



# Wideband Technique

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# Narrowband v/s Wideband Technique



- In the recent years, the radio telescopes are upgraded with **Wideband receivers** and associated backends to obtain precise ToAs
- Profile evolution and ISM effects (Dispersion Measure and Scattering) makes it difficult to attain better precision with traditional narrowband method
- **Pennucci et al. 2018**, first provided an algorithm for the simultaneous measurement of DMs and ToAs from wideband pulsar data, called the "wideband timing technique".

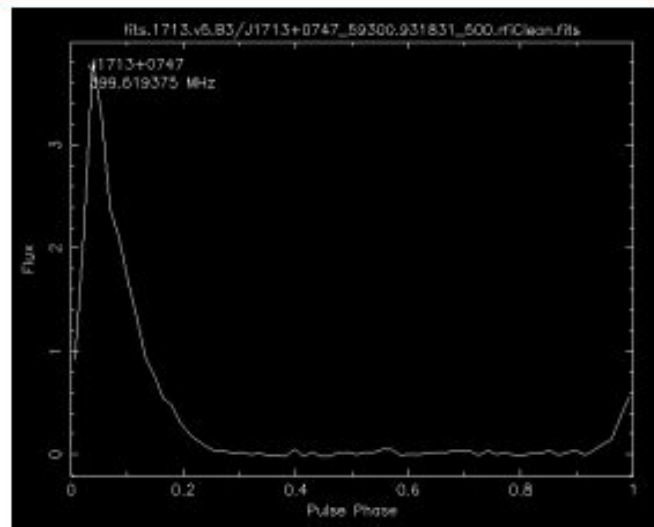
## Introduction - Narrowband

High S/N profiles are added and then are scrunched in frequency and time.

We fit gaussians to it and create an analytical template (**paas**).

This analytical template is used to generate ToAs by cross correlating the template with the folded profile of each epoch (**pat**).

Dispersion measure (DM) is obtained using frequency resolved ToAs over several epochs and then fitting for DM.



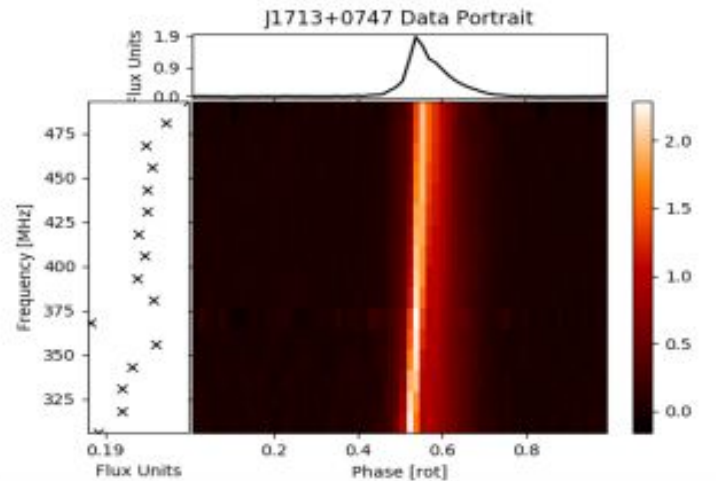
$$\chi^2(\phi, a) = \sum_k \frac{|d_k - a t_k e^{-2\pi k \phi}|^2}{\sigma_k^2}$$

# Introduction - Wideband

Averaging of the data for each pulsar is performed, while maintaining frequency resolution to arrive at a high S/N mean “portrait” (a collection of aligned mean pulse profiles across a contiguous frequency band). ([ppalign](#))<sup>1 2</sup>

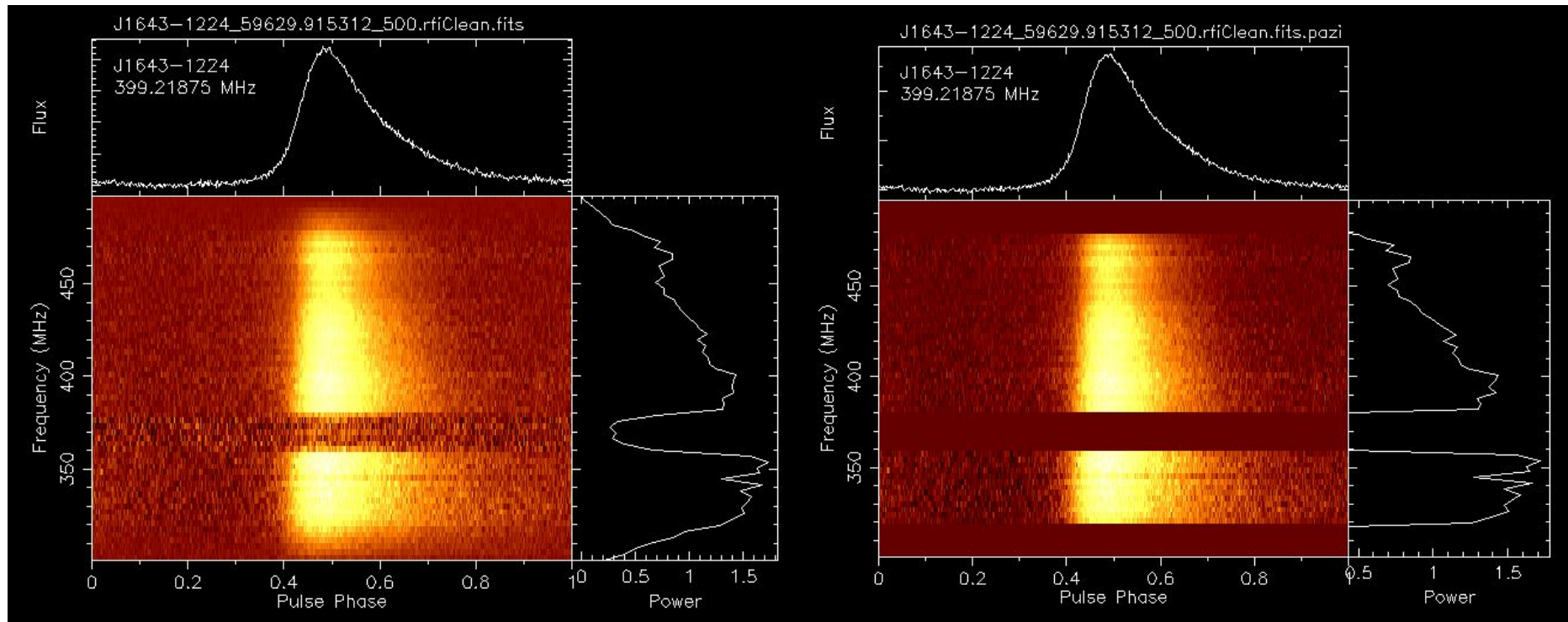
A principal component analysis is performed on the average portrait.

The most significant eigenvectors and the mean profile are smoothed to become noise-free basis functions (“eigenprofiles”). ([ppspline](#))



$$\chi^2 = \sum_{n,k} \frac{|d_{nk} - a_n t_{nk} e^{-2\pi i k \phi_n}|^2}{\sigma_n^2}.$$

# Removing RFI channels





## ppalign

- Aligns multiple data files and add them to create a “portrait”
- In InPTA with uGMRT we are only using single data file to create portrait
- It is stored as port file (.port)

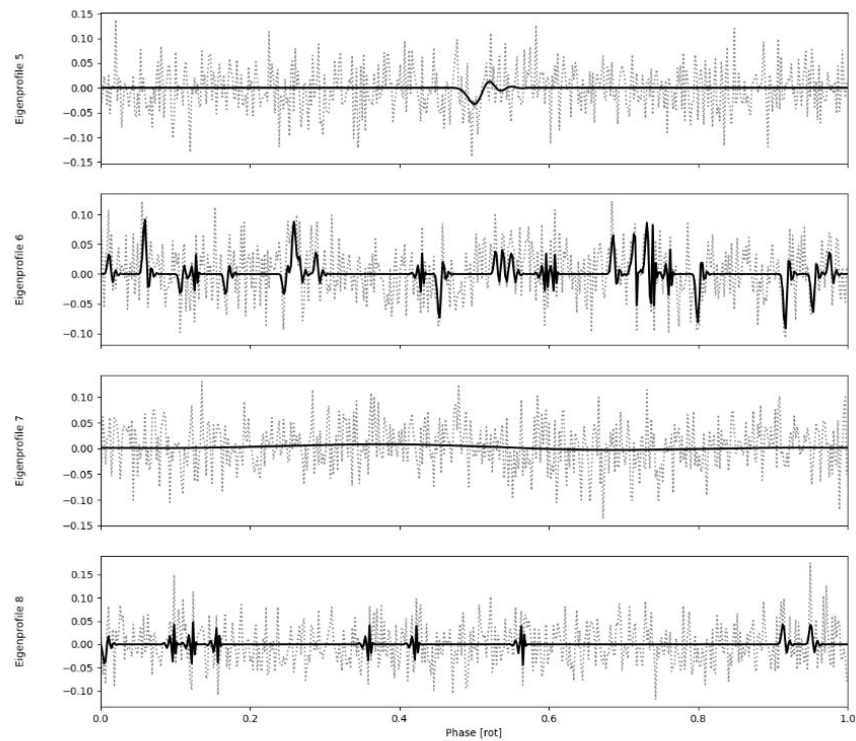
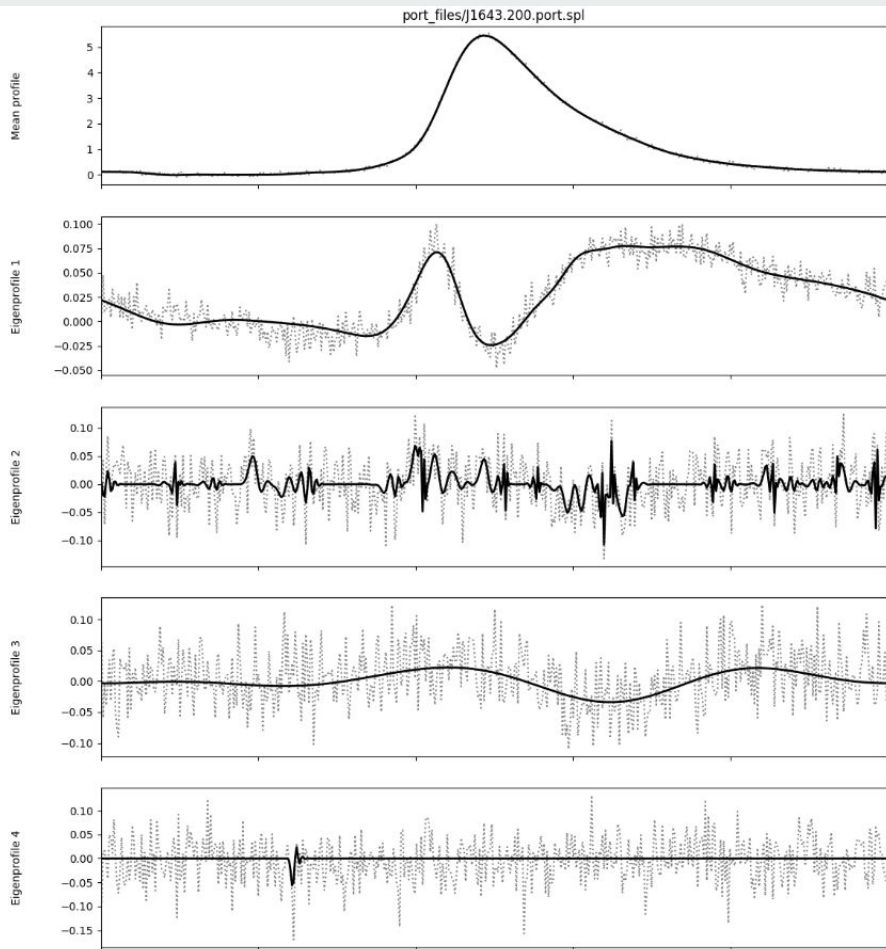


## pp spline

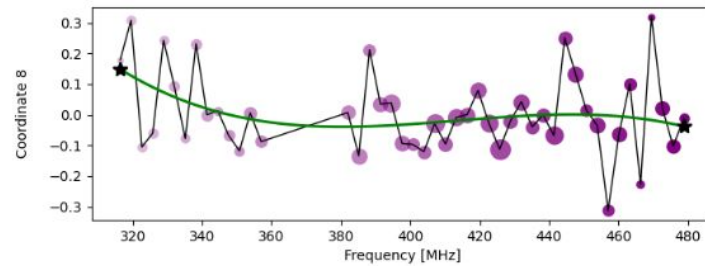
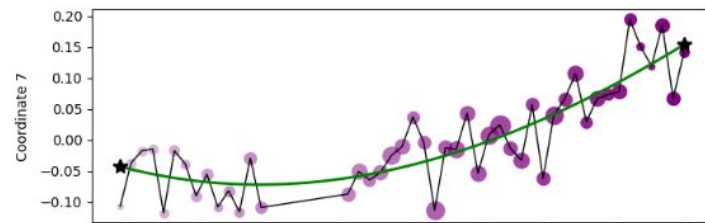
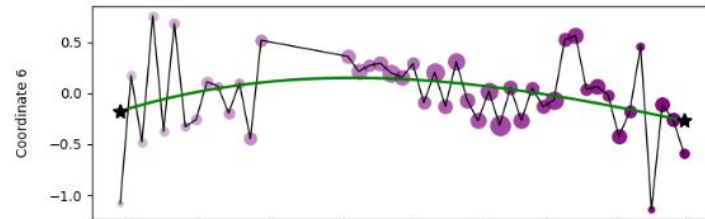
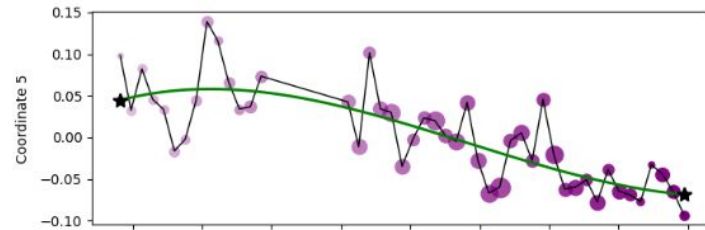
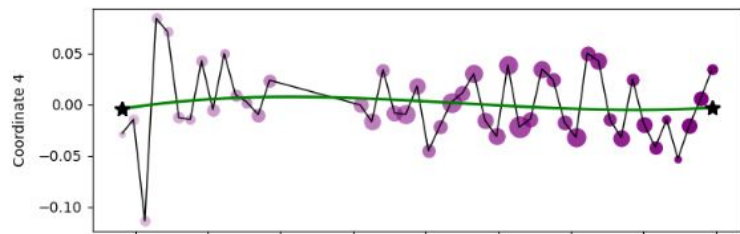
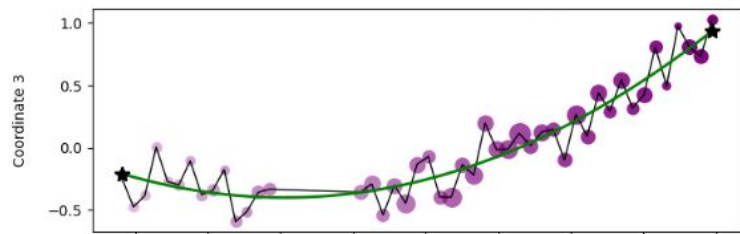
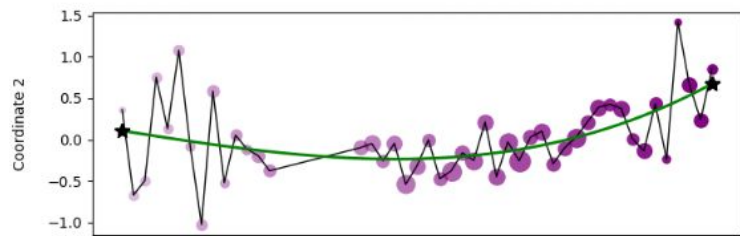
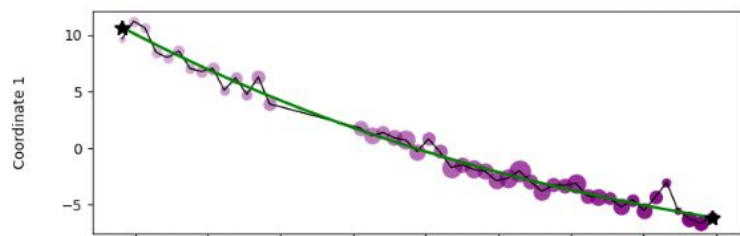
- ppspline performs PCA on our portrait
- It decomposes our data into eigenvectors or eigen profiles
- It produces an analytical template

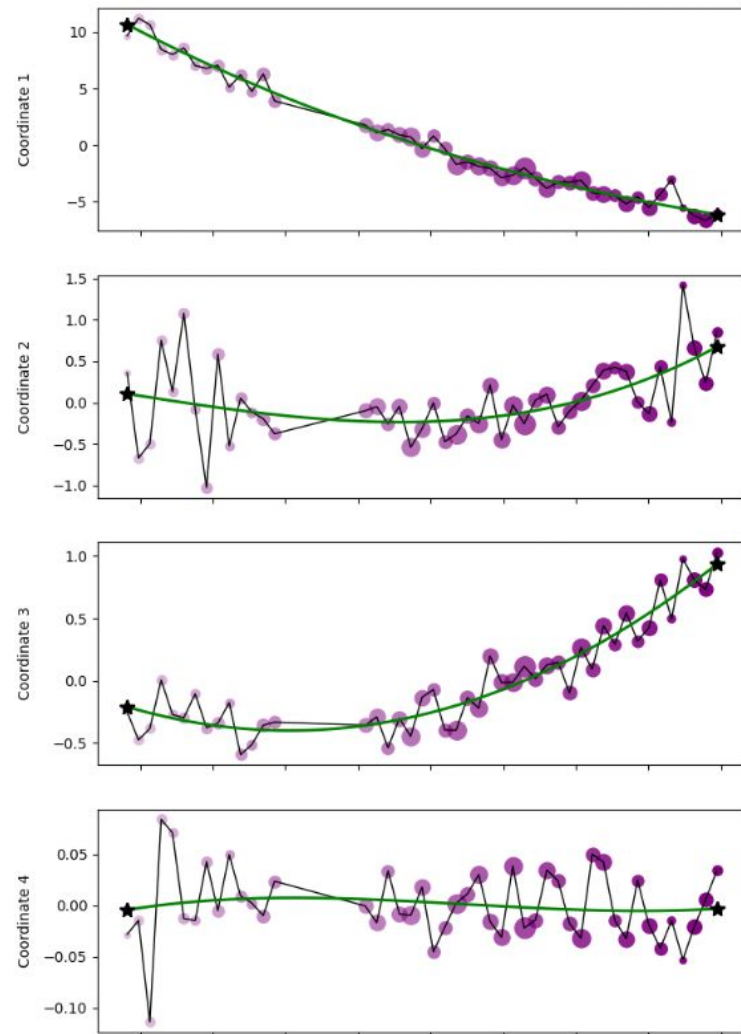
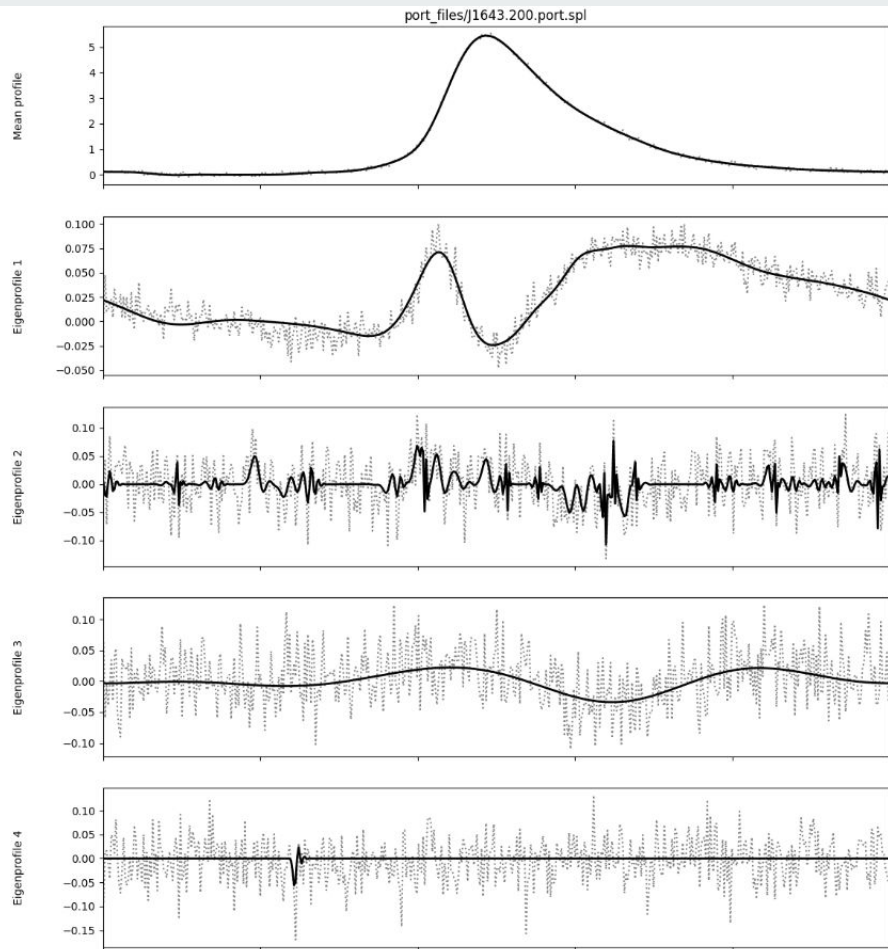
$$T(\nu) = \sum_{i=1}^{n_{\text{eig}}} B_i(\nu) \hat{e}_i + \tilde{p}.$$

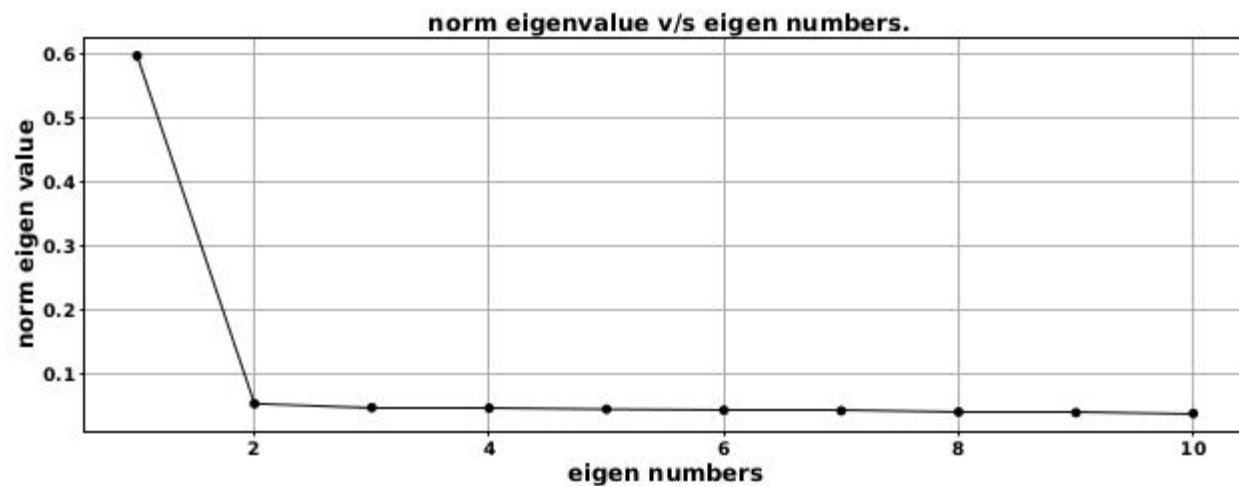
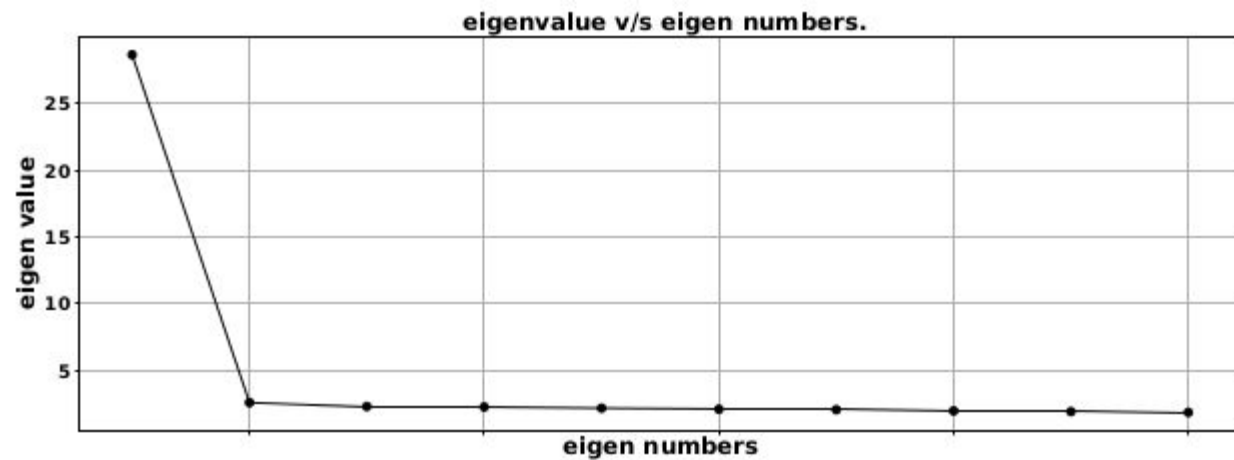


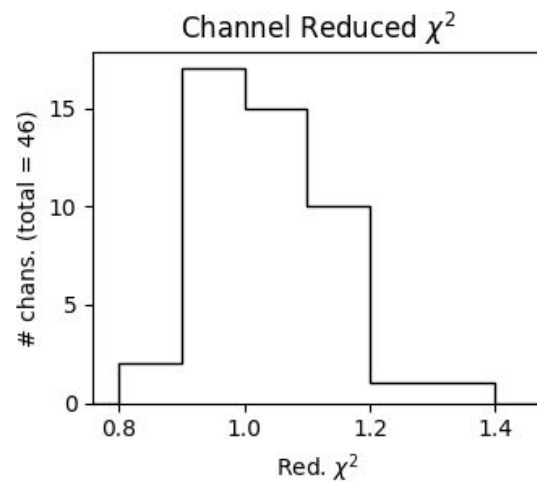
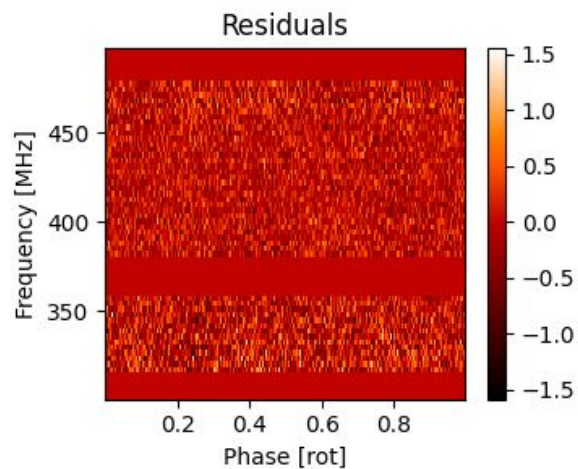
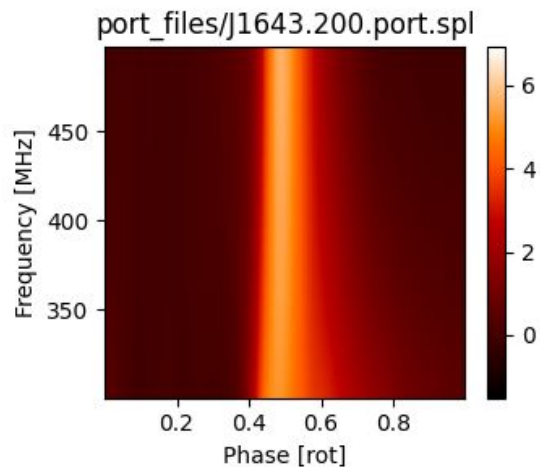
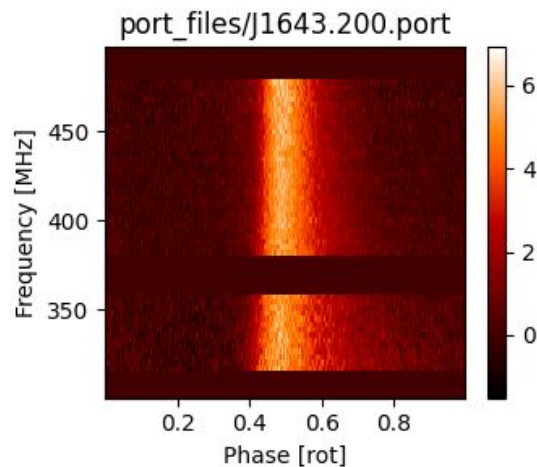














## pptoas

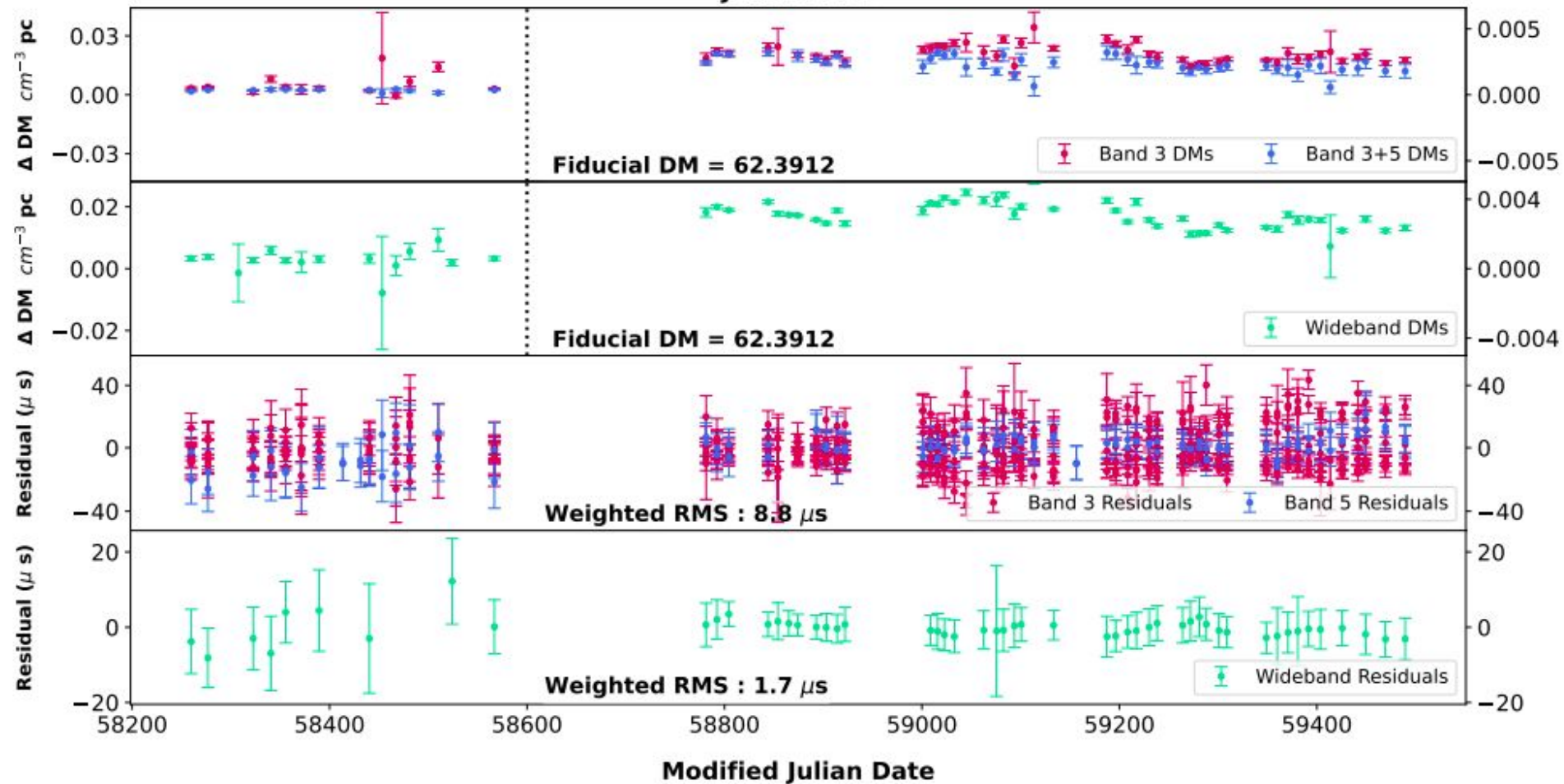
- Pptoas cross correlate the analytic template with data in fourier domain
- It does global fit for dm and TOA i.e. only one TOA per epoch
- It creates a tim file which provides all the TOA and DM for each epoch

$$\chi^2(\phi_n, a_n) = \sum_{n,k} \frac{|d_{nk} - a_n p_{nk} e^{-2\pi i k \phi_n}|^2}{\sigma_n'^2}. \quad \phi_n = \phi_{\text{ref}}^\circ + \frac{K \times \text{DM}}{P_s} (v_n^{-2} - v_{\text{ref}}^{-2}),$$



J1643/data\_200bw/J1643-1224\_58781.324655\_500.rfiClean.fits 428.09437203 58781.341990540106443 0.816 gmrt -pp\_dm 62.3951239 -pp\_dme 0.000267  
2 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 2999.766 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 431.523 -gof 1.238  
J1643/data\_200bw/J1643-1224\_58792.348162\_500.rfiClean.fits 397.31820864 58792.365543652314069 0.720 gmrt -pp\_dm 62.3946182 -pp\_dme 0.000113  
0 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3000.102 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 510.192 -gof 1.281  
J1643/data\_200bw/J1643-1224\_58804.373146\_500.rfiClean.fits 398.49711261 58804.392222071303465 0.458 gmrt -pp\_dm 62.3945001 -pp\_dme 0.000073  
4 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3299.869 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 797.579 -gof 1.600  
J1643/data\_200bw/J1643-1224\_58844.132989\_500.rfiClean.fits 416.08135893 58844.152060024433903 0.457 gmrt -pp\_dm 62.3952944 -pp\_dme 0.000106  
5 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3299.743 -fratio 1.653 -tmplt J1643/  
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5 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3300.078 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 526.203 -gof 1.183  
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7 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3300.078 -fratio 1.653 -tmplt J1643/  
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6 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3300.078 -fratio 1.653 -tmplt J1643/  
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3 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3300.078 -fratio 1.653 -tmplt J1643/  
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port\_files/J1643.spl -snr 694.112 -gof 1.396  
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2 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3300.078 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 600.564 -gof 1.277  
J1643/data\_200bw/J1643-1224\_59000.868037\_500.rfiClean.fits 414.12801749 59000.886979284932955 1.263 gmrt -pp\_dm 62.3944144 -pp\_dme 0.000230  
2 -be GWB -fe uGMRT\_B3 -f uGMRT\_B3\_GWB -nbin 128 -nch 64 -nchx 64 -bw 196.875 -chbw 3.125 -subint 0 -tobs 3281.569 -fratio 1.653 -tmplt J1643/  
port\_files/J1643.spl -snr 285.097 -gof 1.173

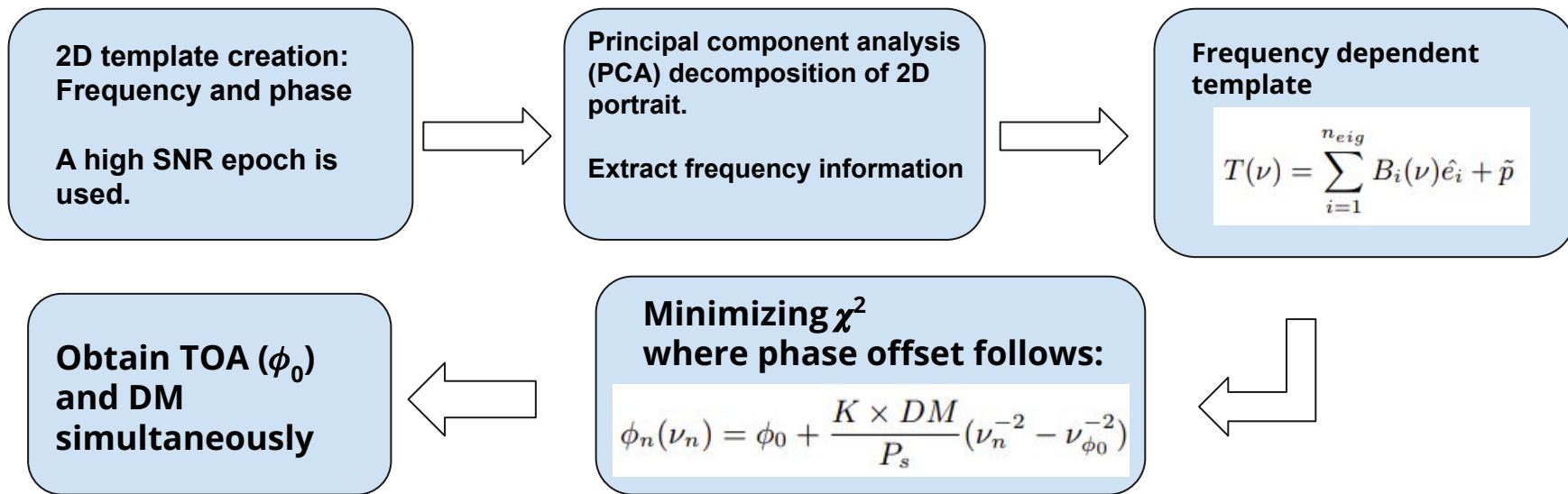
# J1643-1224





# Summary

- Wideband technique (Pennucci et. al. 2014, Pennucci 2019, Nobleson et. al. 2022)





**Thank You**