

The magneto-thermal instability (MTI)

Discovered by Steren Balbus in 2000, 2001.

(a man of many instabilities).

Numerically explored by Parrish & Jim Stone 2007.

Extended by Quataert 2008.

[Clearly to be a named chair in Oxford, Princeton or IAS
you need to study this instability].

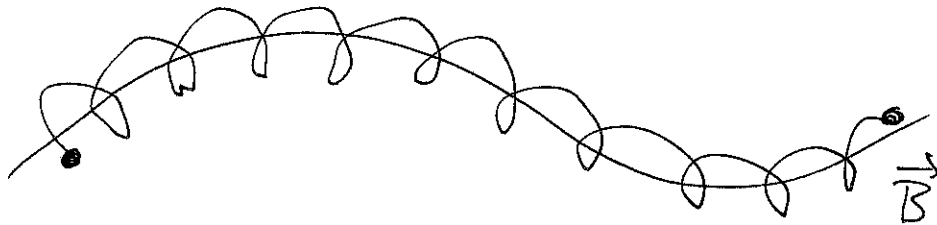
Still being actively studied now Perrone & Latter 2020, 2022
(Cambridge).

Current Astrophysical motivation is in understanding
Galaxy clusters, one of the most famous problems
with which is the "cooling flow problem" \Rightarrow in brief,
Galaxy clusters are hotter than they should be given the
naïve cooling time and their age.

To model the MTI we will need to add one
non-ideal ~~piece~~ piece of Physics, the flow of heat.
Described by vector \vec{Q} , and modifies entropy equation

$$\frac{\gamma}{\gamma-1} \frac{D}{Dt} \ln(P e^{-\delta}) = - \vec{\nabla} \cdot \vec{Q}$$

Galaxy clusters are full of dilute plasmas,
 meaning that the Larmor radius is much smaller
 than the mean free path.



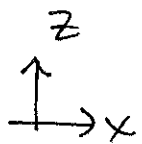
Particle undergoes many revolutions between
 collisions \Rightarrow heat only conducted along field
 lines.

Normal heat conduction $\vec{Q} = -\kappa \vec{\nabla} T$,
 but now modify to

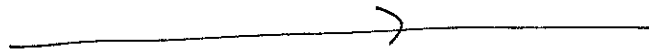
$$\vec{Q} = -\chi \hat{b} (\hat{b} \cdot \vec{\nabla}) T - \psi \vec{\nabla} T$$

I will carry both χ & ψ through to
 "tag" the magnetic field, which will
 help with interpretation.

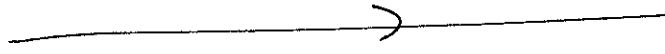
Set up simple equilibrium



$$\vec{g} = -g \hat{z}$$



$$\vec{B}_0 = B_0 \hat{x}$$



B_0

$$P = P_0(z)$$

$$\rho = \rho_0(z)$$

$$\vec{u}_0 = \vec{0}$$

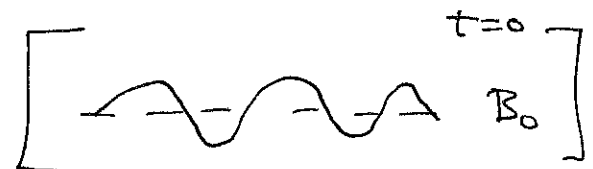
Equilibrium described by momentum equation

$$\vec{0} = -\vec{\nabla} P_0 + \rho_0 \vec{g} + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}_0) \times \vec{B}_0$$

$$\Rightarrow \vec{g} = + \frac{1}{\rho_0} \frac{\partial P_0}{\partial z} \hat{z} \quad \left[\text{so } \frac{\partial P_0}{\partial z} < 0 \right]$$

Perturb with the following

$$\vec{\zeta} = \zeta_z(t) e^{ikx} \hat{z}$$



this means

$$\vec{\nabla}_0 \cdot \vec{\zeta} = \frac{\partial \zeta_z}{\partial z} = 0$$

$$\begin{aligned} \Rightarrow \mathcal{L} &= -\vec{\nabla}_0 \cdot (\rho_0 \vec{\zeta}) = -(\vec{\zeta} \cdot \vec{\nabla}) \rho_0 \\ &= -\zeta_z(t) e^{ikx} \frac{\partial \rho_0}{\partial z} \end{aligned}$$



Magnetic field perturbation

$$\frac{\partial \vec{B}}{\partial t} = \vec{\nabla} \times (\vec{u} \times \vec{B}) \Rightarrow \delta \vec{B} = \vec{\nabla} \times (\vec{\xi} \times \vec{B}_0)$$

$$\begin{aligned} \delta \vec{B} &= B_0 \xi_z(t) \vec{\nabla} \times (e^{ikx} \hat{y}) \\ &= B_0 \xi_z(t) (ik) e^{ikx} \hat{z} \end{aligned}$$

* Heat flux perturbation

$$\begin{aligned} \delta \vec{Q} &= -\chi \hat{b} (\hat{b} \cdot \vec{\nabla}) T - \chi \hat{b} (\delta \hat{b} \cdot \vec{\nabla}) T \\ &\quad - \chi \hat{b} (\hat{b} \cdot \vec{\nabla}) \delta T - \chi \vec{\nabla} \delta T \end{aligned}$$

Note that $(\hat{b} \cdot \vec{\nabla}) T = 0$ as $\hat{b} = \hat{x}$, $T = T(z)$

so

$$\begin{aligned} \delta \vec{Q} &= -\chi \hat{b} (\delta \hat{b} \cdot \vec{\nabla}) T - \chi \hat{b} (\hat{b} \cdot \vec{\nabla}) \delta T \\ &\quad - \chi \vec{\nabla} \delta T \end{aligned}$$

Employ the "Boussinesq" approximation, where we perturb slowly [$t \gg 1/c_s$] so pressure equilibrium maintained, $\Rightarrow \delta P = 0$.

$$\rho \frac{D\vec{u}}{Dt} = -\vec{\nabla} P + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}) \times \vec{B} + \rho \vec{g}$$

Perturb

$$\rho_0 \frac{\partial^2 \vec{\xi}}{\partial t^2} = \frac{1}{\mu_0} (\vec{\nabla} \times \delta \vec{B}) \times \vec{B}_0 + \frac{1}{\mu_0} (\vec{\nabla} \times \vec{B}_0) \times \delta \vec{B} + \delta \rho \vec{g}$$

$$\frac{\partial^2 \vec{\xi}}{\partial t^2} = \frac{1}{\mu_0 \rho} (\vec{\nabla} \times \delta \vec{B}) \times \vec{B}_0 - \frac{\delta \rho}{\rho_0} \vec{g}$$

$$\text{Use } \vec{g} = + \frac{1}{\rho_0} \frac{\partial P_0}{\partial z} \hat{z}$$

$$\text{and } \delta \vec{B} = ik B_0 \vec{\xi}$$

$$\frac{\partial^2 \vec{\xi}}{\partial t^2} = \frac{1}{\rho^2} \delta \rho \frac{\partial P_0}{\partial z} \hat{z} + \frac{ik B_0^2}{\mu_0 \rho_0} (\vec{\nabla} \times \vec{\xi}) \times \hat{x}$$

$$= \frac{1}{\rho_0^2} \delta \rho \frac{\partial P_0}{\partial z} \hat{z} - \frac{k^2 B_0^2}{\mu_0 \rho_0} \vec{\xi}$$

Finally, perturb entropy

$$\frac{P}{\gamma-1} \left[\frac{\partial}{\partial t} + (\vec{u} \cdot \vec{\nabla}) \right] \ln(P e^{-\gamma}) = -\vec{\nabla} \cdot \vec{Q}$$

Perturb [$\delta P = 0$]

$$\frac{P_0}{\gamma-1} \left[-\frac{\gamma}{P_0} \frac{\partial}{\partial t} \delta P + \frac{\partial \mathcal{E}_z}{\partial t} \frac{\partial}{\partial z} \ln(P_0 \rho_0^{-\gamma}) \right] = -\vec{\nabla} \cdot \delta \vec{Q}$$

We now have everything in place to derive
a dispersion relation.

Ideal gas

$$\frac{P}{\rho} = \frac{N k_B T}{m_p} \Rightarrow \frac{\delta T}{T} = - \frac{\delta \rho}{\rho} \quad [\delta P = 0].$$

Manipulations

Induction + Momentum

$$\frac{\partial^2 \xi_z}{\partial x^2} + k^2 v_A^2 \xi_z = \frac{1}{\rho_0^2} \delta \rho \frac{\partial \rho_0}{\partial z}$$

Entropy:

$$\frac{\rho_0}{\gamma-1} \left[-\frac{\gamma}{\rho_0} \frac{\partial}{\partial x} \delta \rho + \frac{\partial \xi_z}{\partial x} \frac{\partial}{\partial z} \ln(\rho_0^{-\gamma}) \right]$$

$$= + \vec{\nabla}_0 \left[\chi \hat{b} (\hat{b} \cdot \vec{\nabla}) \delta T + \chi \hat{b} (\delta \hat{b} \cdot \vec{\nabla}) T + \psi \vec{\nabla} \delta T \right]$$

$$\text{Define } \chi' \equiv \frac{(\gamma-1)\chi}{P}, \quad \psi' \equiv \frac{(\gamma-1)\psi}{P}$$

$$-\frac{\gamma}{\rho_0} \frac{\partial}{\partial x} \delta \rho + \frac{\partial \xi_z}{\partial x} \frac{\partial}{\partial z} \ln(\rho_0^{-\gamma}) = \chi' \left[\vec{\nabla}_0 (\hat{b} [\hat{b} \cdot \vec{\nabla}] \delta T) + \vec{\nabla}_0 (\hat{b} [\delta \hat{b} \cdot \vec{\nabla}] T) + \psi' \nabla^2 \delta T \right]$$

both $\vec{\nabla}_a \hat{b} = 0$ and $\vec{\nabla}_b \hat{a} = 0$

leaving

$$-\frac{\gamma}{\rho_0} \frac{\partial}{\partial t} \delta \rho + \frac{\partial \zeta_z}{\partial t} \frac{\partial}{\partial z} \ln(\rho_0^{-\gamma})$$

$$= \kappa' \frac{\partial}{\partial x} \left[\frac{\partial}{\partial x} \delta T \right]$$

$$+ \kappa' \frac{\partial}{\partial x} \left[i k \zeta_z e^{i k x} \frac{\partial}{\partial z} T \right]$$

$$+ \chi' \frac{\partial^2}{\partial x^2} \delta T$$

we $\delta T = -\frac{T_0}{\rho_0} \delta \rho$

$$-\frac{\gamma}{\rho_0} \frac{\partial}{\partial t} \delta \rho + \frac{\partial \zeta_z}{\partial t} \frac{\partial}{\partial z} \ln(\rho_0^{-\gamma})$$

$$= + \frac{T_0 \kappa'}{\rho_0} k^2 \delta \rho - \kappa' k^2 \zeta_z \frac{\partial T}{\partial z}$$

$$\equiv + k^2 \chi' \frac{T_0}{\rho_0} \delta \rho$$

Now assume $e^{-i\omega t}$

Momentum:

$$-\omega^2 \zeta_z + k^2 v_A^2 \zeta_z = \frac{1}{\rho_0^2} \delta p \frac{\partial \rho_0}{\partial z}$$

Entropy:

$$i\omega \gamma \frac{\delta p}{\rho_0} = -i\omega \zeta_z \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})$$

$$= \frac{T_0}{\rho_0} \chi' k^2 \delta p - \chi' k^2 \zeta_z \frac{\partial T_0}{\partial z} + k^2 \psi' \delta p / \rho_0$$

Use $\zeta_z = \frac{1}{k^2 v_A^2 - \omega^2} \frac{\delta p}{\rho_0^2} \frac{\partial \rho_0}{\partial z}$

$$\Rightarrow i\omega \gamma \frac{\delta p}{\rho_0} = \frac{i\omega \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})}{k^2 v_A^2 - \omega^2} \cdot \frac{\delta p}{\rho_0^2} \frac{\partial \rho_0}{\partial z}$$

$$= T_0 \chi' k^2 \frac{\delta p}{\rho_0} - \frac{\chi' k^2 \frac{\partial T}{\partial z}}{k^2 v_A^2 - \omega^2} \frac{\delta p}{\rho_0^2} \frac{\partial \rho_0}{\partial z} + k^2 \psi' T_0 \delta p / \rho_0$$

Dispersion

$$i\omega\gamma(k^2 v_A^2 - \omega^2) - i\omega \frac{1}{\rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})$$

$$= k^2 T_0 (\chi' + \psi') (k^2 v_A^2 - \omega^2) - \chi' k^2 \frac{\partial T_0}{\partial z} \frac{1}{\rho_0} \frac{\partial \rho_0}{\partial z}$$

$$\Rightarrow i\omega^3 - \frac{k^2 T_0 (\chi' + \psi')}{\gamma} \omega^2$$

$$+ i\omega \left(\frac{1}{\gamma \rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma}) - k^2 v_A^2 \right)$$

$$+ \frac{k^2 T_0}{\gamma} \left[(\chi' + \psi') k^2 v_A^2 - \frac{\chi'}{\rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial \ln T_0}{\partial z} \right] = 0$$

Take weak field limit

Take zero field limit $V_A \rightarrow 0, \chi' \rightarrow 0$.

$$i\omega^3 - \frac{k^2 T_0}{\gamma} \psi' \omega^2 + \frac{i\omega}{\gamma \rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma}) = 0$$

$$- (i\omega)^3 + \frac{k^2 T_0}{\gamma} \psi' (i\omega)^2 + (i\omega) \frac{1}{\gamma \rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma}) = 0$$

$$i\omega = \frac{k^2 T_0 \psi'}{\gamma} \pm \sqrt{\frac{k^4 T_0^2 \psi'^2}{\gamma^2} + \frac{1}{\gamma \rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})}$$

Consider long wave length modes $k \rightarrow 0$

$$\omega = \pm i \sqrt{\frac{1}{\gamma \rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})}$$

$$\text{if } \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma}) < 0$$

then $\omega \in \mathbb{C}$

\Rightarrow unstable.

Take weak field limit

$$v_A \rightarrow 0 \quad \kappa' \neq 0, \quad \psi' \rightarrow 0.$$

$$i\omega^3 - \frac{\kappa^2 T_0}{\gamma} \kappa' \omega^2 + i\omega \frac{1}{\rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma})$$

$$- \frac{\kappa^2 T_0}{\gamma} \frac{\kappa'}{\rho_0} \frac{\partial \rho_0}{\partial z} \frac{\partial \ln T_0}{\partial z} = 0$$

There are the same entropy gradient instabilities available for $k \rightarrow 0$,
but what if $\frac{\partial S}{\partial z} > 0$?

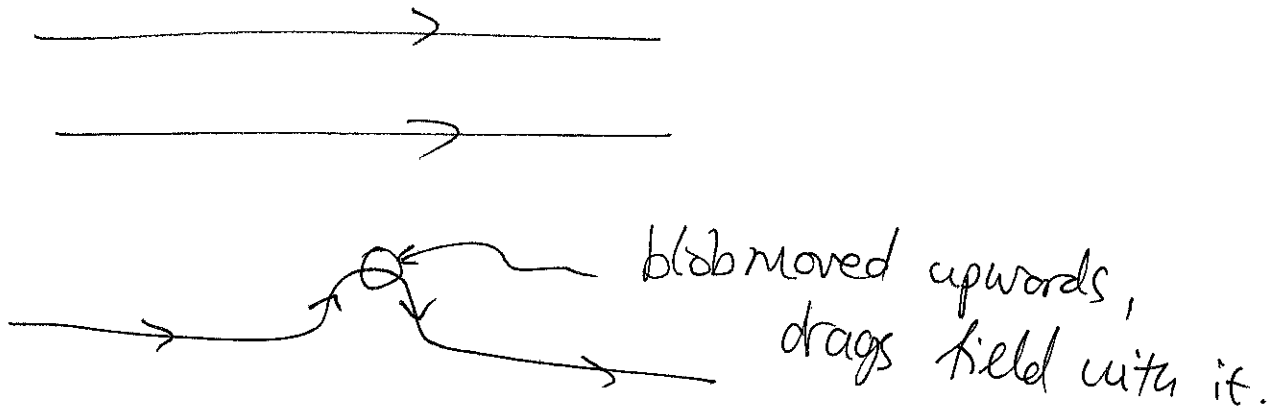
Consider small $|\omega|$, then

$$\omega \simeq -i \kappa' k^2 \frac{\partial T}{\partial z} \left[\frac{\partial}{\partial z} \ln(\rho_0 \rho_0^{-\gamma}) \right]^{-1}$$

if $\frac{\partial T}{\partial z} < 0 \Rightarrow -i\omega$ is real
& positive.

\Rightarrow instability.

What is the physics?



temperature only conducts along field so
perturbation isothermal.

but $\frac{dT}{dz} < 0$ so gas now hotter than
surroundings, so buoyant.

\Rightarrow keeps rising.