We're given the van der Waals equation of state in (33) and the free energy F in (38).

Relevant equations:

$$\begin{split} U &= F + \sigma \tau \\ \sigma &= -\frac{\partial F}{\partial \tau}_{V,N} \\ p &= -\frac{\partial F}{\partial V}_{\tau,N} \end{split}$$

We're also given pressure in (39)

$$p = \frac{N\tau}{V - Nb} - \frac{N^2a}{V^2}$$

(a)

$$F_{ ext{vdW}} = -N au \left[\ln\!\left(rac{n_Q(V-Nb)}{N}
ight) + 1
ight] - rac{N^2a}{V}$$

Note that n_Q depends on au, so let's use n_Q' independent of au

$$\begin{split} F_{\text{vdW}} &= -N\tau \left[\ln \left(\tau^{3/2} \frac{n'_Q(V - Nb)}{N} \right) + 1 \right] - \frac{N^2 a}{V} \\ \sigma &= -\frac{\partial F}{\partial \tau} = N + N \frac{\partial}{\partial \tau} \tau \ln \left(\tau^{3/2} \frac{n'_Q(V - Nb)}{N} \right) \\ &= N + N \frac{\partial}{\partial \tau} \left[\tau \ln \left(\tau^{3/2} \right) + \tau \ln \left(\frac{n'_Q(V - Nb)}{N} \right) \right] \\ &= N + N \ln(\tau^{3/2}) + \frac{3}{2} + N \ln \left(\frac{n'_Q(V - Nb)}{N} \right) \\ &= N + N \ln \left(\frac{n_Q(V - Nb)}{N} \right) + N \times 3/2 \\ &= N \left(\ln \left[\frac{n_Q(V - Nb)}{N} \right] + 5/2 \right) \end{split}$$

(b)

$$\begin{split} U = F + \sigma \tau &= N\tau \left[\ln \left[\frac{n_Q(V - Nb)}{N} \right] + \frac{5}{2} - \ln \left[\frac{n_Q(V - Nb)}{N} \right] - 1 \right] - \frac{N^2 a}{V} \\ &= \frac{3}{2} N\tau - N^2 a / V \end{split}$$

(c) From p above

$$p=rac{N au}{V-Nb}-rac{N^2a}{V^2} \ H(au,p,V)=rac{3}{2}N au-rac{N^2a}{V}+pV$$

How do we get $N\tau$ out of pV? It's in the equation of state...

```
In [46]:
    import sympy as sp
p, N, V, a, b, tau = sp.symbols('p N V a b tau')
```

eos =
$$(p+N**2*a/V**2)*(V-N*b)$$

eos.expand()

$$\begin{array}{ll} \text{Out[46]:} & -\frac{N^3ab}{V^2} + \frac{N^2a}{V} - Nbp + Vp \end{array}$$

Since we know the result we're going for, we can add in N au and subtract out the above from H

$$H(au,p,V)=rac{5}{2}N au-rac{2N^2a}{V}+Nbp-rac{N^3ab}{V^2}$$

To finish, we just have to check some equalities:

$$Nbp - \frac{N^3ab}{V^2} = \frac{N^2b\tau}{V}$$
$$-\frac{2N^2a}{V} - \frac{N^3ab}{V^2} = \frac{-2Nap}{\tau}$$

For both of these, plugging in our formula for p and simplifying will yield our equation of state

$$\tau = \frac{pV}{N} - \frac{Na}{V}$$

proving the equality.

(a) The only complexity of this problem compared to the example is finding $Z=\sum_n e^{\frac{-n\hbar\omega+\epsilon_0}{\tau}}$, only with a new n

$$egin{align} Z &= \sum_{n_x} \sum_{n_y} \sum_{n_z} e^{rac{-\sqrt{n_x^2 + n_y^2 + n_z^2 \hbar \omega + \epsilon_0}{ au}}{ au}} \ &= rac{1}{8} \int_0^\infty 4\pi n^2 e^{rac{-n\hbar \omega + \epsilon_0}{ au}} dn \ &= rac{\pi e^{\epsilon_0/ au}}{2} iggl[rac{2 au^3}{\hbar^3 \omega^3} iggr] = \pi e^{rac{\epsilon_0}{ au}} iggl(rac{ au}{\hbar \omega} iggr)^3 \end{split}$$

We are trying to find the pressure when the solid and gas are in equilibrium

$$\lambda_g = \lambda_s$$

As in (10.30), we'll use the ideal gas approximation for λ_g

$$\lambda_g = rac{n}{n_Q} = rac{p}{ au n_Q} = rac{p}{ au} igg(rac{2\pi\hbar^2}{M au}igg)^{3/2}$$

For λ_{s} , we need to assume a small solid volume and apply (10.28)

$$egin{align} \mu_s &= F_s + p v_s \simeq F_s = - \ln Z \ \lambda_s &\equiv e^{\mu_s/ au} = rac{1}{Z} = rac{e^{-\epsilon_0/ au}}{\pi} igg(rac{\hbar \omega}{ au}igg)^3 \ \end{aligned}$$

This gives our pressure as

$$egin{align} \lambda_g &= \lambda_s \ rac{p}{ au} igg(rac{2\pi\hbar^2}{M au}igg)^{3/2} &= rac{e^{-\epsilon_0/ au}}{\pi} igg(rac{\hbar\omega}{ au}igg)^3 \ p &\simeq igg(rac{M}{2\pi}igg)^{3/2} rac{\omega^3}{ au^{1/2}} e^{-\epsilon_0/ au} \ \end{aligned}$$

(b) Using (10.18) we have

$$egin{align} L = au^2 rac{d}{d au} \ln p = au^2 rac{d}{d au} iggl[\ln iggl(rac{e^{-\epsilon_0/ au}}{ au^{1/2}} iggr) + iggl(\ln iggl(rac{M}{2\pi} iggr)^{3/2} \omega^3 iggr) iggr] \ &= rac{2\epsilon_0 - au}{2} = \epsilon_0 - rac{1}{2} au \end{split}$$

The question actually asks for us to explain *why* this is the case. We're able to use (10.18) thanks to the various assumptions we made earlier. Latent heat decreases with temperature because

Rather than recalculating the phonon energy at low temperatures, we'll use our previous result

$$U(\tau) \simeq \frac{3\pi^4 N \tau^4}{5(k_B \theta)^3} \tag{4.46}$$

$$C_V = \frac{12\pi^4 N}{5} \left(\frac{\tau}{k_B \theta}\right)^3 \tag{4.47}$$

The free energy density is

$$F \equiv U - \tau \sigma$$

$$f = \frac{F}{V} = \frac{U}{V} - \tau \frac{\sigma}{V}$$
(3.35)

leaving σ as the last value to find. We actaully have a formula for entropy from heat capacity

$$\sigma(\tau) - \sigma(0) = \int_0^{\tau} \frac{C_V(\tau)}{\tau} d\tau$$

$$\sigma(\tau) = \frac{12\pi^4 N}{5(k_B \theta)^3} \int_0^{\tau} \tau^2 d\tau = \frac{4\pi^4 N \tau^3}{5(k_B \theta)^3}$$

$$f = \frac{3\pi^4 N \tau^4}{5V(k_B \theta)^3} - \frac{4\pi^4 N \tau^4}{5V(k_B \theta)^3}$$
(6.39)

Now to clean things up

$$egin{align} heta &= rac{\hbar v}{k_B} \left(6 \pi^2 n
ight)^{1/3} \ f &= -rac{n \pi^4 au^4}{5 (6 \pi^2 n \hbar^3 v^3)} = -rac{\pi^2 au^4}{30 v^3 \hbar^3} \end{split}$$

(b) Since α is stable at low temperature, we know $U_{\alpha}(0) < U_{\beta}(0)$, and $U_{\beta}(0) - U_{\alpha}(0)$ is positive. We know that τ_c^4 must be positive, so v_{β} must be less than v_{α} .

The only energies here are the ground state energy U(0) and phonon energy f_t so the transition occurs when

$$U_{\alpha}(0) + f_{\alpha}(\tau) = U_{\beta}(0) + f_{\beta}(\tau)$$

These are equal at $au= au_c$, so we know a solution exists, and exists only when $v_{eta} < v_{lpha}$

(c) Since we have phonon entropy and au_c , we need only find $\Delta\sigma$

$$egin{align} L &= au \Delta \sigma = \sigma_{eta}(au_c) - \sigma_{lpha}(au_c) \ &= \left[rac{2\pi^2 au_c^4}{15\hbar^3} \left(v_{eta}^{-3} - v_{lpha}^{-3}
ight)
ight] \ &= rac{2\pi^2}{15\hbar^3} \left[rac{30\hbar^3}{\pi^2} rac{U_{eta}(0) - U_{lpha}(0)}{v_{eta}^{-3} - v_{lpha}^{-3}} \left(v_{eta}^{-3} - v_{lpha}^{-3}
ight)
ight] \ &= 4 \left[U_{eta}(0) - U_{lpha}(0)
ight] \end{split}$$