

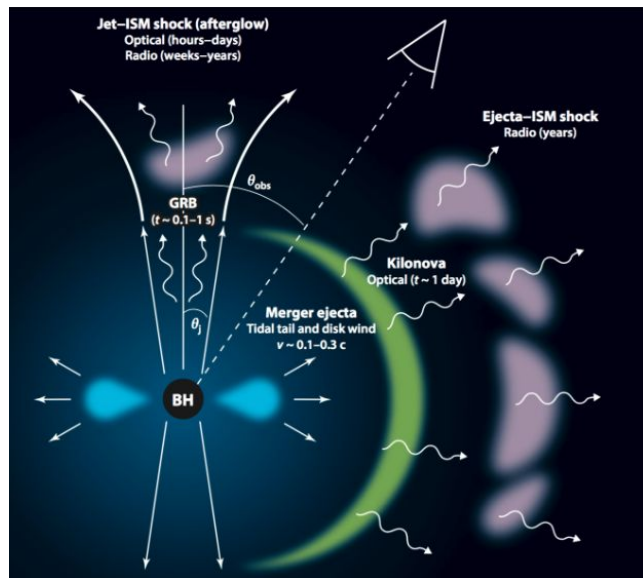
HUNTING KILONOVAE USING WENDELSTEIN AND HET

This is a collective effort!

Gravitational Wave Sources

- Stretches(Squeezes) in spacetime for gravity to comply with Special Relativity
- Major sources: CBCs(Compact Binary Coalescences)- BBH, BHNS, BNS

Optical Counterparts of Gravitational Waves



Metzger 2019

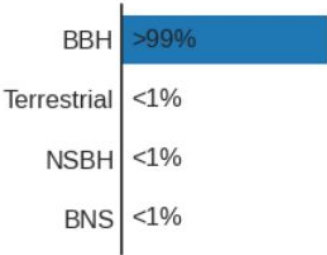
- When NS involved \rightarrow GW+EM Radiation
- BBH \rightarrow Flare in accretion disk
- Kilonovae (optical), short GRBs, radio, X-Ray (jet/afterglow)
- Focus: optical region (kilonovae)
- Formed from the radioactive decay of r-process elements
- Optical timescales: hrs-days

LIGO Data

S240428dr
Log Messages
Full Event Log

S240

Log Image



FAR (yr⁻¹)

t₀

t_{end}

Submitted ▾

Links

Volume rendering of
[bayestar.multiorder.fits,2](#)

— Submitted by LIGO/Virgo EM Follow-Up on April 29,
2024 13:05:40 UTC

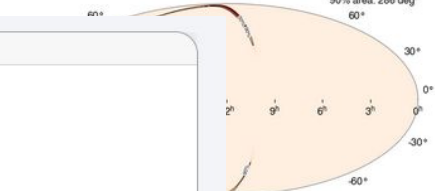
Sky Localization

Log Image

event ID: G478659
distance: 765±177 Mpc

Log Image

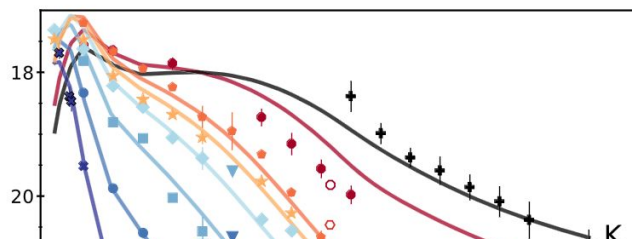
event ID: G478659
50% area: 61 deg²
90% area: 286 deg²



[bayestar.multiorder.fits,2](#)

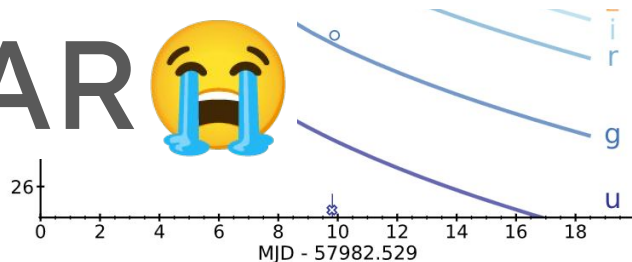
— Submitted by LIGO/Virgo EM Follow-Up on April 29,
2024 13:05:17 UTC

GW170817

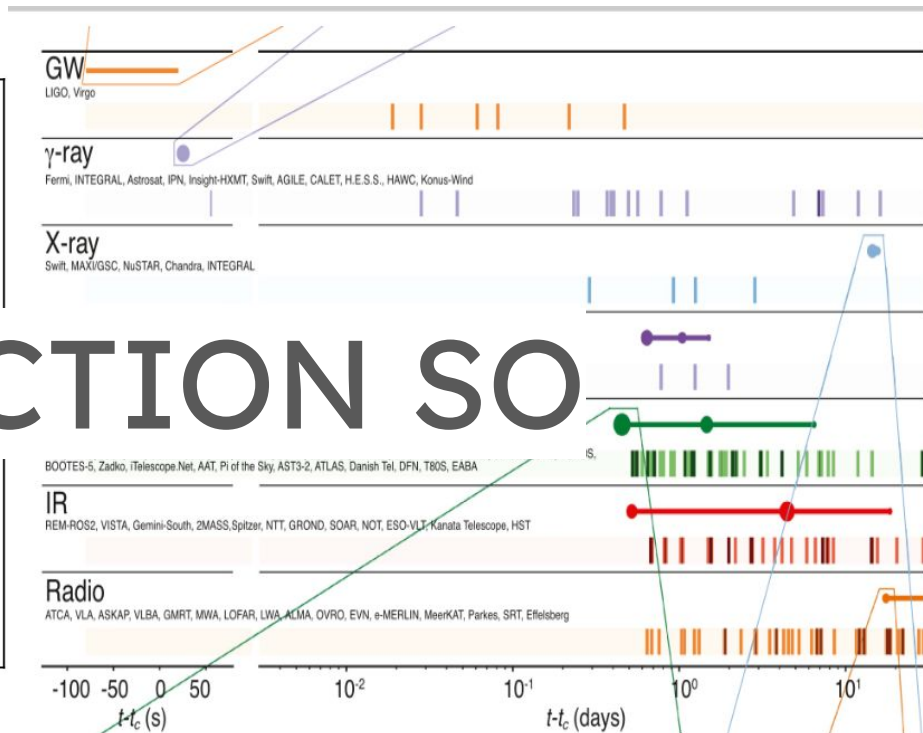


ONLY DETECTION SO

FAR



Cowperthwaite+17



Abbott+17

Implications

- R-process nucleosynthesis(Cowperthwaite+17)
- Standard siren measurements of H_0 (Palmese+23)
- X-Ray+GW→ Properties of ISM (Makhathini+21)
- Tests of GR(Abbott+17) and Modified Gravity Theories(Boran+17)
- Studying the Neutron Star EOS(Margalit & Metzger 17)

Target Selection

Match DESI Data Catalogs to 99% credible region according to GW Localisation+Luminosity(for M_*) or H- β line(for SMBH Mass)

Scheduling

Optimal scheduling algorithm to maximise total probability covered(3KK)

Data Reduction and Difference Imaging

Reduce raw frames from Wendelstein, select template, subtract from the science image to get difference image(look for change in brightness)

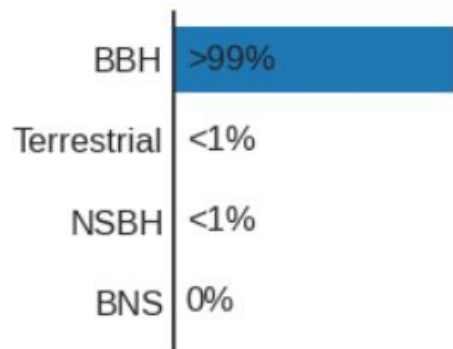
Light Curves and Spectra

Is the source brightening? AB? Rise/fall time? Colour? Kind of galaxy? Redshift?

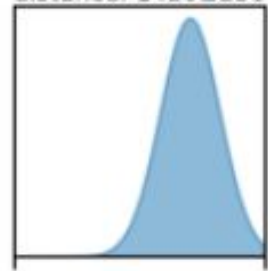
Archival Light Curves

Query past ZTF/ATLAS Data and veto sources with high variability

S240615dg

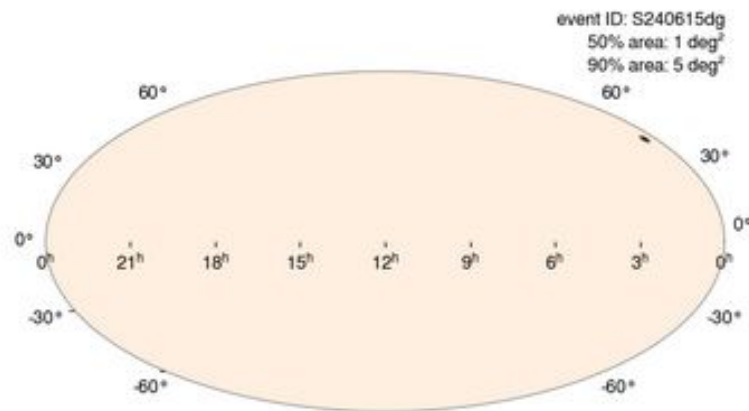


event ID: S240615dg
distance: 1420 ± 236 Mpc



0 2022
Mpc

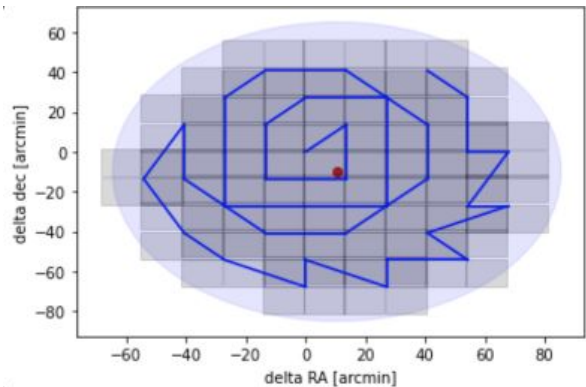
$z=0.2-0.4$




```

1 # QS0 selection for event S240615dg at 1718477904.1148193
2 # Read skymap: /pscratch/sd/j/jgassert/target_selection_data/ligo/S240615dgbayestar.multiorder.fits,2
3 # Used dataset from /pscratch/sd/j/jgassert/target_selection_data/desi/qsos_merged_20240409_20240608.fits
4 # Search area is 17.769420144705176deg^2 large for 0.99 cut
5 # 71 objects in the target list
6 RA, DEC, Z_TRUE, P_QS0, Z_FROM
7 7.179459048264719, 45.91464565960424, 0.27156494312443247, 0.16411154769880884, 1
8 7.013216727505004, 46.77107160396502, 0.33728331327438354, 0.11471919199047348, 1
9 8.962203023010018, 46.52034135057316, 0.3337754460867908, 0.10113340453122312, 1
10 4.807895948753608, 45.439252844594805, 0.26218098402023315, 0.07263620959745845, 1
11 4.6015190220769995, 45.48223697070823, 0.25780320167541504, 0.058420432031132334, 1
12 7.1828391250351595, 45.0305331589576, 0.3342839181423187, 0.051582283263895023, 1
13 7.069876271424079, 43.65307807631715, 0.2910935992686088, 0.04660606328668852, 1
14 6.172420785549909, 46.138430839035216, 0.4767281711101532, 0.04425459420973769, 1
15 6.747933623239949, 44.38260037219568, 0.3512095510959625, 0.034818282484867076, 1

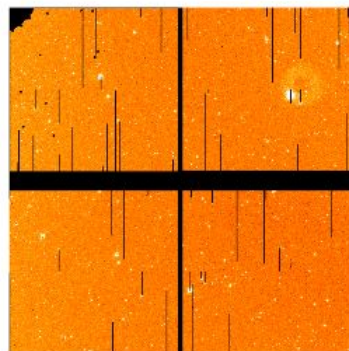
```



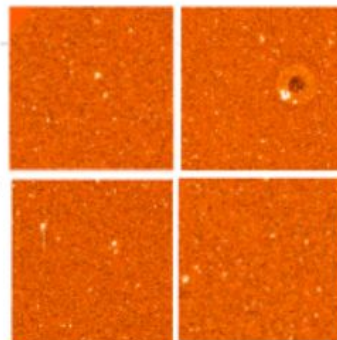
Data Reduction



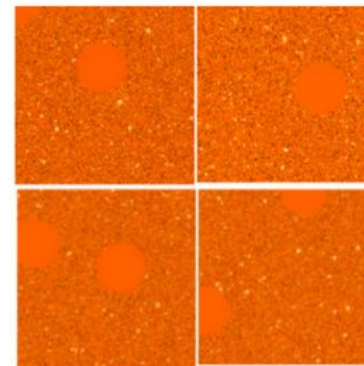
Raw Frames



Reduced Frames



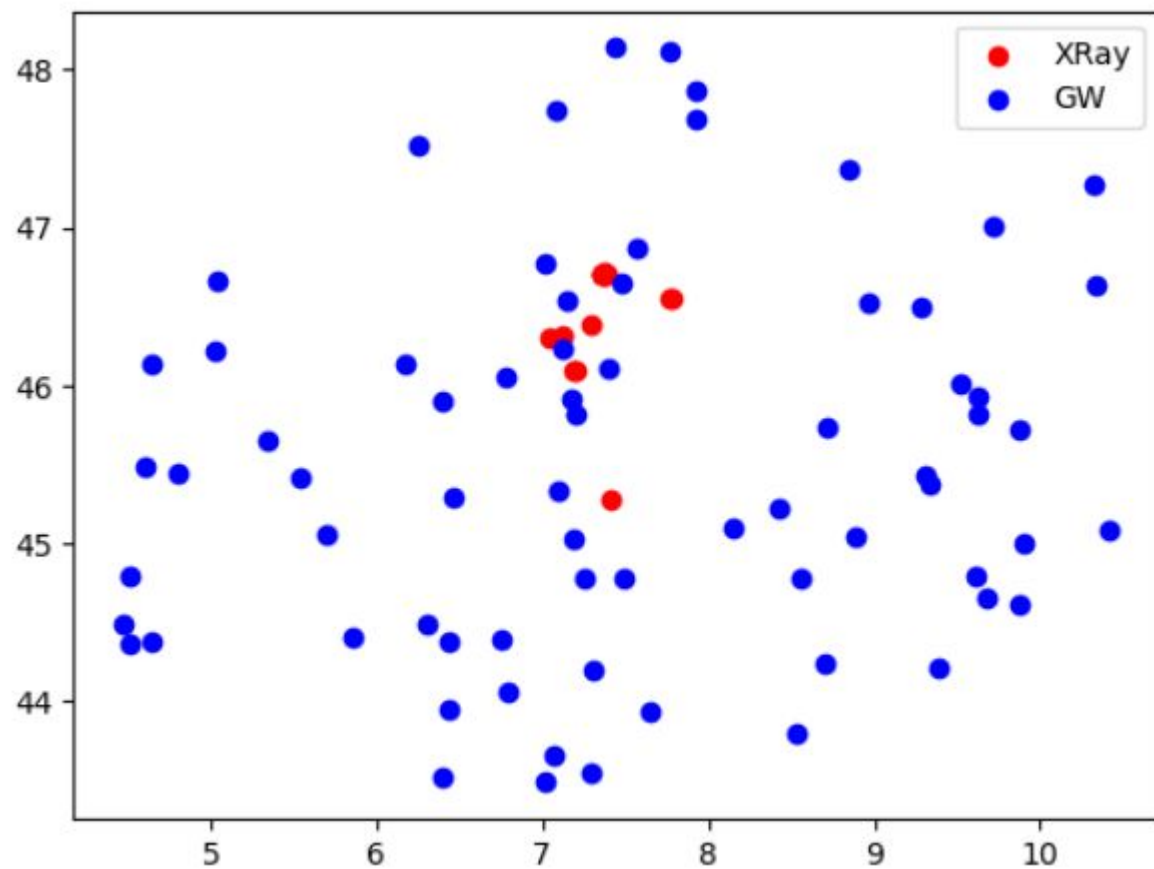
Zero Point, Sky
Subtraction



Masked Images

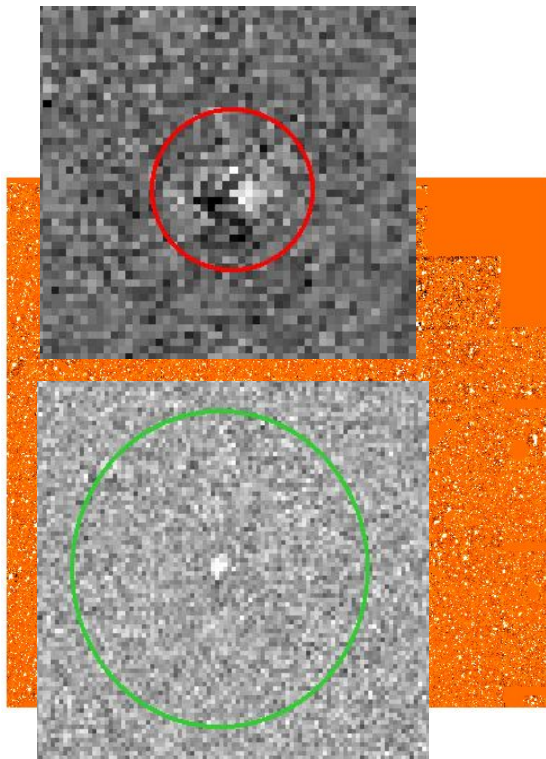
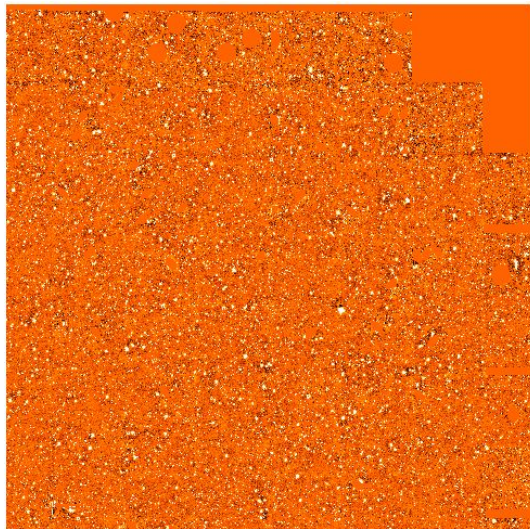
Swift-XRT Sources

- 27 sources of Rank 3: “uncatalogued X-ray sources, however they are not brighter than previous upper limits, so do not stand out as likely counterparts to the GW trigger”
- Crossmatched SExtractor sources within 90% error region of X-Ray sources
- Rejected stars
- Crossmatched this list with the host galaxy candidates
- Found only one match with $\delta=2.88''$ (pretty big!) and $P_{\text{QSO}}=0.05$

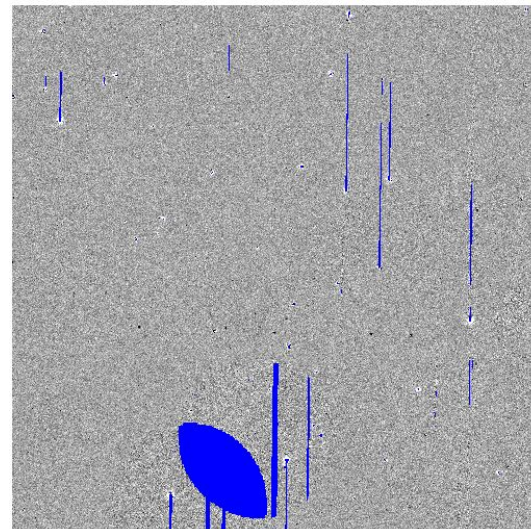


Difference Imaging

Science



Difference



- ❑ Extragalactic?
- ❑ Is it a bird, is it a plane?
- ❑ Is it a star?
- ❑ Is it nuclear?
- ❑ Is it an AGN?
- ❑ Is it at the correct redshift?

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GAIA

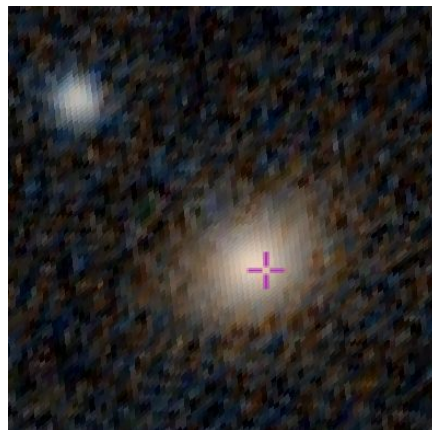
- ❑ Extragalactic?
- ❑ Is it a bird, is it a plane?
- ❑ Is it a star?
- ❑ Is it nuclear?
- ❑ Is it an AGN?
- ❑ Is it at the correct redshift?

NED,
If photo-z, HET Data

The Most Promising Candidate

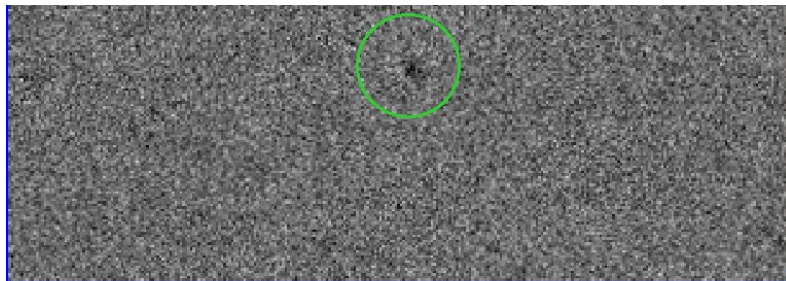
- ✓ Extragalactic? → *Looked like a distant elliptical galaxy*
- ✓ Is it a bird, is it a plane?
- ✓ Is it a star?
- ✓ Is it nuclear?
- ☐ Is it an AGN?
- ☐ Is it at the correct redshift?

No z data on NED
Daniel's photo-z by eye → 0.4
So used HET Data

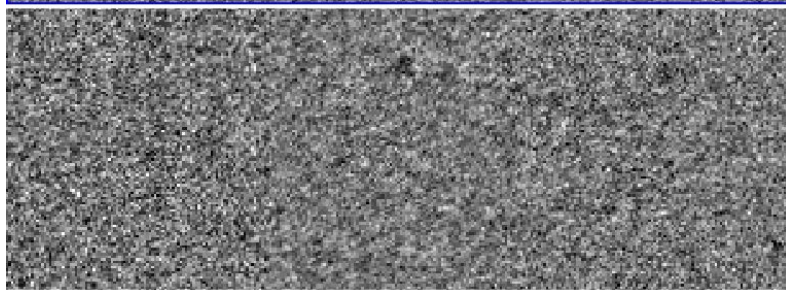
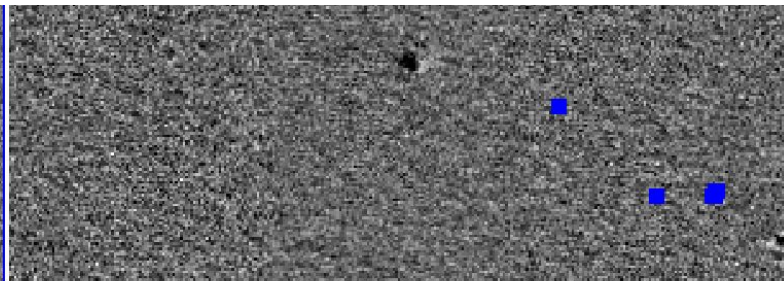


Difference Images

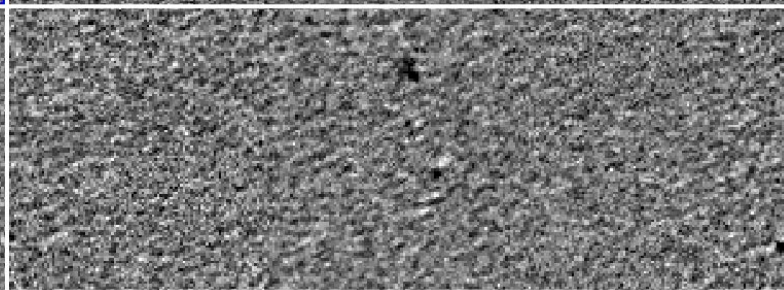
18/06



08/07

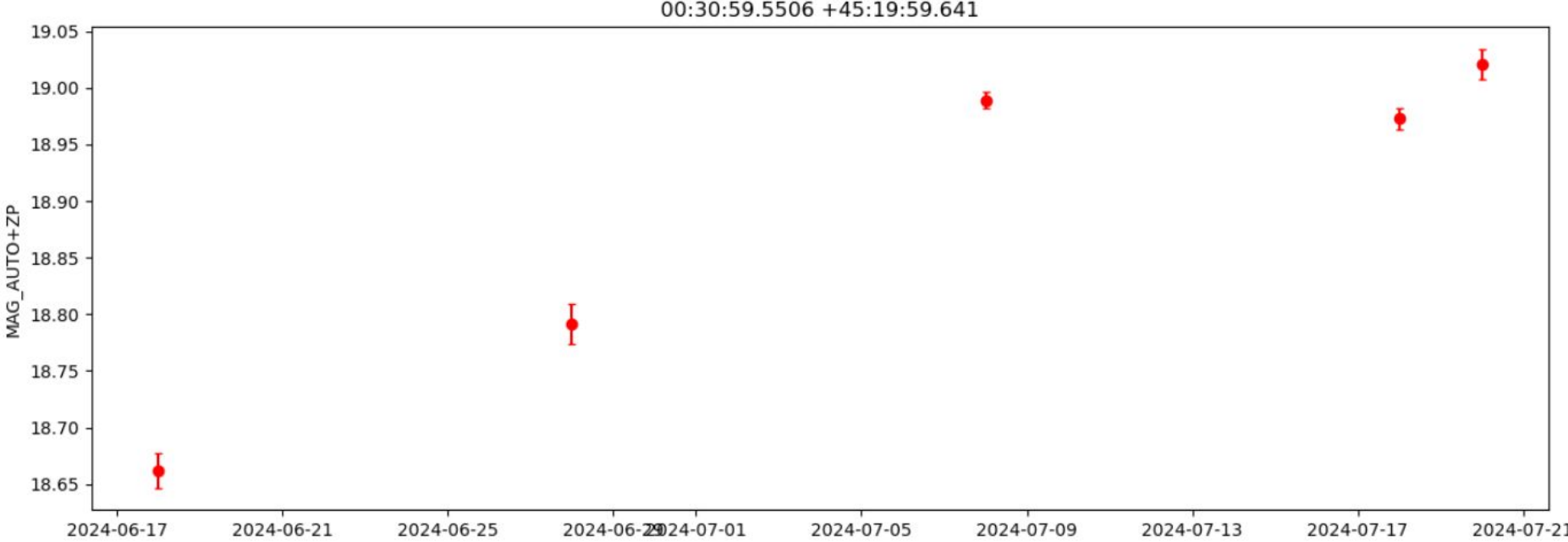


20/07



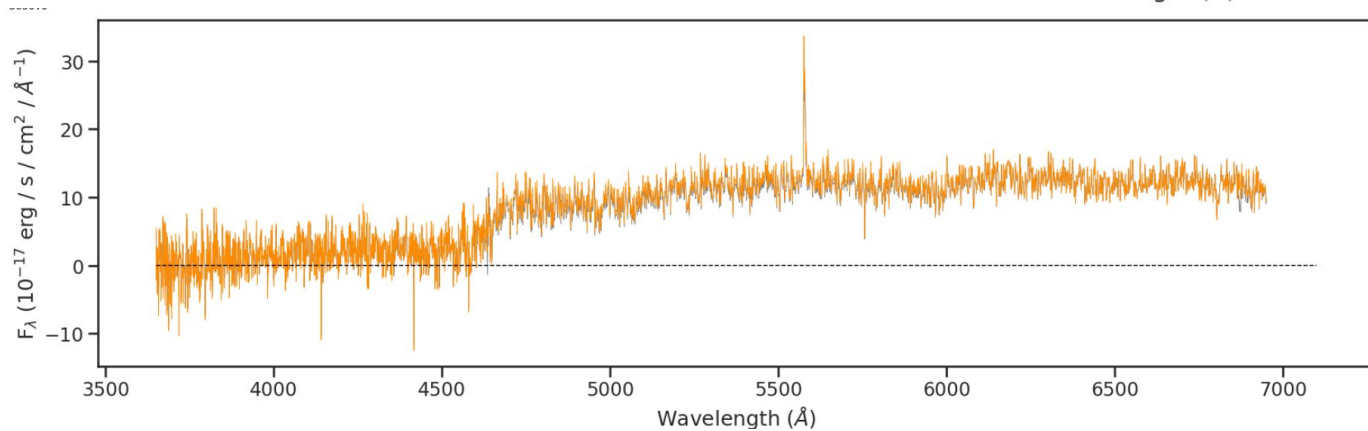
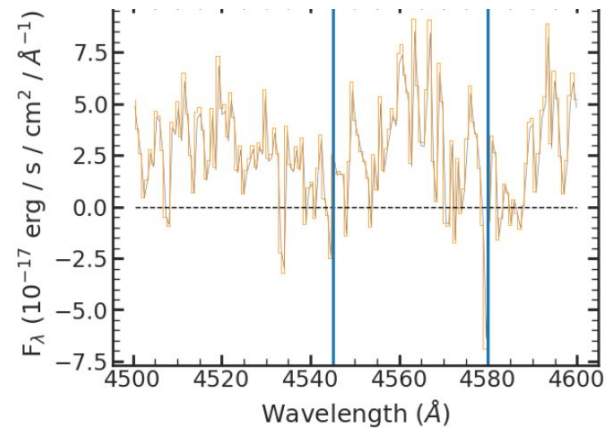
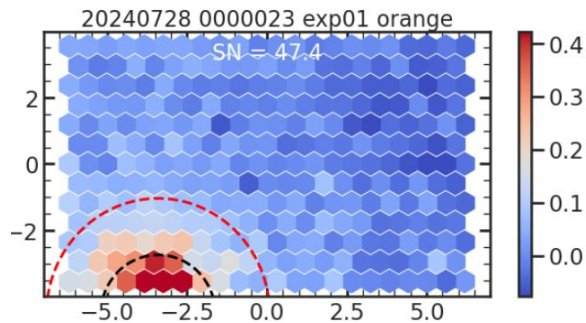
26/07

SExtractor Lightcurve

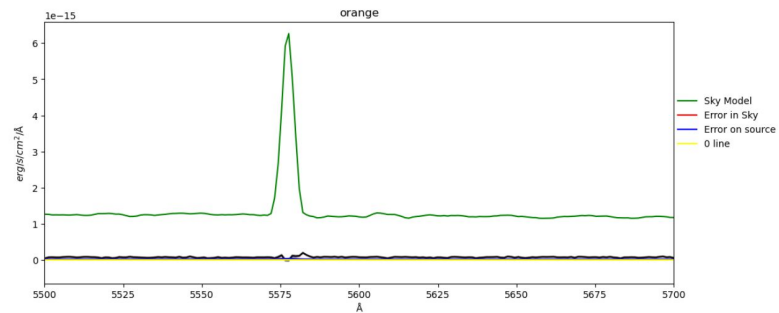
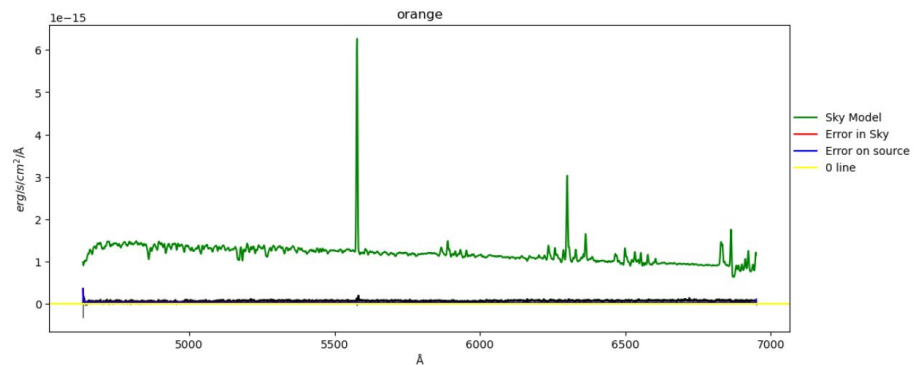
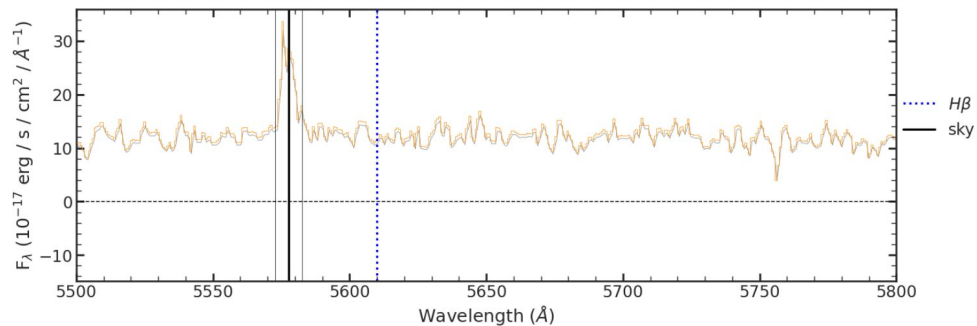


Ca-H,K, $z \approx 0.15$

HET Spectra



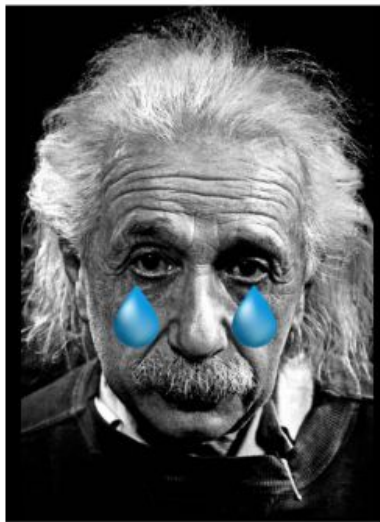
HET Spectra



Spec-z

$z=0.154 \Rightarrow D_{\text{p/w}} 695.5\text{-}757 \text{ Mpc}$

2.8 σ discrepancy from Bilby's $1420 \pm 236 \text{ Mpc}$
So excluded

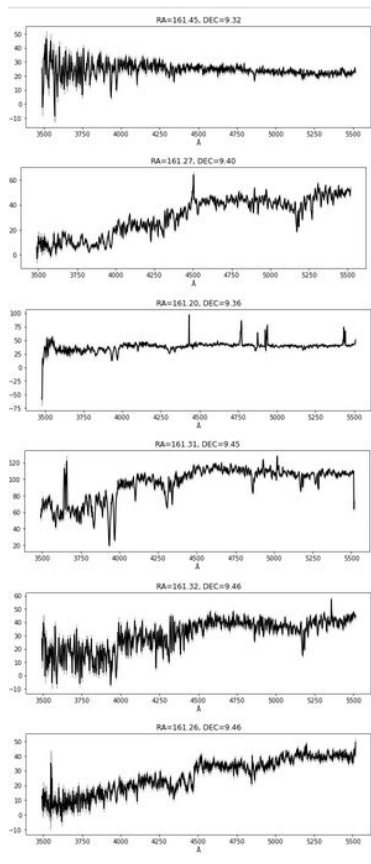


ZTF24aapjmye

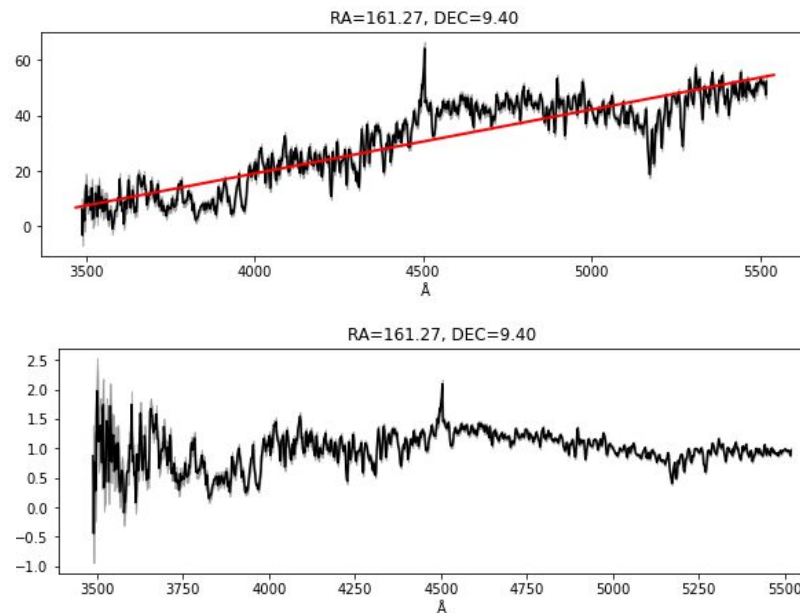
- SnIa event showed sinusoidal behaviour in the lightcurve initially
- Photo-z of the host galaxy was 0.144 ± 0.004
- Confirm using spec-z

Spec-z from VIRUS (ZTF24aapjmye)

Normalise spectrum by
continuum fit

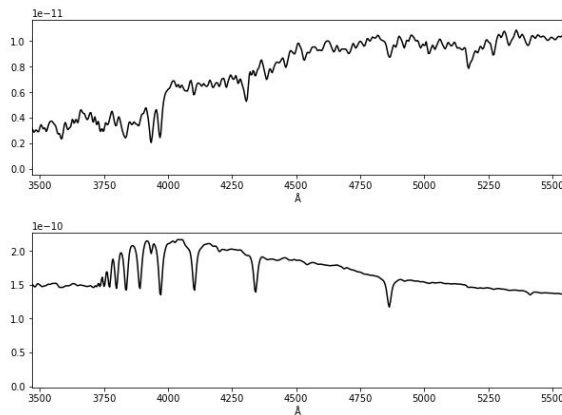
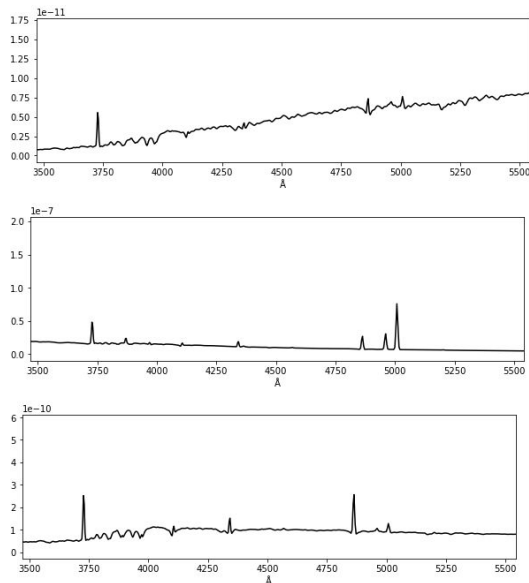


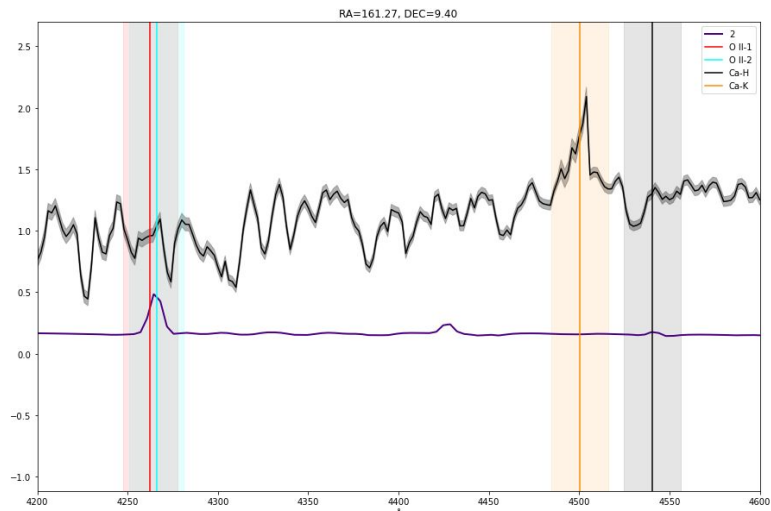
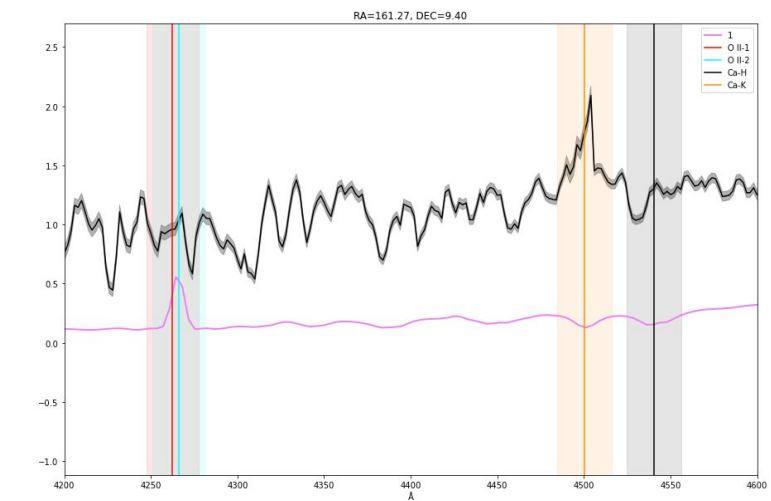
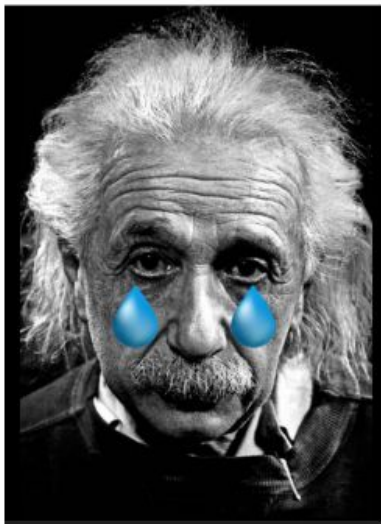
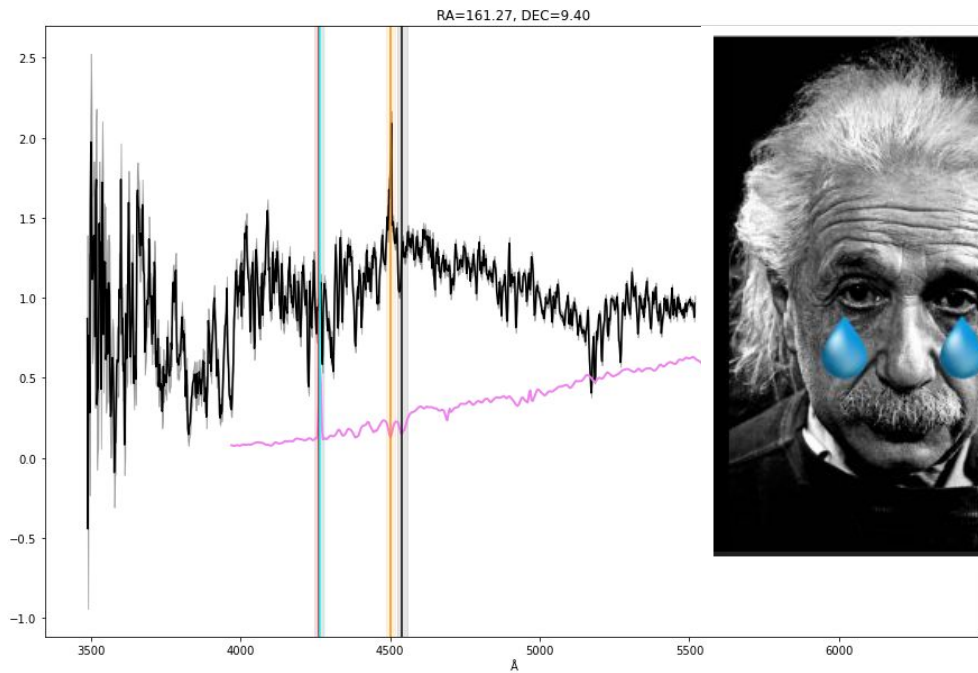
PanSTARRS
sources
within 1.5"



Templates of galaxy spectra(Blanton & Roweis 2007)

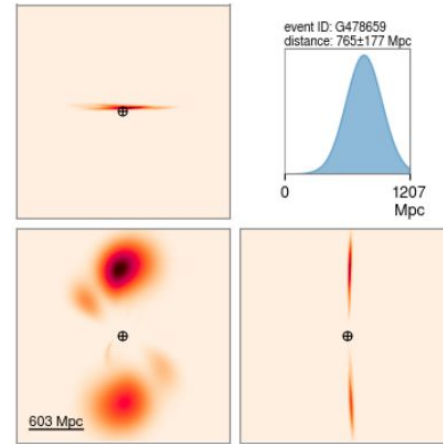
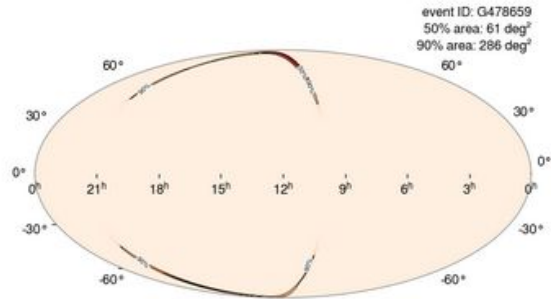
- Selection: $0 < z < 1.5$ observed in UV, IR, optical
- 485 K-corrected galaxy spectra \rightarrow “PCA” restricted to nonnegative templates
- Not model free (SPS Models, emission line models etc) and handles uncertainties



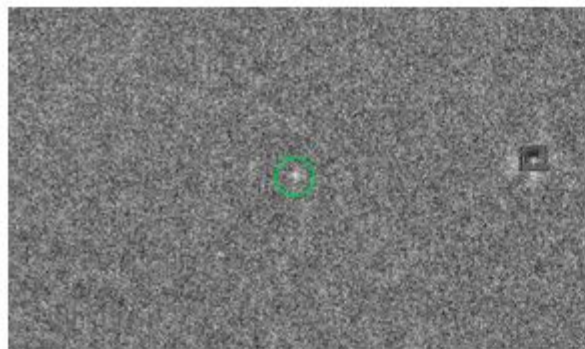


S240428dr

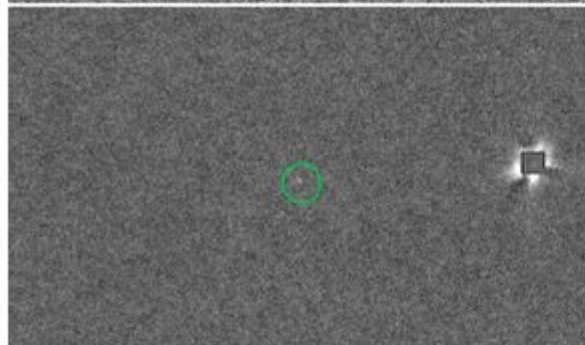
- BBH merger
- 1.3% SMBH Mass weighted probability
- $z=0.2$



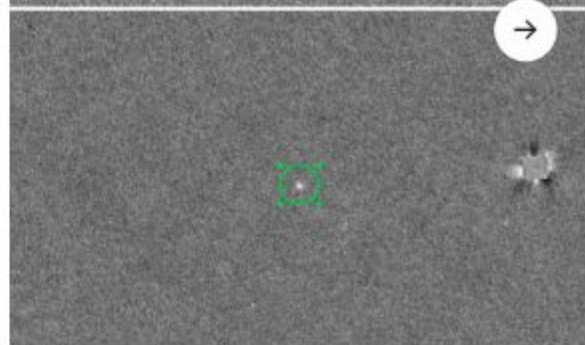
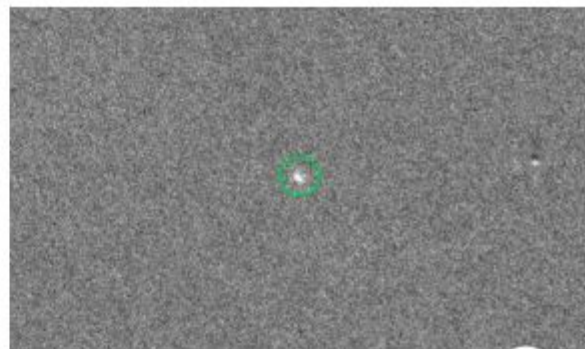
g



i

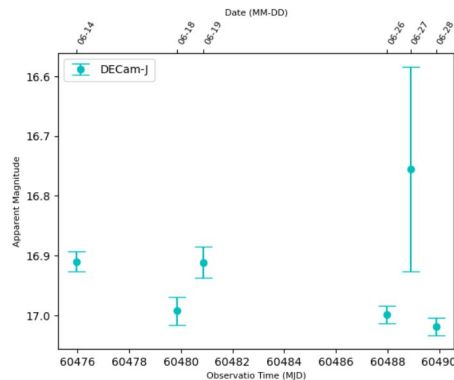
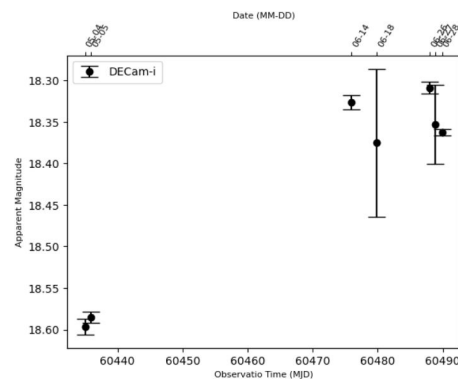
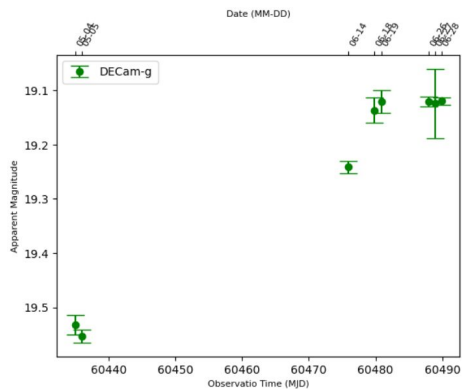


14/06



26/06

LIGHT CURVES

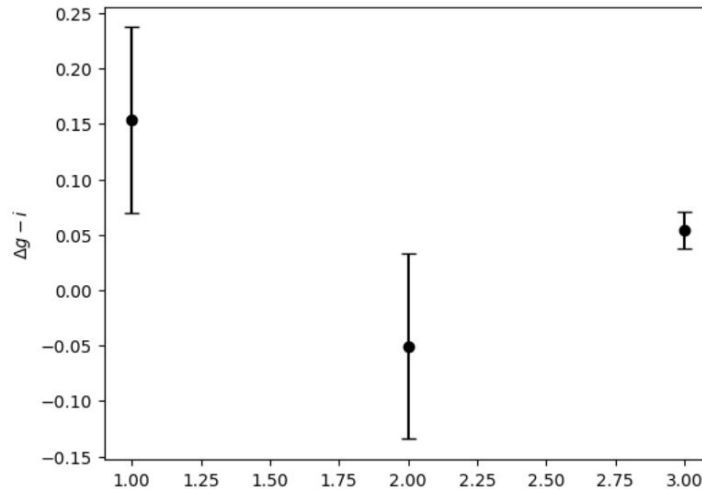
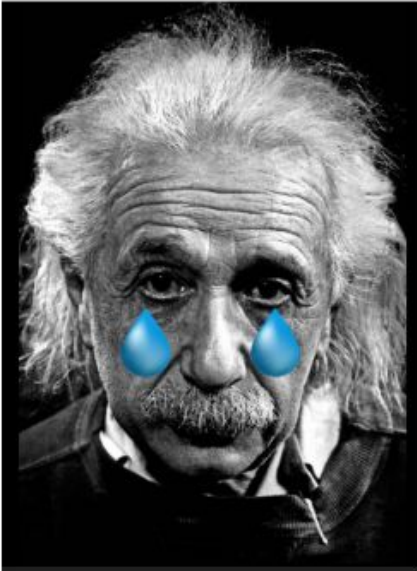


g 464 nm
i 806 nm
J 1220 nm

B/w 14/06 and 28/06,
g-J $\downarrow 0.23(6.6\sigma)$
g-i $\downarrow 0.16(9.3\sigma)$

The AGN has bluened!

Is the added flux from the flare getting the AGN bluer or is the flare itself getting bluer?



No evidence
that the flare is
getting bluer
from
differences in
coadds

Thank you!

Daniel, Arno, Malte, Julian, Ayan, Leo, Julius, Antonella

$$\underline{6-3=6}$$

QUESTIONS???

