

9. Instabilities

We wish to ask the following sensible question:

if we have found an equilibrium which is static [$\vec{v} = \vec{0}$], but is rather general otherwise $P = P_0(\vec{r})$, $\rho = \rho_0(\vec{r})$, $\vec{B} = \vec{B}_0(\vec{r})$, is this ~~perturbation~~ ^{equilibrium} stable to small perturbations?

There are two general approaches to this problem:

- 1.) Solve a specific problem with a linearised normal mode analysis. In other words, write down a specific equilibrium, kick it [$P_0 \rightarrow P_0 + \delta P$, etc.], and look for modal solutions $\propto e^{-i\omega t + i\vec{k} \cdot \vec{r}}$. If $\omega \in \mathbb{C}$, then there may be exponentially growing modes \Rightarrow instability. However, this may be complex for a given P_0, ρ_0, \vec{B}_0 etc.
- 2.) There is a general procedure, "the energy principle", which can tell you whether an equilibrium is stable, without giving you a huge amount of physical information about the instability (or lack thereof).

We will begin with this general overview approach, which is a powerful technique, before showing a particular example of (1). from Astrophysics.

The energy principle

We shall prove, but this may be intuitively obvious, that a way of tackling the instability problem is the following: compute the change in ^{Potential} energy of the fluid resulting from a perturbation. If there is a way in which a perturbation can lower the ^{Potential} energy of the fluid, then this perturbation leads to an instability.

Let's begin.

The total energy in MHD is (see lecture 4)

$$E = \iiint_V \left(\frac{1}{2} \rho u^2 + \frac{B^2}{2\mu_0} + \frac{P}{\gamma-1} \right) dV$$

we will define

$$E \equiv \iiint_V \frac{1}{2} \rho u^2 dV + W.$$

As we saw in the lectures on waves (5+6), all perturbations of an MHD system can be expressed in terms of small displacements $\vec{\xi}$, where our perturbed velocity $\delta \vec{u} = \frac{\partial \vec{\xi}}{\partial t}$.

Then

← (no kinetic energy in equilibrium)

$$E \rightarrow E + \delta E = W_0 + \iiint_V \frac{1}{2} \rho_0 \left| \frac{\partial \vec{\xi}}{\partial t} \right|^2 dV + \delta W_1[\vec{\xi}] + \delta W_2[\vec{\xi}, \vec{\xi}] + \dots$$

where we have expanded to quadratic order, W_0 is our equilibrium potential, and we have split our perturbed potentials into linear $\delta W_1[\vec{\xi}]$ and quadratic $\delta W_2[\vec{\xi}, \vec{\xi}]$ orders.

Energy reduced if $\delta W < 0 \Rightarrow$ care about sign of $\delta W_1, \delta W_2$.

Energy must be globally conserved to all orders

[push volume integral to infinity].

This means we can be clever and work out δW_1 & δW_2

$$\frac{dE}{dt} = \frac{d\delta E}{dt} = \iiint \rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2} \cdot \frac{\partial \vec{\xi}}{\partial t} + \delta W_1 \left[\frac{\partial \vec{\xi}}{\partial t} \right] + \delta W_2 \left[\frac{\partial \vec{\xi}}{\partial t}, \vec{\xi} \right] + \delta W_2 \left[\vec{\xi}, \frac{\partial \vec{\xi}}{\partial t} \right] + \dots = 0.$$

This must be true at all times. Including at $t=0$, where $\vec{\xi} \wedge \frac{\partial \vec{\xi}}{\partial t}$ are independent. (These perturbations can be chosen independently as MHD equations are second order in time for $\vec{\xi}$). Define the "force operator"

$$\vec{F}[\vec{\xi}] \equiv \rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2},$$

then, at $t=0$, calling $\vec{\eta} \equiv \frac{\partial \vec{\xi}}{\partial t}$

$$\iiint_V \vec{\eta} \cdot \vec{F}(\vec{\xi}) dV + \delta W_1[\vec{\xi}] + \delta W_2[\vec{\eta}, \vec{\xi}] + \delta W_2[\vec{\xi}, \vec{\eta}] + \dots = 0.$$

Linear order: $\delta W_1[\vec{u}] = 0 \Rightarrow$ no linear energy perturbations.

Second order:

$$\iiint_V \vec{u} \cdot \vec{F}[\vec{\xi}] dV = -\delta W_2[\vec{u}, \vec{\xi}] - \delta W_2[\vec{\xi}, \vec{u}]$$

Can also set $\vec{u} = \vec{\xi}$ [at $t=0$], this trick

leads to

$$\delta W_2[\vec{\xi}, \vec{\xi}] = -\frac{1}{2} \iiint_V \vec{\xi} \cdot \vec{F}[\vec{\xi}] dV$$

It will transpire that δW_2 can have either sign, and therefore teaches us about instabilities.

Energy principle states

$$\delta W_2[\vec{\xi}, \vec{\xi}] > 0 \quad \forall \vec{\xi} \iff \text{stable equilibrium}$$

So we need to analyse the properties of the "force operator"

$$\vec{F}[\vec{\xi}]$$

Obviously, we will derive the functional form of \vec{F} from the linearised MHD equations, but we can actually do nearly all of our work in generality.

Property 1: $\vec{F}[\vec{\xi}]$ has simple eigenmodes $\vec{\xi}_n$.

By definition

$$\vec{F}[\vec{\xi}] = \rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2}$$

We see that for $\vec{\xi}_n = \vec{\xi}_n(\vec{r}) e^{-i\omega_n t}$

we have

$$\vec{F}[\vec{\xi}_n] = -\rho_0 \omega_n^2 \vec{\xi}_n$$

So $\frac{1}{\rho_0} \vec{F}[\vec{\xi}]$ has eigenmodes $\propto e^{-i\omega_n t}$. \square .

Property 2. $\vec{F}[\vec{\xi}]$ is hermitian (or self adjoint)

We derived

$$\iiint_V \vec{u} \cdot \vec{F}[\vec{\xi}] dV = -\delta W_2[\vec{u}, \vec{\xi}] - \delta W_2[\vec{\xi}, \vec{u}]$$

R.H.S is symmetric in $\vec{u} \leftrightarrow \vec{\xi}$, so

$$\iiint_V \vec{u} \cdot \vec{F}[\vec{\xi}] dV = \iiint_V \vec{\xi} \cdot \vec{F}[\vec{u}] dV \quad \square.$$

Property 3: eigenvalues ω_n^2 are real.

$$\text{As } \vec{F}[\vec{\xi}_n] = -\rho_0 \omega_n^2 \vec{\xi}_n$$

and \vec{F} ~~is hermitian~~ has no complex coefficients then

$$\vec{F}[\vec{\xi}_n^*] = -\rho_0 (\omega_n^2)^* \vec{\xi}_n^*$$

Compute the difference

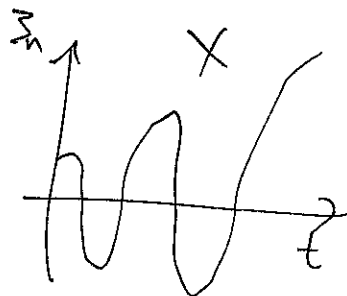
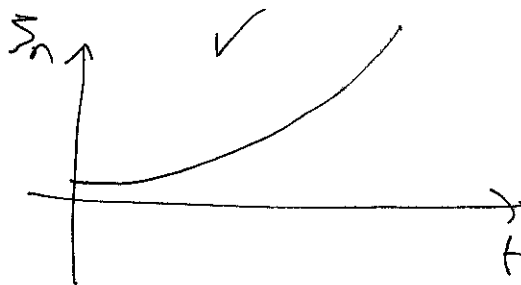
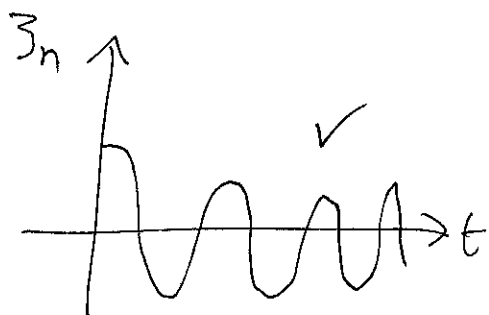
$$\iiint_V \vec{\xi}_n \cdot \vec{F}[\vec{\xi}_n^*] dV - \iiint_V \vec{\xi}_n^* \cdot \vec{F}[\vec{\xi}_n] dV = 0 \quad \leftarrow \text{hermitian!}$$

$$= -[\omega_n^2 - (\omega_n^2)^*] \iiint_V \rho_0 |\vec{\xi}_n|^2 dV$$

$$\text{but } \rho_0 |\vec{\xi}_n|^2 > 0 \Rightarrow \omega_n^2 = (\omega_n^2)^*$$

$$\Rightarrow \boxed{\omega_n^2 \in \mathbb{R}}$$

This means that ω_n is either pure real or pure imaginary \Rightarrow perturbations oscillate or diverge, but don't oscillate with changing amplitude (in MHD)



Property 4: the eigenmodes $\vec{\xi}_n$ are orthogonal

Identical calculation as above, only with $\vec{\xi}_n$ & $\vec{\xi}_m$:

$$\vec{F}[\vec{\xi}_m] = -\rho_0 \omega_m^2 \vec{\xi}_m$$

then

$$-\omega_n^2 \iiint_V \rho_0 \vec{\xi}_m \cdot \vec{\xi}_n dV + \omega_m^2 \iiint_V \rho_0 \vec{\xi}_n \cdot \vec{\xi}_m dV$$

$$= \iiint_V \vec{\xi}_m \cdot \vec{F}[\vec{\xi}_n] dV - \iiint_V \vec{\xi}_n \cdot \vec{F}[\vec{\xi}_m] dV$$

$$= 0$$

$$\Rightarrow -(\omega_n^2 - \omega_m^2) \iiint_V \rho_0 \vec{\xi}_n \cdot \vec{\xi}_m dV = 0$$

$$\Rightarrow \iiint_V \rho_0 \vec{\xi}_n \cdot \vec{\xi}_m dV = \delta_{n,m} \iiint_V \rho_0 |\vec{\xi}_n|^2 dV$$

We can now prove the energy principle.

Proof of energy principle

Write

$$\vec{\xi}(\vec{r}, t) = \sum_n a_n(t) \vec{\xi}_n(\vec{r})$$

Then

$$\delta W_2[\vec{\xi}, \vec{\xi}] = -\frac{1}{2} \iiint_V \vec{\xi} \cdot \vec{F}[\vec{\xi}] dV$$

$$= -\frac{1}{2} \sum_n \sum_m \iiint_V a_n a_m \vec{\xi}_n \cdot \vec{F}[\vec{\xi}_m] dV$$

$$= +\frac{1}{2} \sum_n \sum_m a_n a_m \omega_m^2 \iiint_V \rho_0 \vec{\xi}_n \cdot \vec{\xi}_m dV$$

$$= \frac{1}{2} \sum_n a_n^2 \omega_n^2 \iiint_V \rho_0 |\vec{\xi}_n|^2 dV$$

~~The kinetic term is 0~~
we can define

$$K[\vec{\xi}, \vec{\xi}] \equiv \frac{1}{2} \iiint_V \rho_0 |\vec{\xi}|^2 dV$$

....

$$= \frac{1}{2} \sum_n a_n^2 \iiint_V \rho_0 |\vec{\xi}_n|^2 dV$$

and therefore, our smallest eigenvalue (call it ω_1) satisfies

$$\omega_1^2 = \min_{\vec{\xi}} \left(\frac{\delta W_2[\vec{\xi}, \vec{\xi}]}{K[\vec{\xi}, \vec{\xi}]} \right)$$

as $K > 0$, if $\delta W_2 > 0$ for all possible $\vec{\xi}$, then $\omega_1^2 > 0 \Rightarrow$ all $\omega_n^2 > 0 \Rightarrow$ stable.

However, if $\delta W_2 < 0$ for any $\vec{\xi}$, then at least one of $\omega_n^2 < 0 \Rightarrow$ an instability exists.

Our procedure is therefore clear, write down the

force operator $\vec{F}[\vec{\xi}] = \rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2}$ from the

linearised MHD equations, use this to compute

δW_2 , and this will let us analyse a wide class of instability problems.

Deriving The force operator

Mass: $\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot (\rho \vec{u}) = 0$

Entropy: $\frac{D}{Dt} (\rho e^{-\gamma}) = 0 \Rightarrow \frac{\partial \rho}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \rho + \gamma \rho (\vec{\nabla} \cdot \vec{u}) = 0$

~~Mass~~

Momentum: $\rho \left[\frac{\partial}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \right] \vec{u} = -\vec{\nabla} p + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}) \times \vec{B}$

Induction: $\frac{\partial \vec{B}}{\partial t} = \vec{\nabla} \times (\vec{u} \times \vec{B})$

let's linearise; with $\vec{u} = \vec{\xi}$, $\delta \vec{u} = \frac{\partial \vec{\xi}}{\partial t}$

Mass: $\delta \rho = -\vec{\nabla} \cdot (\rho_0 \vec{\xi})$

Pressure: $\delta p = -(\vec{\xi} \cdot \vec{\nabla}) p_0 = \gamma p_0 (\vec{\nabla} \cdot \vec{\xi})$

Induction: $\delta \vec{B} = \vec{\nabla} \times (\vec{\xi} \times \vec{B}_0)$

Momentum: $\rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2} = -\vec{\nabla} \delta p + \frac{1}{\mu_0} (\vec{\nabla} \times \delta \vec{B}) \times \vec{B}_0 + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}_0) \times \delta \vec{B}$

$$\rho_0 \frac{\partial^2 \vec{z}}{\partial t^2} = \vec{\nabla} [(\vec{z} \cdot \vec{\nabla}) \rho_0] + \gamma \vec{\nabla} [\rho_0 (\vec{\nabla} \cdot \vec{z})] \\ + \frac{1}{\mu_0} (\vec{\nabla} \times [\vec{\nabla} \times (\vec{z} \times \vec{B}_0)]) \times \vec{B}_0 \\ + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}_0) \times [\vec{\nabla} \times (\vec{z} \times \vec{B}_0)]$$

and we therefore have derived our force operator.

Therefore

$$\delta W_2 = - \frac{1}{2} \iiint_V \vec{z} \cdot \vec{F}[\vec{z}] dV \\ = - \frac{1}{2} \iiint_V \vec{z} \cdot (\vec{\nabla} [(\vec{z} \cdot \vec{\nabla}) \rho_0]) dV \\ - \frac{1}{2} \iiint_V \gamma \vec{z} \cdot (\vec{\nabla} [\rho_0 (\vec{\nabla} \cdot \vec{z})]) dV \\ - \frac{1}{2} \iiint_V \frac{1}{\mu_0} \vec{z} \cdot [(\vec{\nabla} \times \delta \vec{B}) \times \vec{B}_0] dV \\ - \frac{1}{2} \iiint_V \vec{z} \cdot (\vec{B}_0 \times \delta \vec{B}) dV$$

there are a myriad number of equivalent formulations of this integral, all based on different numbers of integrations by parts. I will derive the "textbook" version

$$\begin{aligned}
 I_1 &= \iiint_V \vec{E} \cdot (\vec{\nabla} [(\vec{E} \cdot \vec{\nabla}) \rho_0]) dV \\
 &= \underbrace{\iiint_V \vec{\nabla} \cdot (\vec{E} [(\vec{E} \cdot \vec{\nabla}) \rho_0]) dV}_0 \text{ by divergence theorem} - \iiint_V [(\vec{E} \cdot \vec{\nabla}) \rho_0] [\vec{\nabla} \cdot \vec{E}] dV
 \end{aligned}$$

$$\begin{aligned}
 I_2 &= \iiint_V \cancel{\vec{E} \cdot \vec{\nabla}} \gamma \vec{E} \cdot (\vec{\nabla} [\rho_0 (\vec{\nabla} \cdot \vec{E})]) dV \\
 &= \iiint_V \underbrace{\vec{\nabla} \cdot [\gamma \vec{E} [\rho_0 (\vec{\nabla} \cdot \vec{E})]]}_0 - \iiint_V \gamma \rho_0 (\vec{\nabla} \cdot \vec{E})^2 dV
 \end{aligned}$$

$$I_4 = \iiint_V \vec{E} \cdot (\vec{J}_0 \times \delta \vec{B}) dV = - \iiint_V \vec{J}_0 \cdot \underbrace{(\vec{E} \times \delta \vec{B})}_{(\vec{E} \times \delta \vec{B})} dV$$

$$I_3 = \iiint_V \frac{\vec{E}}{\mu_0} [(\vec{\nabla} \times \delta \vec{B}) \times \vec{B}_0] dV$$

$$= \frac{1}{\mu_0} \iiint_V (\vec{\nabla} \times \delta \vec{B}) \cdot (\vec{J} \times \vec{B}_0) dV \quad [\text{vector triple}]$$

$$= -\frac{1}{\mu_0} \iiint_V (\delta \vec{B} \times \vec{\nabla}) \cdot (\vec{J} \times \vec{B}_0) dV \quad [\text{by parts}]$$

$$= -\frac{1}{\mu_0} \iiint_V \delta \vec{B} \cdot (\vec{\nabla} \times (\vec{J} \times \vec{B}_0)) dV$$

$$= - \iiint_V \frac{|\delta \vec{B}|^2}{\mu_0} dV$$

and therefore

$$\delta W_2 = \frac{1}{2} \iiint_V \left[((\vec{J} \cdot \vec{\nabla}) \rho_0) \vec{\nabla} \cdot \vec{J} + \overbrace{\rho_0 (\vec{\nabla} \cdot \vec{J})^2}^{\text{stabilising}} + \vec{J}_0 \times (\vec{J} \times \delta \vec{B}) + \underbrace{\frac{1}{\mu_0} |\delta \vec{B}|^2}_{\text{stabilising}} \right] dV$$

Let's use this as the largest hammer to ever hit a small nail.

Consider adding a constant gravitational field to the momentum equation

$$\rho \left[\frac{\partial}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \right] \vec{u} = - \vec{\nabla} P + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}) \times \vec{B} + \rho \vec{g}$$

then a perturbation gets an extra force term

$$\delta \rho \vec{g} = - \vec{g} (\vec{\nabla} \cdot (\rho_0 \vec{\xi}))$$

and our energy integral has

$$\delta W_2 = \frac{1}{2} \iiint_V \left[((\vec{\xi} \cdot \vec{\nabla}) \rho_0) (\vec{\nabla} \cdot \vec{\xi}) + \gamma \rho_0 (\vec{\nabla} \cdot \vec{\xi})^2 + (\vec{\xi} \cdot \vec{g}) (\vec{\nabla} \cdot (\rho_0 \vec{\xi})) + \vec{J}_0 \times (\vec{\xi} \times \delta \vec{B}) + \frac{|\delta \vec{B}|^2}{\mu_0} \right] dV$$

Let us return to hydrodynamics $\Rightarrow \vec{B}_0 = \vec{J}_0 = \vec{0}$.
 $\delta \vec{B} = 0$.

$$\delta W_2 = \frac{1}{2} \iiint_V \left[(\vec{\nabla} \cdot \vec{\xi}) (\vec{\xi} \cdot \vec{\nabla}) \rho_0 + \gamma \rho_0 (\vec{\nabla} \cdot \vec{\xi})^2 + (\vec{\xi} \cdot \vec{g}) \vec{\nabla} \cdot (\rho_0 \vec{\xi}) \right] dV$$

Interchange instability

$$\vec{g} = -g \hat{z}, \quad \rho = \rho_0(z), \quad P = P_0(z).$$

Vertical momentum equation

$$\vec{0} = -\vec{\nabla} P + \rho \vec{g} \Rightarrow \frac{dP_0}{dz} = -\rho_0 g$$

Perturbation δW_2 satisfies

$$\delta W_2 = \frac{1}{2} \iiint \left[\xi_z \frac{d\rho_0}{dz} (\vec{\nabla} \cdot \vec{\xi}) + \gamma \rho_0 (\vec{\nabla} \cdot \vec{\xi})^2 - g \xi_z \left(\frac{d\rho_0}{dz} \xi_z + \rho_0 \vec{\nabla} \cdot \vec{\xi} \right) \right] dV$$

(use $\rho_0 = -\frac{1}{g} \frac{dP_0}{dz}$)

$$= \frac{1}{2} \iiint \left(2 \xi_z \frac{d\rho_0}{dz} (\vec{\nabla} \cdot \vec{\xi}) + \gamma \rho_0 (\vec{\nabla} \cdot \vec{\xi})^2 - g \xi_z^2 \frac{d\rho_0}{dz} \right) dV$$

$$\therefore \delta W_2 = \iiint_V f(\xi_z, \vec{\nabla} \cdot \vec{\xi}) dV$$

Let's look for the most unstable perturbation, by minimising f . Treat ξ_z & $\vec{\nabla} \cdot \vec{\xi}$ as independent

$$\frac{\partial f}{\partial (\vec{\nabla} \cdot \vec{\xi})} = 2 \xi_z \frac{d\rho_0}{dz} + 2 \gamma \rho_0 (\vec{\nabla} \cdot \vec{\xi}) = 0$$

$$\Rightarrow (\vec{\nabla} \cdot \vec{\xi}) = -\frac{1}{\gamma \rho_0} \frac{d\rho_0}{dz} \xi_z$$

Substituting back

$$\delta W_2 = \frac{1}{2} \iiint_V \left(-\frac{2 \xi_z^2}{\gamma \rho_0} \left(\frac{d\rho_0}{dz} \right)^2 + \frac{1}{\gamma \rho_0} \left(\frac{d\rho_0}{dz} \right)^2 \xi_z^2 - g \xi_z^2 \frac{d\rho_0}{dz} \right) dV$$

simplifying

$$\delta W_2 = \frac{1}{2} \iiint_V \left[-\frac{1}{\rho_0} \left(\frac{d\rho_0}{dz} \right)^2 \bar{z}^2 - g \bar{z}^2 \frac{d\rho_0}{dz} \right] dV$$

$$\text{with } \frac{d\rho_0}{dz} = -\rho_0 g$$

$$\delta W_2 = \frac{1}{2} \iiint_V \left[\frac{\rho_0 g}{\rho_0} \frac{d\rho_0}{dz} \bar{z}^2 - g \bar{z}^2 \frac{d\rho_0}{dz} \right] dV$$

$$= \frac{1}{2} \iiint_V \frac{\rho_0 g \bar{z}^2}{\gamma} \left[\frac{1}{\rho_0} \frac{d\rho_0}{dz} - \frac{\gamma}{\rho_0} \frac{d\rho_0}{dz} \right] dV$$

$$= \frac{1}{2} \iiint_V \frac{\rho_0 g \bar{z}^2}{\gamma} \left[\frac{d}{dz} \ln(\rho_0 \rho_0^{-\gamma}) \right] dV$$

$$\infty \quad \text{unstable iff. } \frac{d}{dz} \ln(\rho_0 \rho_0^{-\gamma}) < 0$$

\Rightarrow entropy decreases upwards.

Highlights the pro's & con's of this method.

- Pro: easy calculation, no underlying solutions of the hydro required
- Con: no insight about why this is unstable.