Lecture 16

 $\frac{\widehat{\theta}_{i}}{\theta_{i}} = \frac{n_{i}}{n_{i}} \quad \text{proportion of people with black hoir}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}} \quad \widehat{\theta}_{i} = \frac{n_{i}}{n_{i}}$   $= \frac{1}{2} \cdot \widehat{\theta}_{i} = \frac{n_{i}}{n$ 

This class has focused mainly on the 3 goals of inference; estimation, testing and confidence sets. We will continue to study these 3 goals but... we will also study some tangential "meta" concepts that are classic.

Here's one such classic "meta concept." Usually you are a) given a dataset X1,..., Xn, then you b) assume a DEP, then you c) define one (or many) inferential target parameters 0, then d) compute  $\hat{\theta} = w(x_1,...,x_n)$  and then you c) make a CI/run a test at size x

Let's examine (b). How do you just "assume a DBP"? Sometimes you really do know the DBP eig. a coin flipped repeatedly is iid Bern (b), a die relled repeatedly is iid uniform discrete. But what about "daily wind speeds at JFK airpert" or "rat survival times" (like on midterm) or "daily percentage returns of the S&P 500"? The DBP's for the last three are very complicated and they're unknown.

What if we wanted to "guess" the DEP's? This is actually a really

big part of what statisticions do. This is called "model fitting." DEP= model. Models you kinda make up and hopefully they're useful for whatever you're doing. Why don't we proceed as follows: let's guess M condidate models / DEPs m= 1,2,..., M and then (1) Pick the "best" model out of my M guesses and maybe (2) provide a weighting.

Score to each of the M guesses (lew scores indicate bad guesses and high scores indicate good guesses). Goal (1) is famous and called "model selection".

In 342, we do a little of this atheoretically. Here we'll do it more theoretically.

Model selection is more fundamental than you realize. It's actually the entire problem of all of science. For example, let's say you have some astronomical data on movement of different planets, stars, etc. You want to fit a model (guess a DGP) for the force on two celestial bodies with masses my my at a distance r from each other (i.e. "growity").

Consider the following models:

Mod 1: F= 6 mins + 6 mins Mod 2: F= 6 mins + 6 mins Mod 3: F= 6, mins e-62 r Newton's Extension Laplace's Extension

which model is the best? We know all these are wrong because Einstein came and disproved them with general relativity.

Let's talk about model selection techniques. Our data X, ymy Xn comes from an unknown DEP, Here are M candidate models:

Model mi fm (x,, x, i &m,, ..., Onka) = Zm (One, ..., Onka ix, ...x)

Why don't we just select the model that has the highest likeliheed?

mx: = argmex { Im (Om, ..., Omr, ; x, ..., xn)} = argmax { lm (Om,, ..., Omkon ) x,,..., xn)} The problem with this is we don't know the values of O for any of the models! So let's nichardize K, + Kz+. + Km times! We'll estimate each of the parameters using MLE's and play them all in and then... mx = argmnx { & ( & MLE . ..., Kn)} You could do this. But... it will not give you the best model. Why? I (FILE OFFE X1,..., Xn) is an estimator for I (Om,,...,Omkmix,,xn) and it's blased ... With many assumptions, you can prove that asymptetically ... meaning as a gets large... Bios [l(fmie fmie · x,...,xn)] = Km >0

There is positive bias (meaning that the log-likelihood would appear higher on average) and this bias increases you use more parameters in your candidate models. This was figured out by H. Akaike, a Taponese statistician, and he published it in 1974. Once you have the bias, you just use it to correct your estimate:

2 (0 m,,..., 0 m, ) ≈ 1 (êm, ,..., êm, ; x,..., xn) - Km

Recall that log likelihood is always negative for discrete DGP's and almost always negative for continuous DGP's. So let's flip it's sign and multiply by two:

[AIC := -28 (8 min , 8 min , x , ..., x n) + 2 Km penalty

Knike's Information (riterion)

(Akaike's Information (notenion)

The "best" log likelihood is the langest i.e. Closest to zero so once negated, the "best" negative log likelihood is the smallest i.e. closest to zero.