

Cosmic Acceleration via Modified Gravity

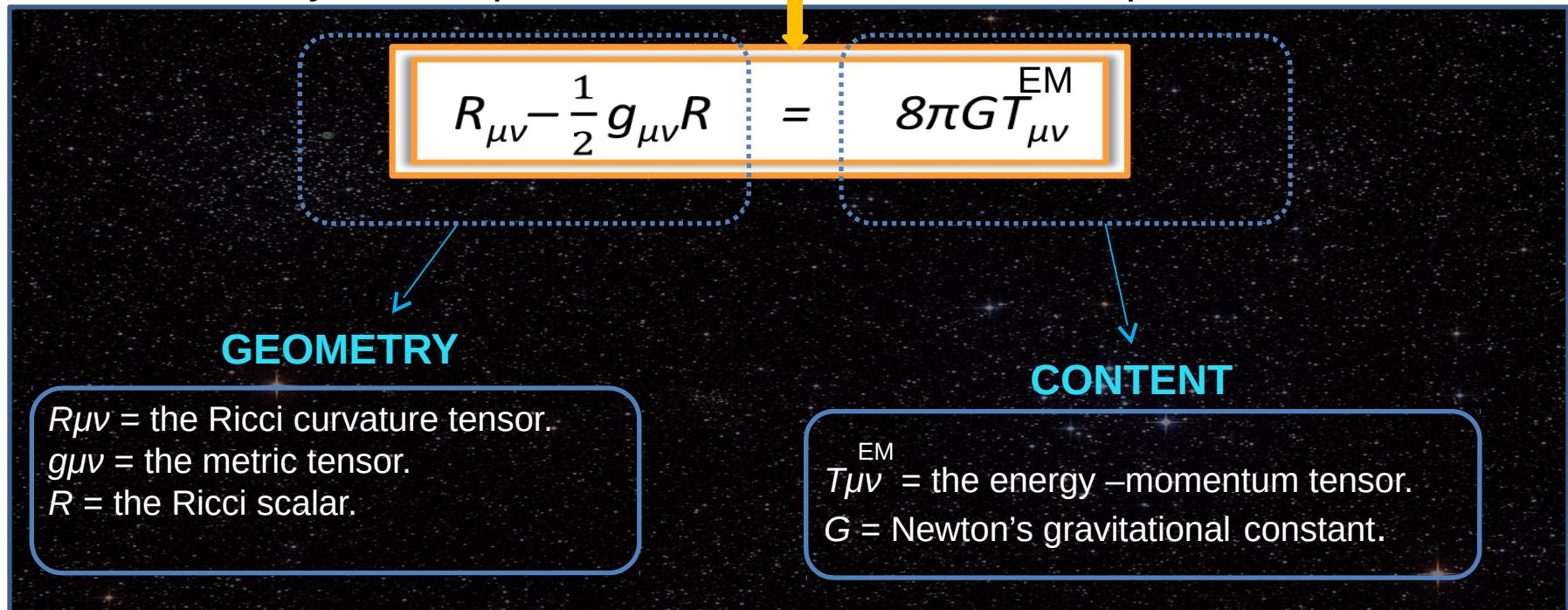
Angela Burden
Supervisor: Andrew Liddle

Presentation Outline

- Background (GR)
- Motivation for modifying gravity.
- $f(R)$ theories of gravity.
- Teleparallel gravity and extensions- $f(T)$.
- Results and future work.
- Discussion.

1. General Relativity

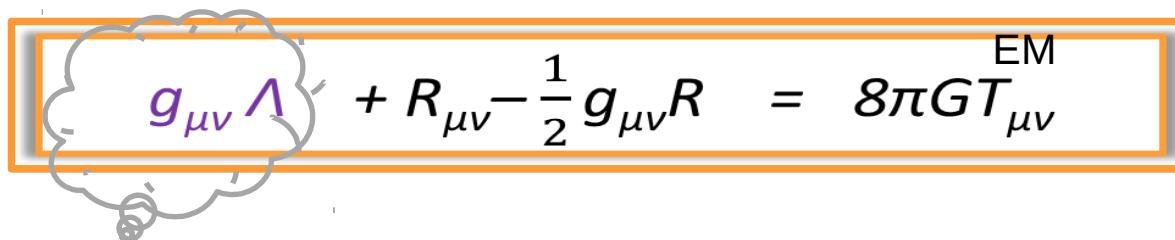
- Einstein's General Relativity formulated in 1915.
- It describes the connection between the curvature of space-time and the matter/radiation content of the Universe.
- The theory is encapsulated in the Einstein field equations

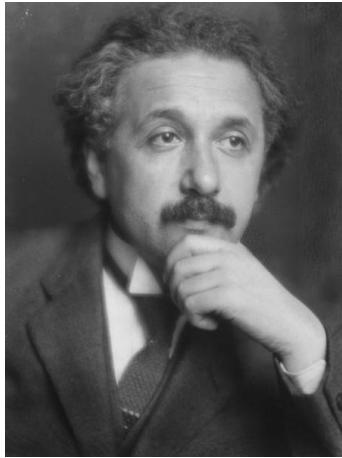


NB - all the equations are in natural units,
Greek indices run from 0-3 over space-time.

2. The Cosmological Constant

- Introduction of the Cosmological Constant Λ to permit static solutions.
- It was included in the geometry side of the equation.


$$g_{\mu\nu} \Lambda + R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi G T_{\mu\nu}^{\text{EM}}$$



- Static solution is unstable to small perturbations.
- BUT, this didn't matter because the Universe was found to be **EXPANDING**.

Ref: LIFE Photo Archive

3. Accelerated expansion of the Universe

- In 1929, Hubble showed that the Universe is **expanding**.
- **AND**, in 1998, Reiss et al., Perlmutter et al. simultaneously discovered the expansion was **accelerating** (Supernovae Type 1a).
- **Dark energy** causing the acceleration? What is it?

Dark energy modelled as a perfect fluid with an equation of state parameter $w \leq -\frac{1}{3}$.

Equation of state \longrightarrow

$$P = w\rho$$

P is the pressure.

w is the equation of state parameter.

ρ is the density of the material.

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = 8\pi G(T_{\mu\nu} + T_{\mu\nu}^{DE})$$

Dark energy on the right hand side.

- The cosmological constant -represents DE with $w=-1$, Λ CDM model.
- **It is a very good model compared with observations at high and low z.**

4. Motivation for modifying GR

- Two problems with Λ CDM :

1. The cosmological constant problem

The vacuum energy predicted by quantum field theory -about 10^{120} times greater than observation..

Observations: $\Lambda \approx 10^{-83} \text{ GeV}^2$

QFT: $\Lambda \approx 10^{38} \text{ GeV}^2$

Durrer et al (2008) arXiv: 0811.4132.

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2. The coincidence problem.

The matter density of the Universe today is of the same order as dark energy density -very unlikely.

Density of matter: $\rho_M \propto \left(\frac{1}{a}\right)^3$

a = scale factor

Density of Λ today: $\rho_\Lambda = \frac{\Lambda}{8\pi G}$

BUT $\rho_\Lambda \approx \rho_M$

POSSIBLE SOLUTION: Avoid using Λ , modify the geometry side of the Einstein equation.

5. The action of GR.

- The action S describes the trajectory of a particle could take between point A and B. $\rightarrow S = \int_A^B ds$
- The infinitesimal line element in GR is: $\rightarrow ds^2 = g_{\mu\nu} dx^\mu dx^\nu$
- In GR, the action is the sum of Einstein-Hilbert Action SEH (geometry) and the matter action SM .
$$S = SEH + SM$$
- The Einstein-Hilbert Action is a function of the Ricci scalar:

$$S_{EH} = \frac{1}{16\pi G} \int \sqrt{-g} R d^4 x$$

R = the Ricci scalar
 g = the determinant of the metric

Modifying the geometry part (SEH) of the action= modification to gravity .

6. $f(R)$ theories of gravity

- The action becomes:

$$S_R = \frac{1}{16\pi G} \int \sqrt{-g} [f(R) + L_M] d^4x$$

- Allows an extra scalar degree of freedom- dubbed the scalaron (Starobinsky, 1980).
- **Couples to matter – acts like a fifth force!**
- To provide late-time acceleration – mass of scalaron must be small.
- **HOWEVER , no fifth force has been observed in the Solar System.**
- The mass of the scalaron evolves with environment i.e. –high mass and short range in solar system–**CHAMELEON MECHANISM.**

More Constraints...

7. $f(R)$ -constraints

Constraint			Region			Reason
Constraint	Region	Reason	Constraint	Region	Reason	
$\frac{df(R)}{dR} > 0$	$R \geq R_0$	To prevent ghosts.	$\frac{df(R)}{dR} > 0$	$R \geq R_0$	To prevent ghosts.	To prevent ghosts.
$\frac{d^2f(R)}{dR^2} > 0$	$R \geq R_0$	To prevent tachyons.	$\frac{d^2f(R)}{dR^2} > 0$	$R \geq R_0$	To prevent tachyons.	
$f(R) = R - 2\Lambda$ (ΛCDM)	High density	To adhere to local gravitational constraints.	$f(R) = R - 2\Lambda$ (ΛCDM)	High density	To adhere to local gravitational constraints.	
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R_0 is the Ricci scalar of the background

- As a result, the model needs to be finely tuned.

Example: $f(R) = R - \mu R_0 \frac{(R/R_0)^{2n}}{(R/R_0)^{2n} + 1}$

Hu and Sawicki (2007),
arXiv:0705.1158
with $n, \mu, R_0 > 0$

8. Teleparallel gravity

- When the Ricci scalar R is replaced with T , the torsion scalar-equivalent to GR.

$$S_T = \frac{1}{16\pi G} \int \sqrt{-g} [T + L_M] d^4x$$

T is the torsion scalar

- Original motivation-** Einstein proposed teleparallel gravity to unite electro-magnetism and gravity.
- Replaces the Christoffel symbols of GR
- with the Weitzenbock connections ...

$$\Gamma^\rho_{\mu\nu} \quad \Gamma^\rho_{\mu\nu} \quad w$$

9. $f(T)$ theories of gravity

- Original motivation lost, BUT renewed interest recently to model late-time acceleration.
- Following the process of $f(R)$ theories, promote to a function of T , $f(T)$, giving the action:

$$S_T = \frac{1}{16\pi G} \int \sqrt{-g} [f(T) + L_M] d^4 x$$

Advantages of $f(T)$ over $f(R)$:

- Not so tightly constrained: $f/T \rightarrow 0$ at early times,
must have $w=-1$ at late times.
- Equations of motion are 2nd order compared to 4th in $f(R)$ theory.

10. Results and future work

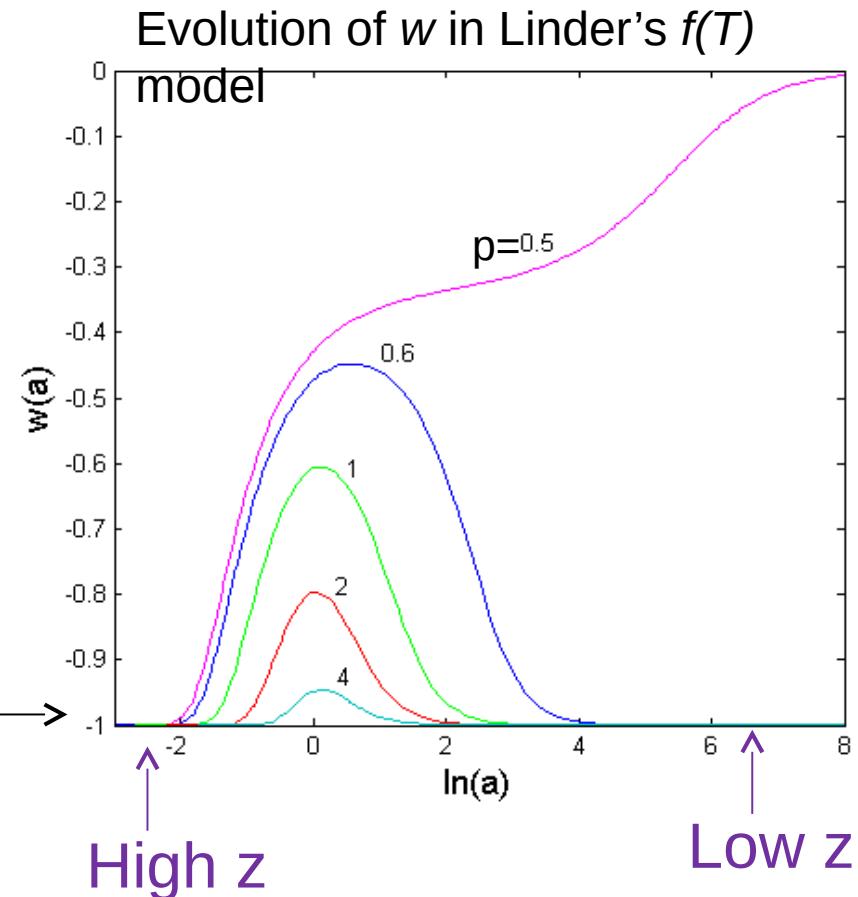
$$f(T) = \frac{1-\Omega_M}{1-(1+p)e^{-p}} T_0 (1 - e^{-p \sqrt{T/T_0}})$$

Linder (2010), arXiv:1005.3039v2

where: T_0 is the torsion scalar today.
 p is an adjustable parameter.
 Ω_M is the matter density

- **Good model** -when $p > 0.5$ as mimics Λ CDM at high and low z .
- **Future work**- to fit observational data from SN1a to this model.

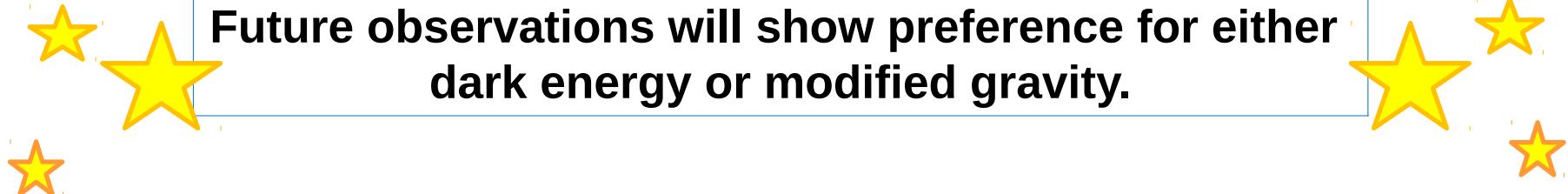
$w=-1$, equivalent to Λ CDM



10. Discussion (and future work)

1. Is it possible to distinguish between dark energy and modified gravity observationally?
 - Not yet!
 - Through the evolution of density perturbations -Dent et al.(2010) arXiv: 1008.1250

Future observations will show preference for either dark energy or modified gravity.



Finally (in the meantime)

2. Are $f(T)$ and $f(R)$ sensible theories?
 - $f(T)$ not Lorentz invariant-Sotiriou et al. (2010), arXiv: 1012.4039,
 - $f(R)$ too tightly constrained, the need for Lorentz invariance-Durrer et al. (2008), arXiv:0811.413.

$$\Gamma^{\rho\mu\nu} = \Gamma^{\rho\nu\mu}$$

Christoffel symbols of GR, symmetric under lower index exchange.

$$T^{\rho\mu\nu} = \Gamma^{\rho\mu\nu} - \Gamma^{\mu\nu\rho}$$

Torsion tensor is the difference between Weitzenbock connections under lower index exchange.

- Kunz et al. (2007) arXiv:0612452, say it is **not** possible to see the difference between dark energy and modified gravity in growth of perturbations.
- Dent et al. (2010) arXiv:1010.2215, say that the equivalent $f(T)$ theory compared to quintessence will leave a signature in the growth of density perturbations.

9. $f(T)$ continued...

- Instead of the metric - the tetrad $e^A = \text{diag} (1, a, a, a)$ where a is the scale factor of the expansion.
 - the Latin indices represent tangent space and range from 0-3.
 - the Greek indices represent space-time and run from 0-3.
- The tetrad are the orthonormal tangent vectors related to the metric by
- The action is varied with respect to the tetrad to retrieve the modified equations of motion.

$$g_{\mu\nu} = \eta^{AB} e_A^\mu e_B^\nu$$

$$T^\rho_{\mu\nu} \equiv e_A^\rho \left(\partial_\mu e_A^\nu - \partial_\nu e_A^\mu \right).$$

$$K^{\mu\nu}{}_\rho \equiv -\frac{1}{2} \left(T^{\mu\nu}{}_\rho - T^{\nu\mu}{}_\rho - T_\rho{}^{\mu\nu} \right)$$

$$S_\rho{}^{\mu\nu} \equiv \frac{1}{2} \left(K^{\mu\nu}{}_\rho + \delta_\rho^\mu T^{\alpha\nu}{}_\alpha - \delta_\rho^\nu T^{\alpha\mu}{}_\alpha \right)$$

$$T = S\rho\mu\nu T\rho\mu\nu$$

Field equations

$$e^{-1} \partial_\mu (e_A^\rho S_A^{\mu\nu}) [1 + f_{,T}(T)] - e_A^\lambda T_{\mu\lambda}^\rho S_{\nu\rho}^\mu [1 + f_{,T}(T)] \\ + e_A^\rho S_\rho^{\mu\nu} \partial_\mu (T) f_{,TT}(T) - \frac{1}{4} e_A^\nu [T + f(T)] = 4\pi G e_A^\rho T_\rho^{\text{EM } \nu}.$$

$$H^2 = \frac{8\pi G}{3}\rho - \frac{f}{6} - 2H^2 f_{,T}$$

Modified Friedman equation

$$(H^2)' = \frac{16\pi G p + 6H^2 + f + 12H^2 f_{,T}}{24H^2 f_{,TT} - 2 - 2f_{,T}}$$

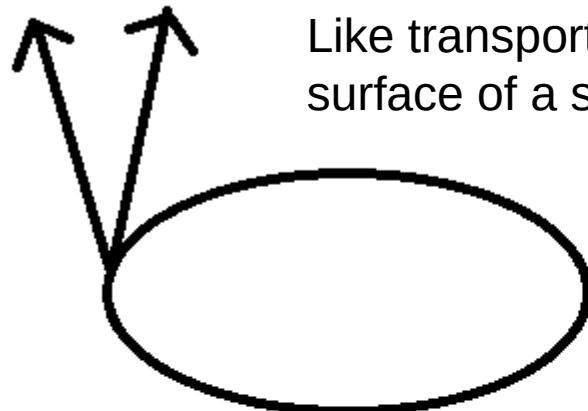
Modified acceleration equation

$$\frac{d \ln(a)}{dT} = \left(-\frac{1}{3}\right) \frac{1 + f_T + 2T f_{,TT}}{T - f + 2T f_{,T}}$$

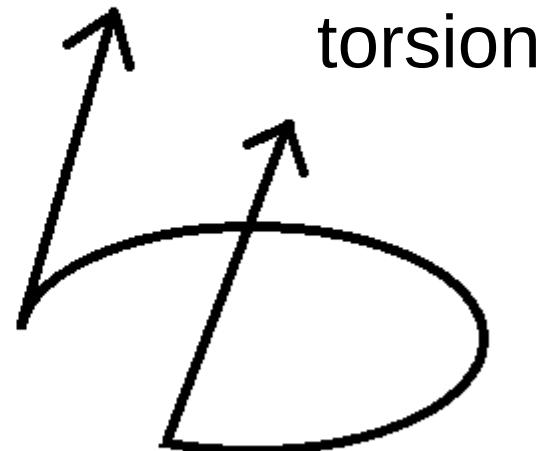
$$w = -\frac{\frac{f}{T} - f_{,T} + 2T f_{,TT}}{(1 + f_{,T} + 2T f_{,TT})(\frac{f}{T} - 2f_{,T})}$$

Curvature and torsion

curvature

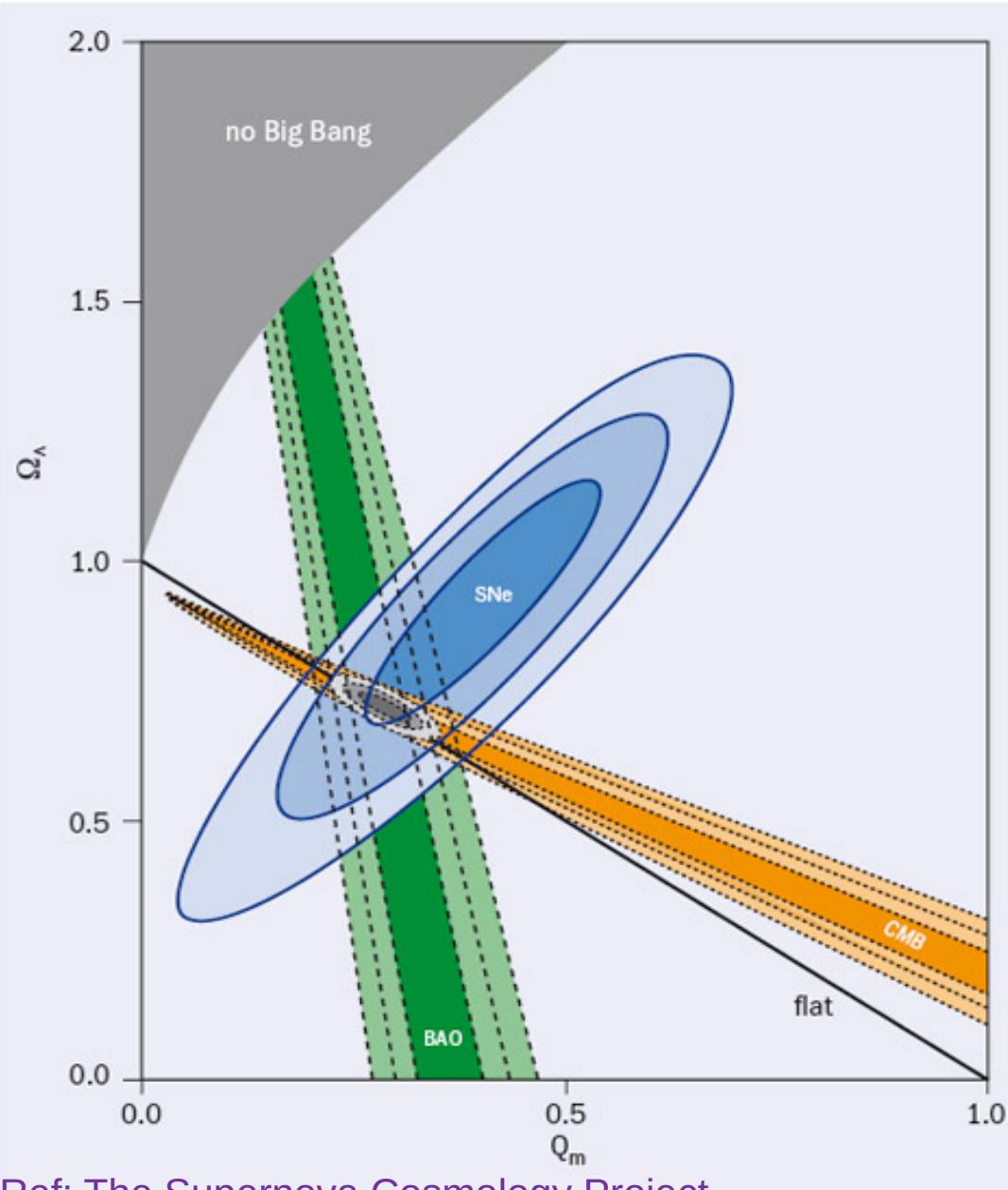


Like transporting a vector over the surface of a sphere



torsion

Torsion proportional to the difference when two vectors are parallel transported along each other.



Ref: The Supernova Cosmology Project

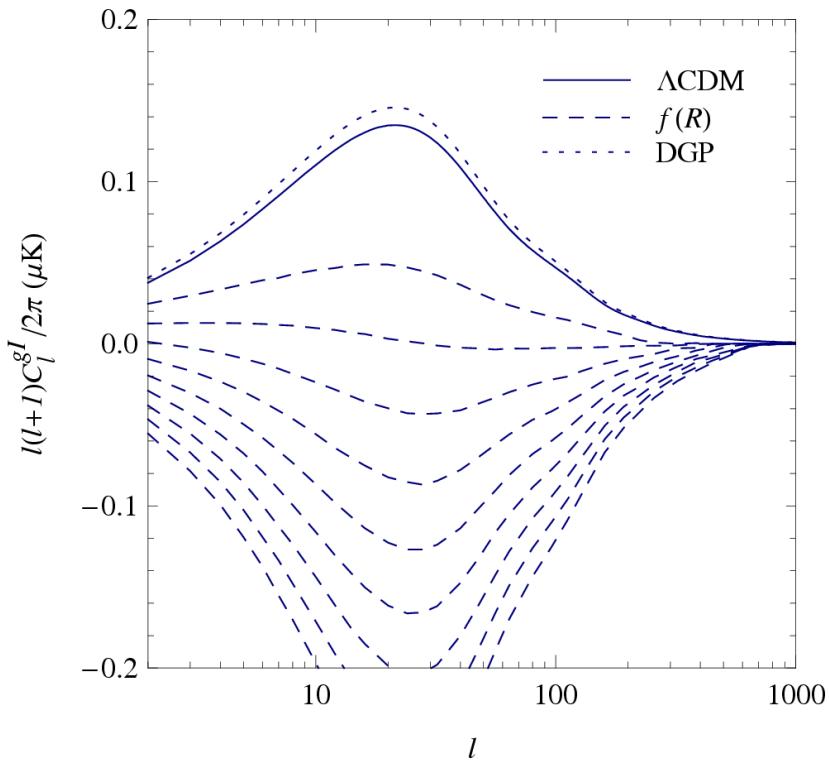
GR tests in the Solar System

- **Precession of Mercury's orbit.**
 - Newton out by approx. 43 arcsecs per century
 - Einstein's GR predicts this very accurately.
- **Bending of light rays close to the Sun**
 - Light from stars that travels very close to the Sun is bent.
- **Radar Echo Delays**
 - Time dilation of radio signals when passing close to a massive body

References

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- [8] Wayne Hu and Sawicki. Models of f(R) Cosmic Acceleration that Evade Solar-System Tests. *Phys. Rev. D*, 76, 064004, 2007.

Distinguishing DE from modified gravity



Galaxy-ISM cross correlations

Figure from : Lombriser (2008) arXiv:

How should teleparallel gravity unite EM and gravity?