

# Excitation spectra for Andreev billiards of box and disk geometries

J. Cserti, A. Bodor, J. Koltai, and G. Vattay

*Eötvös University, Department of Physics of Complex Systems, H-1117 Budapest, Pázmány Péter sétány 1/A, Hungary*

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We study Andreev billiards of box and disk geometries by matching the wave functions at the interface of the normal and the superconducting region using the exact solutions of the Bogoliubov–de Gennes equation. The mismatch in the Fermi wave numbers and the effective masses of the normal system and the superconductor, as well as the tunnel barrier at the interface are taken into account. A Weyl formula (for the smooth part of the counting function of the energy levels) is derived. The exact quantum mechanical calculations show equally spaced singularities in the density of states. Based on the Bohr-Sommerfeld quantization rule a semi-classical theory is proposed to understand these singularities. For disk geometries two kinds of states can be distinguished: states either contribute through whispering gallery modes or are Andreev states strongly coupled to the superconductor. Controlled by two relevant material parameters, three kinds of energy spectra exist in disk geometry. The first is dominated by Andreev reflections, the second, by normal reflections in an annular disk geometry. In the third case the coherence length is much larger than the radius of the superconducting region, and the spectrum is identical to that of a full disk geometry.

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## I. INTRODUCTION

Recent technological advances in manufacturing almost ballistic semiconductors of mesoscopic size [two-dimensional electron gas (2DEG)] coupled to a superconductor has initiated growing interest in considering the phase-coherent transport and the excitation spectrum in hybrid superconducting nanostructures. In these systems the electron can coherently evolve into a hole (and vice versa) at the interface between the semiconductor and the superconductor. This mechanism had been discovered by Andreev.<sup>1</sup> Classically, the particle momentum is conserved and the hole is reflected back in the same direction as the incoming electron. This scattering process is known as retroreflection or Andreev reflection. The overview of the recent progress in this field can be found in several works.<sup>2–5</sup> The Andreev scattering resulting in a discrete spectrum of single-particle excitations of a layer of normal metal in contact with superconductors on both sides was first discussed by Andreev.<sup>1</sup> Based on the Bogoliubov-de Gennes equation<sup>6</sup> (BdG) the excitation spectrum (Andreev states) of a normal ( $N$ ) metal attached to a superconductor ( $S$ ) was first considered by de Gennes and Saint-James.<sup>7</sup> A ballistic normal metal weakly coupled to a superconductor is commonly called an Andreev billiard. Such systems have been extensively studied over the past ten years.<sup>8–15</sup> The bound states were studied, e.g., for SNS junctions.<sup>16–18</sup> A semi-infinite  $N$  region in contact with a semi-infinite  $S$  region in a strong magnetic field was investigated by Hoppe *et al.*<sup>19</sup>

The excitation spectrum of Andreev billiards depends on whether the  $N$  region is classically chaotic or integrable.<sup>9,13–15</sup> It has been shown that a gap opens in the spectrum for chaotic billiards, while for integrable billiards the spectrum is most likely to be gapless.<sup>9</sup> In these works the spectrum was calculated only close to the Fermi level. However, the spectrum shows interesting features throughout the entire energy range below the superconducting gap. The excitation spectrum is determined in this paper for two specific

Andreev billiards, namely, the NS box and disk systems. In NS box systems a rectangular  $N$  region is in contact with the superconductor, while the NS disk system consists of a circular  $S$  region surrounded concentrically by a circular  $N$  region. Both systems are integrable and, in accordance with earlier findings, the density of states (DOS) is gapless. We shall show, however, that for higher energies the DOS has singularities located at equal distances from each other.

One of the central issues in this paper is the investigation of these singularities in the DOS. We calculate the spectra of these systems exactly by solving the BdG equations, taking into account the nonperfect interface (mismatch in the Fermi wave numbers and the effective masses of the normal metal and the superconductor, and tunnel barrier at the interface). For a narrow NS junction it is justified<sup>2,3,5</sup> that the pairing potential can be approximated by a steplike function (zero in the  $N$  region and some constant value in the  $S$  region). Note that the two systems have the common feature that the BdG equation is separable and one of the separated wave function sets (“transverse modes”) is the same in the  $N$  and  $S$  regions. The exact quantization condition can be very simply expressed in terms of a phase  $\Phi_m(\epsilon)$  which is shown to be related to the classical action of an electron moving inside the  $N$  region between two successive bounces at the NS interface [see Eq. (14)].

Several authors<sup>9,10,13–15</sup> have already derived the Bohr-Sommerfeld approximation for the density of states. The DOS can be written in terms of the probability distribution  $P(s)$  of the classical path length  $s$  between two subsequent bounces of the electron at the NS interface. Starting from our exact quantization condition expressed in terms of the phase  $\Phi_m(\epsilon)$  we rederive the commonly used Bohr-Sommerfeld approximation<sup>9,14,15</sup> of the DOS in the case of NS box and disk systems. Moreover, we give an analytical expression for the probability  $P(s)$  in terms of the phase  $\Phi_m(\epsilon)$ .

It is shown that, in the framework of Bohr-Sommerfeld quantization, the singularity of the DOS is a direct consequence of the singular behavior of  $P(s)$ . Using the expres-

sion for  $P(s)$  we can reproduce the counting function very accurately in the entire range of the spectrum and locate the singularities of the DOS for  $NS$  box and disk systems. The probability  $P(s)$  depends on the geometry of the  $N$  region and the location of the  $NS$  interface. We shall show that from a semiclassical point of view some special trajectories of the electron play crucial roles in the singular character of the DOS. Based on our semiclassical analysis a simple formula is given for the location of these singularities. This formula gives an excellent agreement with our numerically exact calculation of the spectrum. Furthermore, it predicts very accurately the number of the singularities and their location in the system studied by de Gennes *et al.*<sup>7</sup> We propose an explanation, based on our theory about the singularities, for the pronounced peaks found by Ihra *et al.*<sup>15</sup> (for  $NS$  systems) in their numerical calculations of the DOS. We believe that the singularities found by Lodder and Nazarov<sup>13</sup> (for Andreev billiards), and Šipr *et al.*<sup>20</sup> (for  $SNS$  systems) are related to some special classical trajectories of the electron for which the probability  $P(s)$  is singular.

The other issue studied in this paper is the Weyl formula for our  $NS$  systems. In the quantization of a normal billiard, the counting function  $N(E)$  gives the number of levels whose energy is less than or equal to  $E$ . The smooth part of the counting function  $N(E)$  is given by the Weyl formula,<sup>21,22</sup> which has already been calculated for billiards of various shapes. We present a Weyl formula for our  $NS$  box and disk systems. As expected, the counting function obtained from the exact (numerical) calculations “oscillates” around the curve given by our Weyl formula.

Our exact quantum mechanical results reveal a more complex energy level structure for  $NS$  disk systems than for box systems. Semiclassically, two kinds of modes exist in the disk geometry. First, the electron can hit the superconductor undergoing an Andreev reflection (hereafter called Andreev states), and second, the electron trajectory may not reach the superconductor (hereafter called whispering gallery modes). The two modes play crucial roles in the energy levels of the system. Contrary to  $NS$  box systems where the DOS is proportional to the energy close the Fermi level, in the case of  $NS$  disk systems it is constant due to the whispering gallery modes. In studying the spectrum of the disk geometry, we have been motivated by Bruder and Imry’s recent work.<sup>23</sup> Based on whispering gallery modes, they proposed a physical picture to interpret the significant paramagnetic effect observed in recent experiments.<sup>24</sup> Thus, a careful analysis of Andreev and whispering gallery states as presented in this paper and the inclusion of the flux induced phase may shed further light on the issue raised by Bruder and Imry.<sup>23</sup> Our work on this problem is in progress.

We also found that, depending on the parameters of the  $NS$  disk system, the spectra can belong to either of three classes. (a) “Mixed phase,” in which Andreev states coexist with whispering gallery modes and the energy-dependent coherence length  $\xi(\varepsilon)$  in the superconductor is much less than the radius  $R_S$  of the  $S$  region, while the DOS is singular. (b) The opposite case, i.e.,  $\xi(\varepsilon) \gg R_S$  when the spectrum is identical to that of a normal billiard whose radius is that of the outer circle. (c) When  $k_F \xi(0)$  is order of 1 (here  $k_F$  is the

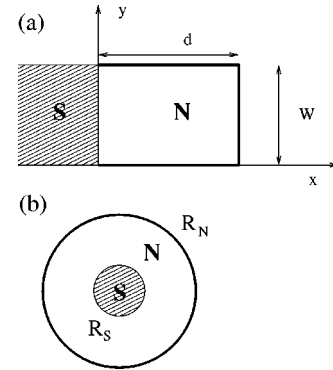


FIG. 1. The normal ( $N$ ) system is in contact with the superconducting ( $S$ ) region. The two geometrical arrangements are (a) box geometry of side lengths  $d$  and  $w$  and (b) disk geometry (two concentric circles of radii  $R_S$  and  $R_N$ ). For the box, the generalized coordinates are  $(x_1, x_2) = (x, y)$ , while for the disk  $(x_1, x_2) = (r, \varphi)$  centered at the origin of the superconducting circle.

Fermi wave number) or the  $NS$  interface is not perfect, and so the energy spectrum corresponds to that of an annular billiard (the circular  $S$  region is cut out). It turns out that there are two relevant parameters to make this classification  $k_F \xi(\varepsilon)$  and  $k_F R_S$ . Using these parameters we sketch a so-called “phase diagram” for the three above classes. It was reported in the works on the paramagnetic effect<sup>24</sup> that the coherence length exceeds the size of the superconducting region. Thus, according to our phase diagram it is possible that the whole  $NS$  disk system behaves as a full normal disk. Then a larger paramagnetic effect can be expected than that predicted by Bruder and Imry who included only the whispering gallery states. However, further work is necessary to clarify this scenario.

The text is organized as follows. In Sec. II a quantization condition (secular equation) is derived from the matching conditions of the wave functions at the interface of the  $NS$  systems. Owing to the symmetry of the BdG equation, the secular equation can be expressed in a very compact form valid both for  $NS$  box and disk systems. In Sec. III the Weyl formula is derived for  $NS$  box and disk systems and compared with the exact results. Our theory for the singularities in the DOS is presented in Sec. IV. We sketch the so-called “phase diagram” for the  $NS$  disk systems in Sec. V, and the conclusions are given in Sec. VI.

## II. SECULAR EQUATION FOR BOX AND DISK

In this section we derive a secular equation determining the energy levels for box and disk geometries shown in Fig. 1. It is possible to treat both problems in a common framework by introducing the generalized coordinates

$$(x_1, x_2) = \begin{cases} (x, y) & \text{for box,} \\ (r, \varphi) & \text{for disk.} \end{cases} \quad (1)$$

The normal region is in contact with a superconducting region at  $x_1 = x_{NS}$  (where  $x_{NS} = 0$  for box and  $x_{NS} = R_S$  for disk).

At the interface  $x_1 = x_{NS}$  the tunnel barrier is modeled by a Dirac delta potential  $V(x_1, x_2) = U_0 \delta(x_1 - x_{NS})$ , as in Ref. 25. The superconducting pairing potential is a constant  $\Delta_0$  in the  $S$  region and zero in the  $N$  region. The self-consistency of the pairing potential is not taken into account just as in Ref. 25. For a planar geometry the difference between the self-consistent gap and the step function is quite small as it was shown by McMillan and Kieselmann.<sup>26</sup> Similar results have been found by Plehn *et al.* in Ref. 27 for superconducting multilayers. For disk geometry the step function approximation is still reliable provided  $R_S$  is larger than the coherence length. However, the Fermi energies (i.e., the energy differences between the band edges of the  $N/S$  region and the chemical potential<sup>28</sup>) and the effective masses in the  $N$  and  $S$  regions are assumed to be different, i.e.,  $E_F^{(N)} \neq E_F^{(S)}$  and  $m_N \neq m_S$ .

The  $NS$  system is described by the BdG equation

$$\begin{pmatrix} H_0 & \Delta \\ \Delta^* & -H_0 \end{pmatrix} \Psi = \varepsilon \Psi, \quad (2)$$

where  $\Psi$  is a two-component wave function,  $H_0 = \mathbf{p}^2/2m - \mu$ , and the Fermi energy is  $\mu = E_F^{(N)}$  in the  $N$  region and  $\mu = E_F^{(S)}$  in the  $S$  region. The energy levels of the Andreev states are the positive eigenvalues  $\varepsilon$  of the Bogoliubov-de Gennes (BdG) equation.<sup>6</sup> In what follows, we consider the energy spectrum below the superconducting gap  $\varepsilon < \Delta_0$ . The two-component wave functions  $\Psi$  in the  $N$  and  $S$  regions can be chosen in the form

$$\Psi_m^{(N)}(x_1, x_2) = \begin{pmatrix} a_m^{(e)} \varphi_m^{(N,e)}(x_1) \\ a_m^{(h)} \varphi_m^{(N,h)}(x_1) \end{pmatrix} \chi_m(x_2), \quad (3)$$

$$\Psi_m^{(S)}(x_1, x_2) = \begin{bmatrix} c_m^{(e)} \begin{pmatrix} \gamma_e \\ 1 \end{pmatrix} \varphi_m^{(S,e)}(x_1) \\ c_m^{(h)} \begin{pmatrix} \gamma_h \\ 1 \end{pmatrix} \varphi_m^{(S,h)}(x_1) \end{bmatrix} \chi_m(x_2), \quad (4)$$

where the “transverse” wave function of the  $m$ th mode is

$$\chi_m(x_2) = \begin{cases} \sqrt{2/W} \sin(m\pi y/W) & \text{for box,} \\ e^{im\varphi} & \text{for disk.} \end{cases} \quad (5)$$

The “transverse” wave functions in the  $N$  and  $S$  regions are the same, and it is possible to separate the variables  $x_1, x_2$  in the BdG equation, i.e.,  $\chi_m(x_2)$  depends only on the coordinate  $x_2$ .

The wave function  $\varphi_m^{(N,e)}(x_1)$  is the electronlike component of the eigenspinor of the BdG equation in the  $N$  region. It satisfies the one-dimensional Schrödinger equation obtained by separating  $\chi_m(x_2)$  in the BdG equation

$$\varphi_m^{(N,e)}(x_1) = \begin{cases} \text{sink}_m^{(e)}(d-x) & \text{for box,} \\ J_m(k_e r) - \frac{J_m(k_e R_N)}{Y_m(k_e R_N)} Y_m(k_e r) & \text{for disk,} \end{cases} \quad (6)$$

where  $k_m^{(e)}(\varepsilon) = k_F^{(N)} \sqrt{1 + \varepsilon/E_F^{(N)} - (m\pi/k_F^{(N)}W)^2}$  for box and  $k_e(\varepsilon) = k_F^{(N)} \sqrt{1 + \varepsilon/E_F^{(N)}}$  for disk. Here  $k_F^{(N)}$  and  $E_F^{(N)}$  are the Fermi wave number and the Fermi energy in the  $N$  region, respectively, while  $J_m(x)$  and  $Y_m(x)$  are the Bessel and Neumann functions of order  $m$ . As we shall see below it is convenient to use real wave functions for  $\varphi_m^{(N,e)}(x_1)$ . The wave functions have been chosen such that they satisfy the Dirichlet boundary conditions at  $x=d$  for box and  $r=R_N$  for disk. Due to the symmetry of the BdG equation the holelike component of the eigenspinor of the BdG equation in the  $N$  region is  $\varphi_m^{(N,h)}(\varepsilon, x_1) = \varphi_m^{(N,e)}(-\varepsilon, x_1)$ , where the explicit dependence of the energy  $\varepsilon$  in the wave functions is emphasized for clarity.

The wave function in the  $S$  region is

$$\varphi_m^{(S,e)}(\varepsilon, x_1) = \begin{cases} \exp(-iq_m^{(e)}x) & \text{for box,} \\ J_m(q_e r) & \text{for disk,} \end{cases} \quad (7)$$

where  $q_m^{(e)}(\varepsilon) = k_F^{(S)} \sqrt{1 + i\eta - (m\pi/k_F^{(S)}W)^2}$  for box and  $q_e(\varepsilon) = k_F^{(S)} \sqrt{1 + i\eta}$  for disk. Here  $k_F^{(S)}$  are the Fermi wave number in the  $S$  region and  $\eta = \sqrt{\Delta_0^2 - \varepsilon^2}/E_F^{(S)}$ , where  $E_F^{(S)}$  is the Fermi energy in the  $S$  region. Note that for box  $\varphi_m^{(S,e)}(x) \rightarrow 0$  as  $x \rightarrow -\infty$  satisfying the boundary condition at  $-\infty$ . On the other hand, only the Bessel function can be chosen for disks, since the Neumann function is singular at the origin. Again, owing to the symmetry of the BdG equation, the hole-like component of the BdG spinor in the  $S$  region is  $\varphi_m^{(S,h)}(\varepsilon, x_1) = [\varphi_m^{(S,e)}(-\varepsilon, x_1)]^*$ . Finally, in Eq. (4)  $\gamma_e = \Delta_0/(\varepsilon - i\sqrt{\Delta_0^2 - \varepsilon^2})$  and  $\gamma_h = \gamma_e^*$ . The coefficients  $a_m^{(e)}, a_m^{(h)}, c_m^{(e)}, c_m^{(h)}$  in Eqs. (3) and (4) are determined from the boundary conditions at the interface of the  $NS$  system (see, e.g., Ref. 28):

$$\Psi_m^{(N)}|_{x_1=x_{NS}} = \Psi_m^{(S)}|_{x_1=x_{NS}},$$

$$\frac{d}{dx_1} \left[ \Psi_m^{(N)} - \frac{m_N}{m_S} \Psi_m^{(S)} \right] \Big|_{x_1=x_{NS}} = \frac{2m_N}{\hbar^2} U_0 \Psi_m^{(S)}|_{x_1=x_{NS}}. \quad (8)$$

The matching conditions yield the following secular equation for the eigenvalues  $\varepsilon$  of the  $NS$  system for fixed mode index  $m$  (for more details see Ref. 29):

$$\text{Im}\{\gamma_e D_m^{(e)}(\varepsilon) D_m^{(h)}(\varepsilon)\} = 0, \quad (9)$$

where

$$D_m^{(e)}(\varepsilon) = \begin{vmatrix} \varphi_m^{(N,e)} & \varphi_m^{(S,e)} \\ [\varphi_m^{(N,e)}]' & Z\varphi_m^{(S,e)} + \frac{m_N}{m_S} [\varphi_m^{(S,e)}]' \end{vmatrix} \quad (10)$$

is a  $2 \times 2$  determinant, and  $D_m^{(h)}(\varepsilon) = [D_m^{(e)}(-\varepsilon)]^*$ . Here  $Z = (2m_N/\hbar^2)U_0$  is the normalized barrier strength, and the prime stands for the derivative with respect to  $x_1$ . All the functions are evaluated at  $x_1 = x_{NS}$ . The energy levels of the  $NS$  systems can be found by solving the secular equation (9) for  $\varepsilon$  at a given quantum number  $m$ . The secular equation (9) is exact in the sense that the usual Andreev approximation is not assumed.<sup>2</sup> The Andreev approximation is valid only for

$\Delta_0/E_F \ll 1$  and quasiparticles whose incident/reflected directions are approximately perpendicular to the interface.<sup>2</sup>

### A. Secular equation for box

To find the energy levels for a box we shall give an explicit form of the secular equation (9). Inserting the wave functions given in Eq. (6) into Eq. (10) we obtain

$$D_m^{(e)}(\varepsilon) = \sin k_m^{(e)} d \left( Z - i \frac{m_N}{m_S} q_m^{(e)} + k_m^{(e)} \cot k_m^{(e)} d \right).$$

Hence, the secular equation given in Eq. (9) becomes

$$\text{Im} \left\{ \gamma_e \left( Z - i \frac{m_N}{m_S} q_m^{(e)} + k_m^{(e)} \cot k_m^{(e)} d \right) \left( Z + i \frac{m_N}{m_S} q_m^{(h)} + k_m^{(h)} \cot k_m^{(h)} d \right) \right\} = 0,$$

where  $k_m^{(h)}(\varepsilon) = k_m^{(e)}(-\varepsilon)$  and  $q_m^{(h)}(\varepsilon) = [q_m^{(e)}(-\varepsilon)]^*$  are the wave numbers of the holelike quasiparticles in the  $N$  and  $S$  regions, respectively. For a given quantum number  $m$  the energy levels can be found by solving the secular equation. Note that the case  $\sin k_m^{(e)} d = 0$  corresponds to the secular equation of the entire box with Dirichlet boundary conditions. Thus, it cannot be zero for the  $NS$  system. Semiclassically, small wave numbers  $k_m^{(e)}$  and  $k_m^{(h)}$  in the  $N$  region correspond to quasiparticles incident at grazing angles on the  $NS$  interface. In this case the Andreev approximation is not valid.<sup>2</sup> However, the secular equation is *still* exact for the energy levels for box geometry.

### B. Secular equation for disk

We now derive the explicit form of the secular equation for disk geometry. Inserting the wave functions given in Eq. (6) into Eq. (10) we obtain

$$D_m^{(e)}(\varepsilon) = \begin{vmatrix} J_m(k_e R_S) - \frac{J_m(k_e R_N)}{Y_m(k_e R_N)} Y_m(k_e R_S) & J_m(q_e R_S) \\ k_e \left[ J'_m(k_e R_S) - \frac{J_m(k_e R_N)}{Y_m(k_e R_N)} Y'_m(k_e R_S) \right] & Z J_m(q_e R_S) + \frac{m_N}{m_S} q_e J'_m(q_e R_S) \end{vmatrix}, \quad (11)$$

where the primes denote the derivatives of the Bessel functions with respect to their arguments. For a given angular momentum quantum number  $m$  the energy levels are the solutions of the secular equation given in Eq. (9).

## III. WEYL FORMULA FOR $NS$ SYSTEMS

For normal systems the counting function  $N(E)$  is the number of states whose energy is less than or equal to  $E$ . The derivative of the counting function with respect to the energy is the density of states. The smooth part of the counting function  $\tilde{N}(E)$  for a cavity was first derived by Weyl<sup>21</sup> (for more details see Ref. 22). The leading term of  $\tilde{N}(E)$  is given by the integral of  $\Theta[E - H_{\text{cl}}(\mathbf{p}, \mathbf{r})]$  over the phase space divided by  $h^2$ , where  $\Theta$  is the Heaviside function and  $H_{\text{cl}}$  is the Hamiltonian of the corresponding classical system. In two dimensions, and excluding the factor 2 for the spin, the smooth part of counting function  $\tilde{N}(E)$  is given by

$$\tilde{N}(E) = \frac{1}{h^2} \int_{H_{\text{cl}}(\mathbf{p}, \mathbf{r}) \leq E} d^2 p d^2 r \Theta[E - H_{\text{cl}}(\mathbf{p}, \mathbf{r})]. \quad (12)$$

For Andreev states of energy  $\varepsilon$  less than  $\Delta_0$  it is not possible to define the classical Hamiltonian  $H_{\text{cl}}$ , therefore the smooth part of the counting function  $\tilde{N}(\varepsilon)$  cannot be derived from Eq. (12). As an alternative method for calculating the DOS one may start from the secular equation.<sup>30,31</sup> In this section we derive a Weyl formula for the  $NS$  systems shown in Fig. 1 using the secular equation (9).

For a given quantum number  $m$  let us introduce the eigenphase  $\Phi_m(\varepsilon)$  for the  $NS$  system

$$[D_m^{(e)}(\varepsilon)]^* / D_m^{(e)}(\varepsilon) = e^{i\Phi_m(\varepsilon)}. \quad (13)$$

Here we have assumed that  $D_m^{(e)}(\varepsilon) \neq 0$ , and in the following it will also be supposed that  $D_m^{(h)}(\varepsilon) \neq 0$ . In Sec. V the case  $D_m^{(e)}(\varepsilon) D_m^{(h)}(\varepsilon) = 0$  will be discussed. It is also assumed that the eigenphases  $\Phi_m(\varepsilon)$  are all monotonic and increasing with  $\varepsilon$ . The secular equation (9) can be further simplified:

$$\Phi_m(\varepsilon) - \Phi_m(-\varepsilon) - 2 \arccos \frac{\varepsilon}{\Delta_0} = 2n\pi, \quad (14)$$

where  $n = 0, \pm 1, \pm 2, \dots$ . Equation (14) is a very simple quantization condition for the  $NS$  system. The solutions of this equation give the energy spectrum  $\varepsilon_{mn}$  below the gap. The quantization condition given above is a convenient starting point to calculate the DOS and the smooth part of the counting function. However, for numerical purposes Eq. (9) may be more suitable. Using Eq. (14) the density of states  $\rho(\varepsilon) = \sum_{m=1}^{M(\varepsilon)} \sum_{n=-\infty}^{\infty} \delta(\varepsilon - \varepsilon_{mn})$  for the  $NS$  system becomes

$$\rho(\varepsilon) = \sum_{m=1}^{M(\varepsilon)} \left| \frac{dF_m(\varepsilon)}{d\varepsilon} \right| \sum_{n=-\infty}^{\infty} \delta[F_m(\varepsilon) - n], \quad (15)$$

where  $2\pi F_m(\varepsilon) = \Phi_m(\varepsilon) - \Phi_m(-\varepsilon) - 2 \arccos(\varepsilon/\Delta_0)$ , and  $M(\varepsilon)$  is the number of “transverse modes” depending on  $\varepsilon$ . Since the  $\Phi_m(\varepsilon)$  are all monotonic and increasing with  $\varepsilon$ , the derivatives  $dF_m(\varepsilon)/d\varepsilon$  are positive, and so the modulus



sign is superfluous in Eq. (15). Applying the Poisson summation formula<sup>32,22</sup> to the second sum one finds

$$\rho(\varepsilon) = \sum_{m=1}^{M(\varepsilon)} \frac{dF_m(\varepsilon)}{d\varepsilon} \left[ 1 + 2 \sum_{k=1}^{\infty} \cos[2\pi k F_m(\varepsilon)] \right]. \quad (16)$$

The DOS can be separated into two parts: the smooth part, i.e., the first term in the square bracket, and the oscillating part, which comes from the terms  $k \geq 1$  in Eq. (16). Here we consider only the smooth part. Then the Weyl formula, i.e., the smooth part of the counting function for the  $NS$  system can be obtained from  $\tilde{N}(\varepsilon) = \int_0^\varepsilon \rho(\varepsilon') d\varepsilon'$ , and one finds

$$\begin{aligned} \tilde{N}(\varepsilon) = & \frac{1}{2\pi} \sum_{m=1}^{M(\varepsilon)} [\Phi_m(\varepsilon) - \Phi_m(-\varepsilon)] \\ & + M(\varepsilon) \left( \frac{1}{2} - \frac{1}{\pi} \arccos \frac{\varepsilon}{\Delta_0} \right). \end{aligned} \quad (17)$$

This is our Weyl formula for the  $NS$  systems. A similar treatment has been mentioned by Schomerus and Beenakker<sup>14</sup> in deriving the Bohr-Sommerfeld approximation of the DOS for Andreev's billiards. The number of transverse modes  $M(\varepsilon)$  is a discontinuous function of  $\varepsilon$ , therefore in Eq. (17) the sum has to be replaced by an integral to get the smoothed version of the counting function.<sup>31</sup> One can see that the Weyl formula is expressed in terms of the eigenphases  $\Phi_m(\varepsilon)$  defined in Eq. (13). Note that our Weyl formula is only valid for those systems in which a common set of the “transverse wave” functions can be separated from the wave functions of the  $N$  and  $S$  regions. Lifting this condition is a possible extension of this problem.

#### A. Weyl formula for box (no mismatch)

We now consider the Weyl formula for a perfect interface, i.e., for  $r_k = 1$ ,  $r_v = 1$ ,  $Z = 0$ . For simplicity, in the case of a perfect interface we shall use the notation  $k_F = k_F^{(N)} = k_F^{(S)}$  and  $E_F = E_F^{(N)} = E_F^{(S)}$ . From Eqs. (10) and (13), and in Andreev approximation ( $\Delta_0/E_F \ll 1$ , i.e.,  $k_m^{(e)} \approx q_m^{(e)}$  except when they appear in an exponent), one finds

$$\Phi_m(\varepsilon) = 2k_m^{(e)} d. \quad (18)$$

Thus, from Eq. (14) the quantization condition becomes  $(k_m^{(e)} - k_m^{(h)})d = n\pi + \arccos(\varepsilon/\Delta_0)$ . Note that this result can also be derived from the Bohr-Sommerfeld quantization rule. In fact,  $\hbar \Phi_m(\varepsilon)$  is the classical action for the electron moving along the  $x$  direction between the superconductor and the wall at  $x = d$ . The numerical calculation reveals that the exact counting function  $N(\varepsilon)$  obtained from Eq. (9) “oscillates” around the Weyl formula given in Eq. (17) (for figure see Ref. 29). To find an analytical expression for the Weyl formula, the sum in Eq. (17) is replaced by an integral and we get

$$\tilde{N}(\varepsilon) = \frac{2}{\pi} \rho^{(N)} E_F g(\varepsilon/E_F) + M_h [1/2 - 1/\pi \arccos(\varepsilon/\Delta_0)], \quad (19)$$

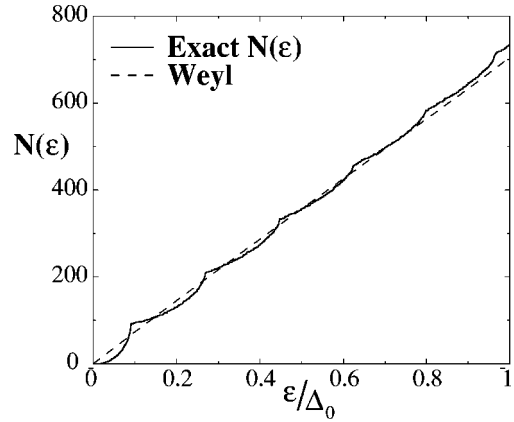


FIG. 2. The exact counting function  $N(\varepsilon)$  (solid line) and the analytical expression of the Weyl formula given in Eq. (19) (dashed line) as functions of  $\varepsilon/\Delta_0$  for the box geometry. The parameters are  $d/W = 3$ ,  $k_F W/\pi = 87.9$ , and  $\Delta_0/E_F = 0.02$ .

where

$$\begin{aligned} g(x) = & \sqrt{2x(1-x)} + (1+x) \arcsin \sqrt{(1-x)/(1+x)} \\ & - \pi/2(1-x), \end{aligned}$$

and  $\rho^{(N)} = 2m/\hbar^2 A/(4\pi)$  is the density of states for the isolated  $N$  region of area  $A = Wd$ . The number of open channels for the holelike quasiparticles is  $M_h = [k_F W/\pi \sqrt{1 - \varepsilon/E_F}]$ , where  $[\dots]$  stands for the integer part. For  $x \ll 1$  (typically  $x < 0.1$ )  $g(x) \approx \pi x$ . Thus, the leading term of the smooth part of the counting function of Andreev states is  $\tilde{N}(\varepsilon) \approx 2\varepsilon \rho^{(N)}$ . Electronlike and holelike quasiparticles make equal contributions to the Weyl formula (19), thereby the factor 2. The exact counting function  $N(\varepsilon)$  and the analytical form of the Weyl formula given in Eq. (19) are plotted in Fig. 2. For parameters see the figure caption. It is seen from the figure that the analytical expression (dashed curve) deviates only for energies close to the gap.

#### B. Weyl formula for disk (no mismatch)

In this section an analytical expression of the Weyl formula is derived for disk geometry. However, most of the results obtained in this section will be used in Sec. IV B, too. Again, we shall take the case of perfect  $NS$  interface, namely, no mismatch and tunnel barrier are assumed ( $r_k = 1$ ,  $r_v = 1$ ,  $Z = 0$ ). To find an analytical expression for the eigenphase  $\Phi_m(\varepsilon)$  defined in Eq. (13) we have to approximate the determinant  $D_m^{(e)}(\varepsilon)$  in Eq. (11). The details of the approximations are presented in Appendix A.

To summarize, according to the approximations of the determinants appearing in the secular equation (9), three ranges of  $m \geq 0$  can be distinguished: for Andreev states  $m < k_h R_S - \sqrt[3]{k_h R_S}$ , for whispering gallery states  $m > k_e R_S + \sqrt[3]{k_e R_S}$ , and for intermediate states  $k_h R_S - \sqrt[3]{k_h R_S} < m < k_e R_S + \sqrt[3]{k_e R_S}$ . Owing to the degeneracy of the  $\pm m$  states the three ranges for  $m \geq 0$  and  $m \leq 0$  are symmetrically located with respect to the state  $m = 0$ . This structure of the energy levels will be called a “mixed phase” (MP). The exact energy levels  $\varepsilon_{nm}$  obtained (numerically) from Eq. (9) are shown as functions of the angular quantum number  $m$  in Fig.

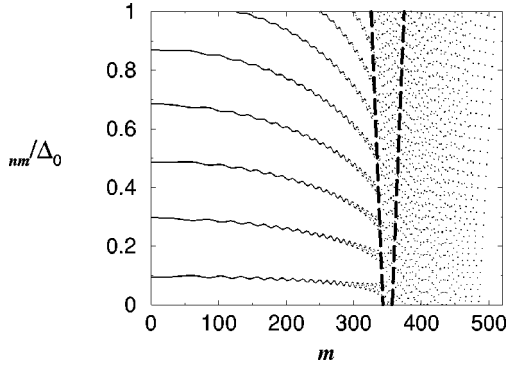


FIG. 3. The exact energy levels  $\varepsilon_{nm}$  (dots) in units of  $\Delta_0$  as a function of the angular momentum quantum number  $m$ . There is no mismatch, and the potential of the tunnel barrier is zero. The parameters are  $R_S/R_N=0.7$ ,  $\Delta_0/E_F=0.1$ , and  $k_F R_S=350$ . The intermediate  $m$  states are located between the two dashed lines. The Andreev states are to the left of the left dashed line, and the whispering gallery modes are to the right of the right one.

3. For each  $m$  the solutions are labeled by the quantum number  $n$  (radial quantum number). Since the pairs  $\pm m$  are degenerate, only  $m \geq 0$  states are plotted in the figure.

The two dashed lines in Fig. 3 separate the three types of states characterized by  $m$ . One can see that the intermediate states indeed occupy a narrow range in  $m$  compared to Andreev and whispering gallery states. It is also clear from the figure that the energy levels for Andreev states “oscillate” with increasing amplitude as  $m$  increases towards the onset of the intermediate states (left dashed line in the figure). Inserting Eq. (A4) into the quantization condition given by Eq. (14) one can approximately calculate the energy levels for Andreev states. Plotting these energy levels in Fig. 3 (not shown in the figure) we found that the exact energy levels “oscillate” around these approximate ones. Similar behavior of the energy levels has been found by Šipr *et al.*<sup>20</sup> For whispering gallery states the approximate secular equation is given by Eq. (A5), and its solutions coincide very accurately with exact energy levels. These approximate energy levels for Andreev and whispering gallery states are those which can be obtained from the semiclassical quantization of the Schrödinger equation describing the radial motion of the electron.<sup>22</sup> Indeed, one can check that the first two terms for  $\Phi_m(\varepsilon)$  in Eq. (A1) multiplied by  $\hbar$  give the *radial action* for the electron moving between the two circles. The last three terms are constant, i.e., independent of  $\varepsilon$  therefore they do not play any role in the dynamics of electron. These trajectories of the electrons correspond to Andreev states. The semiclassical orbits for Andreev and whispering gallery states are shown in Fig. 4.

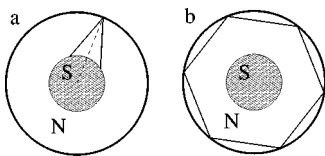


FIG. 4. The corresponding semiclassical orbits for (a) Andreev states and (b) whispering gallery states.

Inserting Eq. (A4) into the Weyl formula (17) one can determine the contributions of the Andreev states to the smooth part of the counting function

$$\tilde{N}_{AS}(\varepsilon) = \frac{1}{\pi} \sum_{|m| < M_{AS}} [\vartheta_m(k_e R_N) - \vartheta_m(k_e R_S) - \vartheta_m(k_h R_N) + \vartheta_m(k_h R_S)] + 2 M_{AS} \left( \frac{1}{2} - \frac{1}{\pi} \arccos \frac{\varepsilon}{\Delta_0} \right), \quad (20)$$

where  $M_{AS} = [k_h R_S - \sqrt[3]{k_h R_S}]$  is the highest  $m$  for the Andreev states.

For whispering gallery modes the eigenphase  $\Phi_m(\varepsilon)$  in Eq. (13) cannot be determined because  $D_m^{(e)}(\varepsilon)$  is approximately zero. In this case the contribution to the smooth part of the counting function will be calculated from the secular equation (A5) for whispering gallery states. It is well known that the zeros of the Bessel functions can be determined to a good approximation by applying Debye’s asymptotic expansion.<sup>22,36</sup> The solution of the secular equation (A5) for  $\varepsilon$  in the case of whispering gallery states is then equivalent to the solution of the following pair of equations:  $F_m^{(e)}(\varepsilon) = \vartheta_m(k_e R_N)/\pi - 3/4 = n_1$  and  $F_m^{(h)}(\varepsilon) = \vartheta_m(k_h R_N)/\pi - 3/4 = n_2$ , where  $n_1, n_2 = 0, \pm 1, \pm 2, \dots$ , and  $\vartheta_m(x)$  is given by Eq. (A2). From these equations one can find the density of states of whispering gallery states much in the same way as from Eq. (15) for the Andreev states

$$\rho_{WGS}(\varepsilon) = 2 \sum_{m=M_{WGS}}^{k_e R_N} \left| \frac{dF_m^{(e)}(\varepsilon)}{d\varepsilon} \right| + 2 \sum_{m=M_{WGS}}^{k_h R_N} \left| \frac{dF_m^{(h)}(\varepsilon)}{d\varepsilon} \right|, \quad (21)$$

where  $M_{WGS} = [k_e R_S + \sqrt[3]{k_e R_S}]$  is the smallest angular momentum quantum number  $m$  for whispering gallery states, and the factor 2 corresponds to the  $\pm m$  degeneracy. The contribution of whispering gallery states to the smooth part of the counting function can be obtained from  $\tilde{N}_{WGS}(\varepsilon) = \int_0^\varepsilon \rho_{WGS}(\varepsilon') d\varepsilon'$  which yields

$$\tilde{N}_{WGS}(\varepsilon) = \frac{2}{\pi} \sum_{m=M_{WGS}}^{k_e R_N} \vartheta_m(k_e R_N) - \frac{2}{\pi} \sum_{m=M_{WGS}}^{k_h R_N} \vartheta_m(k_h R_N). \quad (22)$$

The negative sign in front of the second term comes from the fact that the derivative of  $\vartheta_m(k_h R_N)$  with respect to  $\varepsilon$  is negative.

The counting function  $\tilde{N}(\varepsilon)$  for the NS disk system is the sum of the contribution of Andreev states  $[\tilde{N}_{AS}(\varepsilon)]$  and that of whispering gallery states  $[\tilde{N}_{WGS}(\varepsilon)]$ . We now neglect the contribution of the intermediate states. To find an analytical expression for  $\tilde{N}(\varepsilon)$ , the summation over  $m$  is replaced by an integral in Eqs. (20) and (22) and for  $k_F R_S \gg 1$   $M_{AS} \approx M_{WGS} \approx k_h R_S$  is taken. The last approximation corresponds to replacing the intermediate states by whispering gallery states. This is quite a good approximation, since the number of intermediate states is much smaller than that of the Andreev and whispering gallery states. The resulting integrals can be performed analytically. After some algebra we obtain

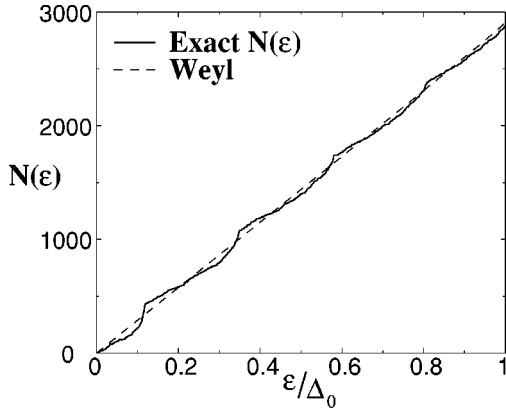


FIG. 5. The exact counting function  $N(\varepsilon)$  obtained from Eqs. (9) (solid line) and  $\tilde{N}(\varepsilon)$  given by Eq. (23) (dashed line) as functions of  $\varepsilon/\Delta_0$  for the disk geometry. There is no mismatch and the potential of the tunnel barrier is zero. The parameters are  $R_S/R_N = 2/7$ ,  $\Delta_0/E_F = 0.05$ , and  $k_F R_S = 100$ .

$$\tilde{N}(\varepsilon) = \frac{k_F^2 R_N^2}{2} \frac{\varepsilon}{E_F} - \frac{k_F^2 R_S^2}{2} g\left(\frac{\varepsilon}{E_F}\right) + 2k_F R_S \sqrt{1 - \frac{\varepsilon}{E_F}} \left( \frac{1}{2} - \frac{1}{\pi} \arccos \frac{\varepsilon}{\Delta_0} \right), \quad (23)$$

where

$$g(x) = \frac{3}{\pi} \sqrt{2x(1-x)} + \frac{1}{\pi} (1+x) \arcsin \sqrt{\frac{1-x}{1+x}} - (1-x) \left( \frac{1}{2} + \frac{2}{\pi} \arccos \sqrt{\frac{1-x}{1+x}} \right).$$

This is our Weyl formula for  $NS$  concentric disk systems. Usually,  $\varepsilon/E_F \ll 1$  and for  $x \ll 1$ ,  $g(x) = x$  in leading order, therefore the first two terms in Eq. (23) yield  $2\rho^{(N)}\varepsilon$ , where  $\rho^{(N)} = 2m/\hbar^2 A/(4\pi)$  is the density of states for a normal annular billiard of area  $A = (R_N^2 - R_S^2)\pi$ . The factor 2 comes from the electron/hole contributions for  $NS$  systems. The higher order corrections in  $g(x)$  become more relevant with increasing  $\varepsilon$ . The last term in Eq. (23) is related to the phase shift due to the Andreev reflection at the  $NS$  interface. It is important to stress that the Weyl formula given in Eq. (23) was derived for the case  $R_S \gg \xi(\varepsilon)$ , i.e., when Andreev states and whispering gallery states coexist (mixed phase). In the opposite case, as we shall see in Sec. V, no Andreev states exist.

In Fig. 5 the exact counting function and the Weyl formula given by Eq. (23) are plotted as functions of  $\varepsilon/\Delta_0$ . For parameters see the figure caption. The singularities of the DOS (derivative of the counting function with respect to  $\varepsilon$ ) will be discussed in the next section.

#### IV. BOHR-SOMMERFELD APPROXIMATION TO THE DENSITY OF STATES

The exact counting function  $N(\varepsilon)$  has *cusps* at some energies as shown Figs. 2 and 5 for box and disk geometries,

and perfect interface. This implies that the DOS of the Andreev states is *discontinuous* here. Moreover, these cusps are at *equal distances*. Similar singularities of the DOS for  $NS$  systems have already been found by de Gennes *et al.*<sup>7</sup> as well as in the numerical works of Richter *et al.*<sup>15</sup> (for  $NS$  systems), Nazarov *et al.*<sup>13</sup> (for Andreev billiard), and Sipr *et al.*<sup>20</sup> (for  $SNS$  systems). In this section we investigate the singularities in the DOS for the  $NS$  systems shown in Fig. 1.

Several authors<sup>9,13-15</sup> have already derived the Bohr-Sommerfeld approximation to the density of states

$$\rho_{BS}(\varepsilon) = M \int_0^\infty ds P(s) \sum_{n=0}^\infty \delta \left[ \varepsilon - \left( n + \frac{1}{2} \right) \frac{\pi \hbar v_F}{s} \right], \quad (24)$$

where  $M$  is the number of modes in the superconducting leads connected to the normal billiard and  $P(s)$  is the classical probability that an electron entering the billiard exits after a path length  $s$ . Starting from our quantization rule given in Eq. (14) we rederive Eq. (24) in this section, and give an explicit expression of the probability  $P(s)$  in terms of  $\Phi_m(\varepsilon)$ . We shall show that  $P(s)$  is singular at some  $s$  resulting in the singularities of the DOS. These singularities correspond to the cusps in the integrated DOS  $N(\varepsilon)$ . It turns out that the simple Bohr-Sommerfeld approximation to the DOS is a good approach to understand the singularities in the DOS provided that  $P(s)$  is approximated correctly for small  $s$ .

The solution of Eq. (14) gives the discrete energy levels for the  $NS$  systems. Since  $\varepsilon \ll E_F$ , one can Taylor expand the left-hand side (LHS) of Eq. (14) in terms of  $\varepsilon$  and find

$$\varepsilon_{n,m} = \frac{(n + \frac{1}{2})\pi}{\Phi'_m(0)}, \quad (25)$$

where  $\Phi'_m(0)$  denotes the derivative of  $\Phi_m(\varepsilon)$  with respect to  $\varepsilon$ , evaluated at the Fermi energy, i.e.,  $\varepsilon = 0$ . For the  $NS$  box systems  $m = 1, 2, \dots, M$ , where the energy dependent  $M$  is the number of “transverse modes.” For the  $NS$  disk systems  $|m| < M$ , where  $M$  is the maximum of the angular momentum quantum number  $m$  for the Andreev states. We consider only the positive energy spectrum, therefore  $n \geq 0$ . Here we take the box geometry. The following expressions for disk geometry can be obtained straightforwardly by taking into account the fact that the set of  $m$  is different in this case. Thus  $M$  has a different meaning for box and disk geometries. We shall always indicate how to convert the box geometry results to the disk geometry case. To get Eq. (25), the phase term in Eq. (14) due to Andreev reflection,  $\arccos(\varepsilon/\Delta_0)$  was approximated by  $\pi/2$ , which is valid for  $\varepsilon \rightarrow 0$ . Later a better approximation will be given by Taylor expanding the phase term, too. We now show that the density of states  $\rho(\varepsilon) = \sum_{n=0}^\infty \sum_{m=1}^M \delta(\varepsilon - \varepsilon_{n,m})$  can be rewritten in the same form as in Eq. (24), which makes it possible to express the probability  $P(s)$  in terms of  $\Phi_m(\varepsilon)$ . Applying the Poisson formula<sup>32,22</sup> for the summation over  $m$ , keeping only the nonoscillating term in  $m$ , and then performing the integral over  $m$  we obtain

$$P(s) = \frac{\Theta(M - m^*)\Theta(m^* - 1)}{M\hbar v_F \left| \frac{\partial \Phi - m'(0)}{\partial m} \right|_{m=m^*}}, \quad (26)$$

where the  $s$ -dependent  $m^*$  satisfies

$$\hbar v_F \Phi - m'(0)|_{m=m^*} = s. \quad (27)$$

$\Theta(x)$  is the Heaviside function (for details see Ref. 29). Note that the probability  $P(s)$  is normalized to 1, i.e.,  $\int_0^\infty P(s) ds = 1$ . For disk geometry the numerator in Eq. (26) has to be replaced by  $\Theta(M - |m^*|)/2$ , and the necessary replacement in Eq. (24) is  $M \rightarrow 2M$ . Our main result is that the probability  $P(s)$  is expressed in terms of  $\Phi_m(\varepsilon)$ , which has already been determined for box and disk geometries of the  $NS$  system [see Eqs. (18) and (A1)].

Performing the integral over  $s$  in Eq. (24) one finds

$$\rho_{BS}(\varepsilon) = \frac{M}{\varepsilon} \sum_{n=0}^{\infty} s_n(\varepsilon) P[s_n(\varepsilon)], \quad (28)$$

where

$$s_n(\varepsilon) = \frac{(n + \frac{1}{2})\pi\hbar v_F}{\varepsilon} = \frac{(n + \frac{1}{2})2\pi}{k_F} \frac{E_F}{\varepsilon}. \quad (29)$$

Similarly, the counting function  $N_{BS}(\varepsilon) = \int_0^\varepsilon d\varepsilon' \rho_{BS}(\varepsilon')$  can be easily found

$$N_{BS}(\varepsilon) = M \sum_{n=0}^{\infty} \int_{s_n(\varepsilon)}^{\infty} P(s) ds. \quad (30)$$

In both Eqs. (28) and (30) we have to replace  $M$  by  $2M$  in the case of  $NS$  disk systems as  $M$  has a different meaning in this case. The above results are valid only for small  $\varepsilon$ . A better approximation of the counting function can be obtained by Taylor expanding the phase shift related to the Andreev reflection  $\arccos(\varepsilon/\Delta_0) \approx \pi/2 - \varepsilon/\Delta_0$ . Carrying out the same procedures as before one finds that the counting function  $N_{BS}(\varepsilon)$  is still given by Eq. (30) after the replacement

$$s_n(\varepsilon) \rightarrow s_n(\varepsilon) - \xi_0 \quad (31)$$

is made in Eq. (29). Here  $\xi_0 = \xi(\varepsilon=0) = \hbar v_F/\Delta_0 = 2E_F/(k_F\Delta_0)$  is the coherence length at the Fermi energy [see Eq. (A3)].

In the following two subsections we shall calculate  $P(s)$  and the counting function  $N_{BS}(\varepsilon)$  for  $NS$  box and disk systems. We shall see that the probability  $P(s)$ , more precisely  $sP(s)$ , is singular at some  $s_{\text{sing}}$  for both  $NS$  systems. Then, if  $s_{\text{sing}}$  is known, one can determine from Eq. (31) the energies where  $P[s_n(\varepsilon)]$ —and consequently [see Eq. (28)] the density of states  $\rho_{BS}(\varepsilon)$ —is singular:

$$\varepsilon_n^{(\text{sing})} = \frac{(n + 1/2)\pi}{1 + s_{\text{sing}}/\xi_0} \Delta_0, \quad (32)$$

which is valid for all  $n$  for which  $\varepsilon_n^{(\text{sing})} < \Delta_0$ . Note that this expression can be applied to those normal systems attached

to a superconductor for which  $sP(s)$  is singular. It is clear from this expression that these singularities are at *equal* distances. The analytical behavior of  $P(s)$ , and thus, the existence of the singularities in the DOS for perfect  $NS$  interface inherently depends on the geometry of the isolated normal billiard. Once  $sP(s)$  is a singular function of  $s$  (because of the shape of the normal region), then the subsequent singularities in the DOS are at equal distances. We shall see that from the exact calculation of the energy eigenvalues for  $NS$  box and disk systems the positions of the singularities in the DOS agree very well with those given by the above derived approximate expression. We would like to emphasize that the second term in the denominator of Eq. (32) resulted from the Taylor expansion of  $\arccos(\varepsilon/\Delta_0)$  which is the phase shift due to the Andreev reflection. We shall see below that this extra term involving the coherence length  $\xi_0$  in Eq. (32) is necessary for getting good agreements between the exact and that of predicted by Eq. (32) for the position of the singularities.

### A. Singularities in DOS for box

In Sec. III A we have calculated  $\Phi_m(\varepsilon)$  for the  $NS$  box system with perfect interface. Inserting  $\Phi_m(\varepsilon)$  given by (18) into Eqs. (26) and (27) one finds

$$P(s) = \frac{4d^2}{s^3 \sqrt{1 - \left(\frac{2d}{s}\right)^2}} \Theta(s - s_{\text{min}}), \quad (33)$$

where  $s_{\text{min}} = 2d/\sqrt{1 - 1/M^2}$  and  $M = k_F W/\pi$  is the number of “transverse modes.” For a large number of transverse modes, i.e.,  $M \gg 1$  we have  $s_{\text{min}} = 2d$ . Notice that the same result can be found for  $P(s)$  by simple geometrical considerations.<sup>33</sup>  $P(s)$  is singular at  $s_{\text{sing}} = 2d$ , therefore the DOS is singular at the energies given by Eq. (32). The DOS can be calculated from Eq. (28) in Bohr-Sommerfeld approximation (for figure see Ref. 29).

From Eq. (29) it is clear that the DOS for small  $\varepsilon$  is dominated by the large  $s$  behavior of  $P(s)$ . According to Eq. (33),  $P(s) \rightarrow 4d^2 s^{-3}$  in this case. Thus,  $\rho_{BS}(\varepsilon)/(2\rho^{(N)}) \rightarrow \varepsilon/(\pi E_T)$ , for  $\varepsilon \rightarrow 0$ , where  $\rho^{(N)} = (2m_N/\hbar^2)A/(4\pi)$  is the DOS for the isolated billiard of area  $A = dW$  and  $E_T = M/(4\pi\rho^{(N)})$  is the Thouless energy defined in Refs. 9,15. For small energies  $\varepsilon$  the DOS is proportional to  $\varepsilon$  in agreement with the findings of Melsen *et al.*<sup>9,10</sup> and Ihra *et al.*<sup>15</sup> However, the slope of  $\rho_{BS}(\varepsilon)/2\rho^{(N)}$  is less than what these authors found. The reason is that the geometries of the billiards they studied are slightly different from our box geometry. Namely, the width  $w$  of their superconducting lead is smaller than the side length of the billiard with which it is in contact. In these geometries, by quantum mechanical calculations, Ihra *et al.*<sup>15</sup> found pronounced peaks in the DOS. These peaks are also related to the singularities in  $P(s)$ .<sup>34</sup> However, to confirm the existence of the singularity in  $P(s)$  one needs to calculate  $P(s)$  on a finer scale than in Fig. 2 of their paper.



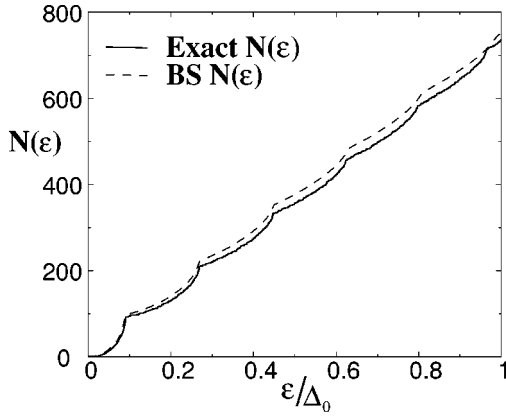


FIG. 6. The exact counting function  $N(\varepsilon)$  (solid line) and  $N_{BS}(\varepsilon)$  given in Eq. (34) (dashed line) as functions of  $\varepsilon/\Delta_0$  for the box geometry. The parameters are the same as in Fig. 2.

To avoid the errors of numerical differentiation, we shall compare the counting function obtained from the quantum mechanical calculations [see Eq. (9)] with the counting function  $N_{BS}(\varepsilon)$  in the Bohr-Sommerfeld approximation. Using Eqs. (30) and (33), and performing the integration, we obtain

$$N_{BS}(\varepsilon) = M \sum_{n=0}^{\infty} f[s_n(\varepsilon)], \quad (34)$$

where

$$f(s) = \begin{cases} 1 & \text{if } s \leq 2d, \\ 1 - \sqrt{1 - 4d^2/s^2} & \text{if } s > 2d, \end{cases}$$

and  $s_n(\varepsilon)$  is given in Eq. (31). The above derived counting function  $N_{BS}(\varepsilon)$  is plotted in Fig. 6 together with the exact counting function.

One can see that the agreement is excellent at low energies (below the second cusp) but for higher energies the exact counting function is smaller than the approximated one. However, the positions of the singularities agree quite well in the two curves. Slight deviations are observed only for the last two cusps. The same was found for other parameter ranges. Thus, our expression (32) for the positions of the singularities in the DOS results in an excellent agreement with exact quantum mechanical calculations. Our result may highlight the origin of the singularities of the DOS. One can see that  $P(s)$  is singular at  $s = 2d$ , which corresponds to the classical orbits when the electron enters the box parallel to the  $x$  axis and is reflected back at the wall. The enhanced return probabilities for these orbits are quite obvious. These are the orbits which result in the singularities of the DOS. It is also clear from this expression that the singularities are at equal distances.

The expression for the positions of the singularities given by Eq. (32) can be applied to other  $NS$  systems, e.g., that studied by de Gennes and Saint-James.<sup>7</sup> We calculated the positions of the singularities from Eq. (32) and found that they are in excellent agreement with those found from the results of de Gennes and Saint-James (for details see Ref. 29). Finally, we mention that Lodder *et al.*<sup>13</sup> and Sipr *et al.*<sup>20</sup>

also found singularities in the DOS. We believe that these singularities are also related to some special classical orbits which give rise to a singular behavior of the probability  $P(s)$ , similarly to the  $NS$  systems studied in this paper.

We also calculated the energy levels in the case of nonperfect  $NS$  interface by solving the secular equation (9) numerically. We have used the same parameters for describing the mismatch in the Fermi energies and the effective masses as in Ref. 35, namely,  $r_k = k_F^{(N)}/k_F^{(S)}$ ,  $r_v = v_F^{(N)}/v_F^{(S)}$ . The strength of the tunnel barrier is given by  $Z$ . We found that for a nonperfect  $NS$  interface the cusps in the counting function are “washed out” (for figure see Ref. 29). This is related to the fact that the normal reflections at the interface are enhanced. The same is true for  $Z \neq 0$ . In Sec. III A we have found that in leading order the Weyl formula is given by  $\tilde{N}(\varepsilon) \approx 2\varepsilon \rho^{(N)}$ , where  $\rho^{(N)} = 2m/\hbar^2 A/(4\pi)$  is the density of states for the isolated  $N$  region of area  $A = Wd$ . The exact counting function can be well approximated by the above Weyl formula (for figure see Ref. 29). This implies that for nonperfect  $NS$  interfaces the effect due to the Andreev reflection is less significant, and the system behaves as a normal metal regarding the energy levels.

### B. Singularities in DOS for disk

The maximum angular momentum quantum number,  $M_{AS}$  for Andreev states has been determined in Sec. III B after Eq. (20). To have a better physical picture of these states we now give an estimation for  $M_{AS}$  based on semiclassical considerations. It is easy to see that the angular momentum (in units of  $\hbar$ ) of the rays shown in Fig. 4(a) is  $m = k_F R_N \sin(\alpha/2)$ , where  $\alpha$  is the angle between the two rays at the outer circle. Rays that are tangent to the inner circle have maximal angular momentum, in which case  $\sin(\alpha/2) = R_S/R_N$ . These are the trajectories which still reach the superconductor so that they belong to the Andreev states. Thus, the maximal angular momentum  $m$  for the Andreev states is  $k_F R_S$ . This is a good estimate for  $M_{AS}$  given after Eq. (20).

In Sec. III B we have seen that for  $R_S \gg \xi(\varepsilon)$  and for perfect  $NS$  interface the energy levels for  $NS$  disk systems can be classified into Andreev and whispering gallery states to a good approximation. We first consider the Andreev states. Inserting  $\Phi_m(\varepsilon)$  given by Eq. (A1) into Eqs. (26) and (27), one finds after some algebra

$$P(s) = \frac{1}{4s^2 R_S} \frac{s_{\max}^4 - s^4}{\sqrt{[s_{\max}^4/s_{\min}^2 - s^2][s^2 - s_{\min}^2]}} \times \Theta(s_{\max} - s) \Theta(s - s_{\min}), \quad (35)$$

where  $s_{\min} = 2(R_N - R_S)$  and  $s_{\max} = 2\sqrt{R_N^2 - R_S^2}$ . The two Heaviside functions come from the factor  $\Theta(M - |m|)/2$  [see the text after Eq. (27)] when we change to the variable  $s$ . Here  $s_{\min}$  and  $s_{\max}$  correspond semiclassically to the smallest and the largest path lengths of the orbits related to the Andreev states. Obviously,  $s_{\min}$  equals to the path length of those orbits for which the trajectory of the electrons between two successive bounces at the  $NS$  interface is along the radius of the two concentric circles. The trajectories of the

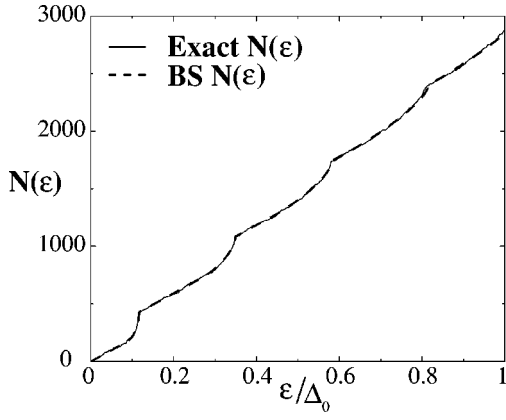


FIG. 7. The exact counting function  $N(\varepsilon)$  (solid line) and  $N_{BS}(\varepsilon)$  given in Eq. (38) (dashed line) as functions of  $\varepsilon/\Delta_0$  for the disk geometry. The parameters are the same as in Fig. 5. Then  $s_{\min}/\xi_0 = 12.5$ .

electrons corresponding to the maximal path length for Andreev states are those which are tangent to the inner circle of radius  $R_S$ . The probability  $P(s)$  is zero for  $s_{\min} > s > s_{\max}$  and normalized to one. Note that in deriving Eq. (35), the solution of Eq. (27) is twofold for  $m^*$  resulting in an extra factor 2 in  $P(s)$ . Notice that the same result can be found for  $P(s)$  by simple geometrical considerations.<sup>33</sup> One can see that  $P(s)$  is singular at  $s = s_{\min}$ . The contributions of the Andreev states to the DOS in Bohr-Sommerfeld approximation is given by Eq. (28), which reads

$$\rho_{AD}(\varepsilon) = \frac{2k_F R_S}{\varepsilon} \sum_{n=0}^{\infty} s_n(\varepsilon) P[s_n(\varepsilon)], \quad (36)$$

where  $P(s)$  is given by Eq. (35) and  $s_n(\varepsilon)$  by Eq. (31). Since there is a maximum path length  $s$  for Andreev states, the sum over  $n$  in Eq. (28) is indeed finite. It can be shown that  $\rho_{AD}(\varepsilon) \rightarrow 0$  as  $\varepsilon \rightarrow 0$ .

Next, we consider the contribution of the whispering gallery states to the DOS. After differentiations with respect to  $\varepsilon$  in Eq. (21) and replacing the summations over  $m$  by integrals the calculations can be carried out analytically (here we

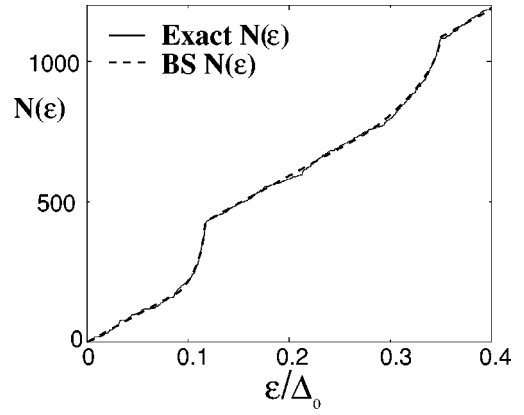


FIG. 8. Enlarged portion of Fig. 7.

used  $M_{WGS} \approx k_e R_S$  and neglected the higher order terms in  $\varepsilon/E_F$ ). It is found that the DOS is *constant* in  $\varepsilon$  and is given by

$$\rho_{WGS} = \frac{(k_F R_N)^2}{2E_F} \left[ 1 - \frac{2}{\pi} \left( \frac{R_S}{R_N} \sqrt{1 - \left( \frac{R_S}{R_N} \right)^2} + \arcsin \frac{R_S}{R_N} \right) \right]. \quad (37)$$

Thus, the total DOS for the NS disk system in Bohr-Sommerfeld approximation is  $\rho_{WGS} + \rho_{AD}(\varepsilon)$ . Since  $P(s)$  is a singular function, the DOS becomes singular at the energies given by Eq. (32) with  $s_{\text{sing}} = 2(R_N - R_S)$ . One can see that the DOS does not tend to zero as  $\varepsilon \rightarrow 0$ , contrary to the box geometry case (for figure see Ref. 29). The nonvanishing DOS is due to the constant contribution of the whispering gallery states given by Eq. (37).

Again, one can determine the counting function  $N_{BS}(\varepsilon)$  by integrating the DOS. Using Eqs. (30) and (35) the integral can be performed analytically, and we have

$$N_{BS}(\varepsilon) = \rho_{WGS} \varepsilon + 2k_F R_S \sum_{n=0}^{\infty} f[s_n(\varepsilon)] \Theta[s_{\max} - s_n(\varepsilon)], \quad (38)$$

where

$$f(s) = \begin{cases} 1 & \text{if } s \leq s_{\min}, \\ 1 - \frac{\sqrt{[s_{\max}^4/s_{\min}^2 - s^2][s^2 - s_{\min}^2]}}{4 s R_S} & \text{if } s_{\min} < s \leq s_{\max}, \end{cases}$$

and  $s_n(\varepsilon)$  is given by Eq. (31). Note that the infinite sum over  $n$  in Eq. (38) is indeed a finite one due to the Heaviside function. Using Eqs. (9) and (11) we calculated the exact energy levels for the NS disk system. In Fig. 7 the exact counting function  $N(\varepsilon)$  and  $N_{BS}(\varepsilon)$  are plotted for perfect interface (no mismatch and zero tunnel barrier). One can see that the agreement between the exact counting function and that obtained from the Bohr-Sommerfeld approximation is

excellent for all energies below the gap. To see the agreement, the details of Fig. 7—up to the first two cusps—are enlarged in Fig. 8. Similarly, very good agreements were found for other parameter ranges of the NS disk systems with perfect interface.

In Fig. 3 one can see that the energy levels (for each radial quantum number  $n$ ) as functions of  $m$  are approximately constant as  $m \rightarrow 0$ . Thus, the DOS should diverge because it

is proportional to the reciprocal of the derivative of  $\varepsilon_{mn}$  with respect to  $m$  (assuming that  $m$  is a continuous variable). The positions of the singularities of the DOS coincide with the energies  $\varepsilon_{mn}$  at  $m=0$  for all possible  $n$ . The angular momentum quantum number  $m=0$  semiclassically corresponds to the radial orbits of the electron between the inner and outer circles. Applying the Bohr-Sommerfeld quantization rule for the electrons and the Andreev reflected holes along the radius one finds  $[k_e(\varepsilon) - k_h(\varepsilon)](R_N - R_S) = n\pi + \arccos(\varepsilon/\Delta_0)$ . Note that this equation can also be derived from the quantization condition given in Eq. (14). The solutions of this equation agree very well with the energies in Fig. 3 at  $m=0$ . Therefore, the singularities of the DOS are related to the classical orbits along the radial directions.

We also calculated the energy levels in the case of a nonperfect  $NS$  interface by solving the secular equation (9) numerically. For a nonperfect  $NS$  interface the cusps in the counting function are “washed out” similarly to the case of  $NS$  box systems (for figure see Ref. 29). This is again related to the fact that the normal reflections at the interface are enhanced. Our numerical results show that the cusps also disappear for  $Z \neq 0$ . Using the Weyl formula derived in Sec. III B we found that  $\tilde{N}(\varepsilon) = 2\rho^{(N)}\varepsilon$  in leading order, where  $\rho^{(N)} = (2m/\hbar^2)(R_N^2 - R_S^2)/4$  is the density of states for normal annular billiard with inner radius  $R_S$  and outer radius  $R_N$ . This Weyl formula is a good approximation for the exact counting function in the case of nonperfect interfaces. The result shows that the role of normal reflections dominates over Andreev reflections, and thus the Weyl formula agrees approximately with that for the circular annular billiard without superconducting region.

## V. “PHASE DIAGRAM” FOR $NS$ DISK SYSTEMS

In our studies of the energy levels of the  $NS$  disk systems we have assumed so far that the energy dependent coherence length  $\xi(\varepsilon) = \hbar v_F / \sqrt{\Delta_0^2 - \varepsilon^2}$  is much smaller than  $R_S$  [see the text after Eq. (A3)]. In this section we shall discuss the case when  $\xi(\varepsilon) \gg R_S$ , i.e., when  $\varepsilon \rightarrow \Delta_0$ . The imaginary part of  $q_e R_S \approx k_F R_S + i R_S / \xi(\varepsilon)$  is then negligible, and  $k_e R_S \approx q_e R_S$ . Thus, in the determinant  $D_m^{(e)}(\varepsilon)$  given in Eq. (11) the arguments of the Bessel functions are almost equal. We now consider the perfect  $NS$  interface ( $Z=0$  and  $m_N = m_S$ ). Subtracting the second column from the first one in the determinant, the ratio  $J_m(k_e R_N) / Y_m(k_e R_N)$  can be factored out. It turns out that the remaining determinant is nonzero for  $\varepsilon < \Delta_0$ . Thus, similarly to the whispering gallery states discussed in Sec. III B, it holds to a very good approximation that  $D_m^{(e)}(\varepsilon)$  (as a function of  $\varepsilon$ ) has zeros whenever  $J_m(k_e R_N) = 0$  for all  $m$ . This corresponds to the energy levels of a circular billiard of radius  $R_N$  for electronlike quasiparticles. The same is true for holelike quasiparticles, i.e., their energy levels are the solutions of  $J_m(k_h R_N) = 0$  for all  $m$ . In summary, we found that for  $\xi(\varepsilon)/R_S \gg 1$  the  $NS$  disk system behaves as a normal full disk, while in the opposite case the energy levels of the system have a “mixed phase” character, in which Andreev states and whispering gallery states coexist (see Sec. III B). We shall call the first case a “full disk” (FD)

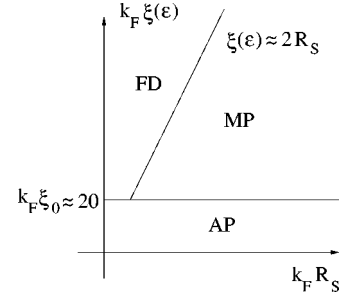


FIG. 9. Schematic phase diagram for the  $NS$  disk systems. The mixed phase (MP) is bounded by the lines  $k_F \xi_0 \approx 20$  and  $\xi(\varepsilon) \approx 2R_S$ . The full disk (FD) phase is bounded by the line  $k_F R_S = 0$ , the line  $k_F \xi_0 \approx 20$  and the line  $\xi(\varepsilon) \approx 2R_S$ . The annular phase (AP) is below the line  $k_F \xi_0 \approx 20$ .

phase. According to our numerical results, it is safe to say that the MP and FD can be distinguished very well by the ratio  $\xi(\varepsilon)/R_S$ . We found mixed phase for  $\xi(\varepsilon)/R_S < 2$  and FD phase for  $\xi(\varepsilon)/R_S > 2$ .

As it is known, the Andreev approximation is valid only for  $\Delta_0/E_F \ll 1$ . If we assume some critical value, say  $\Delta_0/E_F = 0.1$ , then for  $k_F \xi_0 = 2E_F/\Delta_0 < 20$  the Andreev approximation is not valid and normal reflection at the  $NS$  interface dominates over Andreev reflection. The system looks similar to an annular billiard (without superconductor region). We shall call this case an “annular phase” (AP). As we have seen, the nonperfect interface also results in an enhanced normal reflection and the character of the energy levels again corresponds to a circular annular billiard (annular phase).

We calculated the exact energy levels for different parameters and compared with those obtained for the full disk and for circular annular disk. In this way, we can classify the  $NS$  disk systems into the MP, FD, and AP phase regions. The crossovers between the three “phases” are not sharp and depend on the energy, too. It turns out that the three phases can be characterized by two parameters  $k_F R_S, k_F \xi(\varepsilon)$ . These are the relevant parameters for the  $NS$  disk systems and each pair corresponds to one of the phases. Thus, a so-called phase diagram can be given for the different phases. The boundaries of these phases are not so sharp, however. Far from the boundaries the given phase becomes dominating. In Fig. 9 the three different phases are shown.

## VI. CONCLUSIONS

We investigated the energy spectrum of  $NS$  box and disk systems using the BdG equation. Matching the wave functions at the  $NS$  interface we derived a secular equation whose solutions give the energy spectrum of the system. The mismatch in the Fermi wave numbers and the effective masses of the normal system and the superconductor, as well as the tunnel barrier is included in the calculation. Rewriting the secular equation a simple quantization condition was derived. Equation (14) is a central result and serves as a starting point for further theoretical studies in this paper. Using this quantization condition (a) we derived for both  $NS$  systems with perfect interfaces a Weyl formula for the counting func-

tion, (b) by rederiving the commonly used Bohr-Sommerfeld approximation to the DOS we presented an explicit expression for the probability  $P(s)$  in terms of the classical action of the electron moving in the  $N$  region between two successive bounces at the  $NS$  interface. The numerically exact calculations are in very good agreement with our Weyl formula and the Bohr-Sommerfeld approximation. The singularities in the DOS obtained from our exact calculations can be successfully explained by using the correct probability function  $P(s)$ .  $P(s)$  is a singular function of  $s$  both for  $NS$  box and disk systems. According to our theory concerning the singularities, this implies that the singularities of the DOS are located at equal distances from each other. A simple formula is given for the location of the singularities. We demonstrated that it can be applied for the system studied by de Gennes *et al.*,<sup>7</sup> and it may as well give the reason for the significant peaks observed by Ihra *et al.*<sup>15</sup> in their numerical calculations. We show that the singularities in  $P(s)$  are related to some special classical trajectories of the electron moving in the  $N$  region hitting the  $NS$  interface.

In the case of  $NS$  disk systems the DOS is constant as the energy gets close to the Fermi level. This is, at first sight, in contrast to the common belief that for integrable billiards the DOS should go to zero as the energy tends the Fermi level. We pointed out that in this  $NS$  disk system the whispering gallery states give the nonvanishing DOS at the Fermi level. However, the Andreev states still have no contribution to the DOS. In the disk system the Andreev and the whispering gallery states coexist (so-called mixed phase). Depending on the geometry of the disk and the material parameters, we sketch a kind of phase diagram to classify the energy spectra into three classes: (i) mixed phase, (ii) full disk, (iii) annular disk. If the coherence length is much larger than the diameter of the superconducting region, the spectrum corresponds to the full disk phase. The electrons travel thorough the  $S$  region without Andreev reflection. When the Andreev approximation fails, the normal reflection is enhanced at the  $NS$  interface and the system behaves as an annular circular billiard. This is also the case when there is a tunnel barrier at the  $NS$  interface or a mismatch between the effective masses and the Fermi wave numbers of the normal and superconducting regions.

An interesting extension of the study presented in this paper may be the investigation of the role that the geometry of the normal billiard plays in the singularities of the DOS. Under what conditions do such singularities exist? Does the energy spectrum at higher energies show some special structure in chaotic billiards? How can the Weyl formula (derived for  $NS$  box and disk systems) be extended for other integrable and chaotic billiards? Our quantization method can also be applied for a  $NS$  disk in the presence of a magnetic field. In this case, besides Andreev states and whispering gallery states (more specifically edge states) one has to take into account the Landau levels that appear in the spectrum in the presence of strong magnetic fields. These give rise to a much more complex energy spectrum than a zero magnetic field. Another relevant question is the study of the counting function in these systems. Further work along these lines is in progress.

## ACKNOWLEDGMENTS

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## APPENDIX A: CALCULATION OF THE PHASE $\Phi_m(\varepsilon)$

It is clear for symmetry reasons that the eigenstates with quantum numbers  $m$  and  $-m$  are degenerate (except for  $m=0$ ), thus it is enough to consider the states  $m \geq 0$ . In the Andreev approximation ( $\Delta_0/E_F \ll 1$ ) we take  $k_m^{(e)} \approx q_m^{(e)}$  in places where they are multiplied by the Bessel functions in  $D_m^{(e)}(\varepsilon)$  but in the arguments of the Bessel functions we keep them to be different. As we shall see, different methods are necessary to approximate the determinant  $D_m^{(e)}(\varepsilon)$  for different ranges of  $m$ . The Debye asymptotic expansion<sup>36</sup> will be used to approximate the determinant for  $|m| < k_e R_S - \sqrt[3]{k_e R_S}$ . In the intermediate range  $|m - k_e R_S| < \sqrt[3]{k_e R_S}$  one may apply uniform asymptotic expansions<sup>36</sup> for the Bessel functions. For Bessel functions of a complex argument  $z$  with fixed  $m$  the approximate form<sup>36</sup>

$$J_m(z) \approx \sqrt{\frac{2}{\pi z}} \cos\left(z - \frac{1}{2}m\pi - \frac{\pi}{4}\right)$$

for  $z \rightarrow \infty$  will be used. We first calculate the eigenphase  $\Phi_m(\varepsilon)$  for  $|m| < k_e R_S - \sqrt[3]{k_e R_S}$ . Using the Debye asymptotic expansion for the Bessel functions in Eqs. (11) and (13), one finds (for details see Ref. 29)

$$\Phi_m(\varepsilon) = 2[\vartheta_m(k_e R_N) - \vartheta_m(k_e R_S)] + 2k_F R_S - m\pi - \frac{\pi}{2}, \quad (\text{A1})$$

where

$$\vartheta_m(x) = \sqrt{x^2 - m^2} - |m| \arccos \frac{|m|}{x}. \quad (\text{A2})$$

We assumed that  $\Delta_0 \ll E_F$  and  $R_S \gg \xi(\varepsilon)$  (the opposite of the latter case will be discussed in Sec. V), where the coherence length  $\xi(\varepsilon)$  is defined<sup>6</sup> as

$$\xi(\varepsilon) = \frac{\hbar v_F}{\sqrt{\Delta_0^2 - \varepsilon^2}} = \frac{2}{k_F |\eta|}. \quad (\text{A3})$$

Here  $v_F$  is the Fermi velocity. The absolute values of  $m$  in the above expressions are the consequence of degeneracy of the eigenstates  $m$  and  $-m$ .

We now turn to the case  $|m| > k_e R_S + \sqrt[3]{k_e R_S}$ . In this case, according to the analogous Debye asymptotic expansion of the Bessel functions, one finds that  $J_m(k_e R_S)$  can be neglected (it is exponentially small) compared to the second term of the 1,1 element of the determinant in Eq. (11). Similarly,  $J'_m(k_e R_S)$  is also negligible in the determinant. Thus,



the ratio  $J_m(k_e R_N)/Y_m(k_e R_N)$  can be taken out from the determinant. It turns out that the remaining determinant is not equal to zero for  $\varepsilon < \Delta_0$ . Thus, to a very good approximation,  $D_m^{(e)}(\varepsilon)$  (as a function of  $\varepsilon$ ) has zeros whenever  $J_m(k_e R_N) = 0$ . This corresponds to the energy levels of a circular billiard of radius  $R_N$  for electronlike quasiparticles. Note that now  $\Phi - m(\varepsilon)$  cannot be calculated from Eq. (13), since  $D_m^{(e)}(\varepsilon) \approx 0$  for  $|m| > k_e R_S + \sqrt[3]{k_e R_S}$ . In this case we shall use a different method to determine the contributions to the smooth part of the counting function.

To approximate  $D_m^{(e)}(\varepsilon)$  in the intermediate range  $|m - k_e R_S| < \sqrt[3]{k_e R_S}$ , one may apply uniform asymptotic expansions for Bessel functions. Semiclassically, these states are related to the diffraction of the electron in the penumbra of a circular annulus billiard.<sup>37</sup> As a first estimate of the smooth part of the counting function, the energy levels in the intermediate range will be taken into account by the corresponding energy levels of the circular billiard of radius  $R_N$  using Debye's asymptotic expansion. As we shall see, this is a good approximation since the intermediate range of  $m$  is rather narrow compared to the two other ranges discussed above. Using the uniform expansions of the Bessel functions it is straightforward (although algebraically rather tedious) to include the intermediate range more accurately.

Similarly, the method of approximations outlined above can be used to approximate  $D_m^{(h)}(\varepsilon)$  in the secular equation (9). One finds that  $\Phi_m(-\varepsilon)$  is still given by Eq. (A1) after

replacing  $k_e$  by the wave number  $k_h(\varepsilon) = k_e(-\varepsilon)$  of the holelike quasiparticles in the  $N$  region. However, in this case the approximation for  $\Phi_m(-\varepsilon)$  is valid only for  $|m| < k_h R_S - \sqrt[3]{k_h R_S}$ . Finally, from Eq. (A1) one obtains

$$\begin{aligned} \Phi_m(\varepsilon) - \Phi_m(-\varepsilon) = & 2[\vartheta_m(k_e R_N) - \vartheta_m(k_e R_S)] \\ & - 2[\vartheta_m(k_h R_N) - \vartheta_m(k_h R_S)] \end{aligned} \quad (\text{A4})$$

for  $|m| < k_h R_S - \sqrt[3]{k_h R_S}$ . Note that the last three terms in Eq. (A1) cancel out in the difference  $\Phi_m(\varepsilon) - \Phi_m(-\varepsilon)$ .

In a similar way, it can be seen that the zeros of  $D_m^{(h)}(\varepsilon)$  in Eq. (9) coincide with the zeros of  $J_m(k_h R_N)$  for  $|m| > k_h R_S + \sqrt[3]{k_h R_S}$  to a good approximation. Thus, the secular equation (9), determining the energy levels of the  $NS$  disk system, reduces to

$$J_m(k_e R_N) J_m(k_h R_N) = 0 \quad (\text{A5})$$

for  $|m| > k_e R_S + \sqrt[3]{k_e R_S}$ . The energy levels are the same as those of the disk of radius  $R_N$  for electron/holelike quasiparticles. Semiclassically, the states with angular momentum quantum number  $|m| > k_e R_S + \sqrt[3]{k_e R_S}$  are called whispering gallery modes. The wave functions for these states at  $r = R_S$  are negligible, and so no Andreev reflections take place. This is why the superconducting pair potential  $\Delta_0$  does not appear in Eq. (A5).

- <sup>1</sup>A.F. Andreev, Zh. Éksp. Teor. Fiz. **46**, 1823 (1964) [Sov. Phys. JETP, **19**, 1228 (1964)].
- <sup>2</sup>C.J. Lambert and R. Raimondi, J. Phys.: Condens. Matter **10**, 901 (1998).
- <sup>3</sup>C. W. J. Beenakker, in *Mesoscopic Physics, Les Houches Summer School*, edited by E. Akkermans, G. Montambaux, J. L. Pichard, and J. Zinn-Justin (Elsevier Science, Amsterdam, 1995).
- <sup>4</sup>*Mesoscopic Electron Transport*, edited by L. L. Sohn, L. P. Kouwenhoven, and G. Schön (Kluwer Academic Publishers, Dordrecht, The Netherlands, 1996).
- <sup>5</sup>C.W.J. Beenakker, Rev. Mod. Phys. **69**, 731 (1997).
- <sup>6</sup>P. G. de Gennes, *Superconductivity of Metals and Alloys* (Benjamin, New York, 1996).
- <sup>7</sup>P.G. de Gennes and D. Saint-James, Phys. Lett. **4**, 151 (1963).
- <sup>8</sup>I. Kosztin, D.L. Maslov, and P.M. Goldbart, Phys. Rev. Lett. **75**, 1735 (1995).
- <sup>9</sup>J.A. Melsen, P.W. Brouwer, K.M. Frahm, and C.W.J. Beenakker, Europhys. Lett. **35**, 7 (1996).
- <sup>10</sup>J.A. Melsen, P.W. Brouwer, K.M. Frahm, and C.W.J. Beenakker, Phys. Scr., T **69**, 223 (1997).
- <sup>11</sup>A. Altland and M.R. Zirnbauer, Phys. Rev. Lett. **76**, 3420 (1996).
- <sup>12</sup>G.B. Lesovik, A.L. Fauchère, and G. Blatter, Phys. Rev. B **55**, 3146 (1997).
- <sup>13</sup>A. Lodder and Y.V. Nazarov, Phys. Rev. B **58**, 5783 (1998).
- <sup>14</sup>H. Schomerus and C.W.J. Beenakker, Phys. Rev. Lett. **82**, 2951 (1999).
- <sup>15</sup>W. Ihra, M. Leadbeater, J.L. Vega, and K. Richter, Eur. Phys. J. B **21**, 425 (2001).
- <sup>16</sup>I.O. Kulik, Zh. Éksp. Teor. Fiz. **57**, 1745 (1969) [Sov. Phys. JETP **30**, 944 (1970)]; J. Bardeen and J.L. Johnson, Phys. Rev. B **5**, 72 (1972).
- <sup>17</sup>C.W.J. Beenakker, Phys. Rev. Lett. **67**, 3836 (1991); **68**, 1442(E) (1992); C. W. J. Beenakker, in *Transport Phenomena in Mesoscopic Systems*, edited by H. Fukuyama and T. Ando (Springer-Verlag, Berlin, 1992), pp. 235–253.
- <sup>18</sup>P.F. Bagwell, Phys. Rev. B **46**, 12 573 (1992).
- <sup>19</sup>H. Hoppe, U. Zülicke, and G. Schön, Phys. Rev. Lett. **84**, 1804 (2000).
- <sup>20</sup>O. Šipr and B.L. Györfy, J. Phys.: Condens. Matter **8**, 169 (1996); O. Šipr and B.L. Györfy, J. Low Temp. Phys. **106**, 315 (1997).
- <sup>21</sup>H. Weyl, Nachr. Akad. Wiss., 110 (1911).
- <sup>22</sup>M. Brack and R. K. Bhaduri, in *Semiclassical Physics*, edited by D. Pines (Addison-Wesley, Amsterdam, 1997).
- <sup>23</sup>C. Bruder and Y. Imry, Phys. Rev. Lett. **80**, 5782 (1998).
- <sup>24</sup>P. Visani, A.C. Mota, and A. Pollini, Phys. Rev. Lett. **65**, 1514 (1990); A.C. Mota, P. Visani, A. Pollini, and K. Aupke, Physica B **197**, 95 (1994); A.C. Mota, P. Visani, A. Pollini, and K. Aupke, *ibid.* **197**, 95 (1994); F.B. Müller-Allinger and A.C. Mota, Phys. Rev. Lett. **84**, 3161 (2000).
- <sup>25</sup>G.E. Blonder, M. Tinkham, and T.M. Klapwijk, Phys. Rev. B **25**, 4515 (1982).
- <sup>26</sup>W.L. McMillan, Phys. Rev. **175**, 559 (1968); G. Kieselmann, Phys. Rev. B **35**, 6762 (1987).
- <sup>27</sup>H. Plehn, O.-J. Wacker, and R. Kümmel, Phys. Rev. B **49**, 12 140 (1994).

- <sup>28</sup>Y. Asano and T. Kato, J. Phys. Soc. Jpn. **69**, 1125 (2000).
- <sup>29</sup>J. Cserti, A. Bodor, J. Koltai, and G. Vattay, cond-mat/0105472 (unpublished).
- <sup>30</sup>E. Doron and U. Smilansky, Phys. Rev. Lett. **68**, 1255 (1992).
- <sup>31</sup>U. Smilansky, in *Mesoscopic Quantum Physics, Les Houches Summer School*, edited by E. Akkermans, G. Montambaux, J. L. Pichard, and J. Zinn-Justin (Elsevier Science, Amsterdam, The Netherlands, 1995).
- <sup>32</sup>M. Berry, in *Chaos and Quantum Physics, Les Houches Summer School*, edited by M.-J. Giannoni, A. Voros, and J. Zinn-Justin (Elsevier Science, Amsterdam, The Netherlands, 1991).
- <sup>33</sup>Z. Kaufmann (unpublished).
- <sup>34</sup>W. Ihra and K. Richter, Physica E (Amsterdam) **9**, 362 (2001).
- <sup>35</sup>N.A. Mortensen, K. Flensberg, and A.-P. Jauho, Phys. Rev. B **59**, 10176 (1999).
- <sup>36</sup>M. Abramowitz and I. Stegun, *Handbook of Mathematical Functions*, 9th ed. (Dover Publication, New York, NY, 1972).
- <sup>37</sup>H. Primack, H. Schanz, U. Smilansky, and I. Ussishkin, Phys. Rev. Lett. **76**, 1615 (1996); N.C. Snaith and D.A. Goodings, Phys. Rev. E **55**, 5212 (1997).