Dipole DM: Model Definitions

Andre Lessa,^a Jose Zurita,^b

^a Centro de Ciências Naturais e Humanas, Universidade Federal do ABC, Santo André, 09210-580 SP, Brazil

E-mail: andre.lessa@ufabc.edu.br

Contents

1	Dıp	oole Dark Matter Model (DIPOLEDM_FULL_UFO)	1
	1.1	Minimal Couplings to Higgs (DIPOLEDM2_MINIMALH_UFO)	4
	1.0	M': 1 G 1: $D1$ $D2$ DM DM	

1.2 Minimal Couplings to Photon (DIPOLEDM_MINIMALA_UFO) 5

Contents

1 Dipole Dark Matter Model (DIPOLEDM_FULL_UFO)

The Dipole DM model extends the SM by adding three Dirac Fermions (χ_2 , χ_1 and χ_0) and a real scalar (ϕ), all singlets under the SM gauge group. The Lagrangian of the model is given by:

$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{\phi} + \mathcal{L}_{\chi} + \mathcal{L}_{dipole} + \mathcal{L}_{H\chi}, \tag{1.1}$$

where $\mathcal{L}_{\mathrm{SM}}$ represents the SM Lagrangian and

$$\mathcal{L}_{\phi} = (\partial^{\mu}\phi)^{2} - \mu_{2}^{2}|\phi|^{2} - \lambda_{2}\phi^{4} - \lambda_{3}\phi^{2}|H|^{2}, \qquad (1.2)$$

$$\mathcal{L}_{\chi} = i\overline{\chi}_{i} \partial \chi_{i} - \tilde{M}_{ij}\overline{\chi}_{i}\chi_{j} - (y_{\chi})_{ij}\overline{\chi}_{i}\chi_{j}\phi, \qquad (1.3)$$

$$\mathcal{L}_{\text{dipole}} = \frac{(C_{\gamma\chi\chi})_{ij}}{\Lambda} \overline{\chi}_i \sigma^{\mu\nu} \chi_j F_{\mu\nu} , \qquad (1.4)$$

$$\mathcal{L}_{H\chi} = \frac{(C_{H\chi\chi})_{ij}}{\Lambda} \overline{\chi}_i \chi_j |H|^2. \tag{1.5}$$

In the equations above, $\sigma^{\mu\nu} = \frac{i}{2} \left[\gamma^{\mu}, \gamma^{\nu} \right]$ and H represents the Higgs doublet.

The model Lagrangian given in Eq. (1.1) contains the scalar potential:

$$V_{H,\phi} = \mu_1^2 |H|^2 + \lambda_1 |H|^4 + \mu_2^2 \phi^2 + \lambda_2 \phi^4 + \lambda_3 \phi^2 |H|^2.$$
 (1.6)

Assuming that both H and ϕ develop vevs, $\langle \phi \rangle = v_D/\sqrt{2}$ and $\langle H \rangle = v/\sqrt{2}$, we obtain the mass eigenstates h and S:

$$h = (\sqrt{2}H^0 - v)\cos\alpha - (\sqrt{2}\phi - v_D)\sin\alpha,$$

$$S = (\sqrt{2}\phi - v_D)\cos\alpha + (\sqrt{2}H^0 - v)\sin\alpha.$$
(1.7)

Where the mixing angle (α) is given by

$$\tan(2\alpha) \equiv \frac{\lambda_3 v v_D}{\lambda_1 v^2 - \lambda_2 v_D^2} \,. \tag{1.8}$$

and with masses:

$$m_{S,h}^2 = \lambda_1 v^2 + \lambda_2 v_D^2 \mp (\lambda_1 v^2 - \lambda_2 v_D^2) \sqrt{1 + \tan^2(2\alpha)},$$
 (1.9)

Figure 1. Production cross-section for the dark scalar S as a function of its mass.

Using the above equations, we can express the quartic couplings (λ_i) and the mass parameters (μ_i) in terms of the physical masses, mixing angle and vevs:

$$\lambda_1 = \frac{1}{2v^2} \left(\cos^2 \alpha \, m_h^2 + m_S^2 \sin^2 \alpha \right), \tag{1.10}$$

$$\lambda_2 = \frac{1}{2v_D^2} \left(\cos^2 \alpha \, m_S^2 + m_h^2 \sin^2 \alpha \right), \tag{1.11}$$

$$\lambda_3 = \frac{1}{v_D v} \left(m_S^2 - m_h^2 \right) \sin \alpha \cos \alpha , \qquad (1.12)$$

$$\mu_1^2 = -\left(\lambda_1 v^2 + \lambda_3 \frac{v_D^2}{2}\right),\tag{1.13}$$

$$\mu_2^2 = -\left(\lambda_2 v_D^2 + \lambda_3 \frac{v^2}{2}\right) \tag{1.14}$$

Therefore, the parameters of the scalar potential $(\mu_{1,2} \text{ and } \lambda_{1,2,3})$ can be replaced by m_h, m_S, v, v_D and $\sin \alpha$.

Finally, the dark fermion mass matrix M_{ij} is defined in the mass eigenstate basis:

$$\tilde{M}_{ij} = M_i \delta_{ij} - \frac{v_D}{\sqrt{2}} (y_\chi)_{ij} - \frac{v^2}{2\Lambda} (C_{H\chi\chi})_{ij}$$

where M_i are the physical masses.

In addition to the Lagragian in Eq. (1.1), the effective h - G - G and S - G - G couplings induced by a top quark loop were also included as effective operators:

$$\mathcal{L}_{GG\phi} = \frac{g_s^2}{48\pi^2 v} \cos \alpha F(m_h^2/m_t^2) G^{\mu\nu} G_{\mu\nu} h + \frac{g_s^2}{48\pi^2 v} \sin \alpha F(m_S^2/m_t^2) G^{\mu\nu} G_{\mu\nu} S$$

where F(x) is the effective loop function.

Model Parameters

The model parameters are given in Table 1 as well as their naming convention in the UFO model and their default values. The $C_{\gamma\chi\chi}$ and $C_{H\chi\chi}$ matrices are assumed to be real and symmetric.

Feynman rules

The relevant Feynman rules for the BSM particles are given in Table 2 below. Here, f represents any of the SM fermions and q any SM quark.

The production of S at the LHC mainly occurs due to gluon-fusion and is proportional to $\sin \alpha$. In Fig. 1 we show the $\sigma(pp \to S)$ as a function of m_S for $\sin \alpha = 0.2$ and a center of mass energy of 13.6 TeV.

Table 1. Model parameters, their respective names in the UFO model and default values.

Parameter	UFO name	Default Value
Λ	LambdaUV	5 TeV
$(C_{\gamma\chi\chi})_{22}$	Caxx2	0
$(C_{\gamma\chi\chi})_{11}$	Caxx1	0
$(C_{\gamma\chi\chi})_{00}$	Caxx0	0
$(C_{\gamma\chi\chi})_{10}$	Caxx10	0.1
$(C_{\gamma\chi\chi})_{20}$	Caxx20	0.1
$(C_{\gamma\chi\chi})_{21}$	Caxx21	0.1
$(C_{H\chi\chi})_{22}$	Chxx2	0
$(C_{H\chi\chi})_{11}$	Chxx1	0
$(C_{H\chi\chi})_{00}$	Chxx0	0
$(C_{H\chi\chi})_{10}$	Chxx10	0.1
$(C_{H\chi\chi})_{20}$	Chxx20	0.1
$(C_{H\chi\chi})_{21}$	Chxx21	0.1
$(y_\chi)_{22}$	ychi2	1.0
$(y_{\chi})_{11}$	ychi1	1.0
$(y_\chi)_{00}$	ychi0	0
$(y_\chi)_{10}$	ychi10	0
$(y_{\chi})_{20}$	ychi20	0
$(y_\chi)_{21}$	ychi21	0
$\sin \alpha$	sina	0.2
v_D	vevD	1 TeV
m_S	MSd	1.5 TeV
M_0	M0	425 GeV
M_1	M1	500 GeV
M_2	M2	510 GeV

Interaction	Vertex term
$\chi_i \chi_j h$	$\frac{i}{\sqrt{2}}(y_{\chi})_{ij}\sin\alpha - i\frac{v}{\Lambda}(C_{H\chi\chi})_{ij}\cos\alpha$
$\chi_i \chi_j A^{\mu}$	$-irac{1}{\Lambda}(C_{\gamma\chi\chi})_{ij}(\gamma^{\mu}p\!\!\!/-p\!\!\!/\gamma^{\mu})$
$S \chi_i \chi_j$	$-\frac{i}{\sqrt{2}}(y_{\chi})_{ij}\cos\alpha - i\frac{v}{\Lambda}(C_{H\chi\chi})_{ij}\sin\alpha$
Shh	$-i\frac{m_S^2}{2v}\left(1+2\frac{m_h^2}{m_S^2}\right)\left(\coslpha+2rac{v}{v_D}\sinlpha ight)\sin(2lpha)$
$S f \bar{f}$	$-i\frac{m_f}{v}\sin \alpha$
$SW_{\mu}^-W_{\nu}^+$	$2ig^{\mu\nu} \frac{m_W^2}{v} \sin \alpha$
$S Z_{\mu} Z_{\nu}$	$2ig^{\mu\nu}rac{m_Z^2}{v}\sinlpha$
$SG_{\mu}G_{\nu}$	$i\frac{g_s^2}{12\pi^2 v}\sin\alpha F(m_S^2/m_t^2)(p_1^{\mu}p_2^{\nu}-g^{\mu\nu}p_1\cdot p_2)$

Table 2. Feynman rules for the relevant interactions in the full model.

1.1 Minimal Couplings to Higgs (DIPOLEDM2 MINIMALH UFO)

The minimal scenario assumes that χ_0 only couples through the H effective operator and the diagonal entries are zero:

$$(C_{\gamma\gamma\gamma})_{ij} = 0 \tag{1.15}$$

$$(C_{H_{\chi\chi}})_{ij} = 0$$
, if $i = j$ or $i = 2$ or $j = 2$ (1.16)

$$(y_{\chi})_{ij} = 0$$
, if $i = j$ or $i = 1, j = 0$ (1.17)

With the above assumptions the Lagrangian simplifies to:

$$\mathcal{L}_{\chi} = \overline{\chi}_i \left(i \partial - M_i \right) \chi_i - \left[(y_{\chi})_{21} \overline{\chi}_2 \chi_1 \phi + (y_{\chi})_{20} \overline{\chi}_2 \chi_0 \phi + \frac{(C_{H\chi\chi})_{01}}{\Lambda} \overline{\chi}_0 \chi_1 |H|^2 + h.c. \right]$$
(1.18)

resulting in the vertices given in Table 3.

Interaction	Vertex term (Minimal H)
$h \chi_1 \chi_1$	$\frac{i}{\sqrt{2}}(y_\chi)_{11}\sin\alpha$
$S \chi_1 \chi_1$	$-\frac{i}{\sqrt{2}}(y_\chi)_{11}\cos\alpha$
$\chi_1 \chi_0 h$	$-i\frac{v}{\Lambda}(C_{H\chi\chi})_{01}\cos\alpha$
$S \chi_1 \chi_0$	$-i\frac{v}{\Lambda}(C_{H\chi\chi})_{01}\sin\alpha$
$S G_{\mu} G_{\nu}$	$i\frac{g_s^2}{12\pi^2 v}\sin\alpha F(m_S^2/m_t^2)(p_1^\mu p_2^\nu - g^{\mu\nu}p_1 \cdot p_2)$

Table 3. Feynman rules for the relevant interactions in the **minimal** model with couplings to Higgs. The couplings between S and two SM particles are the same as in Table 2 and have been omitted.

In this scenario the heavy fermion can only decay through the effective operator, hence:

$$\Gamma(\chi_1 \to h\chi_0) = \cos^2 \alpha \frac{C_{H\chi\chi}^2}{16\pi} \frac{v^2}{\Lambda^2} M_1 \left[\left(1 + \frac{M_0}{M_1} \right)^2 - \frac{m_h^2}{M_1^2} \right] \sqrt{\lambda \left(1, \frac{M_0^2}{M_1^2}, \frac{m_h^2}{M_1^2} \right)}$$
(1.19)

$$\Gamma(\chi_1 \to S\chi_0) = \sin^2 \alpha \frac{C_{H\chi\chi}^2}{16\pi} \frac{v^2}{\Lambda^2} M_1 \left[\left(1 + \frac{M_0}{M_1} \right)^2 - \frac{m_S^2}{M_1^2} \right] \sqrt{\lambda \left(1, \frac{M_0^2}{M_1^2}, \frac{m_S^2}{M_1^2} \right)}, \quad (1.20)$$

where $\lambda(x, y, z) = x^2 + y^2 + z^2 + 2xy + 2xz + 2yz$.

Note that the decay to $S + \chi_0$ is suppressed by $\sin \alpha$ and will be subdominant as long as $m_h \lesssim m_S$. Furthermore, if $\Delta m = m_{\chi_1} - m_{\chi_0} < m_h$ the decay is kinematically suppressed and takes place through an off-shell Higgs (h^*) . In Fig. 2 we show the branching ratios as a function of Δm for $m_{\chi_1} = 500$ GeV. For the extremely compressed scenario ($\Delta m < 2$ GeV) only the decay to gluons $(h^* \to gg)$ is kinematically allowed, resulting in a 100% BR in this channel. Once decays to τ s are allowed, they become the dominant channel. For $\Delta m \gtrsim 10$ GeV the decays $h^* \to b\bar{b}$ are open and rapidly become the main decay channel up to $\Delta m > m_h$, where the decay is 100% to an on-shell Higgs. Finally, in Fig. 3 we show the decay length as a function of Δm , where we see that it can take a wide range of values depending on the mass compression. In particular, once the $b\bar{b}$ channel is open the decay length falls below 10 m. It is important to notice, however, that the lifetime scales as $\tau \propto (\Lambda/C_{H\chi\chi})^2$ and will be further suppressed for larger values of $\Lambda/C_{H\chi\chi}$.

Figure 2. Branching ratios for χ_1 decays in the minimal H scenario as a function of $\Delta m = m_{\chi_1} - m_{\chi_0}$. Only channels with a branching ratio larger than 1% are shown.

Figure 3. Proper decay length for χ_1 in the minimal H scenario as a function of $\Delta m = m_{\chi_1} - m_{\chi_0}$.

1.2 Minimal Couplings to Photon (DIPOLEDM_MINIMALA_UFO)

The minimal scenario assumes that χ_0 only couples through the $F^{\mu\nu}$ effective operator and the diagonal entries are zero:

$$(C_{\gamma\gamma\gamma})_{ij} = 0 , \text{ if } i = j$$

$$(1.21)$$

$$(C_{H\chi\chi})_{ij} = 0 \tag{1.22}$$

$$(y_{\chi})_{ij} = 0$$
, if $i = 0$ or $j = 0$ (1.23)

With the above assumptions the Lagrangian simplifies to:

$$\mathcal{L}_{\chi} = i\overline{\chi}_{i} \partial \chi_{i} - \tilde{M}_{ij}\overline{\chi}_{i}\chi_{j} - (y_{\chi})_{11}\overline{\chi}_{1}\chi_{1}\phi + \frac{(C_{\gamma\chi\chi})_{01}}{\Lambda} (\overline{\chi}_{0}\sigma^{\mu\nu}\chi_{1} + h.c.) F_{\mu\nu}$$
 (1.24)

resulting in the vertices given in Table 4.

Interaction	Vertex term (Minimal A)
$h \chi_1 \chi_1$	$\frac{i}{\sqrt{2}}(y_\chi)_{11}\sin\alpha$
$S \chi_1 \chi_1$	$-\frac{i}{\sqrt{2}}(y_\chi)_{11}\cos\alpha$
$\chi_1 \chi_0 A^{\mu}$	$-irac{1}{\Lambda}(C_{\gamma\chi\chi})_{01}(\gamma^{\mu} ot\!\!/-p\!\!\!/\gamma^{\mu})$
$S G_{\mu} G_{ u}$	$i\frac{g_s^2}{12\pi^2 v}\sin \alpha F(m_S^2/m_t^2)(p_1^\mu p_2^ u - g^{\mu u}p_1 \cdot p_2)$

Table 4. Feynman rules for the relevant interactions in the **minimal** model with couplings to photons. The couplings between S and two SM particles are the same as in Table 2 and have been omitted.

In this scenario the heavy fermion can only decay through the effective operator, hence:

$$\Gamma(\chi_1 \to \gamma \chi_0) = \frac{C_{\gamma \chi \chi}^2}{2\pi} \frac{M_1^2}{\Lambda^2} M_1 \left(1 - \frac{M_0^2}{M_1^2} \right)^3$$
 (1.25)

Since this is the only allowed channel it has always 100% branching ratio and the lifetime for the compressed scenario is given by:

$$c\tau = 1.55 \times 10^{-6} \text{ m} \left(\frac{0.1}{C_{\gamma\chi\chi}}\right)^2 \left(\frac{\Lambda}{10 \text{ TeV}}\right)^2 \left(\frac{1 \text{ GeV}}{\Delta m}\right)^3,$$
 (1.26)

where we have assumed $M_0 \lesssim M_1$, so $\Delta m^2 = M_1^2 - M_0^2 \simeq 2M_1\Delta m$. In Fig. 4 we show the decay length as a function of Δm . As we can see, even for $\Lambda = 100$ TeV the decay is prompt as long as $\Delta m > 1$ GeV.

