

Dipole DM: Model Definitions

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1 Dipole Dark Matter Model (DIPoleDM_FULL_UFO)

The Dipole DM model extends the SM by adding two Dirac Fermions (χ_1 and χ_0) and a real scalar (ϕ), all singlets under the SM gauge group. The Lagrangian of the model is given by:

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \mathcal{L}_\phi + \mathcal{L}_\chi + \mathcal{L}_{\text{dipole}} + \mathcal{L}_{H\chi}, \quad (1.1)$$

where \mathcal{L}_{SM} represents the SM Lagrangian and

$$\mathcal{L}_\phi = (\partial^\mu \phi)^2 - \mu_2^2 |\phi|^2 - \lambda_2 \phi^4 - \lambda_3 \phi^2 |H|^2, \quad (1.2)$$

$$\mathcal{L}_\chi = i \bar{\chi}_i \not{\partial} \chi_i - \tilde{M}_{ij} \bar{\chi}_i \chi_j - (y_\chi)_{ij} \bar{\chi}_i \chi_j \phi, \quad (1.3)$$

$$\mathcal{L}_{\text{dipole}} = \frac{(C_{\gamma\chi\chi})_{ij}}{\Lambda} \bar{\chi}_i \sigma^{\mu\nu} \chi_j F_{\mu\nu}, \quad (1.4)$$

$$\mathcal{L}_{H\chi} = \frac{(C_{H\chi\chi})_{ij}}{\Lambda} \bar{\chi}_i \chi_j |H|^2. \quad (1.5)$$

In the equations above, $\sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu]$ and H represents the Higgs doublet.

The model Lagrangian given in Eq. (1.1) contains the scalar potential:

$$V_{H,\phi} = \mu_1^2 |H|^2 + \lambda_1 |H|^4 + \mu_2^2 \phi^2 + \lambda_2 \phi^4 + \lambda_3 \phi^2 |H|^2. \quad (1.6)$$

Assuming that both H and ϕ develop vevs, $\langle \phi \rangle = v_D/\sqrt{2}$ and $\langle H \rangle = v/\sqrt{2}$, we obtain the mass eigenstates h and S :

$$\begin{aligned} h &= (\sqrt{2}H^0 - v) \cos \alpha - (\sqrt{2}\phi - v_D) \sin \alpha, \\ S &= (\sqrt{2}\phi - v_D) \cos \alpha + (\sqrt{2}H^0 - v) \sin \alpha. \end{aligned} \quad (1.7)$$

Where the mixing angle (α) is given by

$$\tan(2\alpha) \equiv \frac{\lambda_3 v v_D}{\lambda_1 v^2 - \lambda_2 v_D^2}. \quad (1.8)$$

and with masses:

$$m_{S,h}^2 = \lambda_1 v^2 + \lambda_2 v_D^2 \mp (\lambda_1 v^2 - \lambda_2 v_D^2) \sqrt{1 + \tan^2(2\alpha)}, \quad (1.9)$$

Using the above equations, we can express the quartic couplings (λ_i) and the mass parameters (μ_i) in terms of the physical masses, mixing angle and vevs:

$$\lambda_1 = \frac{1}{2v^2} (\cos^2 \alpha m_h^2 + m_S^2 \sin^2 \alpha), \quad (1.10)$$

$$\lambda_2 = \frac{1}{2v_D^2} (\cos^2 \alpha m_S^2 + m_h^2 \sin^2 \alpha), \quad (1.11)$$

$$\lambda_3 = \frac{1}{v_D v} (m_S^2 - m_h^2) \sin \alpha \cos \alpha, \quad (1.12)$$

$$\mu_1^2 = - \left(\lambda_1 v^2 + \lambda_3 \frac{v_D^2}{2} \right), \quad (1.13)$$

$$\mu_2^2 = - \left(\lambda_2 v_D^2 + \lambda_3 \frac{v^2}{2} \right) \quad (1.14)$$

Therefore, the parameters of the scalar potential ($\mu_{1,2}$ and $\lambda_{1,2,3}$) can be replaced by m_h, m_S, v, v_D and $\sin \alpha$.

Finally, the dark fermion mass matrix \tilde{M}_{ij} is defined in the mass eigenstate basis:

$$\tilde{M}_{ij} = M_i \delta_{ij} - \frac{v_D}{\sqrt{2}} (y_\chi)_{ij} - \frac{v^2}{2\Lambda} (C_{H\chi\chi})_{ij}$$

where M_i are the physical masses.

In addition to the Lagrangian in Eq. (1.1), the effective $h - G - G$ and $S - G - G$ couplings induced by a top quark loop were also included as effective operators:

$$\mathcal{L}_{GG\phi} = -\frac{1}{4} \cos \alpha F(m_h^2/m_t^2) G^{\mu\nu} G_{\mu\nu} h - \frac{1}{4} \sin \alpha F(m_S^2/m_t^2) G^{\mu\nu} G_{\mu\nu} S$$

where $F(x)$ is the effective loop function.

Model Parameters

The model parameters are given in Table 1 as well as their naming convention in the UFO model and their default values. The $C_{\gamma\chi\chi}$ and $C_{H\chi\chi}$ matrices are assumed to be real and symmetric.

Feynman rules

The relevant Feynman rules for the BSM particles are given in Table 2 below. Here, f represents any of the SM fermions and q any SM quark.

Table 1. Model parameters, their respective names in the UFO model and default values.

| Parameter | UFO name | Default Value |
|-----------------------------|----------|---------------|
| Λ | LambdaUV | 5 TeV |
| $(C_{\gamma\chi\chi})_{11}$ | Caxx1 | 0 |
| $(C_{\gamma\chi\chi})_{00}$ | Caxx0 | 0 |
| $(C_{\gamma\chi\chi})_{10}$ | Caxx10 | 0.1 |
| $(C_{H\chi\chi})_{11}$ | Chxx1 | 0 |
| $(C_{H\chi\chi})_{00}$ | Chxx0 | 0 |
| $(C_{H\chi\chi})_{10}$ | Chxx10 | 0.1 |
| $(y_\chi)_{11}$ | ychi1 | 1.0 |
| $(y_\chi)_{00}$ | ychi0 | 0 |
| $(y_\chi)_{10}$ | ychi10 | 0 |
| $\sin \alpha$ | sina | 0.2 |
| v_D | vevD | 1 TeV |

Table 2. Feynman rules for the relevant interactions in the full model.

| Interaction | Vertex term |
|-----------------------|---|
| $\chi_i \chi_j h$ | $\frac{i}{\sqrt{2}}(y_\chi)_{ij} \sin \alpha - i \frac{v}{\Lambda}(C_{H\chi\chi})_{ij} \cos \alpha$ |
| $\chi_i \chi_j A^\mu$ | $-i \frac{1}{\Lambda}(C_{\gamma\chi\chi})_{ij}(\gamma^\mu \not{p} - \not{p} \gamma^\mu)$ |
| $S \chi_i \chi_j$ | $-\frac{i}{\sqrt{2}}(y_\chi)_{ij} \cos \alpha - i \frac{v}{\Lambda}(C_{H\chi\chi})_{ij} \sin \alpha$ |
| $S h h$ | $-i \frac{m_S^2}{2v} \left(1 + 2 \frac{m_h^2}{m_S^2}\right) \left(\cos \alpha + 2 \frac{v}{v_D} \sin \alpha\right) \sin(2\alpha)$ |
| $S f \bar{f}$ | $-i \frac{m_f}{v} \sin \alpha$ |
| $S W_\mu^- W_\nu^+$ | $2i g^{\mu\nu} \frac{m_W^2}{v} \sin \alpha$ |
| $S Z_\mu Z_\nu$ | $2i g^{\mu\nu} \frac{m_Z^2}{v} \sin \alpha$ |

2 Minimal Dipole Model (DIPoleDM_MINIMAL_UFO)

The minimal scenario assumes that χ_0 only couples through the effective operators and these have zero entries in the diagonal:

$$(C_{\gamma\chi\chi})_{ij} = (C_{H\chi\chi})_{ij} = 0, \text{ if } i = j \quad (2.1)$$

$$(y_\chi)_{ij} = 0, \text{ if } i = 0 \text{ or } j = 0 \quad (2.2)$$

With the above assumptions the Lagrangian simplifies to:

$$\mathcal{L}_\chi = \bar{\chi}_i (i\not{\partial} - \tilde{M}_i) \chi_i - (y_\chi)_{11} \bar{\chi}_1 \chi_1 \phi, \quad (2.3)$$

$$\mathcal{L}_{\text{dipole}} = \frac{(C_{\gamma\chi\chi})_{01}}{\Lambda} \bar{\chi}_0 \sigma^{\mu\nu} \chi_1 F_{\mu\nu} + h.c., \quad (2.4)$$

$$\mathcal{L}_{H\chi} = \frac{(C_{H\chi\chi})_{01}}{\Lambda} \bar{\chi}_0 \chi_1 |H|^2 + h.c.. \quad (2.5)$$

resulting in the vertices given in Table 3.

Table 3. Feynman rules for the relevant interactions in the **minimal** model.

| Interaction | Vertex term (Minimal) |
|-----------------------|--|
| $\chi_1 \chi_1 h$ | $\frac{i}{\sqrt{2}}(y_\chi)_{11} \sin \alpha$ |
| $S \chi_1 \chi_1$ | $-\frac{i}{\sqrt{2}}(y_\chi)_{11} \cos \alpha$ |
| $\chi_1 \chi_0 h$ | $-i \frac{v}{\Lambda} (C_{H\chi\chi})_{01} \cos \alpha$ |
| $S \chi_1 \chi_0$ | $-i \frac{v}{\Lambda} (C_{H\chi\chi})_{01} \sin \alpha$ |
| $\chi_1 \chi_0 A^\mu$ | $-i \frac{1}{\Lambda} (C_{\gamma\chi\chi})_{01} (\gamma^\mu \not{p} - \not{p} \gamma^\mu)$ |