Getting used to telescope and astrophotography DSLR

YOUR VERY NEXTDOOR

 Digital SLR (single-lens reflex)



Figure 1.3. A more elaborate view of what's inside a DSLR. Note computer circuitry ("DIGIC II") at right. (Canon USA.)

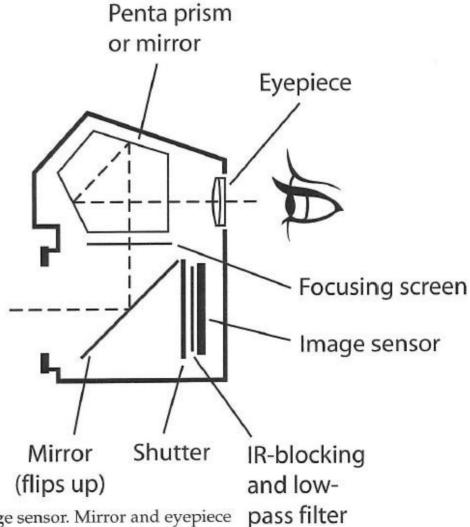
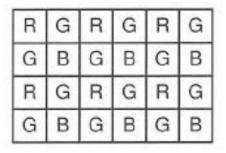


Figure 1.2. A DSLR is a single-lens reflex with a digital image sensor. Mirror and eyepiece allow you to view the image that will fall on the sensor when the mirror flips up and the shutter opens.

How color is recorded



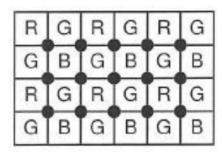


Figure 2.4. Left: Bayer matrix of red, green and blue filters in front of individual sensor pixels. Right: Dots mark "virtual pixels," points where brightness and color can be computed accurately.

- Invented by Dr. Bryce Bayer of Kodak, in 1975
- Green pixels outnumber red and blue because the eye is more sensitive to fine detail in the middle part of the spectrum
- Brightness and color are calculated by combining readings from red, green and blue pixels

- Basic settings
 - Shutter speed and Manual focusing
 - The camera's mode dial to M
 - Setting shutter speed with the thumbwheel
 - Focusing manually
 - ISO (International Standard Organization)
 - Exposure time



Figure 3.2. For manual control of shutter and aperture, set mode dial to M.

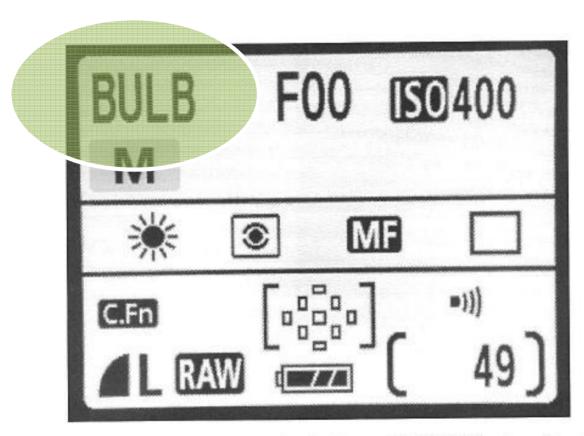


Figure 3.1. Display panel of a Canon XTi (400D) set up for deep-sky photography. F00 means the lens is not electronic and the camera cannot determine the f-stop. C.Fn means custom functions are set.

- Basic settings
 - Shutter speed and Manual focusing
 - ISO (International Standard Organization)
 - Exposure time
 - Bright object, shorter exposure time
 - Fainter object, longer exposure time
 - Trial and error
 - Starting with 3 minutes at ISO 400
 - Underexposed? Then switching 6 minutes at ISO 800

Types of Coupling

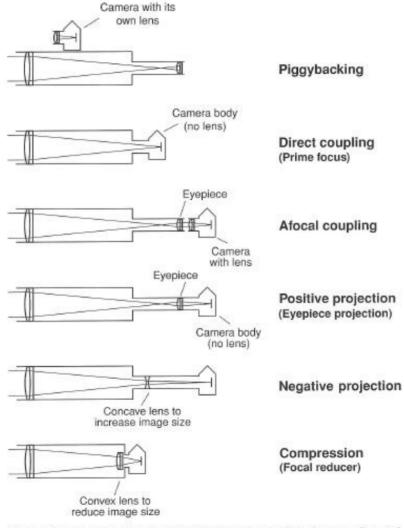


Figure 5.2. Ways of coupling cameras to telescopes. Piggybacking, direct coupling, and compression are main modes for deep-sky work.

• Fitting it all together

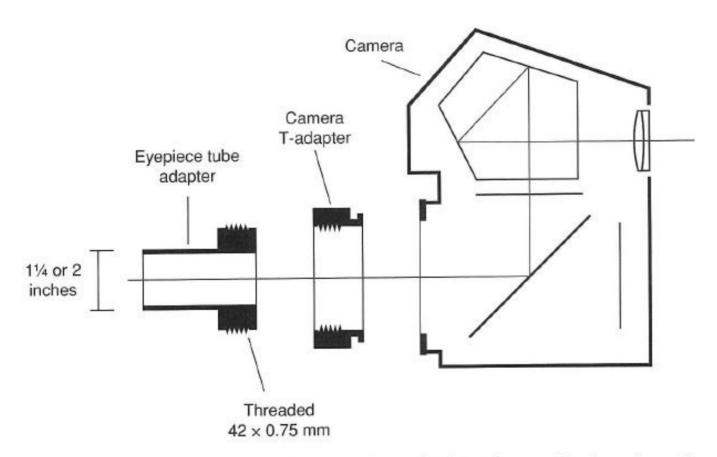
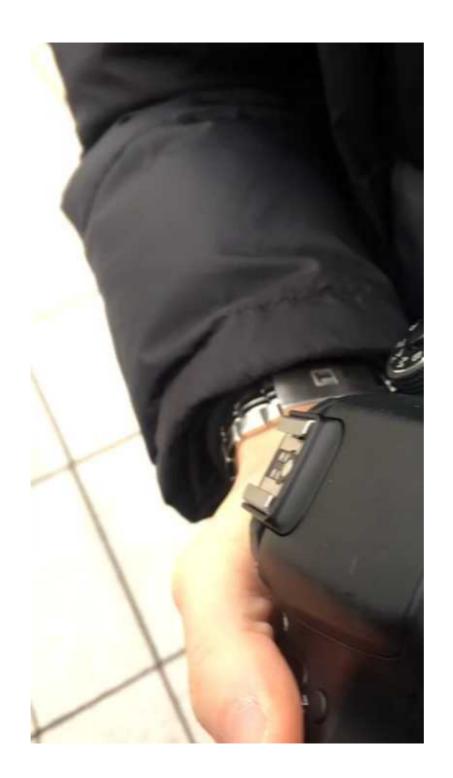


Figure 5.4. Simplest camera-to-telescope adapter fits into telescope in place of eyepiece.



Skills

- Planning observations, making measurements, observations, analyse and interpret data
 - Months, years or even decades on research projects
 - Writing reports, publishing papers and giving presentation at conferences
 - Essential to have an accessible, reliable and complete record of the work
 - Keeping records is crucial
 - Weather conditions prevent observations
 - Equipment breaks down
 - Variety of different approaches to analysing the data
 - Part of the investigation are unsuccessful so that different approaches have to be deviced

- Using an observatory notebook
- Advantages
 - To produce a report of project
 - To compare the records when you get different results from other students for a project
 - To refer back to records of similar techniques or similar data analysis to one that you have done previously → You may avoid having to start again from scratch

- How to keep a record of your observatory work
 - Use a bound notebook or ring-binder
 - Make your record clear and concise
 - Record all relevant information
 - Title; date
 - Record details of observations
 - Time; target, filter...
 - Space out your entries; Correcting errors in your notebook
 - Data analysis
 - Conclusions and critical reflections

• A typical observing log from part of a night's observing session

Telescope Meade 12" LX200 CCD camera SBIG ST-8 Spectrograph n/a Notes Observers: a Smith & B Jones, weather: no cloud, temperature ~ 6°C, numidity ~ 25%								
File No.	Time/UT	Target	RA/h m s	Dec/°′"	Filter or Slit width	Exposure/s	Airmass	Comments
128	2134	bias	i i	-	-	0	-	last bias frame
129	21.36	m39	21 32	+48 26	В	120	1.04	open cluster
130	21.39	m39	21 32	+48 26	٧	90	1.03	open cluster
131	21.43	lan 112	20 12	+00 19	В	40	1.23	Standards, Landolt field 112
132	21.45	lan 112	20 42	+00 19	V	45	1.24	Standards, Landolt field 112

Title:						
	•					
Date:						
Telescope:	Name: (Pegasus	, WilliamA,	WilliamB,	WilliamC,	Takahashi)
	Diameter:		Focal length	:		
Camera:						
Observer:						
Weather:						
			Obj_Date_TelNamex. Jupiter_20140		servers.jpg	
Time	Target	Exposures		FileName		Comments
Printed Photo:						

Date: 2015.03.20,

Site: KSA 전문대.

Photometry/Spectroscopy/image:

Telescope: Pegasus.

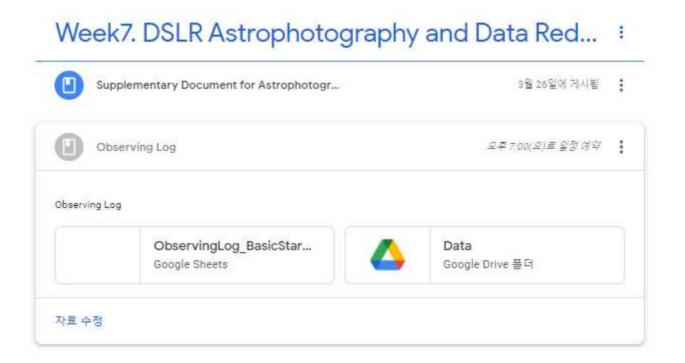
Camera: Gannon @ 700-201310

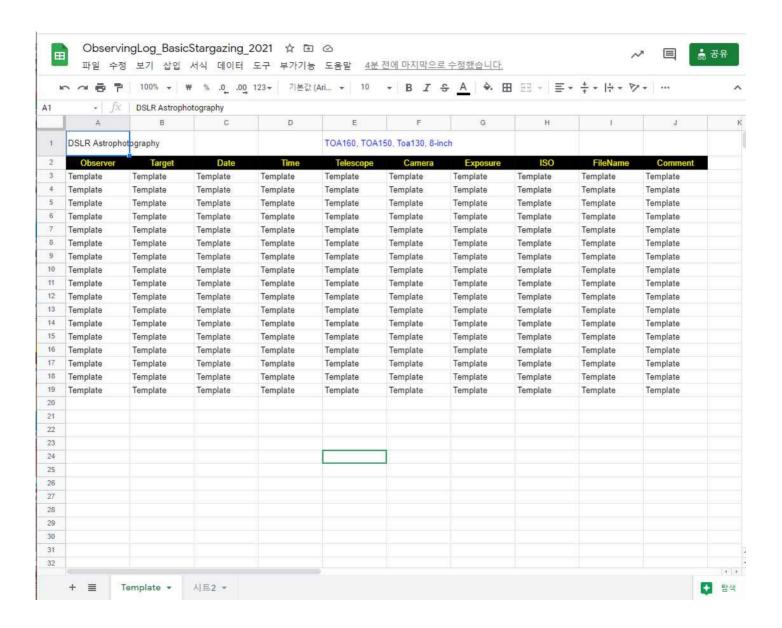
Spectrograph:

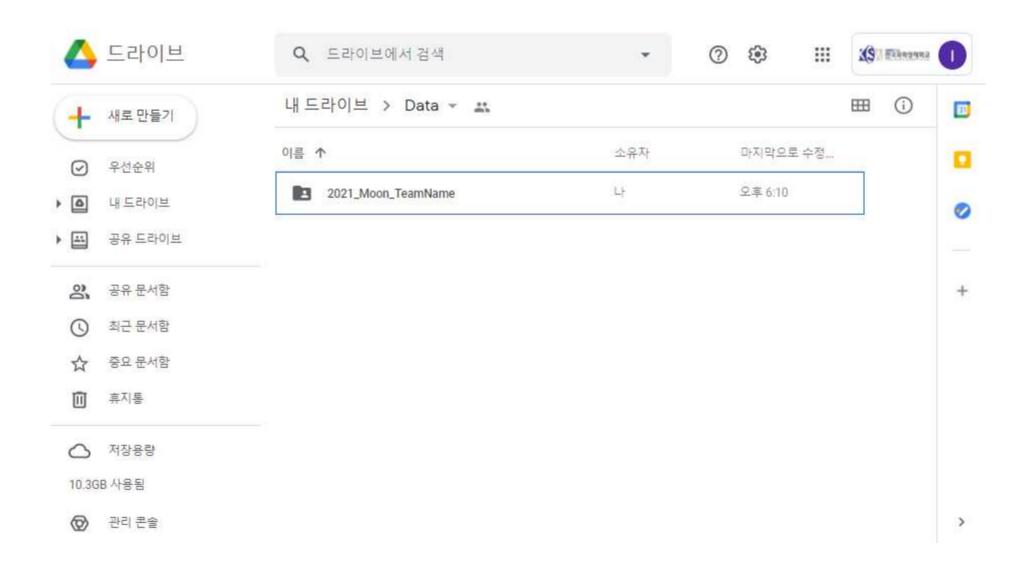
Observers: 박정찬 양정우 표성준

Weather: 맟음

FileNo.	Time	Target	RA/hms	Dec/deg.	Filter or Slit width	Exposure	Comments
Suptrer satell	rte - 2015030	22.2045 _ 박정찬야?	14年はそ			2 (5)	ISO 7 400 / 4개 위성확인
4							allisto.
-							6 Europa
							Garrnede
-							
							1
1						4	
				2	015. 3.22 20:4	15	







Target Today

The Moon (There will be two more in collaboration)

For determining eccentricity of the Moon's orbit

Target Today

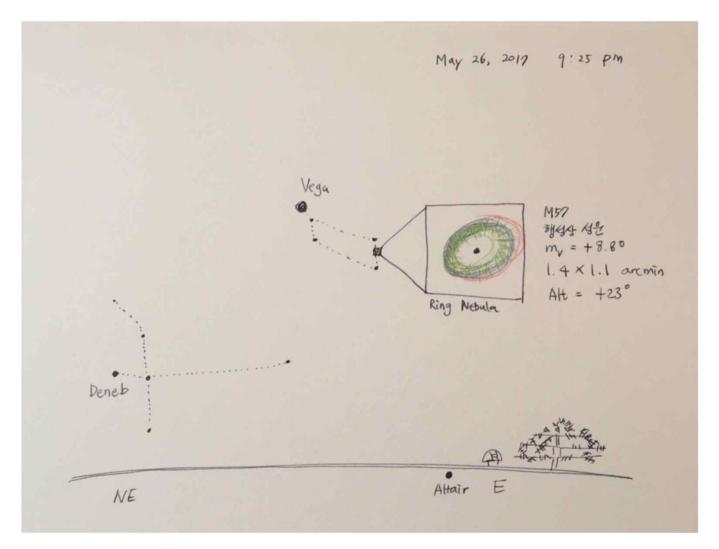
Today is good condition to take astrophotography because of no moon



Early Evening: Orion Nebula

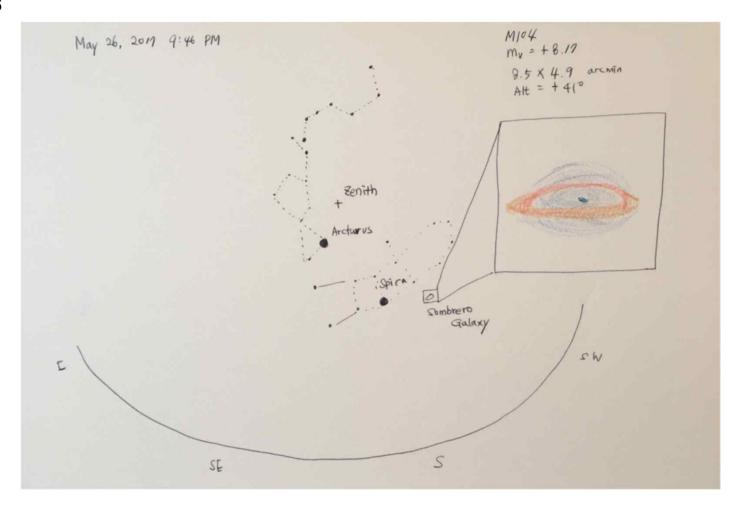
M3 M44 M51 M104 M101 M81 M97

대상 사진 촬영 예,



대상 사진 촬영

예,



Camera Setup

관측 카메라 셋업

You have English Document



1.1 DSLR 카메라 세팅

Cannon 70D

메뉴얼 모드 (M) 를 선택하여 메뉴얼 상태의 메뉴를 선택할 수 있게 한다 (그림 1.1).

raw image 생성 선택 raw image 와 좋은 화질로 저장될 수 있도록 선택 (그림 1.2) 센서 클리닝 센서 클리닝 OFF. FLAT 영상을 얻어 더 효율적으로 제거할 것임 (그림 1.3) 노이즈 감소 노이즈 감소 기능 OFF. DARK 영상을 얻어 더 효율적으로 제거할 것임 (그림 1.4) 날짜 확인 관측날짜와 시간의 기록 중요함 (그림 1.5)

LCD 밝기 조정 최대 밝기로 설정. 어두운 별을 확인하기에 좋음 (그림 1.6) 화이트 밸런스 화이트 밸런스 태양광으로 셋업 (그림 1.7)



Figure 1.2: raw image 생성 선택



Figure 1.3: 센서 클리닝 해제



Figure 1.4: 노이즈 감소 기능 해제



Figure 1.5: 날짜와 시간 설정



Figure 1.6: LCD 밝기 조정



Figure 1.7: 화이트 밸런스 - 태양광

In Practice

- Today's objective:
 Taking photo of the Moon
- Using Basket for Accessories
- Make sure the best focused image and copy the file (then, clear the files on Camera for the next use)
- Write the logbook on Google Classroom
- Upload the file with proper filename on Google Drive (As it is is the best)
- Important!! Safety first
- Class plan