

**End Sem PHY690M. 30 April, 2024. Time : 3 Hours.**

**Please show all workings.**

**1a)** In flat spacetime, consider an event that occurs at the origin of a Cartesian system.

Write the defining relation for the corresponding null hypersurface and hence construct the normal vector  $k_\alpha$ . Write a set of parametric equations defining the hypersurface, clearly identifying the intrinsic coordinates.

From the parametric equations compute the tangent vectors.

Finally, construct the two metrics on the null hypersurface.

[2+2+2+2]

**1b)** On a three-dimensional timelike hypersurface, define the intrinsic covariant derivative of a three vector  $A_a$  as a suitable projection of the covariant derivative  $A_{\alpha;\beta}$ . Now show that this is indeed equal to the covariant derivative constructed out of the three-dimensional induced metric. [2]

**2)** Set up the problem of Oppenheimer-Snyder collapse where the internal and external metrics are given by  $ds_-^2 = -d\tau^2 + a^2(\tau)(d\chi^2 + \sin^2 \chi d\Omega^2)$  and  $ds_+^2 = -f dt^2 + f^{-1} dr^2 + r^2 d\Omega^2$ . Calculate a) the induced metric and b) the extrinsic curvature on both sides. Clearly state any assumptions that you make. [2+2]

**Useful info:** If  $n_\alpha$  is the unit normal to a hypersurface and is defined by  $n_\alpha = \frac{\epsilon \Phi_{,\alpha}}{|g^{\mu\nu} \Phi_{,\mu} \Phi_{,\nu}|^{1/2}}$  with  $\Phi$  being the defining relation for the hypersurface and  $\epsilon = \pm 1$  for timelike/spacelike hypersurface, then the extrinsic curvature is defined as  $K_{ab} = n_{\alpha;\beta} e_a^\alpha e_b^\beta$ .

**3)** In the Lagrangian formulation of GR, set up the Einstein-Hilbert action. Vary this action and indicate how it necessitates the Gibbons-Hawking boundary term. Now include the matter part and varying this, define the energy-momentum tensor. [4+2]

**4)** Motivate and draw the Penrose diagram for Minkowski space.

[5]