pos}(p)e^{-iEt} + \phi^-(p)u_{neg}(p)e^{iEt}.$$

Finally, taking the inverse Fourier transform of the wavefunction in momentum space yields the time dependent wavefunction in position space, written as

$$\psi(x,t) = \int_{-\infty}^{\infty} e^{ipx} (\hat{\phi}^+(p)u_{pos}(p,t) + \hat{\phi}^-(p)u_{neg}(p,t)) dp.$$

This method can be extended to higher dimensional cases by adding extra spatial dimension.

Appendix C. Artifact Description

We maintain all our implementation using git version control system which is available in https://gitlab.com/resundermann/dirac as an open source. We use Matlab to analyze our data and plot the result. Description of our repository is listed in below

- applications: This holds the main driver functions for the 2D, 3D, and 4D codes.
- include: Folder for the header files of each main and src.
- miscell: This is our miscellaneous folder, it contains documentation, old MATLAB code references and scripts we used in our testing. This is currently re-organizing.
- src: This folder contains the functions called in the main functions, including the Dirac functions that build the Dirac matrix, in addition to the functions that separate it into smaller matrices, handled in the Dirac and ShorteMat scripts respectfully.
- tools: currently holds maxima files used in integration for the 2D, 3D, and 4D Dirac codes.

Our repository provide self-contained document for building the code and development work flow.

The code has the functionality to change the number of elements in each dimension at run time. These are the x, y, z, t and angle parameters. Space dimensions default to 4, time 3, angle 30. To change any of these parameters use -variable value. For example, to change x to 10 when running the 4D code the option is "./RotationDirac4D -x 10".

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