Internal Gravity Waves & Cloud Evolution in Substellar Atmospheres

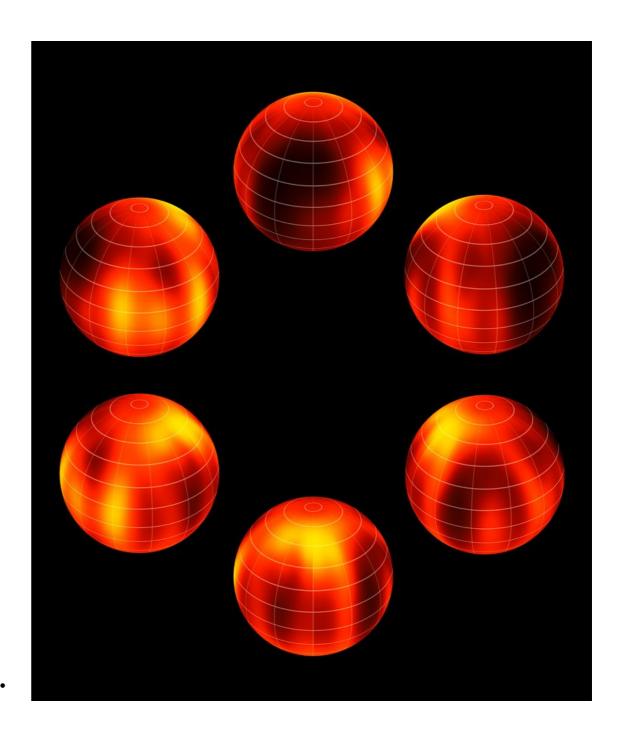
NUMERICAL SIMULATIONS & FUTURE CAPABILITIES

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Clouds in Brown Dwarfs

- Photometric variability observed in brown dwarfs, mapped (Crossfield+ 2014)
- Current atmosphere models suggest variability is due to dust clouds (Helling+ 2008)
- ► Models suggest internal gravity waves impact cloud evolution (Freytag+ 2010)

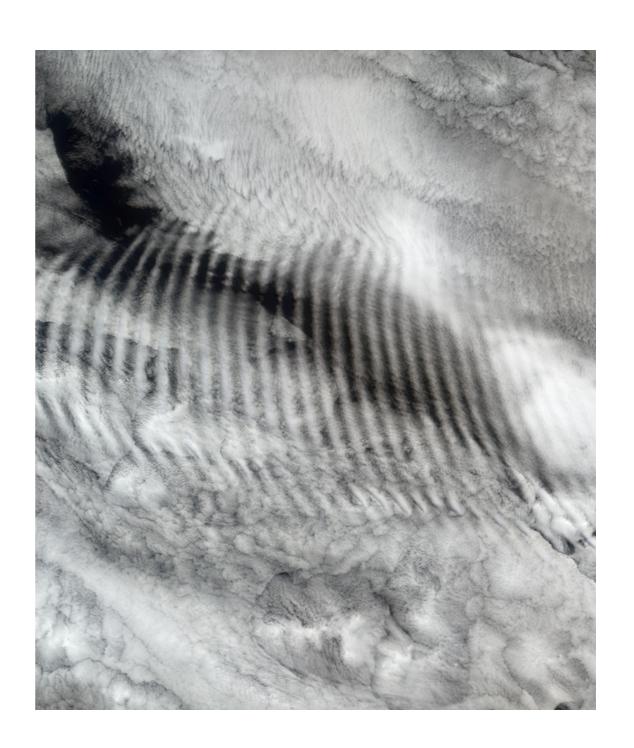
Cloud map of Luhman-16B Crossfield et al. 2014, Nature, 505, 654.



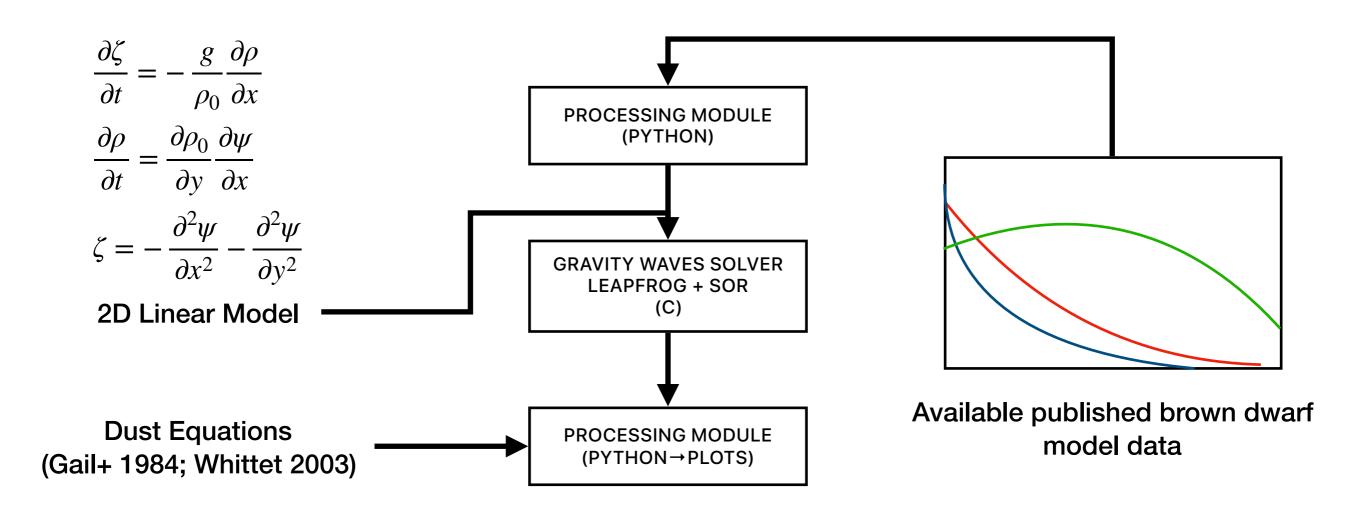
Internal Gravity Waves

- Density or velocity perturbations create atmospheric oscillations;
- ▶ IGW lead to banded structures in clouds on Earth & Solar System Planets
- ► Can similar structures teach us about brown dwarf atmospheres?

Gravity Waves Ripple over Marine Stratocumulus Clouds NASA/GSFC/LaRC/JPL, MISR Team



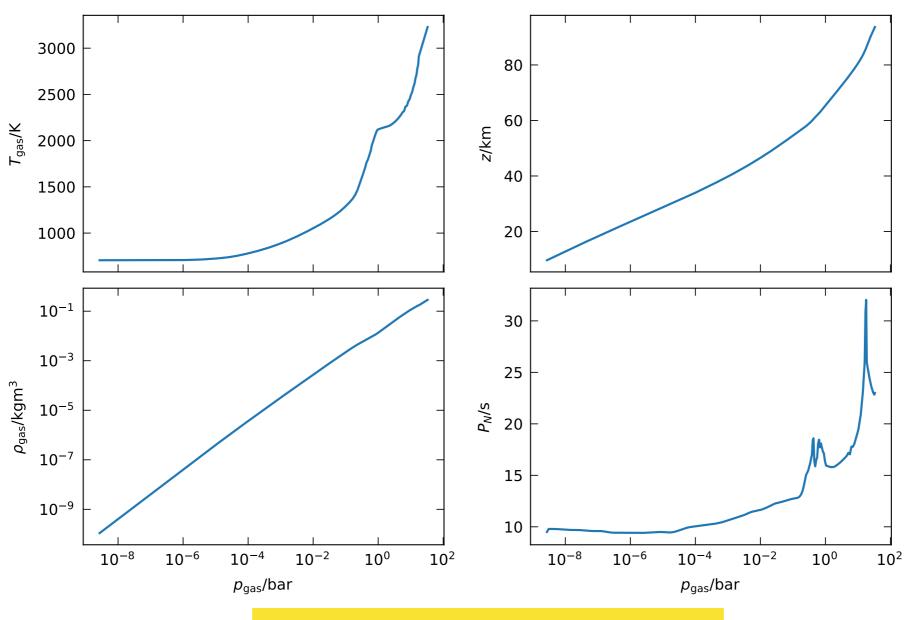
Numerical Simulations/Methods



$$\gamma = \frac{f_x \rho_{\text{gas}} v_{\text{rel},x}}{4\rho_{\text{dust}}}$$

$$J_* = \frac{n_x}{\tau} Z \exp\left[(N_* - 1) \ln S - \left(\frac{T_{\Theta}}{T} \right) \frac{N_* - 1}{(N_* - 1)^{1/3}} \right]$$

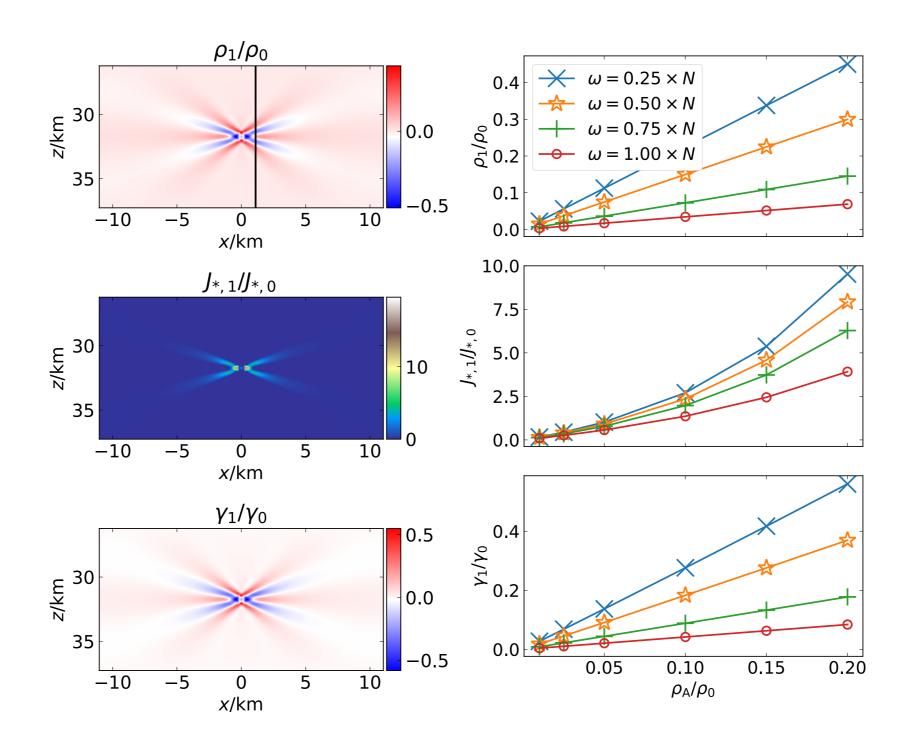
Numerical Simulations/Data



 $T_{\rm eff} = 1500 \, \text{K}, \log g = 5.0$

(Stark+ 2013; Rodríguez-Barrera+ 2018)

Numerical Simulations/Results



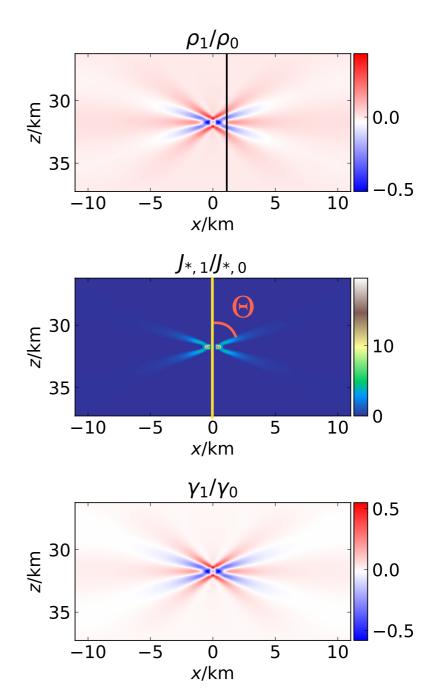
Core Findings

- X-shaped patterns, with banded areas of perturbed pressure, density and temperature
- ► Adiabatic process:
 - ▶ Lower ρ_{gas} , T_{gas} \rightarrow higher S_x
 - → stronger nucleation rate
 - ▶ Higher ρ_{gas} , T_{gas} → higher $v_{rel,x}$, ρ_{gas}
 - → faster mantle growth
- ► Gravity waves→banded areas of faster cloud formation

Next Steps

Linking (future) observable characteristics (magnitude, periodicity) to atmospheric density profile

$$\omega^{2} = -\frac{g}{\rho_{0}} \frac{\partial \rho_{0}}{\partial y} \cos^{2} \Theta$$
$$\rho_{0} \propto f(\Delta y, \Delta m, g, \omega, J_{*})$$



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Thank you! Questions?

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