

Physical Optics (Diffraction)

1(a) Given that 12^{th} order maximum is seen in a Young's double slit experiment with Na-light of wavelength $\lambda_1 = 6000 \text{ \AA}$

We know the condition for n^{th} order ^{interference} maxima is $d \sin \theta = n\lambda$, for normal incidence, $d =$ distance between two slits & θ is the angle made by the n^{th} order maxima with the normal of the slit at the middle of the source.

If we replace the monochromatic source with $\lambda_2 = 4800 \text{ \AA}$ then at the same place say m^{th} order maxima is seen whose condition is $d \sin \theta = m\lambda_2$

$$\therefore d \sin \theta = n\lambda_1 = m\lambda_2 \text{ yields } m = \frac{n\lambda_1}{\lambda_2} = \frac{12 \times 6000}{4800} = 15$$

So 15^{th} order interference maxima will be visible with 4800 \AA source at the same place where 12^{th} order interference maxima was seen through a 6000 \AA source.

(b) Given that the central maxima contains 9 fringes, meaning 4 equidistant bright bands on both side of the central bright line.

As the central fringe is bright, we have $\frac{2b}{a} + 1 = 9$ where $d = a + b =$ distance between two slits & a is the slit width. $\therefore \frac{2b}{a} = 8$ or $b = 4a$ & $d = 5a$

From the condition of missing order $\frac{d}{a} = \frac{n}{\beta}$, $n = n^{\text{th}}$ interference maxima & β is β^{th} diffraction minima, so $n = 5\beta$, so 5, 10, 15 - etc order of interference maxima are absent

which corresponds to 1, 2, 3 ... orders of diffraction dark bands. This means that within first & second minima (diffraction dark band) only 6th, 7th, 8th & 9th order interference maxima (fringes) will be visible.

We know the condition for n^{th} order interference maxima is $d \sin \theta_n = n \lambda$. $\therefore \theta_n = \sin^{-1} \left(\frac{n \lambda}{d} \right) = \sin^{-1} \left(\frac{n \lambda}{5a} \right)$.

So the angles where fringes will appear will be

$$\theta_6 = \sin^{-1} \left(\frac{6 \lambda}{5a} \right), \theta_7 = \sin^{-1} \left(\frac{7 \lambda}{5a} \right), \theta_8 = \sin^{-1} \left(\frac{8 \lambda}{5a} \right) \text{ and } \theta_9 = \sin^{-1} \left(\frac{9 \lambda}{5a} \right)$$

respectively.

(2) (a) Given, for the m^{th} order (say) diffraction spectra of a wavelength $\lambda_1 = 540 \text{ nm} = 540 \times 10^{-9} \text{ m}$ superimposes with the $(m+1)^{\text{th}}$ order of another wavelength $\lambda_2 = 405 \text{ nm} = 405 \times 10^{-9} \text{ m}$ at normal incidence of a plane transmission grating. d being the grating element.

From the condition of secondary maxima, we have

$$d \sin \theta = m \lambda_1 = (m+1) \lambda_2 \quad \therefore m = \frac{\lambda_2}{\lambda_1 - \lambda_2} = \frac{405}{540 - 405} = 3$$

Given, angle of diffraction $\theta = 30^\circ$.

$$\therefore d = \frac{m \lambda_1}{\sin \theta} = \frac{3 \times 540 \times 10^{-9}}{1/2} = \underline{\underline{324 \times 10^{-8} \text{ m}}}$$

So the grating element is $324 \times 10^{-8} \text{ m}$.

(b) Given for a plane transmission grating,

slit width $a = 0.001 \text{ mm}$

opaque space $b = 0.002 \text{ mm}$ & wavelength of monochromatic

source $\lambda = 500 \text{ nm}$.

As $\sin 90^\circ = 1$ so the maximum angle of diffraction is 90° .
 Let m be the maximum number of order of spectrum that can be observed.

from the condition of secondary maxima, we have

$$d \sin \theta = m\lambda \quad \text{where grating element } d = a + b = 0.003 \text{ mm}$$

$$\therefore m = \frac{d}{\lambda} = \frac{0.003 \times 10^{-3}}{500 \times 10^{-9}} = \underline{6}$$

Now we need to account for the missing grating spectra that happens for m^{th} principal maxima (condition $d \sin \theta = m\lambda$) to coincide with n^{th} minimum intensity (condition $a \sin \theta = n\lambda$)

$$\therefore \frac{m\lambda}{d} = \frac{n\lambda}{a} \quad \therefore m = \frac{a+b}{a} n = \frac{0.003}{0.001} n = 3n$$

So 3^{rd} , 6^{th} order will be missing from the grating spectra and only upto 5^{th} order spectrum is visible.

(c) Given, the number of rulings of the grating $N = 820/\text{cm}$.

$$\therefore \text{Grating element } d = \frac{1}{N} = \frac{1}{820} \text{ cm.}$$

Wavelength of Na-D lines $\lambda_{D_1} = 5890 \text{ \AA}$, $\lambda_{D_2} = 5896 \text{ \AA}$

Let the least-width of the grating be w .

We know that the resolving power of a grating is given by

$$\frac{\lambda}{d\lambda} = Nn, \quad n = \text{order of the spectrum; } d\lambda = \lambda_{D_2} - \lambda_{D_1} \text{ \AA.}$$

$$= 2. \quad M = \# \text{ of rulings in the grating.}$$

$$\therefore M = \frac{\lambda}{n d\lambda} = \frac{(\lambda_{D_1} + \lambda_{D_2})/2}{2(\lambda_{D_2} - \lambda_{D_1})} = \frac{5893}{2 \times 6} \approx \underline{491}$$

$$\therefore \text{Least-width of the grating } w = Md = 491 \times \frac{1}{820} \text{ cm}$$

$$= \underline{\underline{0.599 \text{ cm.}}}$$

to resolve the Na-D line spectra.

(d) Given, the number of rulings of the grating $N = 3000/\text{cm}$

$$\therefore \text{Grating constant } d = \frac{1}{N} = \frac{1}{3000} \text{ cm} = 3.333 \times 10^{-4} \text{ cm}$$

To obtain the direction of the 1st order ($m=1$) of the Na-D lines $\lambda_{D_1} = 5890 \text{ \AA}$ and $\lambda_{D_2} = 5896 \text{ \AA}$ with average wavelength

$$\lambda = \frac{1}{2}(\lambda_{D_1} + \lambda_{D_2}) = 5893 \text{ \AA}, \text{ we know the condition of secondary maxima } d \sin \theta = m\lambda, \text{ we have}$$

$$\theta = \sin^{-1} \left(\frac{m\lambda}{d} \right) = \sin^{-1} \left(\frac{1 \times 5893 \times 10^{-8}}{3.333 \times 10^{-4}} \right) = \underline{\underline{10.18^\circ}}$$

This is the required direction of the 1st order grating spectra.

Now from the relation for resolving power $\frac{\lambda}{d\lambda} = mN$, where N is the number of rulings in the grating and m is the order of the spectrum, we have

$$N = \frac{\lambda}{m d\lambda} = \frac{5893}{1 \times 6} = 982.16 \approx \underline{\underline{983}}$$

$$\text{So the least-width of the grating is } Nd = \frac{983}{3000} \text{ cm} = \underline{\underline{0.32 \text{ cm}}}$$

(e) Given, least-width of grating $Nd = 4 \text{ cm}$ and number of rulings $N = 4000 \text{ lines/cm}$. Wavelength of monochromatic source $\lambda = 5900 \text{ \AA}$, $N = \#$ of rulings in the grating.

$$\therefore \text{Grating constant } d = \frac{1}{N} = \frac{1}{4000} \text{ cm}$$

$$\therefore Nd = 4 \Rightarrow \frac{N}{4000} = 4 \Rightarrow N = 16000$$

While order of spectrum $m=1$, the required resolving power $\frac{\lambda}{d\lambda} = mN = 1 \times 16000 = \underline{\underline{16000}}$.

$$\text{while } \left(\frac{\lambda}{d\lambda} \right)_{\text{Na}} = \frac{5893}{6} = \underline{\underline{982}} \text{ which is way smaller than}$$

the computed above resolving power, so this grating can separate the Na-doublet.

(f) Width of one transmission grating G_1 is $w_1 = M_1 d_1 = 3 \text{ cm}$ with # of rulings $M_1 = 3000$. Width of another grating G_2 is $w_2 = M_2 d_2 = 2 \text{ cm}$ with # of rulings $M_2 = 2000$.

Resolving power of a grating $\frac{\lambda}{\Delta\lambda} = mM$

So for the same order ratio of R.P. of G_1 & G_2 is $mM_1 : mM_2$
 $= 3000 : 2000 = \underline{\underline{3:2}}$.

(3)(a) Given, two points on Moon is resolved by a telescope of diameter $D = 500 \text{ cm} = 5 \text{ m}$. Most sensitive wavelength of light to eye is $\lambda = 5500 \text{ \AA} = 5500 \times 10^{-10} \text{ m}$, and distance of moon to surface of Earth $d = 3.8 \times 10^5 \text{ km} = 3.8 \times 10^8 \text{ m}$.

\therefore The angular separation (inverse of resolving power) of the two points is $d\theta = \frac{1.22\lambda}{D} = \frac{1.22 \times 5500 \times 10^{-10}}{5} = 1.342 \times 10^{-7} \text{ rad}$

So from the geometry of arclength, separation between the points on Moon will be $d d\theta = 3.8 \times 10^8 \times 1.342 \times 10^{-7} \text{ m}$
 $= \underline{\underline{51 \text{ m}}}$.

(b) Given, the resolving power (R.P.) of eye is $\frac{1}{1'} \text{ arc}$. Let D_o , D_{ep} , D_e be the diameters of the telescope objective, telescope eyepiece and human eye. Wavelength $\lambda = 6000 \text{ \AA}$.

Given, the magnification $\frac{D_o}{D_{ep}} = 80$ where $D_o = 8 \text{ cm}$

$\therefore D_{ep} = 0.1 \text{ cm} = 1 \text{ mm}$. We know R.P. of eye $\frac{1}{d\theta_e} = \frac{D_e}{1.22\lambda}$

$$\therefore \frac{D_e}{1.22\lambda} = \frac{1}{1'} = \frac{1}{60 \times 180} \text{ rad.} \quad \therefore D_e = \frac{1.22 \times 6000 \times 10^{-10}}{\pi / (60 \times 180)}$$

$$= \underline{\underline{2.516 \text{ mm}}}$$

For telescope, the angular resolving power is $d\theta_T = \frac{1.22\lambda}{D_o}$

$$= \frac{1.22 \times 6000 \times 10^{-8}}{8} \text{ rad} = \frac{1.22 \times 6000 \times 10^{-8}}{8} \times \frac{180}{\pi}^\circ =$$

$$\frac{1.22 \times 6000 \times 10^{-8}}{8} \times \frac{180}{\pi} \times 3600 \text{ sec} = \underline{\underline{1.887''}}$$

while $d\theta_e = 1' = 60''$ and magnification = 80, so normally human eye's resolving power is $\frac{60''}{80} = \underline{\underline{0.75''}}$ which is way smaller than the Telescope's resolving power $1.887''$.

Therefore, the resolvable angular separation of stars using this refracting telescope is $1.887''$.

(4) We know that n^{th} dark ring in Newton's ring interference experiment is $r_n = \sqrt{n\lambda R}$ where λ is wavelength of monochromatic source and R is the radius of curvature of the planoconvex lens = 2m.

So radius of the 1^{st} dark ring which is also the radius of the 1^{st} dark ring of the zone plate is $r_1 = \sqrt{2\lambda}$.

Now we know that the focal length of m^{th} order of a zone-plate is $f_m = \frac{r_m^2}{m\lambda}$. So at 1^{st} order, $f_1 = \frac{r_1^2}{\lambda} = \frac{2\lambda}{\lambda} = \underline{\underline{2\text{m}}}$.

∴ The first focal length of the zone plate is 2 metres.