

Clouds in The Cloud: Supplementary Material

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Abstract. In [1] we describe simulations and an appearance consistency measure that we used. The essence is in the paper, while in this supplemental document we give more details and results.

1 Simulations

In Sec. 8 of [1] we describe simulations, which enabled us to quantitatively assess performance of 3D cloud recovery. Here we give more details.

An atmosphere over $8\text{km} \times 8\text{km}$ was produced using off-the-shelf UCLA large eddy simulation (LES) [2]. It created a field of liquid water content (LWC) per voxel, which represented clouds between heights of 500m to 1500m. Fig. 1(a) shows a $2.5 \times 2.5 \times 1.5\text{km}^3$ subset of this atmosphere. The LWC field then affected radiative transfer,

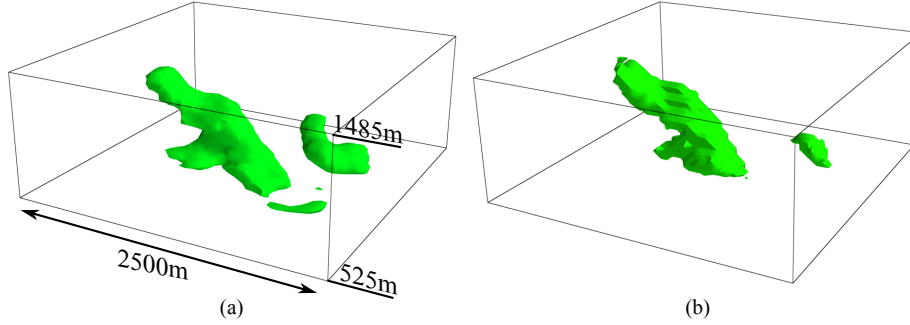


Fig. 1. (a) The water liquid content distribution, showing clouds in the atmosphere. This is a major input to SHDOM calculations. (b) Reconstructed clouds.

calculated using the discrete ordinate spherical-harmonic method (SHDOM) [3]. The solar zenith angle was set to $\frac{\pi}{4}$.

Figs. 2(a,b) show images from two out of 100 synthesized cameras. The atmosphere above the cameras is marked by red rectangles in Figs. 2(a,b). The applied static Sun-blocker is tailored to solar trajectories in the latitude of Copenhagen. Figs. 2(c,d) show images using the simulated Sun-blocker. We reconstructed the cloud distribution in the atmosphere above the cameras. Fig. 1(b) shows one reconstruction result, as described in [1].

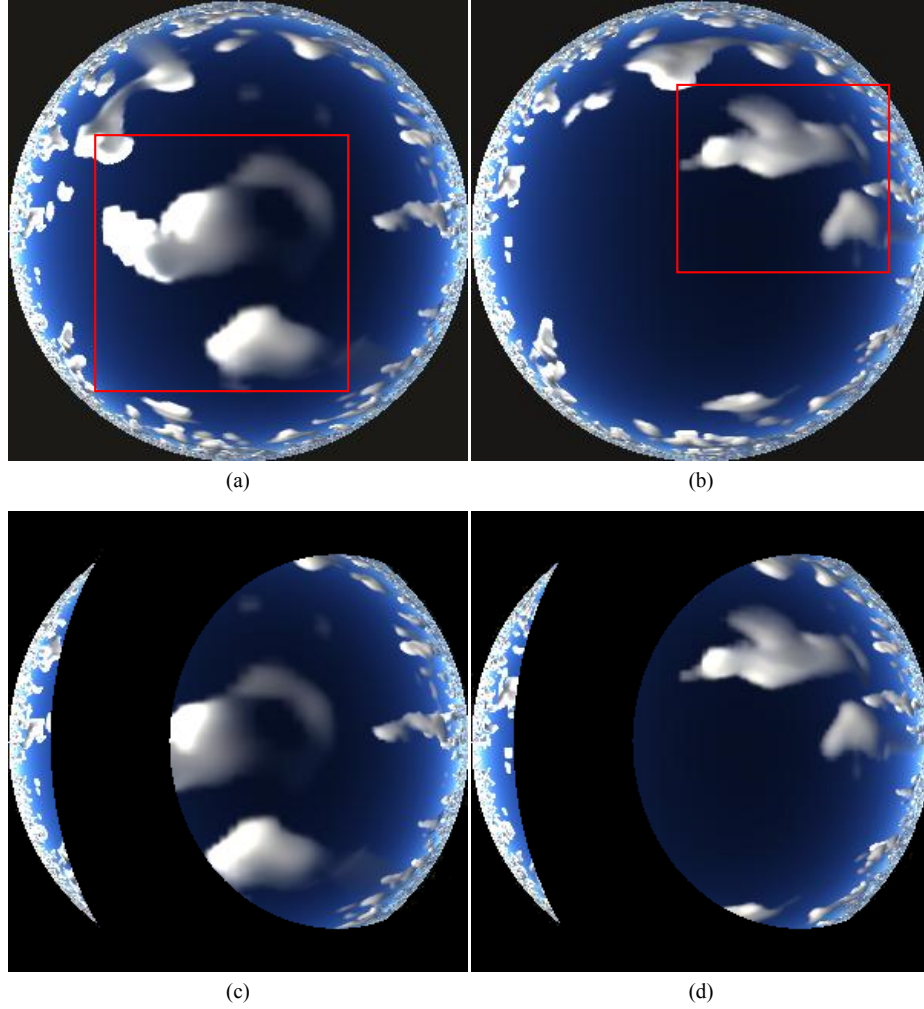


Fig. 2. Images synthesised using SHDOM and used in the reconstruction algorithm: (a,b) Two images (out of 100) of the same scene from different cameras. The red-marked rectangles define a region of interest, right above the network. (c,d) The same images, taken under a static sun-blocker.

2 Appearance Consistency Score

In Sec. 8 of [1], we describe a matching criterion we used to demonstrate space carving of clouds. Here we give more details about it, to ease reproducibility.

Basic space carving is usually based on pointwise photo-consistency. As we explain in Sec. 8 of [1], in clouds photo-consistency is sometimes not sufficiently discriminative. Dense SIFT [4] can be effective for non-rigid image matching. Similarly, dense SIFT can be incorporated into a space-carving scheme, to represent structural appearance of a projected voxel. Voxel k projects to a subset of the rays $\rho_k \subset \mathcal{R}$. For each ray $r \in \rho_k$, we measure the radiance $\hat{I}_t^\chi(r)$ (after correcting radiometric non-uniformities), where $\chi \in \{R, G, B\}$ is the color-channel index. This forms a photometric vector $\mathbf{v}^{\text{phot}}(r, t) = [\hat{I}_t^R(r), \hat{I}_t^G(r), \hat{I}_t^B(r)]$, which is scaled globally to a unit bound.

In addition, each ray r corresponds to a pixel \mathbf{x} . From a patch around \mathbf{x} , a 128-bin SIFT descriptor is extracted [5], yielding a vector $\mathbf{v}^{\text{SIFT}}(r, t)$. Each such vector is normalized such that $\|\mathbf{v}^{\text{SIFT}}(r, t)\|_1 = 1$. Overall, the feature vector we use is $\mathbf{v}(r, t) = [\mathbf{v}^{\text{phot}}(r, t), \mu \mathbf{v}^{\text{SIFT}}(r, t)]$. Here μ determines the weight of SIFT vs. photometric descriptors. We used $\mu = 1/40$ throughout the experiment. Notice that the SIFT vector contains 40 times more values, so this choice of μ gives essentially equal power to the photometric and structural properties of the patch.

Element q of $\mathbf{v}(r, t)$ is $v_q(r, t)$. The values of this element, for all rays that intersect voxel k , form the set $\mathcal{V}_q(k, t) \equiv \{v_q(r, t)\}_{r \in \rho_k}$. Across viewpoints, the measured variance in this set is $\text{VAR}[\mathcal{V}_q(k, t)]$. Appearance-consistency is based on summing the variances in the different components of the feature vectors:

$$P_k = \exp \left(- \sum_{q=1}^{131} \{ \text{VAR}[\mathcal{V}_q(k, t)] \} / \sigma^2 \right). \quad (1)$$

Here σ^2 is a global scale parameter which is set to $\sigma^2 = 0.1$ throughout our experiments.

References

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