## $N_{ m eff}$ priors

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## Abstract

Try to understand the effect of priors on different parameters on the  $N_{\rm eff}$  errors with a Fisher matrix approach. Where should you invest time and money if you really care about the value of  $N_{\rm eff}$ ?

## INTRODUCTION

## **THEORY**

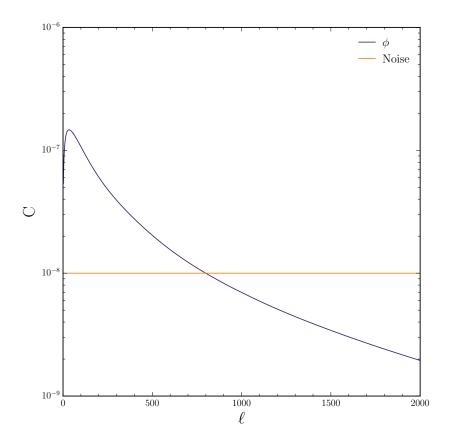


FIG. 1. Lensing potential power spectrum for our fiducial cosmology together with the lensing reconstruction noise  $N^{\phi}$  used in this work.

$$s \left[ \mu \text{K.arcmin} \right] \equiv \frac{\text{NET} \left[ \mu \text{K.} \sqrt{s} \right] \times \sqrt{f_{sky} \left[ \text{arcmin}^2 \right]}}{\sqrt{N_{\text{det}} \times Y \times \Delta T \left[ \text{s} \right]}}.$$
 (1)

$$F_{ij} \equiv -\left\langle \frac{\partial^2 \log \mathcal{L}}{\partial \theta_i \partial \theta_j} \bigg|_{\theta = \theta_0} \right\rangle \tag{2}$$

$$\mathbf{C}_{\ell} \equiv \begin{pmatrix} C_{\ell}^{TT} + N_{\ell}^{TT} & C_{\ell}^{TE} & C_{\ell}^{Td} \\ C_{\ell}^{TE} & C_{\ell}^{EE} + N_{\ell}^{EE} & 0 \\ C_{\ell}^{Td} & 0 & C_{\ell}^{dd} + N_{\ell}^{dd} \end{pmatrix}.$$
(3)

$$\sigma_i \equiv \sigma(\theta_i) = \sqrt{(\mathbf{F}^{-1})_{ii}} \tag{4}$$

$$F_{H_0H_0} \to F_{H_0H_0} + \frac{1}{(1\% \times H_{0,fid})^2},$$
 (5)

$$T_{\nu} = \left(\frac{4}{11}\right)^{1/3} T_{\gamma} \tag{6}$$

 $\mathbf{DATA}$ 

RESULTS

CONCLUSIONS

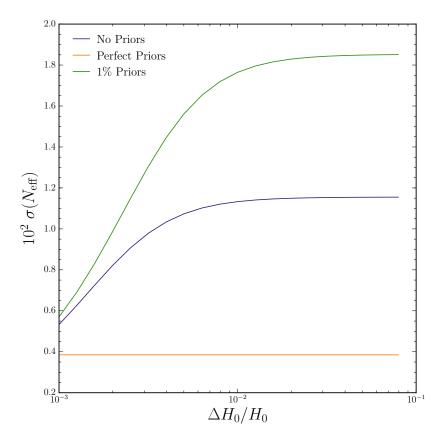


FIG. 2.

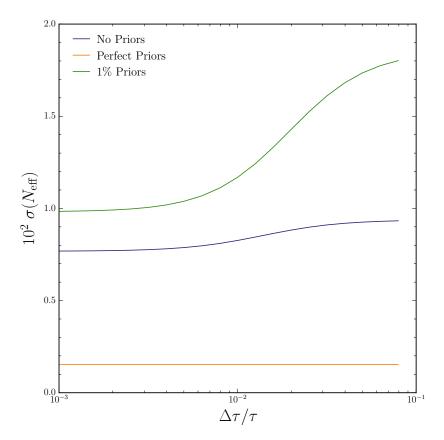


FIG. 3.