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Chapter 1

Preliminaries

1.1 Quick review of Special Relativity

Here we expose a quick review of Special Relativity in order to set the notations.

Fundamental principles of Special Relativity are followings:

- (i) All inertial reference frames are physically equivalent. There is no way to distinguish between different inertial frames in the sense that there is no preferred one.
- (ii) There exists universal (dimensional) constant: $c \simeq 3 \times 10^8 \text{m/s}$, i.e. the speed of massless particles.

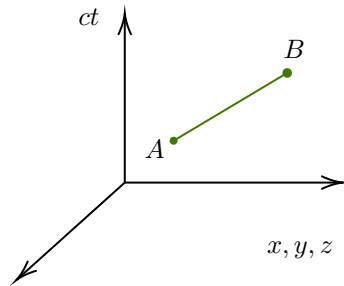
In order to implement these features basic ingredients are

- (i) Space and Time form a unique concept called **spacetime**.
- (ii) A spacetime is a collection of points called **event**.
- (iii) Each inertial frame is associated with a set of *space time coordinates*. Each events is specified through coordinate system of a fixed initial frame.

$$x^\mu = (x^0, x^1, x^2, x^3) \equiv (x^0, x^i) \equiv (ct, x, y, z) \equiv (ct, \mathbf{x})$$

Usually x, y, z are assumed to be Cartesian coordinates.

Given 2 events A and B in spacetime



their distance is $\Delta x^\mu = x_B^\mu - x_A^\mu$. We introduce the (squared) *Minkowski distance*

$$\Delta s^2 = \eta_{\mu\nu} \Delta x^\mu \Delta^\nu \quad \text{where} \quad \eta_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$$

where $\eta_{\mu\nu}$ is *Minkowski metric*. This induces the *line element*

$$ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$$

This element is scalar quantity and therefore does not depends on the specific inertial frame.

The quantity Δs^2 has an intrinsic meaning

$$\begin{cases} \Delta s^2 > 0 & : \quad \Delta x^\mu \text{ is } \textit{space-like} \text{ vector} \\ \Delta s^2 = 0 & : \quad \Delta x^\mu \text{ is } \textit{time-like} \text{ vector} \\ \Delta s^2 < 0 & : \quad \Delta x^\mu \text{ is } \textit{light-like/null} \text{ vector} \end{cases}$$

Space-like vector means that exists different frames where two events are simultaneous. Time-like vector means that exists different frames where two events have same space coordinates but they happen at different times. Light-like vectors means that two events may be connected by a light signal.

- (iv) Allowed transformations for spacetime vectors must preserve the line element: $\Delta \tilde{s}^2 = \Delta s^2$. These transformations are the *Poincaré Transformations*

$$x^\mu \rightarrow \tilde{x}^\mu = \Lambda^\mu{}_\nu x^\nu + a^\mu \quad \text{with} \quad \Lambda^\rho{}_\mu \Lambda^\sigma{}_\nu \eta_{\rho\sigma} = \eta_{\mu\nu}$$

Once we have reformulated notions of space and time, we have to reformulate law of physics in such a way they does not depends on the reference frame.

Trajectories of point like-particles are associated to curved *wordlines* in space time and described evolution of events. Mathematically they are described by maps from \mathbb{R} into a set of four functions: $\lambda \in \mathbb{R} \rightarrow x^\mu(\lambda)$. Near if we consider nearby events separated by infinitesimal shift we can obtain infinitesimal variation of coordinates:

$$dx^\mu(\lambda) \equiv x^\mu(\lambda + d\lambda) - x^\mu(\lambda) = \frac{dx^\mu(\lambda)}{d\lambda} d\lambda$$

Since no particles can move at a speed higher then light this implies that ds^2 must be time-like. Notice that choice of parameter λ is free. One possible choice of this parameter is the (*differential*) *proper time*:

$$d\tau \equiv \sqrt{-ds^2} = d\lambda \sqrt{-\eta_{\mu\nu} \dot{x}^\mu(\lambda) \dot{x}^\nu(\lambda)} = c dt \sqrt{1 - \frac{v^2}{c^2}} \equiv \frac{c dt}{\gamma}$$

where third step holds if $\lambda \equiv t$. If we define $\beta \equiv v/c$ ¹, then $\gamma = 1/\sqrt{1 - \beta^2}$ is called *Lorentz factor*. Notice that last step implies time dilatation at higher velocities. For $\lambda = t$ we obtain

$$\tau = c \int dt \sqrt{1 - \frac{v^2}{c^2}}$$

i.e. with this definition the proper time has dimension of a length $[\tau] = L$. Physically the proper times it's the time measured by a clock moving along the trajectory.

Proper time allow us to define a vector called *4-velocity* that can be identified as relativistic generalization of velocity. Namely:

$$u^\mu(\tau) = \frac{dx^\mu(\tau)}{d\tau} = \left(\gamma, \gamma \frac{\mathbf{v}}{c} \right)$$

Notice

$$u^\mu u_\mu = -1$$

i.e. is a time-like vector. Moreover, this vector has only three degrees of freedom, since one component is fixed by previous propriety.

¹This is the speed in natural units, i.e. in units of β . If we set $c = 1$ then $v = \beta$.

Now we can define the generalization of acceleration, *4-acceleration*, as follows

$$\alpha^\mu(\tau) = \frac{du^\mu(\tau)}{d\tau}$$

Notice that, as we expected, 4-acceleration is orthogonal to 4-velocity

$$u_\mu \alpha^\mu = 0$$

and this implies that α^μ is a space-like since it is orthogonal to a time-like vector.

This proves a relativistic generalization of distance, speed and acceleration. Also laws of dynamic can be generalized, in particular if we define the *four-force* f^μ as the generalization of force we can obtain the *Relativistic Second Newton's law*:

$$m\alpha^\mu \equiv \frac{dp^\mu}{d\tau} = f^\mu$$

where we used four acceleration or equivalently the generalization of newtonian momentum, *4-momentum*,

$$p^\mu \equiv mcu^\mu = \left(\frac{E}{c}, \mathbf{p} \right)$$

For example, for Lorentz force

$$\mathbf{F}_L = e \left(\mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{B} \right)$$

can be generalizzated in a manifestly covariant way into^{II}

$$f_L^\mu = \frac{e}{c} F^{\mu\nu} u_\nu \quad \text{with} \quad F^{\mu\nu} = \begin{pmatrix} 0 & E_1 & E_2 & E_3 \\ -E_1 & 0 & B_3 & -B_2 \\ -E_2 & -B_3 & 0 & B_1 \\ -E_3 & B_2 & -B_1 & 0 \end{pmatrix}$$

where $F^{\mu\nu}$ is the *EM-Tensor*.

We can also rewrite Maxwell equations into two covariant equation

$$\partial_\mu F^{\mu\nu} = -\frac{4\pi}{c} j^\nu \quad , \quad \partial_{[\mu} F_{\nu\rho]} = 0$$

where the former, inhomogeneous, shows the *4-current* $j^\mu = (c\rho, \mathbf{j})$. The second equation, homogeneous, exhibits total antisymmetrized indexes^{III}. Each of these equations contains 2 independent equations.

We can conclude saying that all possible interactions can be written in a covariant way, except from gravitation. In order to include this force General Relativity has been developed.

^{II}Here is evident that this formula does not change under Poincaré transformations.

^{III} $\partial_{[\mu} F_{\nu\rho]} = \partial_\mu F_{\nu\rho} + \partial_\rho F_{\mu\nu} + \partial_\nu F_{\rho\mu}$.