

$$\bar{U} = Y \dot{q} + \bar{z}$$

$$\bar{U}_r = \sum_{s=1}^n Y_{rs} \dot{q}_s + z_r$$

Kinematical differential equations (modeler chooses)

$$\begin{bmatrix} \bar{v}_1 \\ \bar{v}_2 \\ \bar{v}_3 \end{bmatrix} = \begin{bmatrix} A \\ \\ \end{bmatrix} \begin{bmatrix} u_1 \\ u_2 \\ u_3 \end{bmatrix} + \begin{bmatrix} B \\ \\ \end{bmatrix}$$

### Partial Velocities

$$\bar{V} = \sum_{r=1}^n \bar{V}_r u_r + \bar{V}_t$$

for  $n$  G.S.  $\rightarrow$   $r$ -th holonomic partial velocity

if motion constraint

$P = n - m$   
 $\rightarrow$  independent G.S.

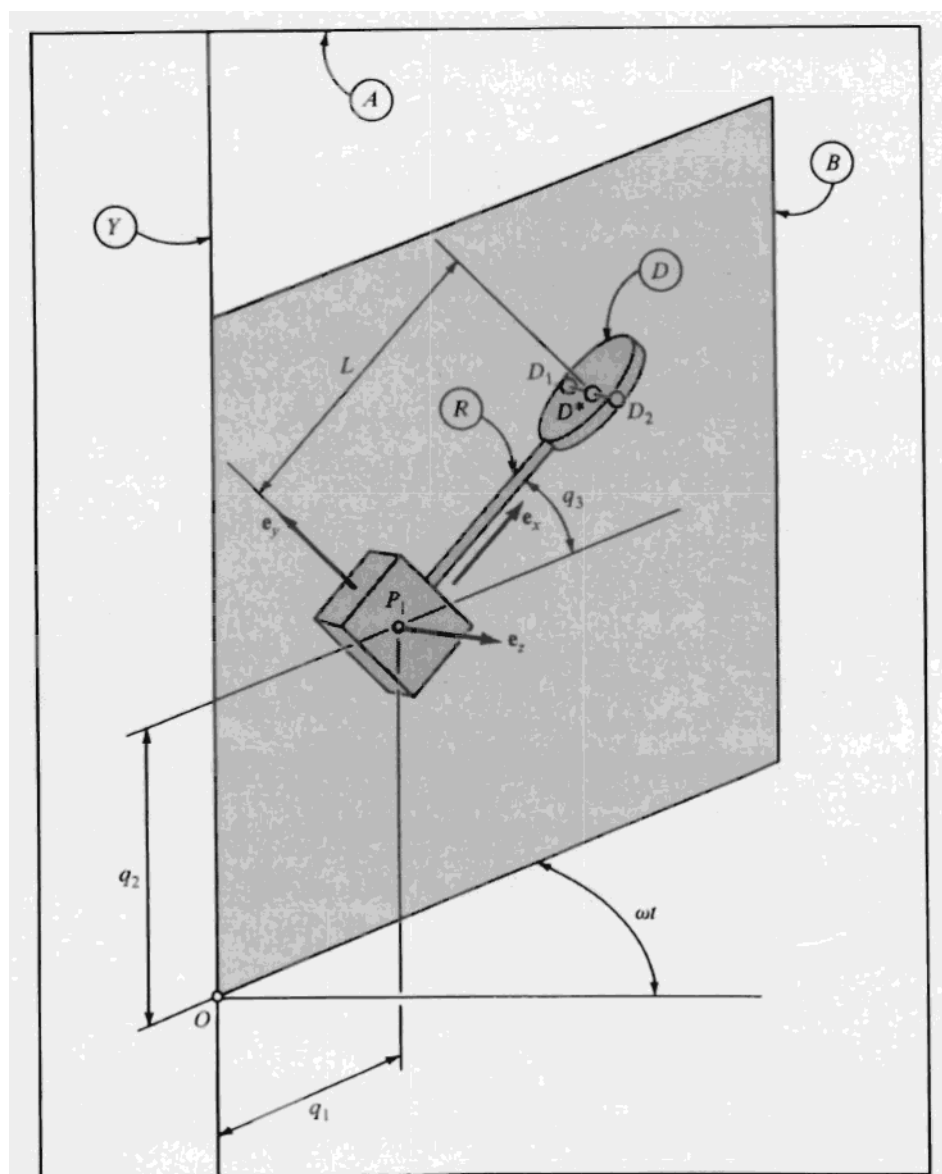
### Nonholonomic

$$\bar{V} = \sum_{r=1}^P \tilde{V}_r u_r + \tilde{V}_t$$

$r$ -th nonholonomic partial velocity

$$\bar{\omega} = \sum_{r=1}^P \tilde{\omega}_r u_r + \tilde{\omega}_t$$

$r$ -th nonholonomic angular vel



**Figure 2.13.1**

3 G.C.'s  $q_1, q_2, q_3$

$$^A \overline{V}^P = u_1 \hat{e}_x + u_2 \hat{e}_y - \omega_1 \hat{e}_z$$

$$u_1 = \dot{e}_1$$

$$u_1 \equiv i_1$$

$$A - \hat{D}^T = A - \hat{D}_1 + A - \omega \times L \hat{e}_x$$

$$\frac{A}{\omega} \underline{F} = \frac{A-B}{\omega} \underline{F} = \omega s_3 \hat{e}_x + \omega c_3 \hat{e}_y + u_3 \hat{e}_z$$

$$u_3 = \dot{q}_3$$

$$A \frac{1}{V} D^* = u_1 \hat{e}_x + (u_2 + L u_3) \hat{e}_y - \omega (q_1 + L c_3) \hat{e}_z$$

NO lateral slip constraint

$${}^B \bar{V}^{\delta*} \cdot \hat{e}_y = 0 \Rightarrow u_2 + L u_3 = 0$$

$$u_2 + L u_3 = 0$$

$$u_3 = \frac{-u_2}{L} \leftarrow \text{ind. GS.}$$

We get to choose  
which are  
ind. and dep.

## Partai Vels

$$A \bar{V} D^+ = e_x$$

$$A - D^T V = \hat{e}_y$$

$$A \frac{1}{V} \frac{dV}{dt} = \frac{1}{\rho} \frac{d\rho}{dt} = -\omega (q_1 + \frac{1}{2} q_2)$$

$$A - \frac{1}{V_+} = -\omega (q_1 + L C_3 \epsilon_2)$$

Dep. GS.

$$A \sim \delta^+ \frac{1}{V_L} = -\omega (q_1 + L C_3 \hat{e}_2)$$

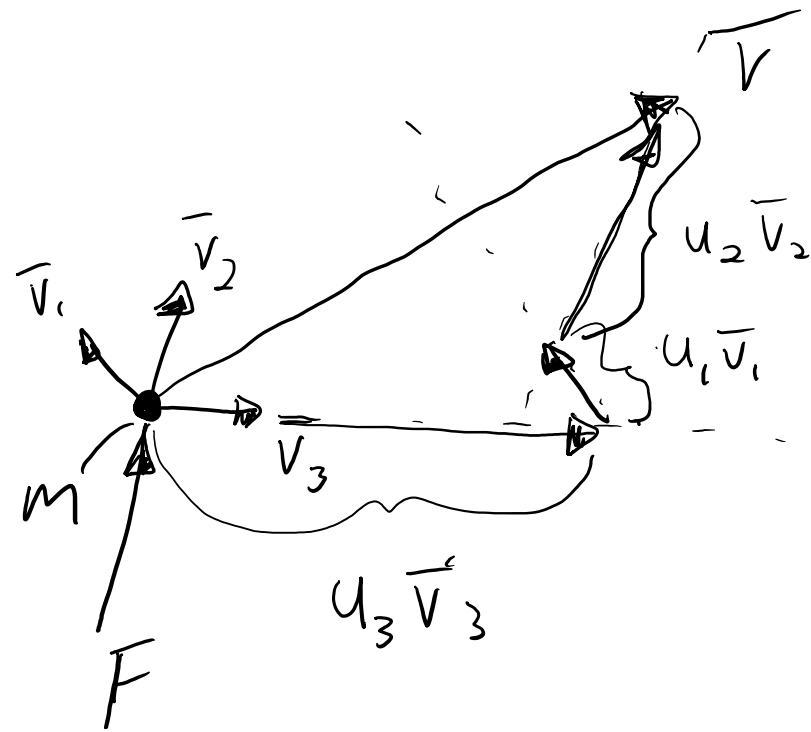
$$A \overline{\omega}^T = \hat{e}_2$$

$$A \frac{\omega}{\omega_t} = \omega_{s_3} \hat{e}_x + \omega_{c_3} \hat{e}_y$$

$$E \otimes_A \mathbb{O} = \mathbb{O}$$

$$E \sim \frac{A}{2} = - \frac{\hat{e} z}{L}$$

$$\mathbb{E} \hat{\omega}_t^A = \omega_{s_3} \hat{e}_x + \omega_{c_3} \hat{e}_y$$



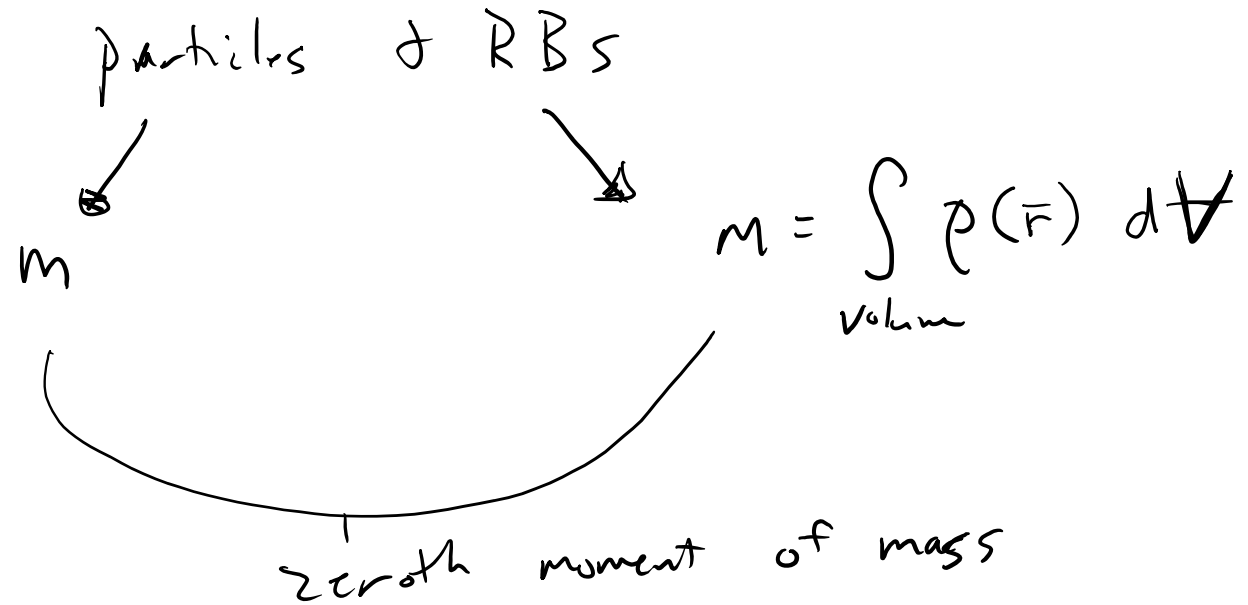
$$\bar{a} = \frac{d\bar{v}}{dt}$$

$$\bar{F}_m = \bar{a}$$

$$\frac{\bar{F}}{m} \cdot \bar{v}_i \Rightarrow \text{the component of force that causes motion in the } u_i \text{ G.S.}$$

$\downarrow$   
 direction associated with  $u_i$

if a partial velocity is  $\perp$  to  $\bar{F}$ , then  $\bar{F}$  doesn't cause any motion in the  $r^{\text{th}}$  direction  
 so if  $V_r \perp \bar{F}$

Mass DistributionMass center of a rigid body (where the mass is on average)

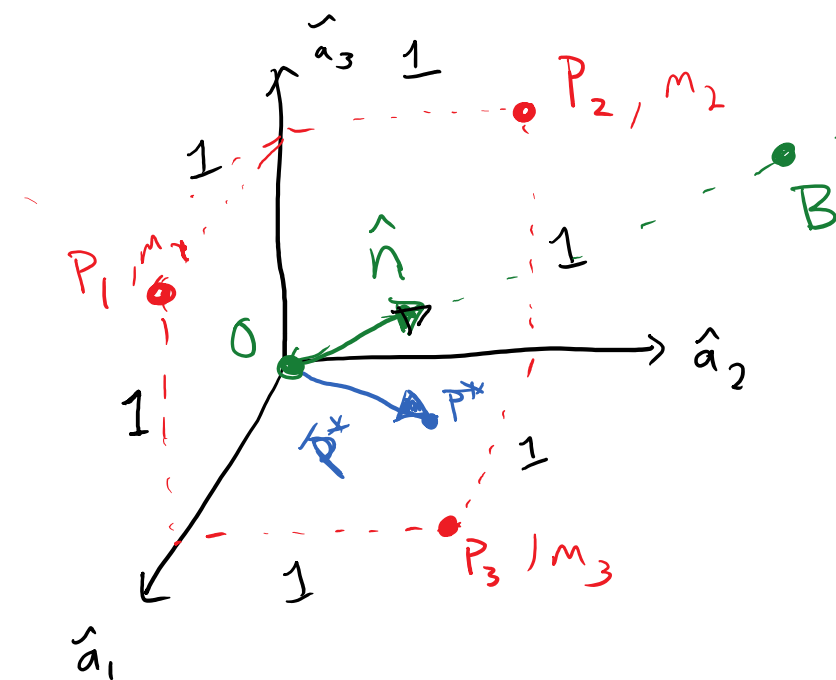
Given a set of particles of mass  $m_1, \dots, m_N$  located at positions  $\bar{r}_1, \dots, \bar{r}_N$ , there is a point  $S^*$  so that  $\sum m_i \bar{r}_i = 0$  where  $\bar{r}_i$  is the vector from  $S^*$  to particle  $i$ .  $S^*$  is the mass center.

The "first moment of mass" is zero relative to the mass center. Or if we choose arbitrary point  $O$ , from which  $\bar{r}_i$  are measured then position vector to  $S^*$  from  $O$ ;  $\bar{p}^*$  is given by

$$\bar{p}^* = \frac{\sum_{i=1}^N m_i \bar{r}_i}{\sum_{i=1}^N m_i} = \frac{\text{first moment}}{\text{zeroth moment}}$$

$\parallel$   
 $m$

Ex 3 particles



Can we vary  $m_3$  some amount to move the mass center to get mass center as close as possible to line  $\overline{OB}$ ?

Now assume  $\hat{n} = \frac{(\hat{a}_1 + \hat{a}_2 + \hat{a}_3)}{\sqrt{3}}$

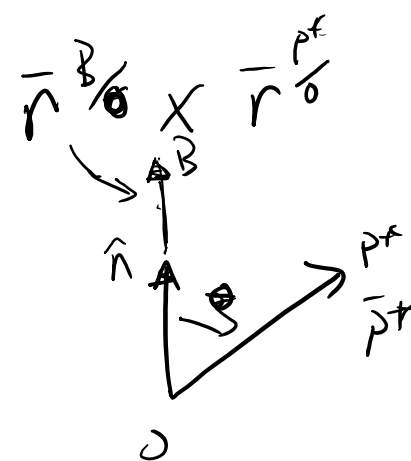
$$D^2 = \frac{2(m^2 - 3\mu m + 3\mu^2)}{3(m + \mu)^2}$$

Let  $\bar{p}^*$  = position vector from O to  $S^*$

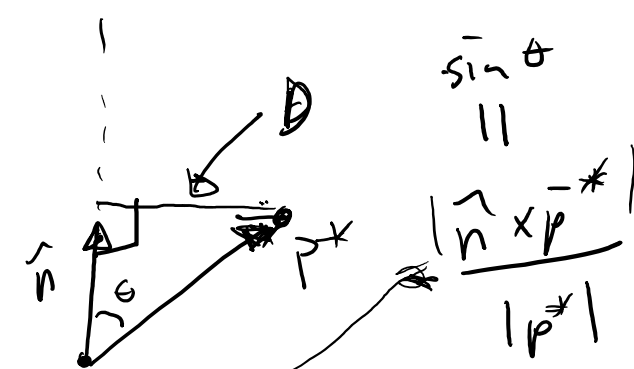
say  $m_1 = m, m_2 = 2m, m_3 = \mu$

$$\bar{p}^* = \frac{\sum_{i=1}^3 \bar{p}_i m_i}{\sum_{i=1}^3 m_i} = \frac{m(\hat{a}_2 + \hat{a}_3) + 2m(\hat{a}_1 + \hat{a}_3) + \mu(\hat{a}_1 + \hat{a}_2)}{m + 2m + \mu}$$

Distance from  $P^*$  to  $\overline{OB}$ ?



$$\hat{n} \times \bar{p}^* = |\bar{p}^*| \sin \theta$$



$$\frac{D}{|\bar{p}^*|} = \sin \theta$$

$$\frac{dD^2}{d\mu} = 0$$

$$D = |\bar{p}^* - (\bar{p}^* \cdot \hat{n}) \hat{n}| = |\hat{n} \times \bar{p}^*|$$

$$\frac{dD}{d\mu} = - \frac{m\sqrt{6}(9m - 5\mu)}{18(m^2 + 2m\mu + \mu^2)\sqrt{3m^2 - 3m\mu + \mu^2}}$$

$$\frac{dD}{d\mu} = 0 \Rightarrow \mu = \frac{5m}{3}$$

$$3(m+n)^2$$

$$du$$

$$\frac{d^2}{du^2} = 0 \Rightarrow \mu = \frac{5m}{3}$$

