

Star Formation in Merging Clusters of Galaxies

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WOC Summit

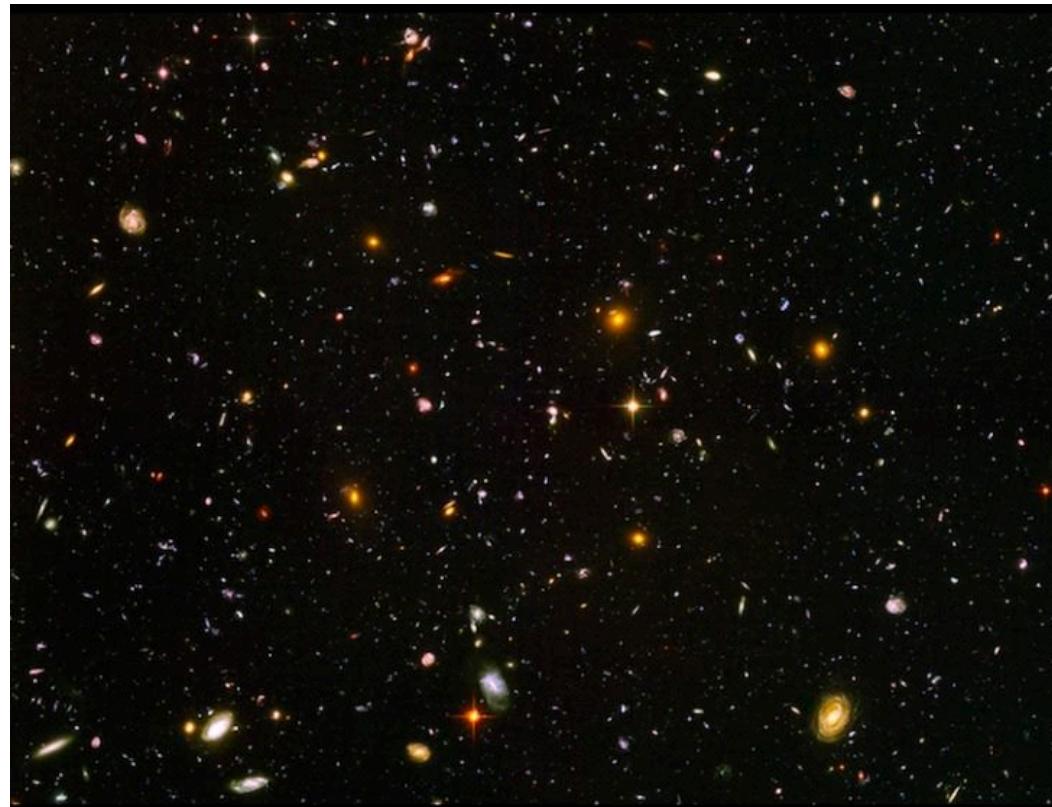
Jan 31, 2015

Roadmap

- ❖ Inspiration Why?
 - ❖ Background What?
 - ❖ Galaxies
 - ❖ Clusters of Galaxies
 - ❖ Techniques How?
 - ❖ Spectroscopy
 - ❖ Models
 - ❖ Conclusion So?

Inspiration

On September 3rd, 2003 The Hubble Space Telescope began pointing its camera at a particularly small and dark patch of sky. After 4 months...

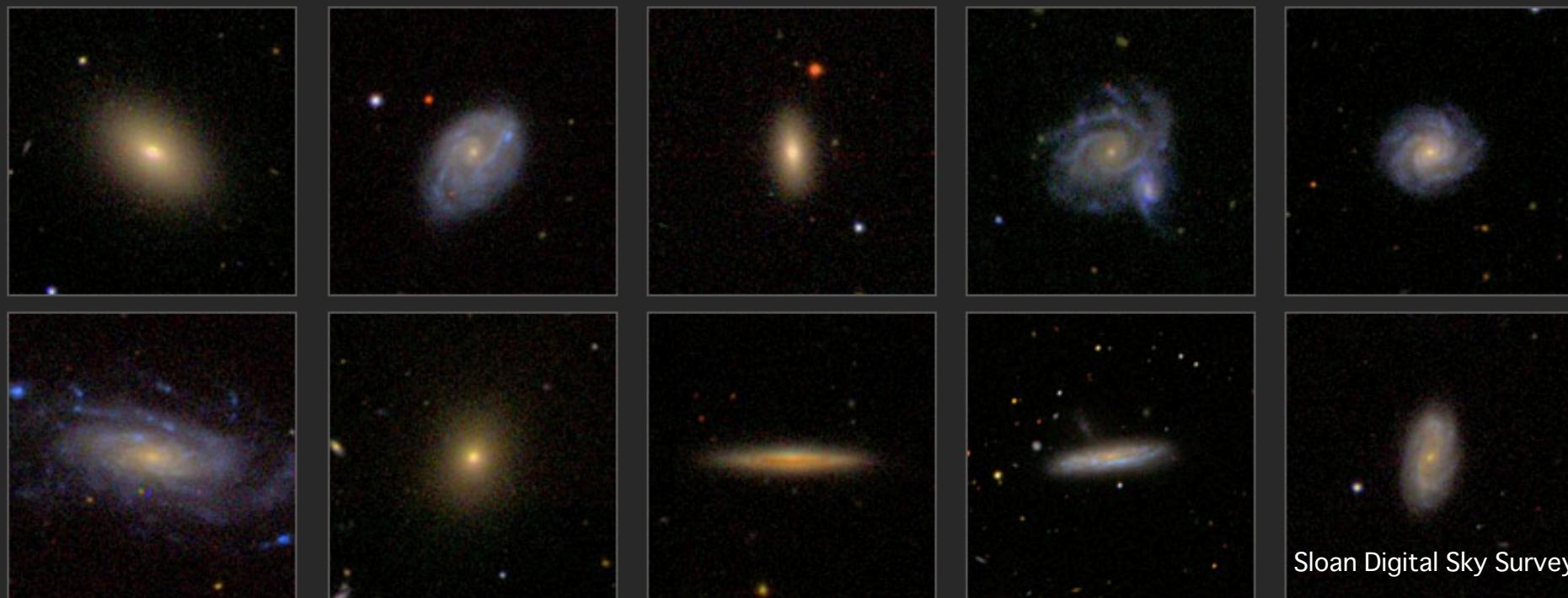


Each dot of light in this image is a galaxy. Each galaxy contains up to 1 trillion stars. Each star may have a system of planets. There are 10 THOUSAND galaxies in this one image alone.

Inspiration

Infinite colors, shapes and sizes!

WHY?

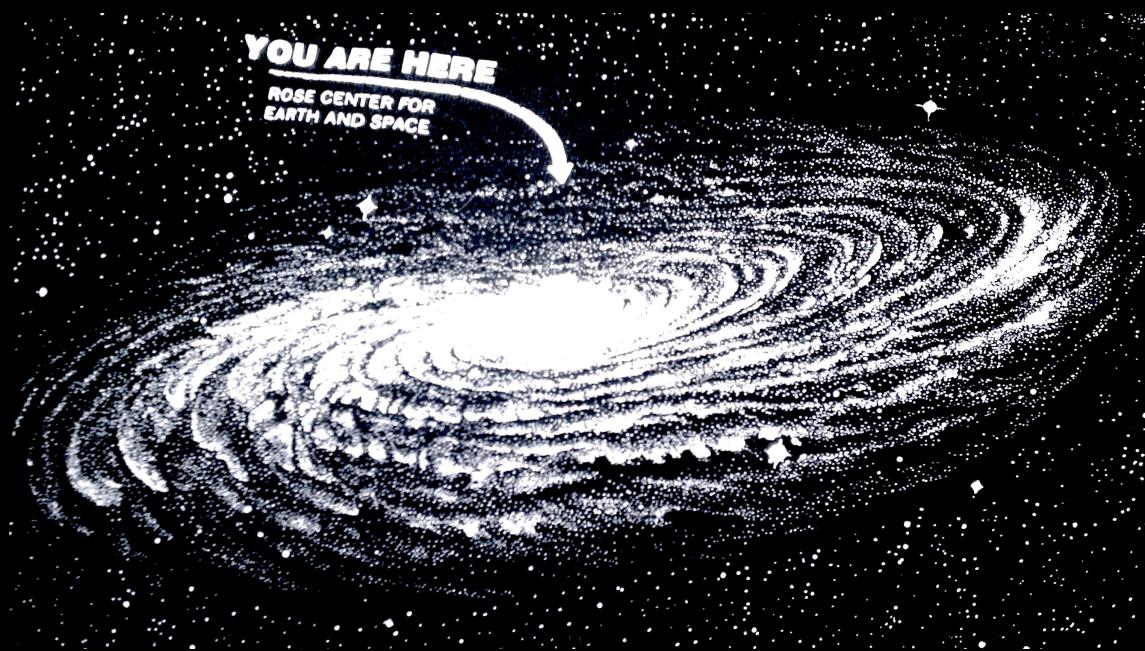


Sloan Digital Sky Survey

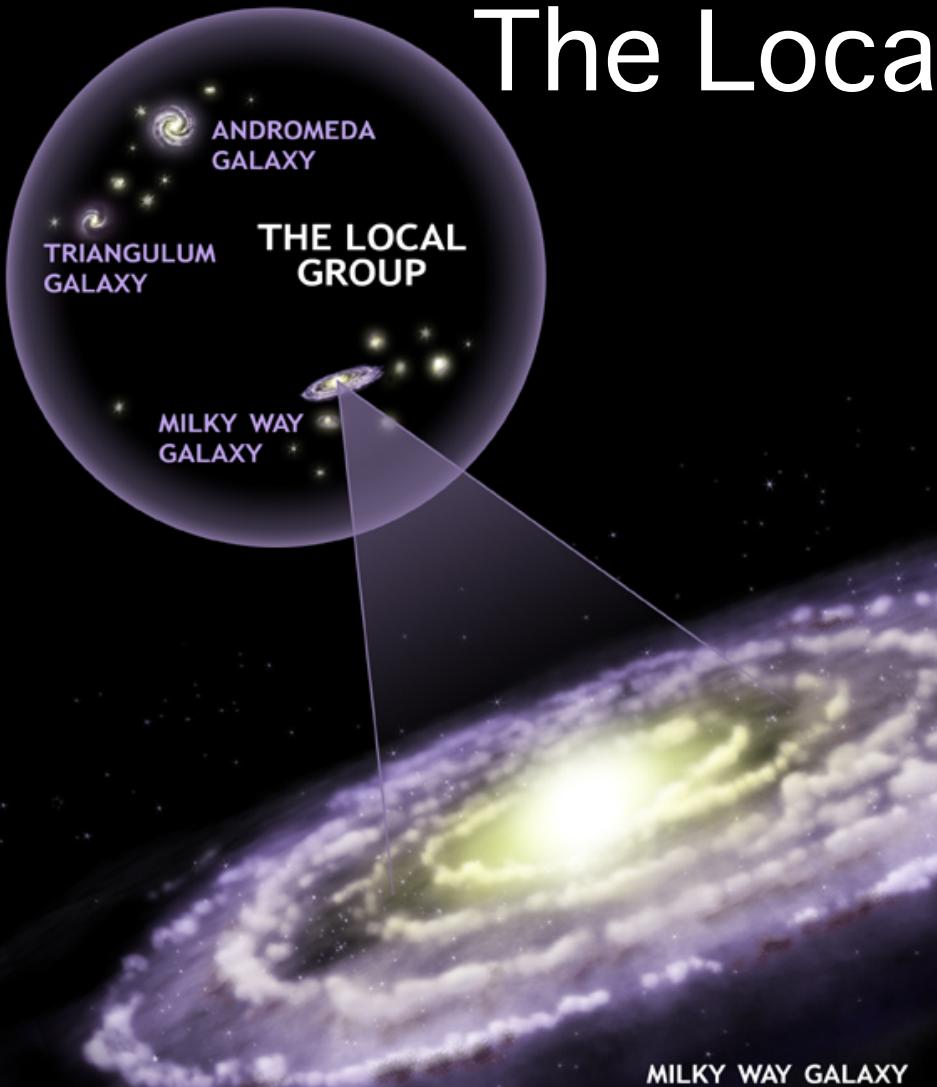
Our Galaxy: The Milky Way



Wally Pacholka / AstroPics.com



Beyond the Milky Way: The Local Group



- The local group is a group of 3 dozen nearby galaxies (3Mly)
- How do we find more?

WAY Beyond the Milky Way: Clusters of Galaxies

- Clusters of galaxies are the largest gravitationally bound objects in the universe
- $M_{200} > 10^{14} M_{\text{solar}}$
- $r_{200} \sim 2 \text{Mpc} \sim 10^{19} \text{ miles}$
- Diverse population
- 50-3000 galaxies
- Dynamic environment



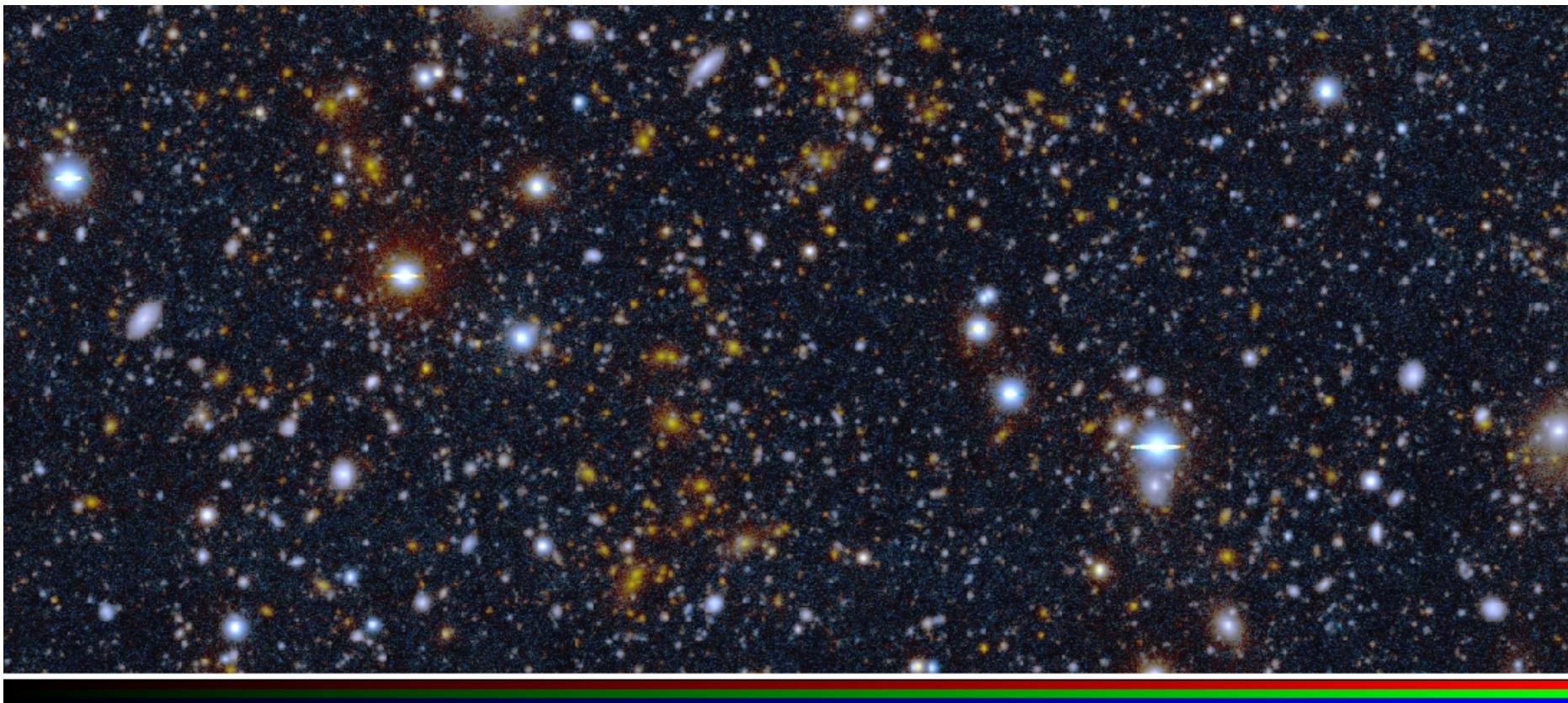
Merging Clusters of Galaxies

- More: Sample size, Diversity, Challenge
- Extreme conditions speeds up evolutionary processes
- Created to relaxed clusters we see today



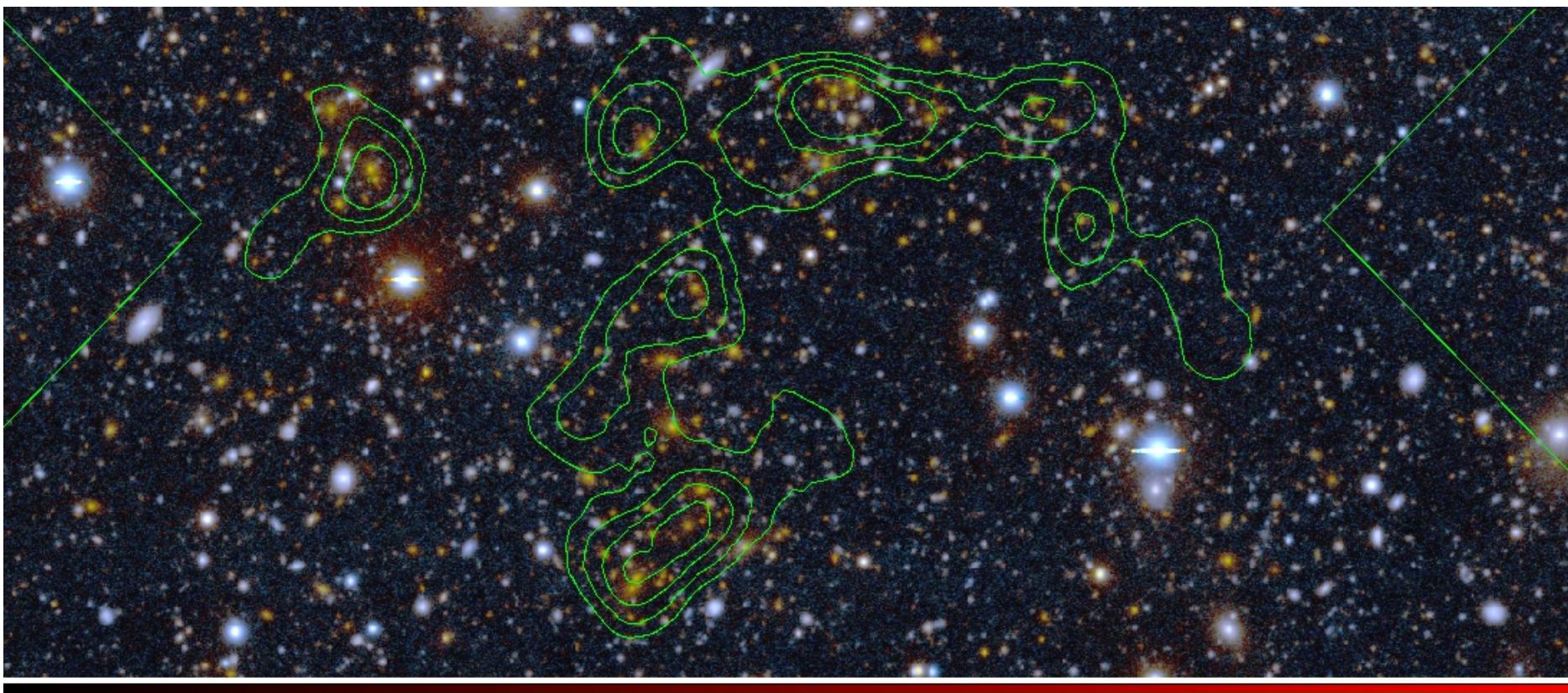
Musket Ball Merger (DLSCL J0916.2+2953)

- $1.9 \times 1.9 \text{ Mpc}^2 \sim 6 \times 6 \text{ million-trillion miles}^2$)
- $z \sim 0.53$ (universe 2Gyr old)



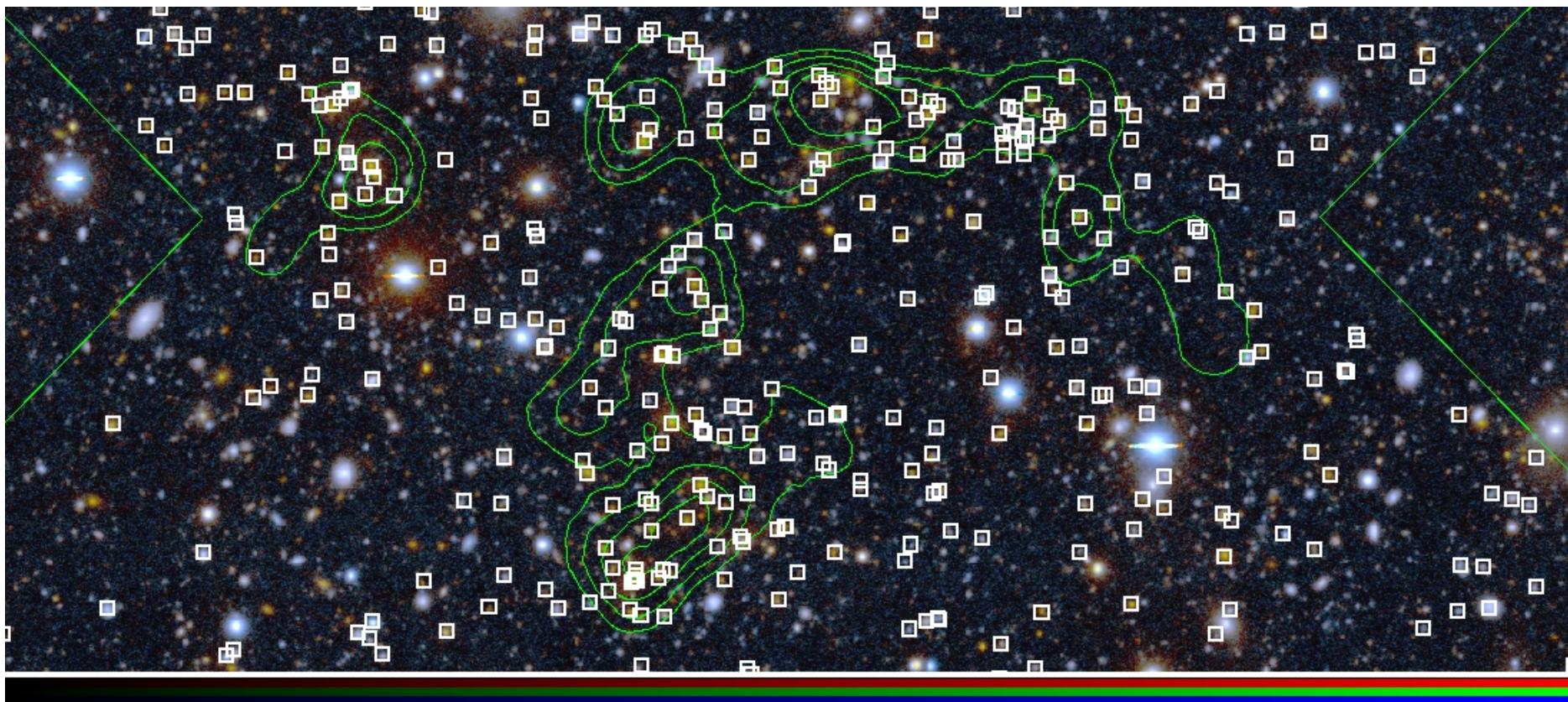
Musket Ball Merger (DLSCL J0916.2+2953)

- $1.9 \times 1.9 \text{Mpc}^2 = 6 \times 6 \text{ million-trillion miles}^2$)
- $z \sim 0.53$ (universe 2Gyr old)
- 0.7Gyr since collision



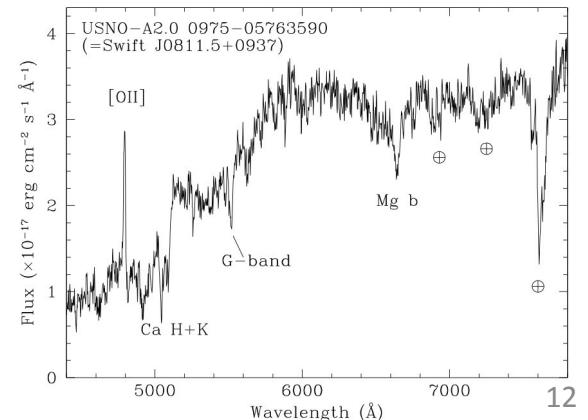
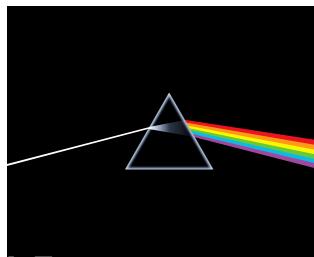
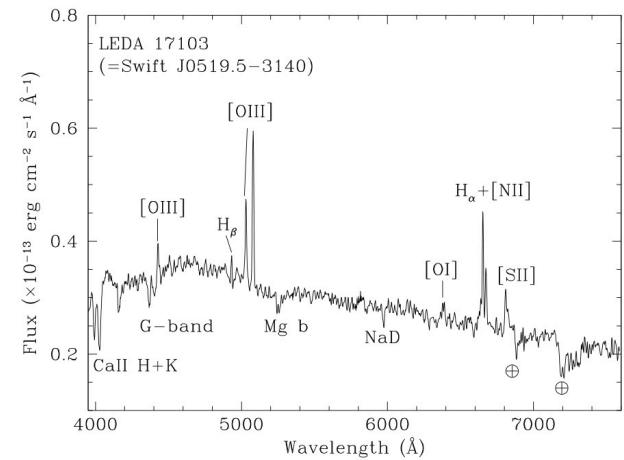
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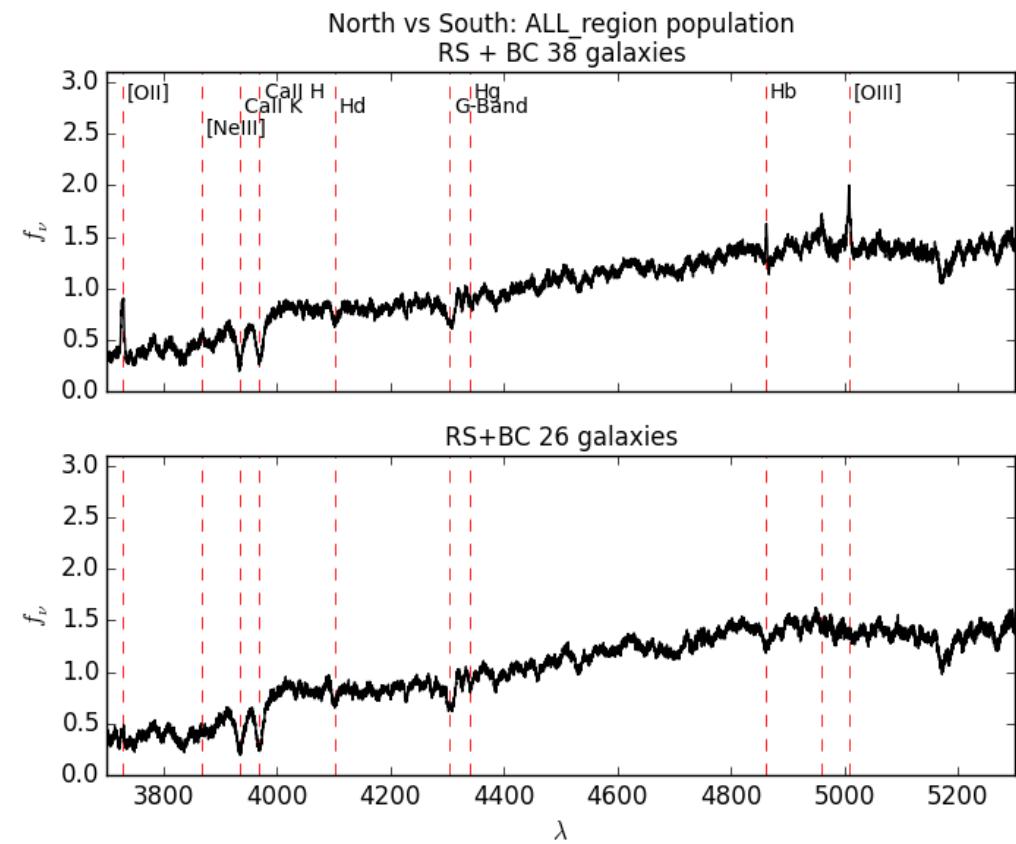


How do we study galaxies?

- Dissect light using spectroscopy.



Composite Spectra: Star Formation in the North and South Clusters

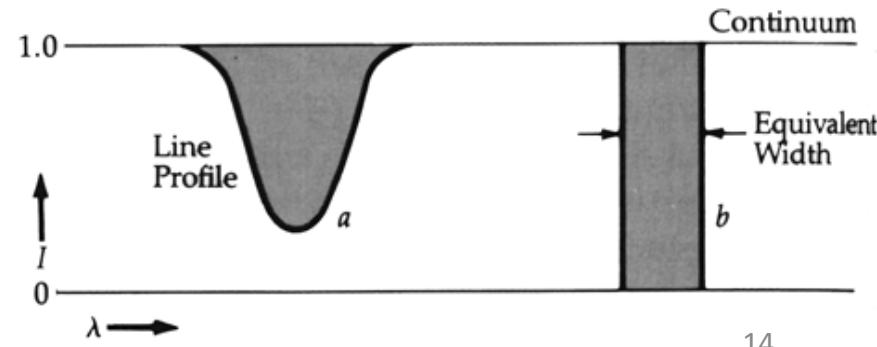
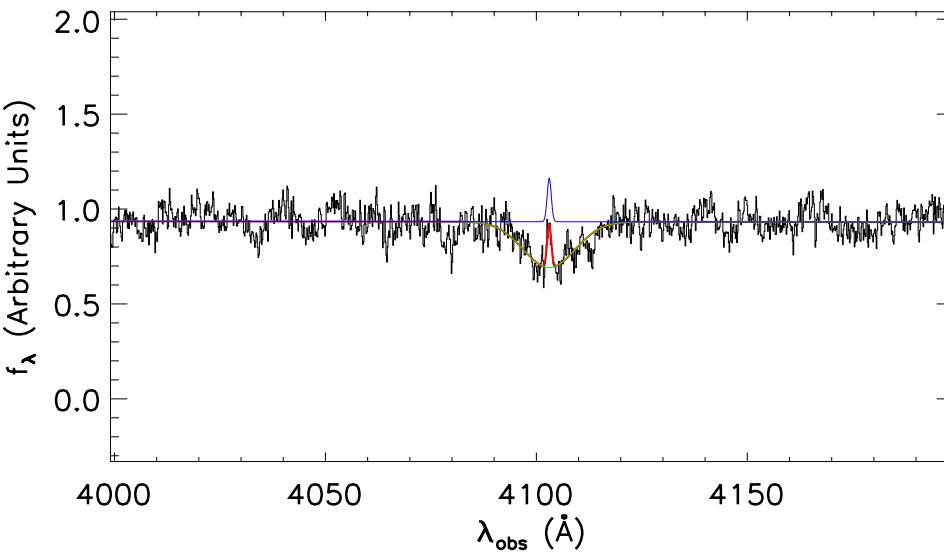


✧ Composites reveal key features of average galaxy in region

How do we measure a spectrum?

- Equivalent Width (EW) is a quantitative measure of line strength
- Allows us to compare strengths among different populations

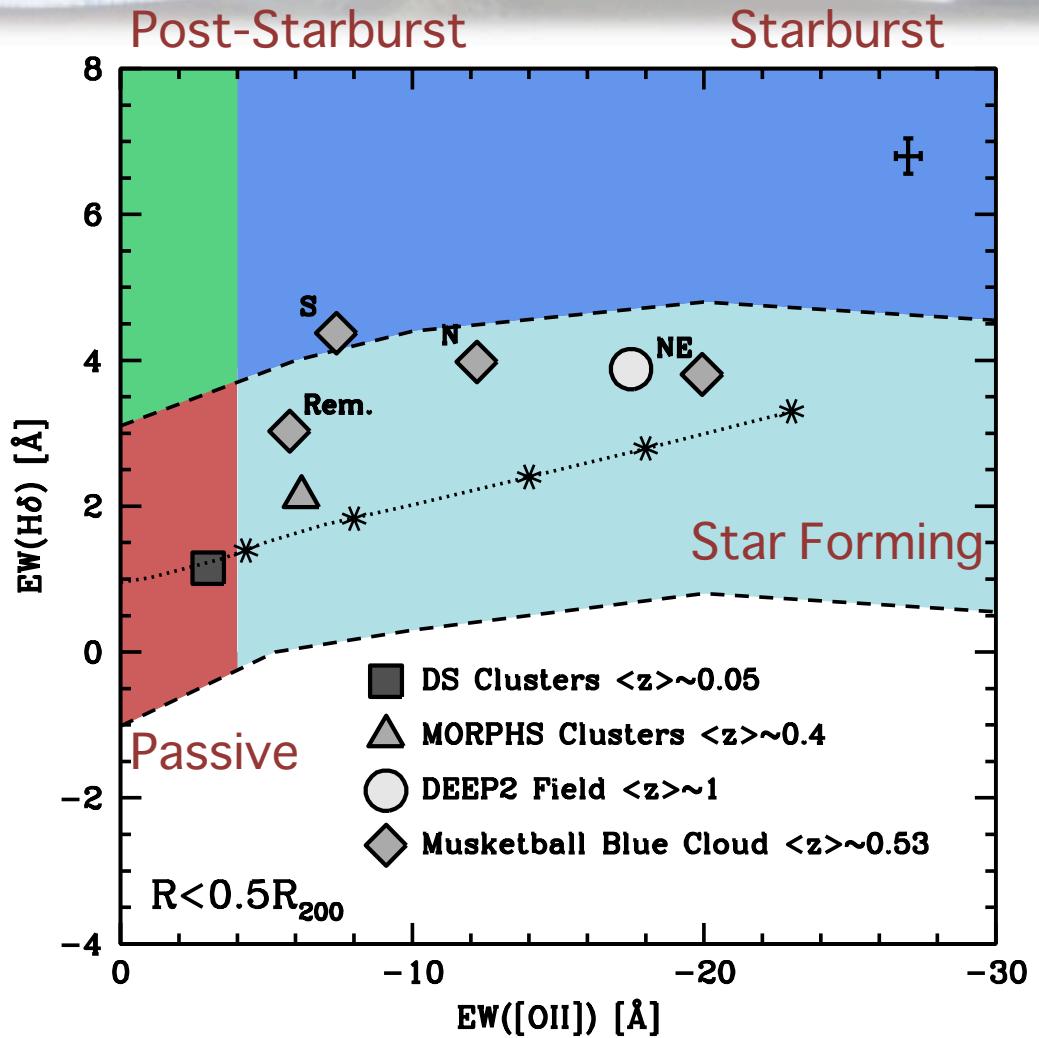
H δ 4101 EW fitted doublet
3.6804323 +/- 0.21035767 Angstroms
From BCcut_comp0.fits



Results: EW(H δ) vs EW([OII])

How much current star formation is happening on average in each region?

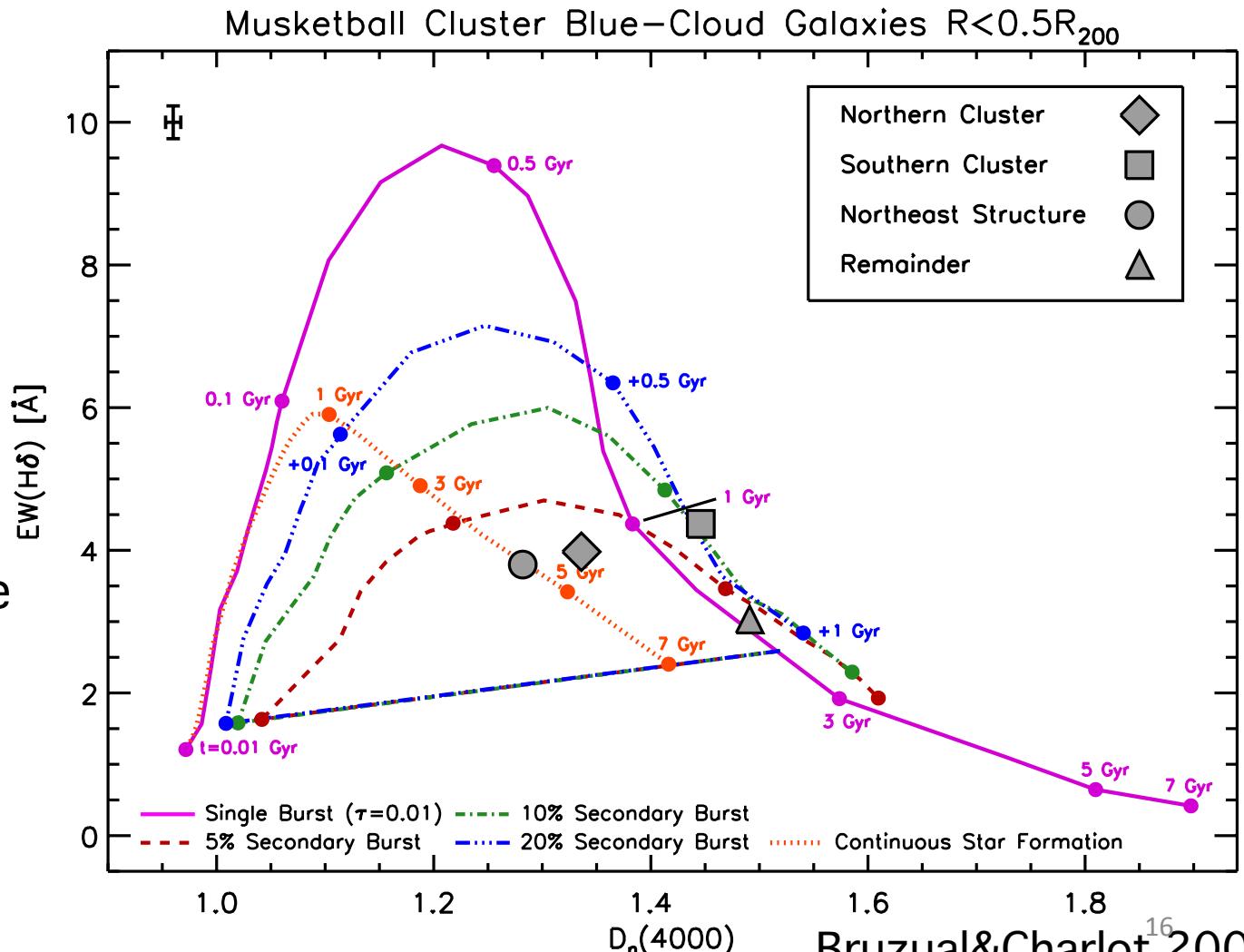
- ✧ N&S clusters well above passive/continuous line, indicating past burst
- ✧ S>N>Superfield



Results: EW(H δ) vs D_n(4000)

How long ago was last starburst?

- ◆ Data consistent with Starburst in South approximately 0.7 Gyrs ago
- ◆ Recall time since merger ~0.7Gyr ago!





Summary

- ✧ Clusters of galaxies are important tools for studying galaxy evolution and star formation
- ✧ Spectroscopy allows us to connect the light we observe to the internal workings of a galaxy
- ✧ Data suggests a starburst in the Southern cluster approximately 0.7 Gyrs ago
- ✧ Starburst coincides with time of cluster merger indicating collision as possible cause



Thank you to the organizers for this opportunity!

Thank you to the audience for listening!

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