Running SeBa

This document explains how to run SeBa for stand-alone use (independent from AMUSE).

a)Simulation of one system:

with the parameters primary mass M=2 solar mass, secondary mass m= 1 solar mass, eccentricy e, orbital separation a=200 solar radii, time T=13500 Myrs, metallicity z=0.001.

./SeBa -M 2 -m 1 -e 0.2 -a 200 -T 13500 -z 0.001 default values e=0, T=13500, z =0.02 (solar)

b)Simulation of multiple system - though specific systems

For example

./SeBa -M 2 -m 1 -e 0.2 -a 200 -T 13500 -z 0.001

./SeBa -M 2.5 -m 1.5 -e 0.5 -a 500 -T 500 -z 0.02

Three options:

- 1) Run code multiple times :-) Not handy for more then 5 systems..
- 2) Use an input file: ./SeBa -I 'SeBa_input.txt'

SeBa_input.txt containts:

a e M m z

for example

200 0.2 2 1 0.001

500 0.5 2.5 1.5 0.02

3) Use a shell script. Big advantage: this method is applicable to all command line programs. For example a file named run.sh, should contain the lines for the example given above:

./SeBa -M 2 -m 1 -e 0.2 -a 200 -T 13500

./SeBa -M 2.5 -m 1.5 -e 0.5 -a 500 -T 500

Check permissions of run.sh file; it should be executable by the owner. If not: type 'chmod 744 run.sh 'in command line. To run the shell script: ./run.sh

c)Simulation of multiple system - random population

Monte Carlo based approach

./SeBa -R -n 200

./SeBa -R -n 250000 -m 0.96 -M 11 -q 1e-4 -Q 1 -A 1e6 -f 4 -T 13500

- -R SeBa generates randomly the initial parameters
- -n number of systems simulated
- -m -M min/max primary mass
- -q -Q min/max mass ratio
- -e -E min/max eccentricy
- -a -A min/max orbital separation

- -T time in Myr in the simulation of the binaries. Same time for all binaries
- -z metalicity of binary stars. Same metalicity for all binaries.
- -N initial ID number of first simulated binary (Default: 0, may come in handy for stitching together production runs)

```
initial parameters are drawn from probably distributions
////
          -x mass function exponent in case of power law [-2.35]
////
          -F/f mass function option: 0) Equal mass
////
                             1) Power-law [default]
////
                            2) Miller & Scalo
////
                            3) Scalo
////
                            4) Kroupa
////
          Option -F requires one of the following strings:
////
                 (mf_Power_Law, Miller_Scalo, Scalo, Kroupa)
////
               -f requires the appropriate interger (see mkmass.C)
////
           -y exponent for a power-law distribution [0] (flat in log)
////
           -G/g Semi major axis option: 0) Equal_sma
////
                              1) Power Law [default]
////
                              2) Duquennoy & Mayor (1987)
////
          Option -G requires one of the following strings:
////
                 (Egual sma, sma Power Law, Duguennov Mayor)
////
               -g requires appropriate interger (see double_star.h)
////
           -v exponent for a power-law distribution
////
           -U/u eccentricity option: 0) Equal eccentricity
////
                            1) Power Law
////
                            2) Thermal distribution [default]
////
          Option -U requires one of the following strings:
                 (Equal_ecc, ecc_Power_Law, Thermal_Distribution)
////
////
               -u requires appropriate interger (see double star.h)
////
           -w exponent for a power-law distribution
////
           -P/p mass ratio option: 0) constant mass ratio
////
                            1) Flat a
////
                            2) Power Law
////
                            3) Hogeveen (1992)
////
          Option -P requires one of the following strings:
                 (Equal_q, Flat_q, qf_Power_Law, Hogeveen)
////
////
               -p requires appropriate interger (see double star.h)
////
```