

# Chapter 1

## Vortices

[Shifman:2012]

The second quantum soliton that we consider is the vortex in  $2 + 1$  dimensions in a model that in high-energy *Abelian Higgs mode* and in  $3 + 1$  dimensions was the first model proposed to make massive gauge fields in a gauge theory without losing gauge-invariance.

In its non-relativistic version in condensed matter it is described by the *Landau-Ginzburg model* and in 3 space dimensions it was proposed as a phenomenological model for superconductors.

Our discussion will be performed in the Lagrangian formalism, starting from the classical model.

### 1.1 Classical treatment

#### The classical Lagrangian

The field content is made of a complex scalar field  $\phi$  (whose complex conjugate is denoted by  $\phi^*$ ) and a  $U(1)$  gauge field  $A_\mu$ . The classical relativistic Lagrangian is

$$\mathcal{L} = -\frac{1}{4e^2}F_{\mu\nu}^2 + |D^\mu\phi|^2 - \lambda(|\phi|^2 - v^2)^2 \quad (1.1)$$

where  $e$  is the electric charge, the covariant derivative is defined by

$$D_\mu\phi = (\partial_\mu - in_e A_\mu)\phi \quad (1.2)$$

and  $n_e$  is the electric charge of  $\phi$  in units of  $e$ .

The non-relativistic Euclidean version replaces  $|D^0\phi|^2$  by a first order term

$$|D^0\phi|^2 \rightarrow \phi^*(\partial_0 - in_e A_0)\phi \quad (1.3)$$

For the model of superconductivity  $\phi$  is a field representing the large distance behaviour of the Cooper pairs generated by phonon attraction and  $n_e \equiv 2$ . The vortices that will be discussed later in fact really appear in nature.

#### Application of vortices

A lattice version of such vortices, called *Abrikosov<sup>I</sup> vortices*, is the equilibrium state of a class of superconductors in the presence of a magnetic field, orthogonal to the surface of the superconductors, whose direction will be denoted by  $z$ . The  $z$ -dependence is then trivial, and in the gauge  $A_z = 0$  the  $3 + 1$  model reduces to a  $2 + 1$  model.

Notice that a typical characteristic of superconductors is the expulsion of the magnetic field (*Meissner effect*), but there are two behaviours of superconducting materials in this respect, called *type I* and *type II*. In type I the magnetic flux is completely expelled from the bulk of the material, whereas in type II it penetrates in the superconductor in tubes whose two-dimensional cross section are the vortices, as shown

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<sup>I</sup>Nobel prize in the 2003.

in fig. 1.1. Each of these tubes contains a flux  $\frac{h}{e}$  and at equilibrium if they are sufficiently many<sup>II</sup> they are arranged in a triangular lattice, the *Abrikosov lattice*.

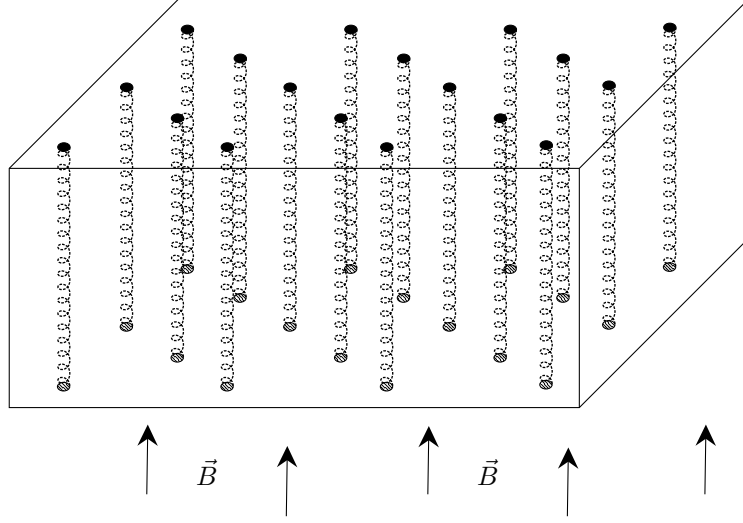


Figure 1.1: Vortices created by the magnetic field in a type II superconductor.

### Gauge symmetry, energy density and Higgs mechanism

The model is invariant under the  $U(1)$  gauge transformation<sup>III</sup>

$$\begin{cases} \phi(x) & \rightarrow e^{i\beta(x)}\phi(x) \\ A_\mu(x) & \rightarrow A_\mu(x) + \frac{1}{in_e}e^{-i\beta(x)}\partial_\mu e^{i\beta(x)} = A_\mu(x) + \frac{1}{n_e}\partial_\mu\beta(x) \end{cases} \quad (1.5)$$

Let us first consider static configurations in the *temporal* gauge  $A_0 = 0$ , so that there is no difference between relativistic and non-relativistic models. The energy is given by

$$\mathcal{E}(\mathbf{A}(\mathbf{x}), \phi(\mathbf{x})) = \int d^2x \frac{1}{4e^2} F_{ij}^2 + |D_i\phi|^2 + \lambda(|\phi|^2 - v^2)^2 \quad (1.6)$$

and it has global minima at

$$\phi(\mathbf{x}) = ve^{i\theta} \quad \text{for } \theta \in [0, 2\pi) \quad , \quad A_\mu(\mathbf{x}) = 0 \quad (1.7)$$

However by gauge invariance this configuration is physically equivalent to

$$\phi(\mathbf{x}) = ve^{i[\theta + \beta(\mathbf{x})]} \quad , \quad A_\mu(\mathbf{x}) = \frac{1}{in_e}e^{-i\beta(\mathbf{x})}\partial_\mu e^{i\beta(\mathbf{x})} \quad (1.8)$$

with  $\beta(\mathbf{x})$  globally defined of compact support (indeed it cannot act on the boundary of the spacetime, otherwise it changes the boundary conditions).

The existence of degenerate global minima (up to gauge equivalence) labelled by  $\theta$  suggests that the global  $U(1)$  symmetry is spontaneously broken. Indeed the degeneracy of the vacuum cannot be regarded as a manifestation of the gauge symmetry of the theory, since gauge symmetries cannot act at boundaries of the spacetime, whereas the action of  $e^{i\theta}$  extends also at the infinity. The presence of spontaneously symmetry breaking of the theory can be shown perturbatively (as in the SM), non-perturbatively in some

<sup>II</sup>In order to increase the number of tubes one can increase the strength of the magnetic field.

<sup>III</sup>Notice that one cannot write

$$e^{-i\beta(x)}\partial_\mu e^{i\beta(x)} = \partial_\mu\beta(x) \quad (1.4)$$

since  $\beta(x)$  may be defined with “jumps” of  $2\pi$  (more formally, we have  $\beta : \mathbb{R}^2 \rightarrow \mathbb{R}/2\pi\mathbb{Z}$ ) hence its derivative is not completely well defined. Instead,  $e^{i\beta(x)}$  is a smooth function, hence its derivative is well defined and smooth as well.

Mi è chiaro che nella trasformazione di gauge di  $A_\mu$  lei voglia rendere esplicita la forma di Maurer-Cartan, ma non sono molto sicuro del perché invece la forma  $\partial_\mu\beta$  dovrebbe essere imprecisa. Nella nota a piè di pagina ho cercato di darci una giustificazione che mi sembrasse plausibile, anche se non completamente corretta dal punto di vista matematico.

specific gauge (e.g. in the Coulomb gauge  $\nabla \cdot \mathbf{A} = 0$ , for which a change of boundary conditions at  $\infty$  is impossible), and using a non-local order parameter (which turns out to have a non zero expectation value independent from the gauge choice).

Due to the *Anderson-Higgs mechanism* (in high energy physics also called *BEH*, due to Brout-Englert-Higgs<sup>IV</sup>), let  $v = \langle \phi \rangle$  be the (real) expectation value of  $\phi$  in the broken symmetry phase, the gauge field  $A_\mu$  acquire a mass gap, whose inverse in condensed matter (in the Landau-Ginzburg model) is called *penetration depth*, given in the quadratic approximation by

$$m_A = \sqrt{2} v n_e \quad (1.9)$$

and also, writing

$$\phi(x) = (v + \chi(x)) e^{i\theta(x)} \quad \text{with } \chi(x), \theta(x) \text{ real fields} \quad (1.10)$$

the Higgs field  $\chi(x)$  acquires a mass gap, whose inverse in condensed matter is called *coherence length*, which in the quadratic approximation is given by

$$m_H = 2\sqrt{\lambda} v \quad (1.11)$$

This can be easily proved: inserting the expansion eq. (1.10) in the Lagrangian eq. (1.1) we get

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4e^2} F_{\mu\nu}^2 + |(\partial_\mu - in_e A_\mu)[(v + \chi) e^{i\theta}]|^2 - \lambda(|(v + \chi) e^{i\theta}|^2 - v^2)^2 \\ &= -\frac{1}{4e^2} F_{\mu\nu}^2 + |iv\partial_\mu\theta + \partial_\mu\chi - in_e A_\mu(v + \chi)|^2 - \lambda(2v\chi + \chi^2)^2 \end{aligned} \quad (1.12)$$

In terms of eq. (1.10) the gauge transformation (1.5) become

$$\begin{cases} \chi(x) & \rightarrow \chi(x) \\ \theta(x) & \rightarrow \theta(x) + \beta(x) \\ A_\mu(x) & \rightarrow A_\mu(x) + \frac{1}{n_e} \partial_\mu \beta(x) \end{cases} \quad (1.13)$$

hence acting on  $\mathcal{L}$  with a transformation of the form eq. (1.13) with  $\beta(x) = -\theta(x)$  we finally get

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4e^2} F_{\mu\nu}^2 + |\partial_\mu\chi - in_e v A_\mu - in_e A_\mu\chi|^2 - \lambda(2v\chi + \chi^2)^2 \\ &= -\frac{1}{4e^2} F_{\mu\nu}^2 + (\partial_\mu\chi)^2 + n_e^2 v^2 A_\mu^2 + 2n_e^2 v A_\mu^2 \chi - 4\lambda v^2 \chi^2 - 4\lambda v \chi^3 - \lambda \chi^4 \\ &= \underbrace{-\frac{1}{4e^2} F_{\mu\nu}^2 + \frac{1}{2} m_A^2 A_\mu^2 + (\partial_\mu\chi)^2 - \frac{1}{2} m_H^2 \chi^2}_{\text{kinetic term}} + \underbrace{2n_e^2 v A_\mu^2 \chi - 4\lambda v \chi^3 - \lambda \chi^4}_{\text{interaction term}} \end{aligned} \quad (1.14)$$

hence the masses of  $A_\mu$  and the  $\chi$  are exactly  $m_A$  and  $m_H$ , as we claimed above. Notice that the previous expression of the Lagrangian, eq. (1.14), holds in the specific gauge where  $\theta$  vanishes, which is usually called *unitary gauge*.

## Finite energy solutions - Vortices

Let us now impose the finiteness of the energy. The last term in eq. (1.6) forces

$$|\phi| \xrightarrow{|\mathbf{x}| \rightarrow \infty} v e^{if(\alpha)} \quad (1.15)$$

where  $\alpha$  is the angle in polar coordinates of  $\mathbf{x}$  at  $\infty$ ,  $\mathbf{x} = |\mathbf{x}| e^{i\alpha}$ , and  $f$  is a real function satisfying  $e^{if(2\pi)} = e^{if(0)}$ . The first term in eq. (1.6) implies that at  $\infty$ ,  $F_{ij} = 0$ , i.e.  $A_i$  is a pure gauge. The second term in eq. (1.6) implies that

$$|e^{if(\alpha)} \partial_j e^{if(\alpha)} - in_e A_j| \xrightarrow{|\mathbf{x}| \rightarrow \infty} 0 \quad (1.16)$$

<sup>IV</sup>Nobel prizes in the 2013.

Qui ho tolto il fattore  $e$ , i quanto non appariva espandendo la lagrangiana rispetto al campo di Higgs.

so  $A_j$ , asymptotically, is the pure gauge

$$A_j = \frac{1}{in_e} e^{-if(\alpha)} \partial_j e^{if(\alpha)} = \frac{1}{in_e} \partial_j \log e^{if(\alpha)} \quad \text{for } |\mathbf{x}| \rightarrow \infty \quad (1.17)$$

As a consequence of eq. (1.15) we have, asymptotically, the following map

$$e(\mathbf{x}) := \frac{\phi(\mathbf{x})}{|\phi(\mathbf{x})|} : \begin{array}{ccc} S^1 & \longrightarrow & S^1 \\ \alpha & \longmapsto & e^{if(\alpha)} \end{array} \quad (1.18)$$

which, if continuous, define a homotopy class in  $\pi_1(S^1) \simeq \mathbb{Z}$ .<sup>V</sup> These maps are classified by the number of times that  $e^{if(\alpha)}$  covers the circle labelled by  $\alpha$ . For example, if  $f(\alpha) = n\alpha$ , the homotopy class of  $e^{if(\alpha)}$  is  $[e^{if(\alpha)}] = n \in \pi_1(S^1)$ .

In fact, using continuous gauge transformations one can always put  $f(\alpha)$  in the form  $n\alpha$  for some  $n \in \mathbb{Z}$ . Let's prove this. From eq. (1.17) we have that

$$\begin{aligned} \lim_{R \rightarrow \infty} \oint_{|\mathbf{x}|=R} A_j dx^j &= \int_0^{2\pi} d\alpha \frac{1}{in_e} \frac{dx^j}{d\alpha} \partial_j \log e^{if(\alpha)} \\ &= \frac{1}{in_e} \int_0^{2\pi} d\alpha \frac{d}{d\alpha} \log e^{if(\alpha)} \\ &= \frac{1}{n_e} (f(2\pi) - f(0)) = 2\pi \frac{n}{n_e} \end{aligned} \quad (1.19)$$

hence we can compute  $n$  starting from  $A_j$ . If  $\beta(x)$  labels a gauge transformation globally defined in  $\mathbb{R}^2$  and its derivative is continuous (i.e.  $\beta \in \mathcal{C}^1$ ) then by Stokes theorem

$$\oint_{|\mathbf{x}|=R} \partial_j \beta(x) dx^j = 0 \quad (1.20)$$

hence, for a given  $f(\alpha)$  and the corresponding  $n$ , we can take  $\beta(x)$  such that

$$\beta(x) \xrightarrow{|\mathbf{x}| \rightarrow \infty} -f(\alpha) + n\alpha \quad (1.21)$$

so that we can replace  $f(\alpha)$  by  $n\alpha$ . Such gauge transformation, even if it acts also at boundaries, do not change the boundary conditions, thanks to eq. (1.20), hence it is allowed in our gauge theory.

Using Stokes theorem one can also derive

$$\lim_{R \rightarrow \infty} \oint_{|\mathbf{x}|=R} A_j \frac{dx^j}{2\pi} = \int_{\mathbb{R}^2} \frac{\epsilon_{ij} F^{ij}}{2\pi} d^2x = \int_{\mathbb{R}^2} \frac{B(x)}{2\pi} d^2x \quad (1.22)$$

where  $B(x)$  is the magnetic field associated to  $F^{ij}$ . Hence  $n$  can also be interpreted as the magnetic flux, and rescaling  $A_\mu$  to give the standard form of the Maxwell free action  $-\frac{1}{4}F_{\mu\nu}^2$ , we get

$$\int B d^2x = n \frac{2\pi}{e} \hbar = n \frac{\hbar}{e} \quad (1.23)$$

where we restored  $\hbar$ . Notice that  $\frac{\hbar}{e}$  is the unit of quantum flux. The number  $n$  is also called *vorticity* and its topological origin is at the basis of the stability of the vortices that we will now discuss in details. The configurations minimizing the energy for  $n = \pm 1$  are called *vortex* and *anti-vortex* respectively. Plotting the complex field  $\phi$  as a vector, we get fig. ??, which shows clearly the analogy between our discussion and well known vortices in water.

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<sup>V</sup>The homotopy class  $\pi_1(S^1)$  is the set of equivalent maps from  $S^1$  to  $S^1$ , where the equivalence is defined as follows: two maps are said equivalent if it is possible to deform one map into the other continuously, or better, if exists a continuous map interpolating between the two initial ones. The set  $\pi_1(S^1)$  is isomorphic to  $\mathbb{Z}$  since to each equivalence class in  $\pi_1(S^1)$  can be associated the number of windings of the map around the second circle.

## The vortex solution

According to previous considerations, we want to see how the vortices are localized (which is a property required for solitons), at least qualitatively. Let us consider the case  $n = 1$ ,  $f(\alpha) = \alpha$ , and in the usual notation for polar coordinates we set  $r \equiv |\mathbf{x}| = \sqrt{x_1^2 + x_2^2}$  and  $\varphi \equiv \alpha$ . Using  $\varphi = \arctan \frac{x_2}{x_1}$  we get the following asymptotic behaviours

$$\begin{aligned} A_i(\mathbf{x}) &= \frac{1}{n_e} \partial_i \varphi = -\frac{1}{n_e} \epsilon_{ij} \frac{x^j}{r^2} \quad \text{for } |\mathbf{x}| \rightarrow \infty \\ \phi(\mathbf{x}) &= v e^{i\varphi} = v \frac{x^1 + ix^2}{r} \end{aligned} \quad (1.24)$$

We can improve our approximation in finite regions of the spacetime introducing some corrections to our fields, in such a way that we can really obtain a solution of the equations of motion, in terms of some functions  $g_A(r)$  and  $g_H(r)$ :<sup>VI</sup>

$$\begin{aligned} A_i(\mathbf{x}) &= \frac{1}{n_e} \partial_i \varphi (1 - g_A(r)) = -\frac{1}{n_e} \epsilon_{ij} \frac{x^j}{r^2} (1 - g_A(r)) \\ \phi(\mathbf{x}) &= v e^{i\varphi} (1 - g_H(r)) = v \frac{x^1 + ix^2}{r} (1 - g_H(r)) \end{aligned} \quad (1.25)$$

We claim that a characteristic feature of vortex configurations is that  $F_{ij}$  is significantly different from 0 in a region of radius  $O(m_A^{-1})$  from the center of the vortex and  $\frac{|\phi|}{v}$  is significantly different from 1 in a region of radius  $O(m_H^{-1})$ , or more precisely

$$\begin{aligned} g_A(r) &\sim e^{-m_A r} \\ g_H(r) &\sim e^{-m_H r} \end{aligned} \quad (1.26)$$

We will show this with some simplifications. Furthermore, continuity at 0 implies that

$$g_A(0) = 1 = g_H(0) \quad (1.27)$$

so that

$$\phi(0) = 0 \quad (1.28)$$

In polar coordinates we can rewrite the first of eq. (1.25) as

$$A_r = 0 \quad \text{and} \quad A_\varphi = \frac{1}{n_e} (1 - g_A(r)) \quad (1.29)$$

Since

$$F_{r\varphi} = \partial_r A_\varphi - \partial_\varphi A_r \quad (1.30)$$

we get (recall that in polar coordinates the metric which we use to raise indices is not the flat one)

$$\partial^r F_{r\varphi} = \partial_r \frac{1}{r} \partial_r A_\varphi \quad (1.31)$$

so that the equation of motion for  $A_\varphi$  become

$$\partial_r \frac{1}{r} \partial_r A_\varphi = 2i \frac{n_e e^2}{r} (\phi^* \partial_\varphi \phi - i n_e \phi^* A_\varphi \phi) \quad (1.32)$$

Introducing our ansatz eq. (1.25) we get

$$r \frac{d}{dr} \frac{1}{r} \frac{d}{dr} g_A(r) - 2n_e^2 e^2 v^2 (1 - g_H(r))^2 g_A(r) = 0 \quad (1.33)$$

and linearizing at large  $r$  (hence  $g_H \sim 0$ ) we get

$$r \frac{d}{dr} \frac{1}{r} \frac{d}{dr} g_A(r) - m_A^2 g_A(r) = 0 \quad (1.34)$$

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<sup>VI</sup>Such functions depends only on  $r$  in such a way that we preserve the behaviour of the vortex.

Putting  $g_A(r) = r^\gamma \tilde{g}_A(r)$  we have

$$(2\gamma - 1)r^{\gamma-1} \frac{d\tilde{g}_A(r)}{dr} + r^\gamma \left[ \frac{d^2 \tilde{g}_A(r)}{dr^2} - m_A^2 \tilde{g}_A(r) \right] = 0 \quad (1.35)$$

We see that the piece in the square brackets give us the exponential behaviour, and in order to compensate  $r^{\gamma-1}$  and  $r^\gamma$  we necessarily have  $\gamma = \frac{1}{2}$ , so that the first factor cancels. Finally we have

$$g_A(r) \sim \sqrt{r} e^{-m_A r} \quad (1.36)$$

It worth to notice that  $g_A(r)$  decay exponentially with the penetration depth  $m_A^{-1}$ .  
Doing the same for the  $\phi$ , we have the following equation of motion:

$$\Delta_A \phi + m_e^2 e^2 \phi (|\phi|^2 - v^2) = 0 \quad (1.37)$$

where the covariant laplacian  $\Delta_A$  (recall that  $A_r = 0$ ) can be written in polar coordinate as

$$\Delta_A = \frac{1}{r} \frac{d}{dr} r \frac{d}{dr} + \frac{1}{r^2} \left( \frac{\partial}{\partial \varphi} - i A_\varphi \right)^2 \quad (1.38)$$

so that introducing the ansatz eq. (1.25) we get

$$\frac{1}{r} \frac{d}{dr} r \frac{d}{dr} g_H(r) + \frac{1}{r^2} g_A^2(r) (1 - g_H(r)) + 2\lambda v^2 (1 - g_H(r)) (-2g_H(r) + g_H^2(r)) = 0 \quad (1.39)$$

Again linearizing at large  $r$ , we can neglect non-linear terms in  $g_A$  and  $g_H$ , so that the equation of motion become

$$\frac{1}{r} \frac{d}{dr} r \frac{d}{dr} g_H(r) - m_H^2 g_H(r) = 0 \quad (1.40)$$

and, inserting  $g_H(r) = r^\gamma \tilde{g}_H(r)$ , we finally get the behaviour of  $g_H$ :

$$g_H(r) \sim \frac{1}{\sqrt{r}} e^{-m_H r} \quad (1.41)$$

Hence both the magnetic field associated to  $A_\mu$  and the scalar field  $\phi$  approach the vacuum outside of a compact region of radius  $O(m_A^{-1}, m_H^{-1})$  around the center of the vortex, which therefore behaves as an “extended particle”.

### Interactions among vortices - BPS limit

In general these vortices interact among each other, but for  $m_H = m_A$ , i.e.  $\lambda = \frac{1}{2} n_e^2 e^2$  a clever inequality (again of the Bogomol’nyi type, sometimes called *Bogomol’nyi-Prasad-Sommerfield (BPS)* in this context) show that they are free, non-interacting.

To prove such inequality, let us start by rewriting the energy density as

$$\begin{aligned} \mathcal{E}(\mathbf{A}(\mathbf{x}), \phi(\mathbf{x})) &= \int d^2x \frac{1}{2} \left[ \frac{B}{e} + n_e e (|\phi|^2 - v^2) \right]^2 + |(D_1 + iD_2)\phi|^2 \\ &\quad + \left( \lambda - \frac{n_e^2 e^2}{2} \right) (|\phi|^2 - v^2)^2 + n_e e B v^2 - i\partial_j (\epsilon_{ij} \phi^* D_i \phi) \end{aligned} \quad (1.42)$$

Indeed using

$$\begin{aligned} |(D_1 + iD_2)\phi|^2 &= |D_i \phi|^2 + i\epsilon_{ij} \phi^* D_j D_i \phi + i\partial_j (\epsilon_{ij} \phi^* D_i \phi) \\ &= |D_i \phi|^2 - n_e e B |\phi|^2 + i\partial_j (\epsilon_{ij} \phi^* D_i \phi) \\ &= |D_i \phi|^2 - n_e e B (|\phi|^2 - v^2) + i\partial_j (\epsilon_{ij} \phi^* D_i \phi) - n_e e B v^2 \end{aligned} \quad (1.43)$$

we get from eq. (1.42)

$$\begin{aligned} \mathcal{E} &= \frac{1}{2e^2} B^2 + \frac{n_e^2 e^2}{2} (|\phi|^2 - v^2)^2 + B n_e (|\phi|^2 - v^2) + |D_i \phi|^2 - n_e e B (|\phi|^2 - v^2) \\ &\quad + i\partial_j (\epsilon_{ij} \phi^* D_i \phi) - n_e e B v^2 + n_e e B v^2 - i\partial_j (\epsilon_{ij} \phi^* D_i \phi) + \left( \lambda - \frac{n_e^2 e^2}{2} \right) (|\phi|^2 - v^2)^2 \\ &= \frac{1}{2e^2} B^2 + |D_i \phi|^2 + \lambda (|\phi|^2 - v^2)^2 \end{aligned} \quad (1.44)$$

hence eq. (1.42) really coincides with eq. (1.6).

Notice that

$$\int_{\mathbb{R}^2} d^2x \partial_j (\epsilon_{ij} \phi^* D_i \phi) = \int_{S_\infty^1} \epsilon_{ij} \phi^* D_i \phi dx^j = 0 \quad (1.45)$$

hence

$$\mathcal{E} = \int d^2x \frac{1}{2} \left[ \frac{B}{e} + n_e e (|\phi|^2 - v^2) \right]^2 + |(D_1 + iD_2)\phi|^2 + \left( \lambda - \frac{n_e^2 e^2}{2} \right) (|\phi|^2 - v^2)^2 + n_e e B v^2 \quad (1.46)$$

and

$$\text{if } \lambda \geq \frac{n_e^2 e^2}{2} \quad \text{then} \quad \mathcal{E} \geq n_e e v^2 \int d^2x B(x) \quad (1.47)$$

In particular for the *BPS limit* or *Bogomol'nyi limit*, where  $\lambda = \frac{n_e^2 e^2}{2}$ , i.e.  $m_A = m_H$ , and

$$\begin{cases} \frac{1}{e} |B| + n_e e (|\phi|^2 - v^2) = 0 \\ (D_1 + iD_2)\phi = 0 \end{cases} \quad (1.48)$$

then the energy is proportional to the vorticity:

$$\mathcal{E} = n_e e v^2 \int d^2x B(x) = 2\pi n v^2 \quad (1.49)$$

where we used the previous identity

$$\int d^2x B(x) n_e e = 2\pi n \quad (1.50)$$

Since  $\mathcal{E} \sim n$  this means that vortices are free in the BPS limit.

The configurations saturating the *BPS bound* (in this case  $m_A = m_H$ ) are called *BPS states* and even in higher dimensions, where instead of vortices forming world-lines we have higher dimensional vortex defects, they play an important role in supersymmetric theories, since they preserve some supersymmetry generators and guarantees a special stability of the system (for instance, in our case, BPS states are non-interacting).

From eq. (1.47) we see that for  $\lambda > \frac{n_e^2 e^2}{2}$  the energy density is greater than in the non-interacting case, hence vortices repels each others, whereas for  $\lambda < \frac{n_e^2 e^2}{2}$  the energy density is smaller and vortices attract each others.

For the BPS solution

$$|B| = n_e e^2 (v^2 - |\phi|^2) \quad (1.51)$$

while for  $\lambda < \frac{n_e^2 e^2}{2}$  it turns out that

$$|B| \leq c(v^2 - |\phi|^2) \quad (1.52)$$

for some constant  $c$ . If vortices attracts each other, the value of  $|B|$  increases and at some point we have a violation of eq. (1.52), which implies that vortices cannot exist. Therefore for  $\lambda > \frac{n_e^2 e^2}{2}$  we have superconductors of type II, while for  $\lambda < \frac{n_e^2 e^2}{2}$  we have superconductors of type I.

### Sketch of semi-classical treatment

As for the position  $x_0$  of the kink, the two coordinates of the center of the vortex are the moduli of the vortex, and break the translational invariance of the theory. A semi-classical quantization of the vortex can be obtained by promoting those coordinates to be quantum operators  $\hat{x}_0(t) := (\hat{x}_0^1(t), \hat{x}_0^2(t))$  depending on  $t$ , restoring the translational invariance.

## 1.2 Quantum field theory treatment

Let us now outline briefly the construction of the quantum vortex operator, using the path-integral formalism, analogously to the construction of the kink operator.

### Closed defects

In 3 Euclidean dimensions the vortex is a line defect  $D$ . In the path integral formulation of the partition function  $D$  appears where  $|\phi| = 0$ , which is the *locus of the defect*.

In the partition function, for a typical configuration far away from the locus of a defect we have  $|\phi| \sim v$  and  $A_\mu \sim \partial_\mu \theta$ , which implies  $F_{\mu\nu}$ , hence such configurations are close to the vacuum configurations.

Instead, near the locus of a defect, the configurations should be close to a vortex solution in order to maintain the action close to a minimum.

As for the kink, in fig. 1.2 we pictorially describe a typical field configuration of the improved semi-classical approximation for the partition function.

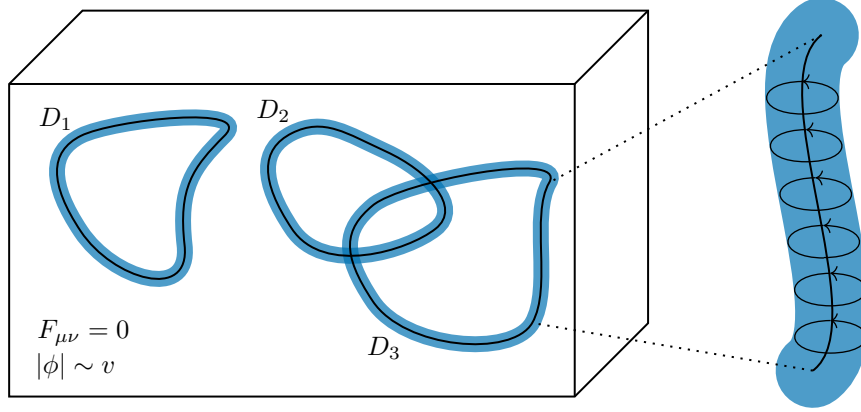


Figure 1.2: Defects (blue closed lines) and their loci (black closed lines). On the right side of the picture, is represented in more detail the vortex associated to the defect  $D_3$ .

In order to understand how to “open” these line defects, let’s first consider the restriction of the field strength  $F$  of the vortex to a 2 dimensional disk crossing transversally the locus of the defect, such that the boundary of the disk identify the region where the configuration approximately approaches the vacuum one, as represented in fig. 1.3.

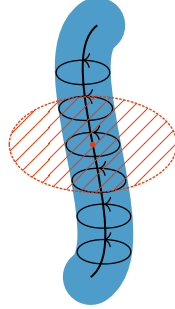


Figure 1.3: Disk crossing the locus of the defect, whose boundary identify the region where the configuration approaches the vacuum.

We then identify the points of the boundary to one-point, obtaining a closed 2-dimensional surface  $S^2$ , as represented in fig. 1.4.

According to our previous discussion,

$$\frac{1}{2\pi} \int_{S^2} F_{\mu\nu} dS^{\mu\nu} = n = \text{vorticity carried by the defect } D \quad (1.53)$$

Let’s see how  $A_\mu$  behaves on such sphere. Using definition eq. (1.18), we get

$$\begin{aligned} D_\mu \phi &= e^{i\theta} \partial_\mu |\phi| + |\phi| e^{i\theta} (e^{-i\theta} \partial_\mu e^{i\theta} - iA_\mu) \\ (D_\mu \phi)^* &= e^{-i\theta} \partial_\mu |\phi| + |\phi| e^{-i\theta} (e^{i\theta} \partial_\mu e^{-i\theta} + iA_\mu) \end{aligned} \quad (1.54)$$



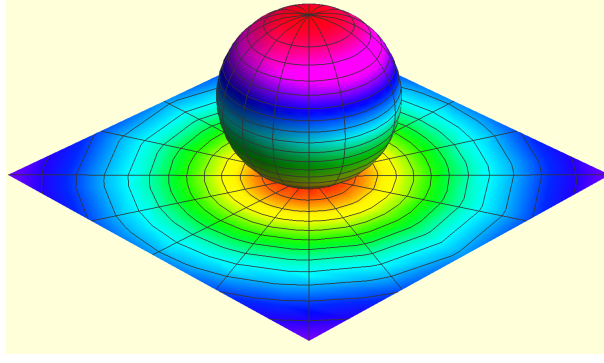


Figure 1.4: Compactification of the plane by identification of the boundaries. Source: <https://demonstrations.wolfram.com/TheRiemannSphereAsAStereographicProjection/>

where we set  $n_e e = 1$  in order to get rid of these coefficients. Hence we obtained

$$A_\mu = \underbrace{\frac{i}{2} \frac{e^{-i\theta} D_\mu \phi - e^{i\theta} (D_\mu \phi)^*}{|\phi|}}_{\text{regular term}} + \underbrace{\frac{i}{2} (e^{i\theta} \partial_\mu e^{-i\theta} - e^{-i\theta} \partial_\mu e^{i\theta})}_{\text{singular term}} \quad (1.55)$$

Come mai il secondo termine può esser singolare?

Recall that  $\theta(x)$  is an angle around the center of the vortex, where  $|\phi| = 0$ . Without loss of generality we can assume that the center of the vortex coincide with the origin of the space. In the sense of distributions, the field strength of the singular term above is

$$\epsilon^{\mu\nu} \partial_\mu \left( \frac{i}{2} (e^{i\theta} \partial_\nu e^{-i\theta} - e^{-i\theta} \partial_\nu e^{i\theta}) \right) = \frac{1}{i} \epsilon^{\mu\nu} \partial_\mu \partial_\nu \log e^{i\theta} = 2\pi \delta(\mathbf{x}) \quad (1.56)$$

since the dependence of  $\theta$  on the space coordinates goes with  $\arctan \frac{y}{x}$ . Hence such field strength is non zero (actually, singular) only in the center of the vortex. Therefore the complete field strength of  $A_\mu$  is

$$F_{\mu\nu}(\mathbf{x}) = \partial_{[\mu} a_{\nu]}(\mathbf{x}) + 2\pi \delta(|\phi|) \epsilon_{\mu\nu} \quad (1.57)$$

so that the support of  $F_{\mu\nu}$  is given by the loci of defects, i.e. where  $|\phi| = 0$ .

We see then that one can identify the locus of the defect in  $F_{\mu\nu}$  in terms of a “singular current” (i.e. “ $\delta$ -like”) in the support where  $|\phi|$  vanishes, which is necessarily closed because the boundary condition at infinity is  $|\phi| = v$ .

Che cos'è  $a_\nu$ ? Forse intendeva  $A_\nu$ ? Inoltre ho visto che nelle note ha modificato la formula data a lezione, volevo chiederle quale fosse la versione corretta (stessa cosa anche per la formula precedente).

## Open defects

If we want to open a defect and  $\mathbf{x}$  is a boundary of the locus of such defect, we need to construct a 2-tensor  $F_{\mu\nu}^{\mathbf{x}}$  such that for a surface  $S_{\mathbf{x}}^2$  centered at  $\mathbf{x}$  and of arbitrarily small radius we have

$$\int_{S_{\mathbf{x}}^2} F_{\mu\nu}^{\mathbf{x}} dx^\mu dx^\nu = 2\pi n \quad (1.58)$$

and then if  $B_{\mathbf{x}}^3$  is the ball centered in  $\mathbf{x}$  whose boundary is  $S_{\mathbf{x}}^2$ , by Gauss theorem we have

$$\int_{S_{\mathbf{x}}^2} F_{\mu\nu}^{\mathbf{x}} dx^\mu dx^\nu = \int_{B_{\mathbf{x}}^3} \epsilon^{\mu\nu\rho} \partial_\mu F_{\nu\rho}^{\mathbf{x}} d^3x = 2\pi n \quad (1.59)$$

which should hold for any arbitrarily small radius, so that

$$\epsilon^{\mu\nu\rho} \partial_\mu F_{\nu\rho}^{\mathbf{x}}(\mathbf{y}) = 2\pi n \delta(\mathbf{x} - \mathbf{y}) \quad (1.60)$$

or equivalently, using  $\epsilon^{\mu\nu\rho} F_{\nu\rho}^{\mathbf{x}} = B_{\mathbf{x}}^\mu$  where  $B_{\mathbf{x}}^\mu$  is the magnetic field centered at  $\mathbf{x}$ ,

$$(\nabla \cdot \mathbf{B}_{\mathbf{x}})(\mathbf{y}) = 2\pi n \delta(\mathbf{x} - \mathbf{y}) \quad (1.61)$$

This is precisely the magnetic field of a *monopole* in  $\mathbb{R}^3$ , i.e. the analogue of a point-like electric charge of charge  $q$  which obeys

$$(\nabla \cdot \mathbf{E}_x)(\mathbf{y}) = q \delta(\mathbf{x} - \mathbf{y}) \quad (1.62)$$

with the magnetic field replacing the electric field.

Therefore we can build open defects constructing line defects with monopoles at the boundaries. Respect to fig. 1.3, this is equivalent to the introduction of two compactified disks (hence they actually are 2-spheres) at the boundaries of the line, describing classical monopoles, as represented in fig. 1.5. Comparing eq. (1.53) and eq. (1.61) we see that such monopoles “provide the right vorticity” to the defect.



Figure 1.5: Representation of an open defect, whose boundaries are described by monopoles (red circles).

### Correlators

Similarly to what we have done for the kink, where we modified the action introducing a new parallel transporter to construct the open defects, now we should insert classical monopoles in the action at the positions in which we want to find the creation or the annihilation of the quantum vortices of our theory. Correlators of the quantum vortex field will have insertion points at the monopoles positions.