(a) Helmholtz free energy $F(au) = - au \ln Z$

$$Z = e^{-0} + e^{-\epsilon/\tau} = 1 + e^{-\epsilon/\tau}$$

$$F(\tau) = -\tau \ln \left(1 + e^{-\epsilon/\tau}\right)$$

That was easy.

(b) Helmholtz free energy $F(au) = U - au \sigma$

In the canonical ensemble, the probability of being in a state is

$$\sigma = -\frac{\partial F}{\partial \tau} = \ln \left(e^{-\epsilon/\tau} + 1 \right) + \frac{\epsilon}{\tau} \frac{e^{-\epsilon/\tau}}{e^{-\epsilon/\tau} + 1}$$

This makes finding the energy easy

$$\begin{split} U(\tau) &= F(\tau) + \tau \sigma \\ &= -\tau \ln \Bigl(1 + e^{-\epsilon/\tau} \Bigr) + \tau \ln \Bigl(e^{-\epsilon/\tau} + 1 \Bigr) + \epsilon \frac{e^{-\epsilon/\tau}}{e^{-\epsilon/\tau} + 1} \\ &= \epsilon \frac{e^{-\epsilon/\tau}}{1 + e^{-\epsilon/\tau}} \end{split}$$

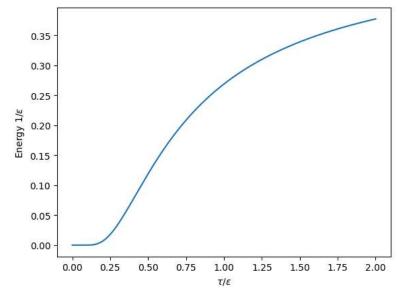
import matplotlib.pyplot as plt
from numpy import exp, linspace

temp = linspace(0, 2, 1000)
energy = exp(-1/temp)/(1+exp(-1/temp))

plt.plot(temp, energy)
plt.xlabel(r'\$\tau/\epsilon\$')
plt.ylabel(r'Energy \$1/\epsilon\$')

C:\Users\andre\AppData\Local\Temp\ipykernel_27408\2397519148.py:5: RuntimeWarning: divide by zero encountered in divide energy = exp(-1/temp)/(1+exp(-1/temp))

Out[133... Text(0, 0.5, 'Energy \$1/\\epsilon\$')



▶ Q2

(a)

$$Z = \sum_{s=0}^{\infty} e^{-\epsilon_s/ au} = \sum_{s=0}^{\infty} e^{-s\hbar\omega/ au}$$

$$\sum_{a=0}^{\infty} e^{a\,b} = -\frac{1}{e^b-1} \approx -\frac{1}{2.71828^b-1} \quad \text{when} \quad e^{\text{Re}(b)} < 1$$

$$Z=rac{1}{1-e^{-\hbar\omega/ au}}$$

(b) Easy! $F=- au \ln Z$

$$F = - au \ln\!\left(rac{1}{1-e^{-\hbar\omega/ au}}
ight) = au \ln\!\left(1-e^{\hbar\omega/ au}
ight)$$

(c) I'm actually going to skip this one and come back. I'd rather do the derivative of σ than do this one.

Using $U=F+ au\sigma$ we have

$$egin{align} U &= au \ln\!\left(1 - e^{\hbar\omega/ au}
ight) - au \ln\!\left(1 - e^{\hbar\omega/ au}
ight) + rac{\hbar\omega}{e^{\hbar\omega/ au} - 1} \ &= rac{\hbar\omega}{e^{\hbar\omega/ au} - 1} \end{split}$$

See derivation of σ below

(d) Here we'll take the derivative of ${\cal F}$

$$\sigma = \frac{\partial F}{\partial \tau}$$

In [71]: from sympy import symbols, ln, exp, diff, latex, Derivative
 hbar, omega, tau = symbols('hbar omega tau')
 print('Helmholtz energy:')
 F = tau*ln(1-exp(-hbar*omega/tau))
 display(F)

sigma = - diff(F, tau)
 print('Entropy:')
 sigma

Helmholtz energy:

$$\tau \log \left(1 - e^{-\frac{\hbar \omega}{\tau}}\right)$$

Entropy:

$$\begin{array}{cc} \text{Out[71]:} & \frac{\hbar \omega e^{-\frac{\hbar \omega}{\tau}}}{\tau \left(1 - e^{-\frac{\hbar \omega}{\tau}}\right)} - \log \left(1 - e^{-\frac{\hbar \omega}{\tau}}\right) \end{array}$$

This works out to be exactly what we're looking for:

$$\begin{split} &\frac{\hbar\omega e^{-\frac{\hbar\omega}{\tau}}}{\tau\left(1-e^{-\frac{\hbar\omega}{\tau}}\right)} \cdot \frac{e^{\hbar\omega/\tau}}{e^{\hbar\omega/\tau}} = \frac{\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} \\ &\sigma = \frac{\hbar\omega e^{-\frac{\hbar\omega}{\tau}}}{\tau\left(1-e^{-\frac{\hbar\omega}{\tau}}\right)} - \ln\left(1-e^{-\frac{\hbar\omega}{\tau}}\right) \\ &= \frac{\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} - \ln\left(1-e^{-\frac{\hbar\omega}{\tau}}\right) \end{split}$$

(e) HW2P3:
$$U=rac{N\hbar\omega}{e^{\hbar\omega/ au}-1}$$

Yes, it is N times my result. This makes sense since it is N oscillators, where this question asks about a single oscillator.

I'll use the result from the answer key

$$\sigma(U, N) = N \ln \left(1 + \frac{U}{N \hbar \omega} \right) + \frac{U}{\hbar \omega} \ln \left(1 + \frac{N \hbar \omega}{U} \right)$$

Let's insert in U(au)

$$\begin{split} \sigma(\tau,N) &= N \ln \left(1 + \frac{\frac{N\hbar\omega}{e^{\hbar\omega/\tau}-1}}{N\hbar\omega} \right) + \frac{\frac{N\hbar\omega}{e^{\hbar\omega/\tau}-1}}{\hbar\omega} \ln \left(1 + \frac{N\hbar\omega}{\frac{N\hbar\omega}{e^{\hbar\omega/\tau}-1}} \right) \\ &= N \ln \left(1 + \frac{1}{e^{\hbar\omega/\tau}-1} \right) + \frac{N}{e^{\hbar\omega/\tau}-1} \ln \left(1 + e^{\hbar\omega/\tau} - 1 \right) \\ &= N \ln \left(1 + \frac{1}{e^{\hbar\omega/\tau}-1} \right) + \frac{N\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} \\ &= N \ln \left(\frac{e^{\hbar\omega/\tau}}{e^{\hbar\omega/\tau}-1} \right) + \frac{N\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} \\ &= -N \ln \left(\frac{e^{\hbar\omega/\tau}}{e^{\hbar\omega/\tau}-1} \right) + \frac{N\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} \\ &= -N \ln \left(\frac{e^{\hbar\omega/\tau}-1}{e^{\hbar\omega/\tau}-1} \right) + \frac{N\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} \\ &= N \left(\frac{\hbar\omega/\tau}{e^{\hbar\omega/\tau}-1} - \ln \left(1 - e^{-\frac{\hbar\omega}{\tau}} \right) \right) \end{split}$$

Yippie! Entropy is also just multiplied by N since each oscillator isn't interacting and entropy adds like energy

(a) Adding energy -> ${\cal U}$ goes up.

$$F = U - \tau \sigma$$

$$\Delta F = \Delta U - \tau \Delta \sigma$$

The system does no work, so $\Delta F=0$

$$\Delta U = \tau \Delta \sigma$$
$$U = k_B T(\ln g)$$

Going from g_0 to $2g_0$ states:

$$U_f - U_i = k_B T(\ln 2g_0) - k_B T(\ln(g_0))$$

 $\Delta U = k_B T \ln 2 = 2.87 \times 10^{-21}$

In [123... from scipy.constants import k
U = k*300*ln(2)

U.evalf()

Out[123... $2.87097888507872 \cdot 10^{-21}$

(b) $\Delta U/ au = \Delta\sigma$

$$S = k_b \sigma = k_b \frac{U}{k_b T} = \frac{U}{T}$$

 $S = 9.57 \times 10^{-24}$

In [127... S = U/300 S.evalf()

 ${\tt Out[127...} \quad 9.56992961692908 \cdot 10^{-24}$

(a) We can find probabilities using ${\cal Z}$

$$Z = 1 + e^{-a/ au} + e^{-3a/ au}$$
 $P(1) = rac{e^{-a/ au}}{1 + e^{-a/ au} + e^{-3a/ au}}$

(b) Using a summation:

$$U = 0 + aP(1) + 3aP(2) = \frac{ae^{-a/ au} + 3ae^{-3a/ au}}{1 + e^{-a/ au} + e^{-3a/ au}}$$

(c)

$$\tau$$
 τ
 τ

high temp: all energies are equally likely

```
In [60]: a = symbols('a')
e0 = 1
e1 = exp(-a/tau)
e2 = exp(-3*a/tau)

Z = e0 + e1 + e2
(e0/Z).replace(tau,a*100).evalf()
```

 ${\tt Out[60]:}\ \ 0.337781308846565$

Return to microcanonical

(a)
$$\sigma(U) = \ln(g(U))$$
 easy

$$\sigma(U) = \ln\left(e^{\sqrt{NU/U_0}}\right) = \sqrt{NU/U_0}$$

(b)
$$1/ au = d\sigma/dU$$

$$\begin{split} \frac{1}{\tau} &= \frac{\partial}{\partial U} \sqrt{\frac{N}{U_0}} \sqrt{U} \\ &= \sqrt{\frac{N}{U_0}} \frac{\partial}{\partial U} \sqrt{U} = \frac{1}{2} \sqrt{\frac{N}{UU_0}} \\ &U(\tau) = \frac{N\tau^2}{4U_0} \end{split}$$

(c)

Note that conventional T and S have different formulas

$$au = k_B T \ \sigma = rac{S}{k_B}$$

au has units of energy, and N is unitless

$$\begin{split} \tau &= 4.14 \times 10^{-21} J \\ U(\tau) &= 258.19 J \\ S(U) &= 1.7 J/K \\ F &= U - \tau \sigma = -258.19 J \end{split}$$

Interesting how $F=-{\cal U}$

U=258.19 J Entropy =1.7212910700645034 J/K Helmholtz free energy=-258.194 J

If σau has units joules, then σ must be unitless (boltzmann constant J/K)