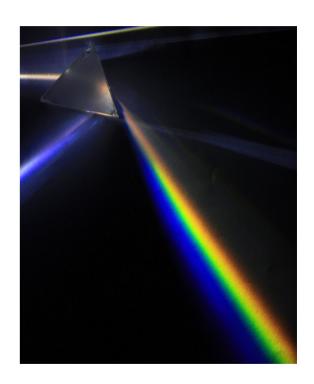
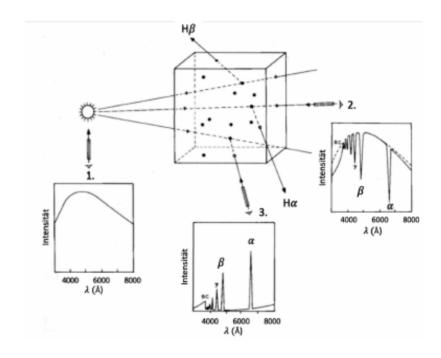
Spectroscopy

- decomposition of light in contributions of different wavelength
- the beginnings: I. Newton's optical experiments (1666–1672)



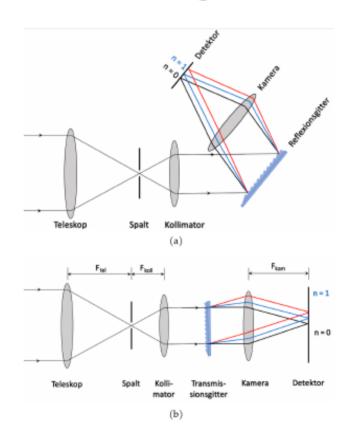
What can we see?

- continuous spectra (black-body)
- emission spectra
- absorption features



Main Components of a Spectrograph

- entrance pupil
 - slit in focal plane
 - fiber
- collimator → parallelizing light
- dispersive element
 - prism
 - grating
 - grism
- camera + detector
 - focusing + imaging + intensity measurement



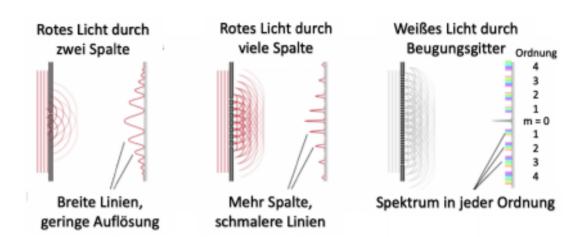
Interference pattern of a diffraction grating

 transmissive or reflective periodic structure (with spacing d)
 → light is being diffracted

$$d(\sin\theta_i - \sin\theta_m) = m\lambda.$$

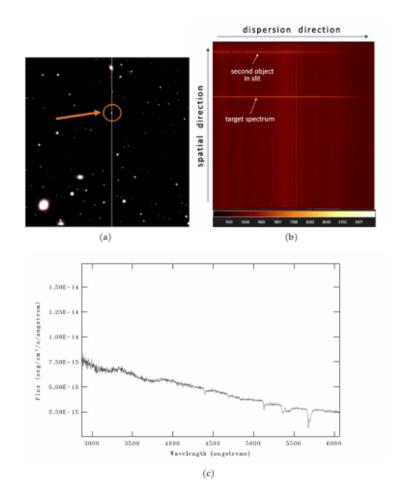
 Resolution R (increasing with order number and smaller spacing):

$$R = \lambda/\Delta\lambda$$



From Raw Data to 1D spectra

- bias correction (varying offset over detector)
- flat-fielding (pixel-to-pixel variations in eff.)
- removal of cosmic rays
- tracing the spectrum on image
- extraction of the spectrum → 1D (flux vs. pixel)
- wavelength calibration → 1D (flux vs. wavelength)
- flux calibration (instr. response)



Data Sets - Osservatorio Astrofisico di Asagio

- Copernico telescope
- 1.82 m
- Asiago Object Spectrograph and Camera (AFOSC)
- Resolution: 311
- Grism #4
 - → Central wavelength (580 nm)





- Galileo telescope
- 1.22 m
- Boller & Chivens Spectrograph

- observations in Sep. 2018
- Wavelength range around 350-800 nm
- Low resolution
- Calibration frames + 15 different objects

What You are expected to do

- full data reduction of one or several objects with IRAF (Image Reduction and Analysis Facility)
- evaluation of the resulting spectra (and identification of object type)

Material:

- raw calibration and object frames
- reference sheet for IRAF
- excerpts from Magister-thesis of V.
 Pfeiffer (reference for procedures and IRAF)
- IRAF web ressources:
 - IRAF.net
 - http://joshwalawender.github.io/IRA Ftutorial/IRAFintro 01.html
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