

Geophysical Fluid Dynamics

Both atmosphere & ocean governed
the eqns. of GFD.

Derivation based on

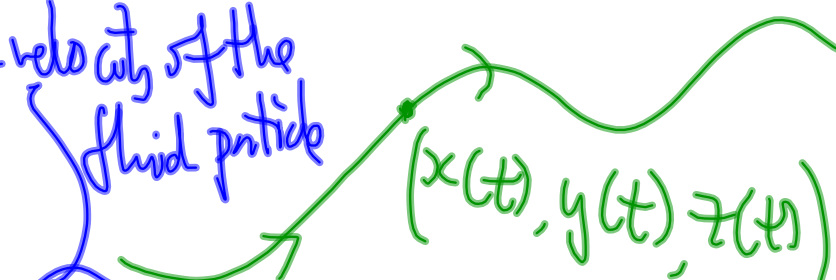
① Newton's law of motion $F = ma$

② Two changes of reference-frame.

Equations govern the motion of a fluid particle: @

$$(x, y, z) \in \mathbb{R}^3$$

for velocity of the fluid particle



(Eqns.) give $u = \frac{dx}{dt}$, $v = \frac{dy}{dt}$, $w = \frac{dz}{dt}$

3 other key physical quantities:

Density: $\rho = \rho(x, y, z, t)$

Pressure: $p = p(x, y, z, t)$

Temperature: $T = T(x, y, z, t)$

Ocean: Salinity: $S = S(x, y, z, t)$

Atmos: Humidity: $q = q(x, y, z, t)$

Frames of Reference:

① Lagrangian
 line @ fluid particle
 & flow with it

vs. Eulerian.
 fixed position
 watching fluid go by

② Rotating
 rotation w/ Earth

vs. Inertial
 absolute frames
 watch the Earth
 rotate.

vs: Eulerian & Rotating,
 Laws of physics: Lagrangian & inertial

Transformations:

Lagrangian \rightarrow Eulerian

Consider a scalar quantity

$$f = f(x, y, z, t)$$

$\left(\frac{d}{dt}\right)_L$ = derivative wrt time in Lagrangian frame

$\left(\frac{d}{dt}\right)_E$ = " " " " Eulerian frames

$$f = f(x, y, z, t)$$

$(x(t), y(t), z(t)) \leftarrow$ fluid particle
(moving in time)

Consider $f = f(x(t), y(t), z(t), t)$

$$\left(\frac{df}{dt} \right)_L = \frac{\partial f}{\partial x} \frac{dx}{dt} + \frac{\partial f}{\partial y} \frac{dy}{dt} + \frac{\partial f}{\partial z} \frac{dz}{dt} + \left(\frac{df}{dt} \right)_E$$

$$= \mathbf{U} \cdot \nabla f + \left(\frac{df}{dt} \right)_E$$

$\mathbf{U} = (u, v, w)$

Inertial \rightarrow Rotating.

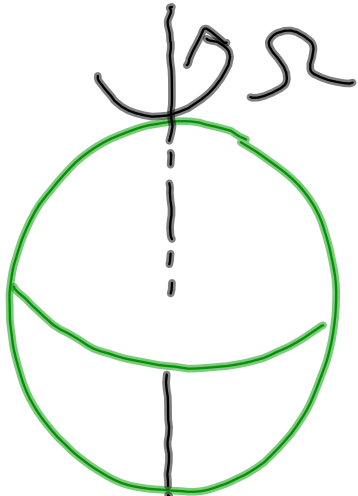
$$\mathbf{F} = \mathbf{F}(x, y, z, t) \in \mathbb{R}^3$$

e.g. $\vec{F} = (u, v, w)$

$$\left(\frac{d\mathbf{F}}{dt}\right)_{\vec{I}} = \left(\frac{d\mathbf{F}}{dt}\right)_{\vec{R}} + \boldsymbol{\Omega} \times \vec{F}$$

$\boldsymbol{\Omega} = |\boldsymbol{\Omega}|(0, 0, 1)$

$7.29 \times 10^{-5} \text{ s}^{-1}$



Newton's Law:

$$F = ma$$

Since thinking is a continuum limit of point (masses) of fluid, look at eqns. in terms of mass / unit volume, i.e. density.

$$\rho \left(\frac{d}{dt} U_I \right)_{L,I} = \text{sum of forces acting on fluid particle}$$

Forces on RHS:

$$- F_{\text{gravity}} - F_{\text{pressure}} + F_{\text{friction}}$$

\uparrow \uparrow \uparrow
 gravitational force of Earth pressure force of ambient fluid none mysterious!

Note: signs are conventional

Friction forces:

- ① internal friction i.e., molecular viscosity is negligible for atmosphere and ocean.
- ② A non-turbulent viscosity term is usually added to account for sub-grid-scale turbulent effects.
- ③ Friction from boundaries:
 - Ⓐ wind @ surface of oceans
 - Ⓑ friction from sea floor.

$$\mathbf{r} = (x, y, z)$$

$$\text{Want } \left(\frac{d\mathbf{r}}{dt} \right)_{\mathbf{I}} = \left(\frac{d\mathbf{r}}{dt} \right)_{\mathbf{R}} + \boldsymbol{\Omega} \times \mathbf{r}$$

$$\begin{aligned} \left(\frac{d}{dt} \mathbf{v}_{\mathbf{I}} \right)_{\mathbf{I}} &= \frac{d}{dt} \left(\mathbf{v}_{\mathbf{R}} + \boldsymbol{\Omega} \times \mathbf{r} \right)_{\mathbf{I}} \\ &\hat{=} \left(\frac{d}{dt} \mathbf{v}_{\mathbf{R}} \right)_{\mathbf{I}} + \frac{d}{dt} (\boldsymbol{\Omega} \times \mathbf{r})_{\mathbf{I}} \end{aligned}$$

$$= \left(\frac{dV_R}{dt} \right)_R + 2\Omega \times V_R + \Omega \times \Omega \times r$$

N's 2nd law:

$$\rho \left(\left(\frac{dV_R}{dt} \right)_R + 2\Omega \times V_R + \Omega \times \Omega \times r \right) = -F_g - F_p + F_f$$

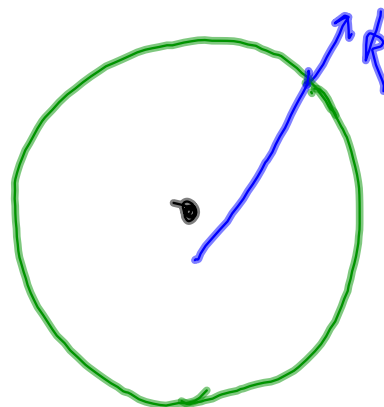
drop the subscript i

$$\rho \left(\frac{dV}{dt} + \underbrace{2\Omega \times V}_{\text{Coriolis}} + \underbrace{\Omega \times \Omega \times r}_{\text{Centrifugal force}} \right) = \Sigma \text{ forces.}$$

In eqs. of GFD, $\Omega \times \Omega \times r$ not present.



surface of
Earth



To a good approx:

Force due to gravity = $-g\rho k + 2\Omega \times \Omega \times r$
 account for oblateness of Earth. Thus the
 centrifugal term is neglected.

$$\rho \left(\frac{dU}{dt} + 2\Omega \times U \right) = -g\rho k - \nabla p + G$$

frictional force

$$\text{But } \frac{dU}{dt} = \left(\frac{dU}{dt} \right)_L = \frac{\partial U}{\partial t} + U \cdot \nabla U$$

$$\Rightarrow \left[\frac{\partial U}{\partial t} + U \cdot \nabla U + 2\Omega \times U = -gk - \frac{\nabla p}{\rho} + \tilde{G} \right]$$

$\tilde{G} = G/\rho$

$U = (u, v, w)$ This is 3 eqns. Actually
cons. of mom.

Problem: 3 eqns. in 5 unknowns.

↓
(u, v, w, p, ρ)

Need 2 more eqns.

Conservation of mass $\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \mathbf{U} = 0$

Conservation of energy \leftarrow brings in

Need eqn. of state \leftarrow temp \rightarrow density, pressure, temperature.