

# Computational Astrophysics

E. Larrañaga

Observatorio Astronómico Nacional  
Universidad Nacional de Colombia

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# Elliptic Partial Differential Equations

The simplest elliptic partial differential equation is the linear Poisson equation.

For the Newtonian gravitational potential  $\Phi$ , this is

$$\nabla^2 \Phi = 4\pi G \rho . \quad (1)$$

# Poisson Equation with Spherical Symmetry

Consider a spherically symmetric mass distribution. The Laplacian  $\nabla^2$  reduces to

$$\begin{aligned}\nabla^2[\cdot] &= \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} [\cdot] \right) , \\ &= \frac{2}{r} \frac{\partial}{\partial r} [\cdot] + \frac{\partial^2}{\partial r^2} [\cdot] .\end{aligned}\tag{2}$$

Since the PDE reduces to just one variable,  $r$ , we write it as a second-order ODE

$$\frac{d^2\Phi}{dr^2} + \frac{2}{r} \frac{d\Phi}{dr} = 4\pi G\rho . \quad (3)$$

This equation is reduced to the system

$$\begin{aligned} \frac{d\Phi}{dr} &= z , \\ \frac{dz}{dr} + \frac{2}{r} z &= 4\pi G\rho . \end{aligned} \quad (4)$$

In order to solve inside the matter distribution, we will use inner and outer boundary conditions,

$$\begin{aligned}\left. \frac{d\Phi}{dr} \right|_{r=0} &= 0 , \\ \Phi(R_{\text{surface}}) &= -\frac{GM(R_{\text{surface}})}{R_{\text{surface}}} .\end{aligned}\tag{5}$$

In practice, one sets  $\Phi(r = 0) = 0$  and integrates out. Then, the result is corrected by adding a constant term to obtain the outer boundary condition, which is the analytic value for the gravitational potential outside a spherical mass distribution.

Remember that this added constant doesn't affect the field value because of gauge invariance!

# Matrix Method

In the matrix method, we discretize Poisson Equation with centered differences. Then, for an interior (non-boundary) grid point  $i$  we have

$$\begin{aligned}\left. \frac{\partial \Phi}{\partial r} \right|_i &\approx \frac{1}{2\Delta r} (\Phi_{i+1} - \Phi_{i-1}) , \\ \left. \frac{\partial^2 \Phi}{\partial r^2} \right|_i &\approx \frac{1}{(\Delta r)^2} (\Phi_{i+1} - 2\Phi_i + \Phi_{i-1}) .\end{aligned}\tag{6}$$

Then, Poisson equation is

$$\frac{1}{r_i} \frac{(\Phi_{i+1} - \Phi_{i-1})}{\Delta r} + \frac{(\Phi_{i+1} - 2\Phi_i + \Phi_{i-1})}{(\Delta r)^2} = 4\pi G \rho_i .\tag{7}$$



# Matrix Method

## **Inner boundary condition:**

Because of the  $1/r$  term in the Laplacian, it is clear that the behavior at  $r = 0$  must be regularized.

It is possible to use an expansion or we can stagger the grid so that there is no point exactly at  $r = 0$ , for example by moving the entire grid over by  $0.5\Delta r$ .

**Inner boundary condition:**

At the inner boundary, we have the condition  $\frac{\partial \Phi}{\partial r} = 0$ , so we assume

$$\Phi_{-1} = \Phi_0 . \quad (8)$$

# Matrix Method

Hence, Poisson equation can be written as linear system

$$J\Phi = \mathbf{b} , \quad (9)$$

where  $\Phi = (\Phi_0, \dots, \Phi_{n-1})^T$  (for a grid with  $n$  points labeled 0 to  $n - 1$ ) and  $\mathbf{b} = 4\pi G(\rho_0, \dots, \rho_{n-1})^T$ .

# Matrix Method

The matrix  $J$  has tri-diagonal form and can be explicitly given as:

(a)  $i = j = 0$ :

$$J_{00} = -\frac{1}{(\Delta r)^2} - \frac{1}{r_0 \Delta r} , \quad (10)$$

(b)  $i = j$ :

$$J_{ij} = \frac{-2}{(\Delta r)^2} , \quad (11)$$

(c)  $i + 1 = j$ :

$$J_{ij} = \frac{1}{(\Delta r)^2} + \frac{1}{r_i \Delta r} , \quad (12)$$

(d)  $i - 1 = j$ :

$$J_{ij} = \frac{1}{(\Delta r)^2} - \frac{1}{r_i \Delta r} . \quad (13)$$

# Matrix Method

After finding the solution to the linear system, it is necessary to correct for the analytic outer boundary value

$$\Phi(R_{\text{surface}}) = -\frac{GM}{R_{\text{surface}}}. \quad (14)$$

# Next Class

Poisson Equation