

# Assignments

Paper: Orbital mechanics  
MCD560

Name: Amit Mondal

Admission no. : 23DP0044

Prob. 1. solution of the universal Kepler's equation using Newton's method

Prob. 2. Calculation of the Lagrange coefficients  $f$  and  $g$  and their time derivatives

Prob. 3. Gauss's method of preliminary orbit determination with iterative improvement

Prob. 4. calculation of the state vector of the spacecraft/planets for Bi-elliptic transfer

```

1: #include <stdio.h>
2: #include <math.h>
3:
4: #define EPSILON 1e-10 // Tolerance for convergence
5: #define MAX_ITER 1000 // Maximum number of iterations
6:
7: // Function to calculate f(E)
8: double f(double E, double e, double M) {
9:     return E - e * sin(E) - M;
10: }
11:
12: // Derivative of f(E)
13: double f_prime(double E, double e) {
14:     return 1 - e * cos(E);
15: }
16:
17: // Newton's method for solving Kepler's equation
18: double kepler(double M, double e) {
19:     double E = M; // Initial guess
20:
21:     for (int i = 0; i < MAX_ITER; i++) {
22:         double next_E = E - f(E, e, M) / f_prime(E, e);
23:
24:         // Check for convergence
25:         if (fabs(next_E - E) < EPSILON)
26:             return next_E;
27:
28:         E = next_E;
29:     }
30:
31:     // If no convergence, return NaN
32:     return NAN;
33: }
34:
35: int main() {
36:     double M, e;
37:     printf("Enter mean anomaly (M): ");
38:     scanf("%lf", &M);
39:     printf("Enter eccentricity (e): ");
40:     scanf("%lf", &e);
41:
42:     // Call kepler function
43:     double E = kepler(M, e);
44:
45:     if (!isnan(E)) {
46:         printf("Eccentric anomaly (E) = %lf\n", E);
47:     } else {
48:         printf("Unable to find solution.\n");
49:     }
50:
51:     return 0;
52: }
53:

```

Output:

Enter mean anomaly (M): 30

Enter eccentricity (e): 0.33

Eccentric anomaly (E) = 29.674777

-----

Process exited after 18.06 seconds with return value 0

Press any key to continue . . .

```

1: #include <stdio.h>
2: #include <math.h>
3:
4: #define PI 3.14159265358979323846
5:
6: // Function to calculate the eccentric anomaly E using Newton's method
7: double eccentric_anomaly(double M, double e) {
8:     double E = M; // Initial guess
9:
10:    for (int i = 0; i < 1000; i++) {
11:        double next_E = E - (E - e * sin(E) - M) / (1 - e * cos(E));
12:
13:        // Check for convergence
14:        if (fabs(next_E - E) < 1e-10)
15:            return next_E;
16:
17:        E = next_E;
18:    }
19:
20:    // If no convergence, return NaN
21:    return NAN;
22: }
23:
24: // Function to calculate f and g and their time derivatives
25: void lagrange_coefficients(double M, double e, double a, double mu, double *f, double *g, double *fdot, double *gdot) {
26:     double E = eccentric_anomaly(M, e);
27:     double r = a * (1 - e * cos(E));
28:     double sqrt_a_mu = sqrt(a * mu);
29:
30:     *f = 1 - r / a * cos(E);
31:     *g = r / sqrt_a_mu * sin(E);
32:     *fdot = -sqrt(mu / a) * sin(E) / (1 - e * cos(E));
33:     *gdot = sqrt(a / mu) * (cos(E) - e);
34: }
35:
36: int main() {
37:     double M, e, a, mu;
38:     printf("Enter mean anomaly (M): ");
39:     scanf("%lf", &M);
40:     printf("Enter eccentricity (e): ");
41:     scanf("%lf", &e);
42:     printf("Enter semi-major axis (a): ");
43:     scanf("%lf", &a);
44:     printf("Enter gravitational parameter (mu): ");
45:     scanf("%lf", &mu);
46:
47:     double f, g, fdot, gdot;
48:     lagrange_coefficients(M, e, a, mu, &f, &g, &fdot, &gdot);
49:
50:     printf("f = %lf\n", f);
51:     printf("g = %lf\n", g);
52:     printf("fdot = %lf\n", fdot);
53:     printf("gdot = %lf\n", gdot);
54:
55:     return 0;

```

```
56: }  
57:
```

Output

Enter mean anomaly (M): 30

Enter eccentricity (e): .35

Enter semi-major axis (a): 45

Enter gravitational parameter ( $\mu$ ): 1.33

$f = 1.200124$

$g = -6.088773$

$\dot{f} = 0.158445$

$\dot{g} = -3.128146$

-----

Process exited after 92.95 seconds with return value 0

Press any key to continue . . .

```

1: #include <stdio.h>
2: #include <math.h>
3:
4: #define N_OBSERVATIONS 3 // Number of observations
5: #define G 6.67430e-11    // Gravitational constant
6: #define M_EARTH 5.972e24 // Mass of the Earth
7:
8: typedef struct {
9:     double t; // Time of observation
10:    double r[3]; // Position vector
11: } Observation;
12:
13: double dot_product(double v1[], double v2[]) {
14:     double result = 0;
15:     for (int i = 0; i < 3; i++) {
16:         result += v1[i] * v2[i];
17:     }
18:     return result;
19: }
20:
21: void cross_product(double v1[], double v2[], double result[]) {
22:     result[0] = v1[1] * v2[2] - v1[2] * v2[1];
23:     result[1] = v1[2] * v2[0] - v1[0] * v2[2];
24:     result[2] = v1[0] * v2[1] - v1[1] * v2[0];
25: }
26:
27: void gauss_method(Observation obs[], double r[], double v[], double tol, int max_iter) {
28:     // Initial guess for position vector
29:     double r0[3];
30:     for (int i = 0; i < 3; i++) {
31:         r0[i] = r[i];
32:     }
33:
34:     // Iterative improvement
35:     for (int iter = 0; iter < max_iter; iter++) {
36:         double delta_r[3] = {0};
37:
38:         // Calculate delta_r
39:         for (int i = 0; i < N_OBSERVATIONS; i++) {
40:             double rho_i = sqrt(pow(obs[i].r[0] - r0[0], 2) + pow(obs[i].r[1] - r0[1], 2) + pow(obs[i].r[2] - r0[2], 2));
41:             double f = obs[i].t - dot_product(obs[i].r, r0) / rho_i;
42:             for (int j = 0; j < 3; j++) {
43:                 delta_r[j] += f * (obs[i].r[j] / rho_i);
44:             }
45:         }
46:
47:         // Update r0
48:         for (int i = 0; i < 3; i++) {
49:             r0[i] += delta_r[i];
50:         }
51:
52:         // Check for convergence
53:         double delta_mag = sqrt(pow(delta_r[0], 2) + pow(delta_r[1], 2) + pow(delta_r[2], 2));
54:         if (delta_mag < tol) {
55:             break;

```

```

56:     }
57: }
58:
59: // Final position vector
60: for (int i = 0; i < 3; i++) {
61:     r[i] = r0[i];
62: }
63: }
64:
65: int main() {
66:     // Observations (time in seconds, position vector in meters)
67:     Observation obs[N_OBSERVATIONS] = {
68:         {0, {1500e3, 1500e3, 1000e3}},
69:         {3600, {-1500e3, -1500e3, -1000e3}},
70:         {7200, {0, 0, 0}}
71:     };
72:
73:     // Initial guess for position and velocity vectors
74:     double r[3] = {1000e3, 1000e3, 1000e3};
75:     double v[3] = {0};
76:
77:     // Set tolerance and maximum iterations for Gauss method
78:     double tol = 1e-9;
79:     int max_iter = 1000;
80:
81:     // Perform Gauss method
82:     gauss_method(obs, r, v, tol, max_iter);
83:
84:     // Output final position vector
85:     printf("Final position vector (meters): (%lf, %lf, %lf)\n", r[0], r[1], r[2]);
86:
87:     return 0;
88: }
89:

```



## Output

Final position vector (meters): (23650954.748252, 23650954.748252, 16100636.498835)

-----

Process exited after 0.04542 seconds with return value 0

Press any key to continue . . .

```

1: #include <stdio.h>
2: #include <math.h>
3:
4: #define G 6.67430e-11    // Gravitational constant
5: #define M_EARTH 5.972e24 // Mass of the Earth
6:
7: typedef struct {
8:     double r; // Distance from center of the Earth
9:     double v; // Velocity magnitude
10:    double theta; // Angle (in radians) with respect to the positive x-axis
11: } StateVector;
12:
13: StateVector calculate_bi_elliptic_transfer(double r_initial, double r_final, double delta_v1,
14:     StateVector sv;
15:
16:     // Calculate velocity at the initial orbit
17:     double v_initial = sqrt(G * M_EARTH * (2 / r_initial - 1 / (r_initial + r_final)));
18:
19:     // Calculate state vector after the first burn
20:     double v1x = delta_v1 * cos(theta);
21:     double v1y = delta_v1 * sin(theta);
22:     sv.r = r_initial;
23:     sv.v = sqrt(v_initial * v_initial + delta_v1 * delta_v1);
24:     sv.theta = atan2(v1y, v_initial + v1x);
25:
26:     // Calculate state vector after the second burn
27:     double v_final = sqrt(G * M_EARTH * (2 / r_final - 1 / (r_initial + r_final)));
28:     double v2x = delta_v2 * cos(M_PI - theta);
29:     double v2y = delta_v2 * sin(M_PI - theta);
30:     sv.r = r_final;
31:     sv.v = sqrt(v_final * v_final + delta_v2 * delta_v2);
32:     sv.theta = atan2(v2y, v_final + v2x);
33:
34:     // Calculate state vector after the third burn
35:     double v3x = delta_v3 * cos(theta);
36:     double v3y = delta_v3 * sin(theta);
37:     sv.r = r_final;
38:     sv.v = sqrt(v_final * v_final + delta_v3 * delta_v3);
39:     sv.theta = atan2(v3y, v_final + v3x);
40:
41:     return sv;
42: }
43:
44: int main() {
45:     // Initial and final radii of the orbits (in meters)
46:     double r_initial = 200000; // Initial orbit radius (e.g., 200 km)
47:     double r_final = 400000;   // Final orbit radius (e.g., 400 km)
48:
49:     // Delta v values for the three burns (in m/s)
50:     double delta_v1 = 200; // Delta v for first burn
51:     double delta_v2 = 400; // Delta v for second burn
52:     double delta_v3 = 200; // Delta v for third burn
53:
54:     // Angle (in radians) between initial and final radii
55:     double theta = M_PI / 4; // Example angle of 45 degrees

```

```
56:
57:     // Calculate state vector for bi-elliptic transfer
58:     StateVector sv = calculate_bi_elliptic_transfer(r_initial, r_final, delta_v1, delta_v2, de
59:
60:     // Output state vector
61:     printf("State vector after bi-elliptic transfer:\n");
62:     printf("Radius: %.2f meters\n", sv.r);
63:     printf("Velocity: %.2f m/s\n", sv.v);
64:     printf("Angle (with respect to positive x-axis): %.2f radians\n", sv.theta);
65:
66:     return 0;
67: }
68:
```

## Output

State vector after bi-elliptic transfer:

Radius: 400000.00 meters

Velocity: 36450.93 m/s

Angle (with respect to positive x-axis): 0.00 radians

-----

Process exited after 0.04185 seconds with return value 0

Press any key to continue . . .