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Contents

1	Abstract
2	Introduction
3	Physical theory

1 Abstract

2 Introduction

3 Physical theory

In this project, a solar system will be studied. By solar system, I mean a system where only gravitational forces effect the bodies, and where there is a large mass fixed in origo¹. Newtons gravitational law states that the gravitational force on a body with mass m from another body with mass m and relative position \vec{r} is given by

$$\vec{F}_{\mathrm{G}} = -\frac{GmM}{\left\|\vec{r}\right\|^2}\vec{r}$$

where G is the gravitational constant, $6.67 \cdot 10^{-11} \,\mathrm{N}\,\mathrm{m}^2/\mathrm{s}^2$. The direction of the force is given by the fact that gravity is an attractive force. If one of the objects is the sun, the mass is denoted by M_{\odot} . With the sun placed in origo, r is simply the norm of the position vector of the planet with mass m.

If there are n planets in the solar system, in addition to the sun, the sum of the forces on planet i with mass m_i is

$$\sum ec{F}_i = \sum_{\substack{j=0\j
eq i}}^n rac{Gm_im_j}{\left\|ec{r}_i - ec{r}_j
ight\|^3} \left(ec{r}_i - ec{r}_j
ight)$$

with $m_0 = M_{\odot}$ and $\vec{r}_0 = \vec{0}$. From Newton's second law, we know that $\sum \vec{F}_i = m_i \vec{a}_i$, so the acceleration of planet i is given by

$$\vec{a}_{i} = \sum_{\substack{j=0\\j\neq i}}^{n} \frac{Gm_{j}}{\|\vec{r}_{i} - \vec{r}_{j}\|^{3}} (\vec{r}_{i} - \vec{r}_{j})$$
(1)

¹This is a reasonable approximation, as the mass of the sun is much larger than the masses of the planets.

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As the sun has been fixed to origo, \vec{a}_0 is set to $\vec{0}$.