

# Lecture Narration: Light and Shading

## Slide 1: Light and shading

“Today we begin our discussion on light and shading. This is a 17th-century still life painting where we can observe how light creates brightness and shadow on surfaces. In computer vision, we aim to understand how light is captured by a camera to form an image.”

## Slide 2: Image formation

“The process of image formation is complex. The brightness of an image pixel is determined by multiple factors: the distribution and properties of light sources, sensor properties, exposure, optics, surface reflectance properties, and surface shape and orientation. Together, these elements define what we see in an image.”

## Slide 3: Outline

“In this lecture, we will cover: a brief introduction to radiometry, the in-camera transformation of light, surface reflectance properties, diffuse and specular reflection, shape from shading, and estimating the direction of light sources.”

## Slide 4: Radiometry of image formation

“We start with two key radiometric concepts: *Irradiance*, which is the energy arriving at a surface per unit area, and *Radiance*, which is the energy carried by a ray—the power per unit area per unit solid angle. Understanding the relationship between irradiance  $E$  and radiance  $L$  is essential.”

## Slide 5: Fundamental radiometric relation

“Image irradiance  $E$  is linearly related to scene radiance  $L$ . It is proportional to the lens area and inversely proportional to the squared distance from the lens to the image plane. The irradiance also decreases as the angle  $\alpha$  between the viewing ray and optical axis increases, following a  $\cos^4$  relationship.”

## Slide 6: Fundamental radiometric relation (cont.)

“The equation  $E = \left[ \frac{\pi}{4} \left( \frac{d}{f} \right)^2 \cos^4 \alpha \right] L$  summarizes this relationship. This model is used in calibration tasks, such as calibrating a camera using a flat Lambertian surface.”

## Slide 7: From light rays to pixel values

“Here we see how light is transformed into pixel values. Scene radiance  $L$  becomes sensor irradiance  $E$ , then exposure  $X$ , then analog voltages, and finally digital values  $Z$ . The *camera response function*  $f$  maps irradiance to pixel values—critical for HDR imaging and reflectance estimation.”

## Slide 8: Outline (recap)

“Let’s recap: we’ve covered radiometry and in-camera transformations. Next, we will discuss the reflectance properties of surfaces.”

**Slide 9: Recall: Image formation**

“Again, what determines pixel brightness? Light sources, sensor properties, exposure, optics, surface reflectance, and surface shape. Today, we focus on surface reflectance.”

**Slide 10: What can happen to light when it hits a surface?**

“When light strikes a surface, several things can happen: reflection, transmission, absorption, scattering, and more. We’ll start with basic reflection models.”

**Slide 11: Basic models of reflection**

“Two fundamental reflection types: *specular reflection*, where light reflects sharply around the surface normal, and *diffuse reflection*, where light scatters equally in all directions. Think of a mirror vs. a matte wall.”

**Slide 12: Other possible effects**

“Beyond reflection, light can also be transmitted through a surface, leading to *refraction*. Transparency and translucency are important effects in realistic rendering.”

**Slide 13: Other possible effects (cont.)**

“Another important phenomenon is *subsurface scattering*, where light enters a material, scatters internally, and exits at a different point. This is common in materials like skin, wax, and marble.”

**Slide 14: Other possible effects (cont.)**

“We also have *fluorescence* and *phosphorescence*, where materials absorb light at one wavelength and re-emit it at another, sometimes with a time delay. These effects are used in special lighting and sensing applications.”

**Slide 15: Bidirectional reflectance distribution function (BRDF)**

“The BRDF describes how bright a surface appears from a given viewing direction when lit from a specific incident direction. It is a function of four angles: incident and outgoing zenith and azimuth angles  $\theta_i, \phi_i, \theta_e, \phi_e$ .”

**Slide 16: BRDF (cont.)**

“In simple terms: BRDF is the ratio of outgoing radiance to incoming irradiance. It can be incredibly complex for real-world materials, but we often simplify it for computational efficiency.”

**Slide 17: Diffuse reflectance**

“Diffuse reflectance means light scatters uniformly in all directions. Examples

include brick, matte plastic, and rough wood. This is also called *Lambertian reflectance*.”

**Slide 18: Diffuse reflectance (cont.)**

“Diffuse reflection occurs due to microscopic surface irregularities—*microfacets*—that scatter incoming light randomly. This leads to a uniform appearance from all viewing angles.”

**Slide 19: Diffuse reflectance (cont.)**

“For a fixed incident angle, the BRDF of a diffuse surface is constant. But if we change the incident angle, the total reflected energy changes. This is described by Lambert’s law.”

**Slide 20: Why do we care about diffuse reflectance?**

“Diffuse surfaces appear the same from different camera positions under the same lighting. Specular surfaces, however, change dramatically. This consistency makes diffuse reflectance easier to model and analyze.”

**Slide 21: Diffuse reflectance: Lambert’s law**

“Lambert’s law states:  $I = \rho(S \cdot N) = \rho\|S\| \cos \theta$ , where  $I$  is reflected intensity,  $\rho$  is albedo,  $S$  is the light direction vector, and  $N$  is the surface normal. This is the foundation of shape-from-shading.”

**Slide 22: Outline (continued)**

“We’ve covered reflectance. Now, we move to *shape from shading*: recovering 3D shape from brightness variations in an image.”

**Slide 23: Photometric stereo, or shape from shading**

“Can we reconstruct shape from shading cues? This question has been explored since the Renaissance—here in Luca della Robbia’s relief, shading gives us a strong sense of depth.”

**Slide 24: Photometric stereo**

“Assuming a Lambertian surface, given intensity  $I$ , can we recover light direction  $S$  and surface normal  $N$ ? And can we do it from a single image? The equation  $I = \rho(S \cdot N)$  is our starting point.”

**Slide 25: Shape from shading ambiguity**

“From a single image, shape from shading is inherently ambiguous. The same intensity can result from different combinations of surface orientation, lighting, and albedo.”

**Slide 26: Shape from shading ambiguity (cont.)**

“Humans resolve this ambiguity using prior assumptions, such as that light

comes from above. Contextual cues like atmospheric perspective also help infer shape and distance.”

**Slide 27: Outline (recap)**

“We’re now focused on shape from shading. Next, we’ll see how to solve for shape using multiple images: *photometric stereo*.”

**Slide 28: Review: Lambert’s law**

“Recall:  $I = \rho(S \cdot N)$ . Observed brightness depends on albedo  $\rho$ , light direction  $S$ , and surface normal  $N$ .”

**Slide 29: Photometric stereo assumptions**

“Photometric stereo uses multiple images under different known light directions. Assumptions: Lambertian surface, local shading, known lights, fixed camera, and orthographic projection. Goal: recover shape and albedo.”

**Slide 30: Synthetic example**

“Here’s a synthetic example: from input images under different lights, we recover albedo, normals, and finally a 3D surface model.”

**Slide 31: Image model**

“Known: light source vectors  $S_j$  and pixel intensities  $I_j$ . Unknown: surface normals  $N$  and albedo  $\rho$ . We set up a linear system per pixel.”

**Slide 32: Image model (cont.)**

“Assuming linear camera response, we write  $I_j = k\rho(N \cdot S_j) = g \cdot V_j$ , where  $g = \rho N$  and  $V_j = kS_j$ . This linearizes the problem.”

**Slide 33: Least squares problem**

“We solve for  $g$  using least squares:  $Vg = I$ . Then, albedo  $\rho = \|g\|$  and normal  $N = g/\rho$ . This gives us surface orientation and reflectance at each pixel.”

**Slide 34: Synthetic example (results)**

“Here are the recovered albedo and normal field from the synthetic example. Notice how normals capture surface curvature.”

**Slide 35: Recovering a surface from normals**

“From normals, we compute surface gradients:  $f_x = g_1/g_3$ ,  $f_y = g_2/g_3$ . Then, we integrate these gradients to recover height  $f(x, y)$ .”

**Slide 36: Recovering a surface from normals (cont.)**

“Integration can be done along a path:  $f(x, y) = \int_0^x f_x(s, 0)ds + \int_0^y f_y(x, t)dt + C$ . For robustness, average over many paths.”

**Slide 37: Recovering a surface from normals (cont.)**

“The surface must satisfy integrability:  $\partial_y(g_1/g_3) = \partial_x(g_2/g_3)$ . If not, the estimated normals are not consistent with a valid surface.”

**Slide 38: Surface recovered by integration**

“Here is the final recovered surface from integration. The height map clearly shows the 3D shape reconstructed solely from shading.”

**Slide 39: MP4 preview**

“In this video preview, we see the full pipeline: input images, recovered albedo, normals, and the final 3D model.”

**Slide 40: Limitations of basic shape from shading**

“Classical shape from shading has limitations: orthographic camera, simplistic reflectance, no shadows, no interreflections, missing data, and tricky integration. Modern methods address many of these.”

**Slide 41: Shape from shading today**

“Modern approaches use deep learning to predict surface normals, depth, albedo, and shading from a single image. Models like Stable Diffusion and others are being adapted for inverse rendering tasks.”

**Slide 42: Shape from shading today (cont.)**

“Here we see a comparison: a generated image and its decomposed components—normals, depth, albedo, shading—predicted by recent generative models.”

**Slide 43: Outline (final part)**

“Finally, we will discuss estimating the direction of light sources from a single image—useful for forensics and scene understanding.”

**Slide 44: Finding the direction of the light source**

“Given known surface normals  $N$  and intensities  $I$ , we can solve for light direction  $S$  using a linear system:  $NS = I$ . This requires normals from multiple points.”

**Slide 45: Finding light direction (occluding contour)**

“A special case: points on the occluding contour have  $N_z = 0$ . This simplifies the problem to estimating only the  $S_x$  and  $S_y$  components.”

**Slide 46: Linear system for occluding contour**

“On the occluding contour, the system reduces to  $N_{xy}S_{xy} = I$ . This is easier to solve and provides a good estimate of the projected light direction.”

**Slide 47: Finding the direction of the light source (example)**

“Here’s a real example: from an image with known geometry, we estimate the light source direction. This is useful for virtual object insertion and lighting consistency checks.”

**Slide 48: Application: Detecting composite photos**

“Lighting inconsistency is a telltale sign of photo tampering. By estimating light direction, we can detect if objects in a composite image are lit differently—exposing forgeries.”

**Slide 49: DeepFake detection today**

“Modern DeepFake detection also uses geometric and lighting cues. Here, generative models sometimes make mistakes in shadows, perspective, and light coherence—which can be automatically detected.”