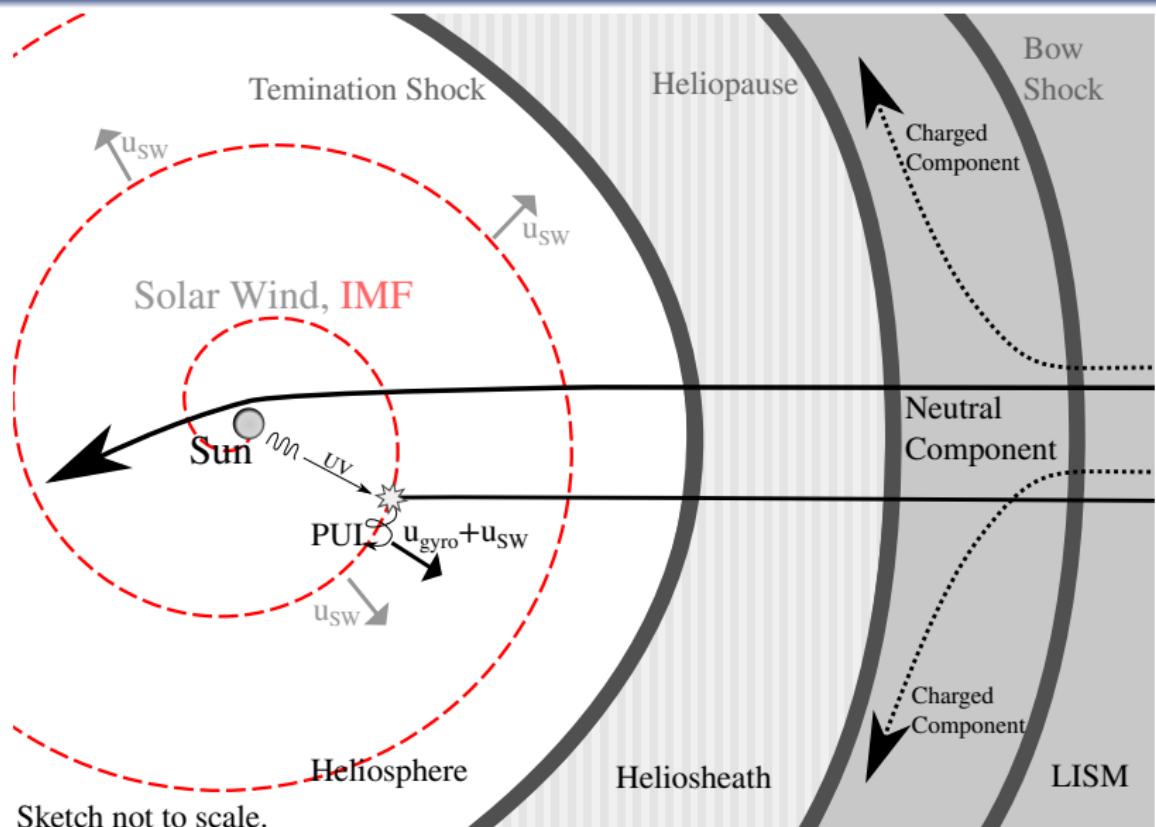


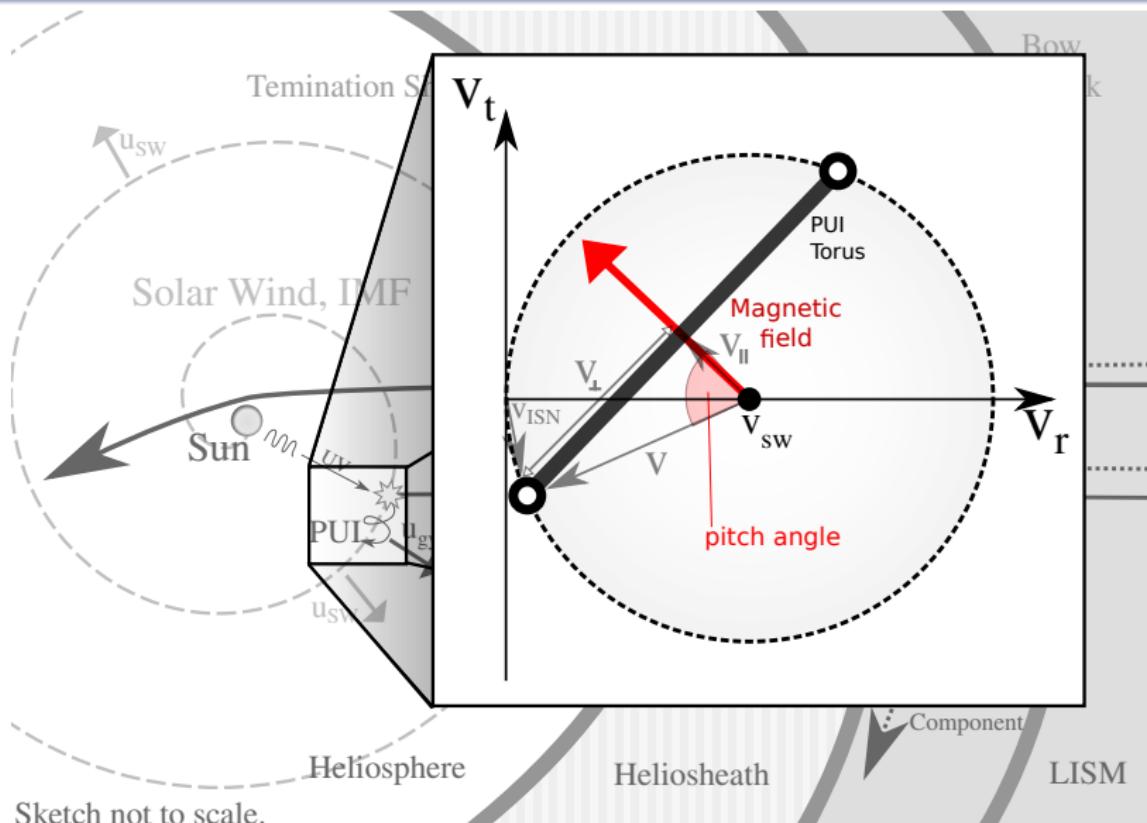
# Kinetic simulations of PUI transport and pitchangle-scattering during turbulent conditions of the solar wind

Duncan Keilbach

## Pickup ions (PUIs)

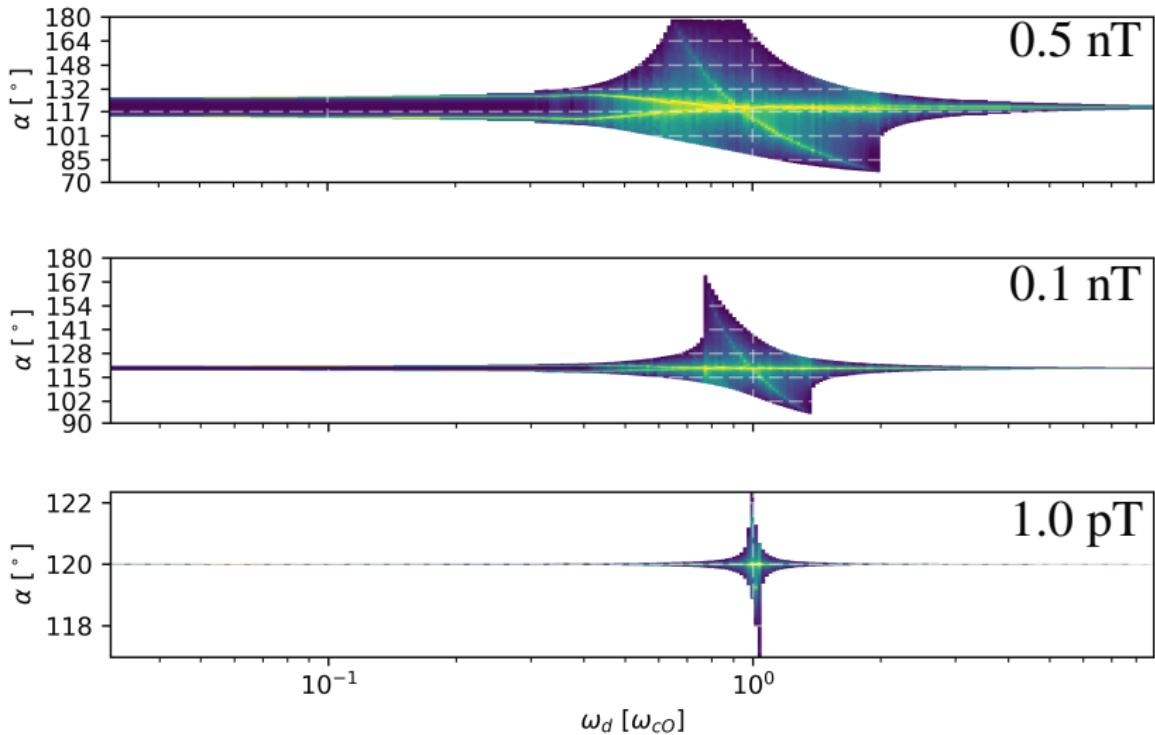


## Pickup ions (PUIs)



Sketch not to scale.

## Wave-modified pitch angle distributions



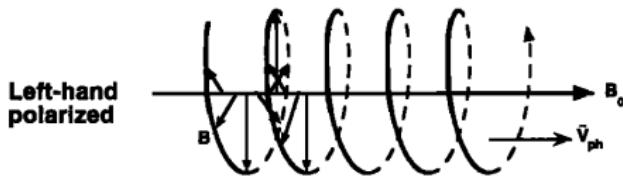
## 1 Alfvenic Waves

## 2 Simulation algorithm and setup

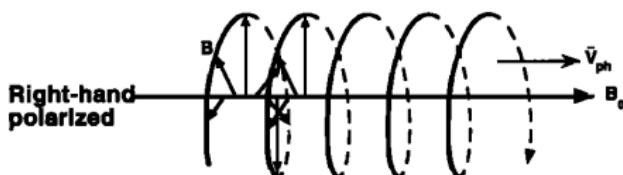
## 3 Results

- Interaction with continuous, monofrequent waves
- Interaction with intermittent waves

## Electromagnetic Wave Polarizations



Left-hand polarized  
ion cyclotron (high  $\omega$ )  
Alfvén mode (low  $\omega$ )



Right-hand polarized  
whistler mode (high  $\omega$ )  
magnetosonic mode (low  $\omega$ )

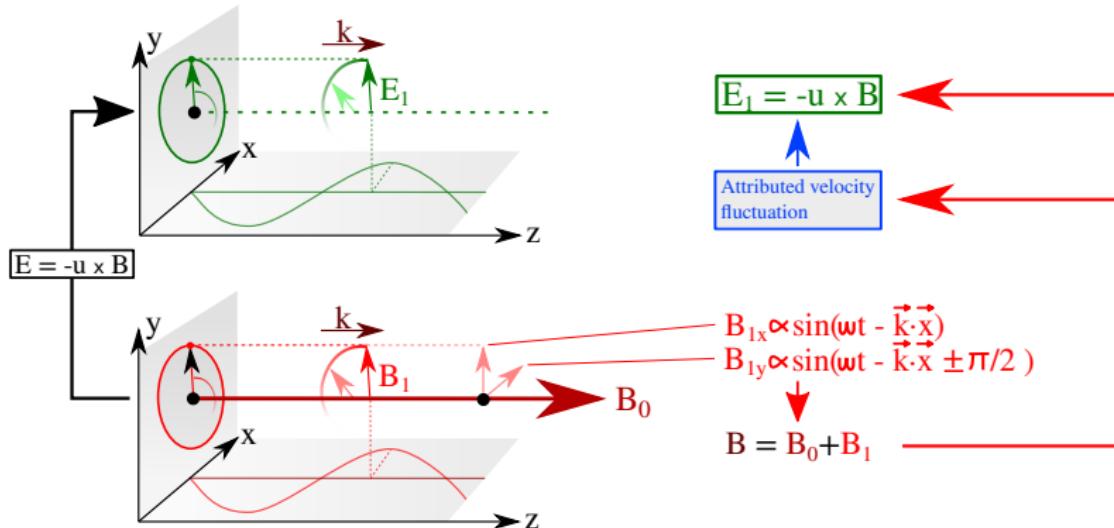
- Circular polarized wave along  $\vec{B}$

$$\rightarrow v_A = \frac{B}{\sqrt{\mu_0 \rho}}$$

$$\bullet \delta \vec{u} = \pm \frac{\delta \vec{B}}{\sqrt{\mu_0 \rho}}$$

$$\bullet \text{Induction: } \vec{E} = -\vec{u} \times \vec{B}$$

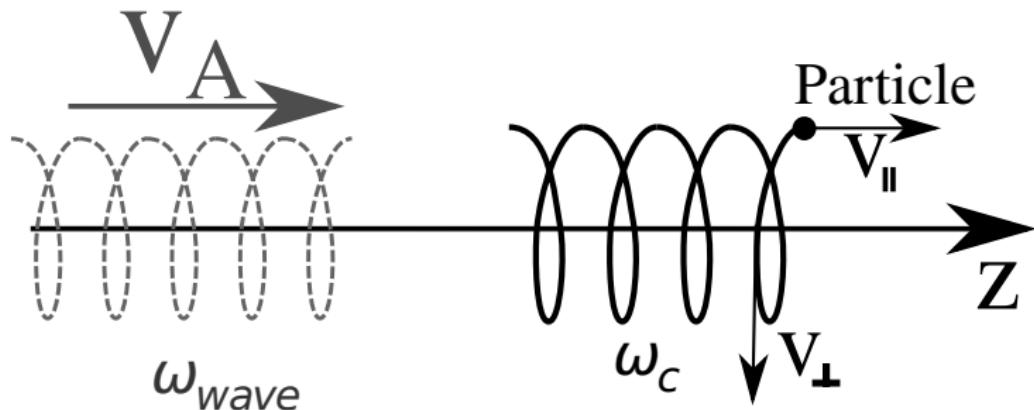
Full MHD-simulation is beyond the scope of this work.  
 ⇒ Analytical expressions for the waves



Dispersion relation:

$$\left( \frac{\omega}{\Omega_p} \right)^2 = (kl_p)^2 + \frac{1}{2}(kl_p)^4 - \frac{1}{2}\sqrt{(kl_p)^2 + 4}$$
,  $l_p = \frac{c}{\omega_p}$

# General principles



**Resonance condition:**  $n \cdot \omega_c = \omega_d := \omega_{wave} - \vec{k} \cdot \vec{v}_{||}$

$$\varepsilon := (v_{||} - v_{ph})^2 + v_{\perp}^2 \text{ is conserved}$$

## 1 Alfvenic Waves

## 2 Simulation algorithm and setup

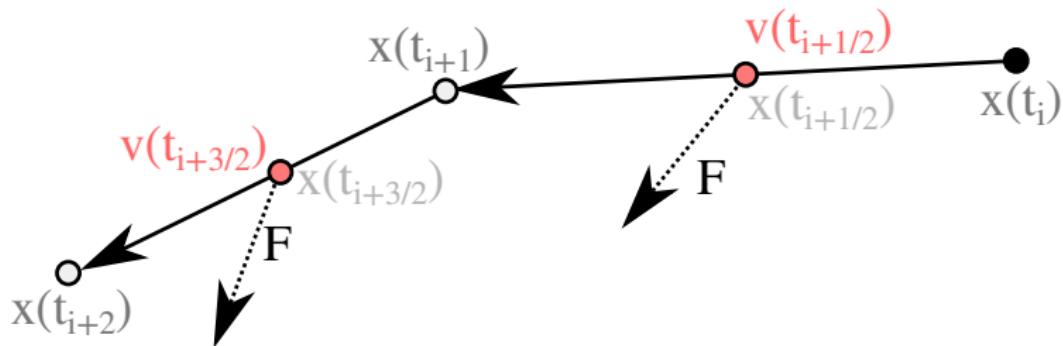
## 3 Results

- Interaction with continuous, monofrequent waves
- Interaction with intermittent waves

## Implementation of the PUI simulation

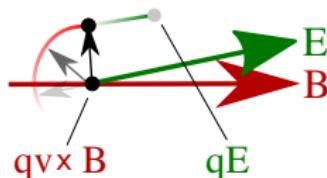
## Timestep-Algorithm (Leapfrog)

$$\frac{\Delta \vec{v}}{\Delta t} \approx \vec{a} = \frac{q}{m} (\vec{E} + \vec{v} \times \vec{B})$$



## Optimization (Boris-Scheme)

- $q\vec{E}$ : Accelleration along  $\vec{E}$
- $q\vec{v} \times \vec{B}$ : Rotation around  $\vec{B}$



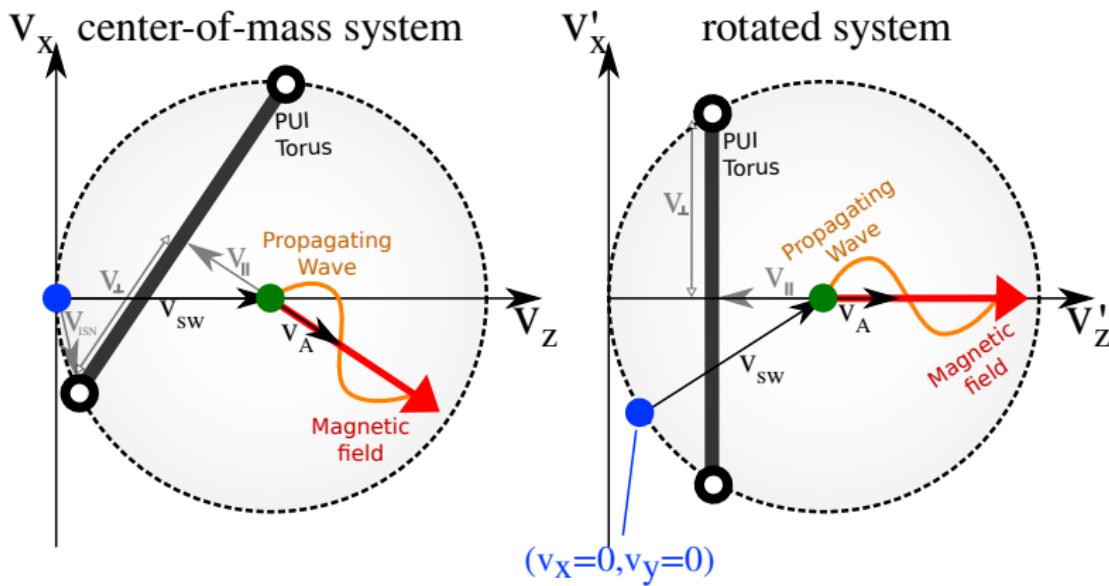
Rotation around  $\vec{B}$

Accelleration along  $\vec{E}$

- ① Add half electric impulse
- ② Rotate around magnetic field (angle:  $\omega_c \cdot \Delta t$ )
- ③ Add another half electric impulse

# Particle coordinate system

## Rotated solar wind bulk system



1 Alfvenic Waves

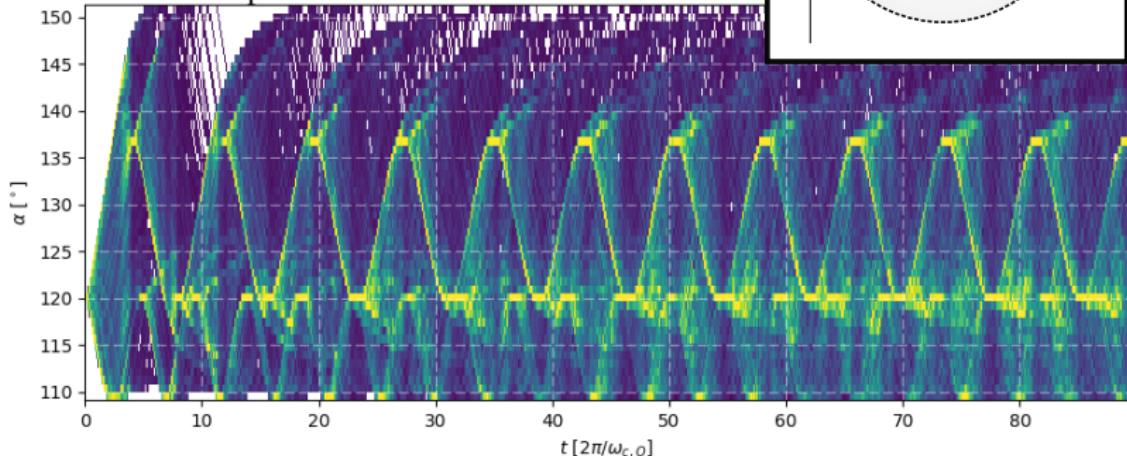
2 Simulation algorithm and setup

3 Results

- Interaction with continuous, monofrequent waves
- Interaction with intermittent waves

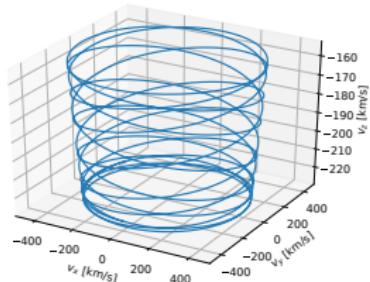
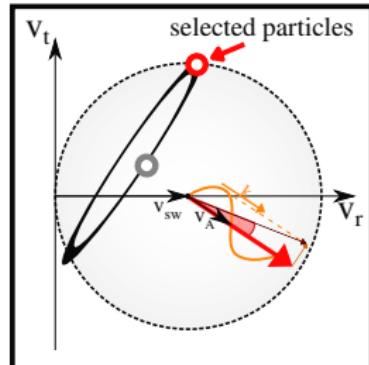
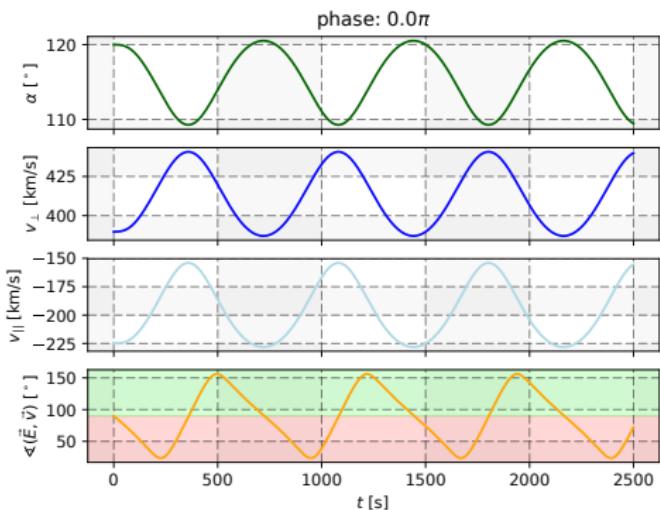
# Boundary Conditions

- Background field: 5 nT
- Wave amplitude: 0.1nT
- Wave frequency: 0.28 of oxygen gyro frequency
- Simulated time: 25 wave cycles
- 500 test particles, initial pitch angle: 120°
- solar wind speed: 450 km/s



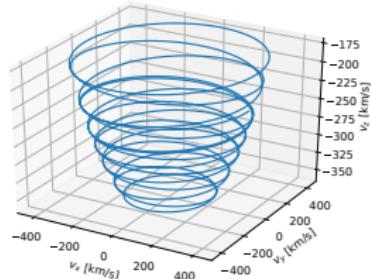
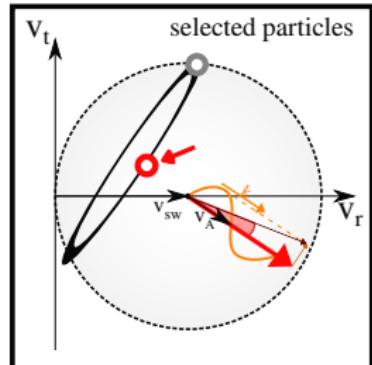
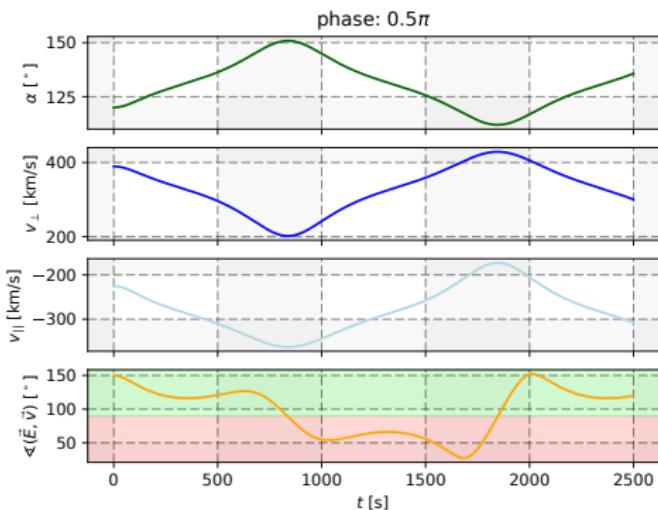
# Boundary Conditions

-2 test particles, initial pitch angle:  $120^\circ$



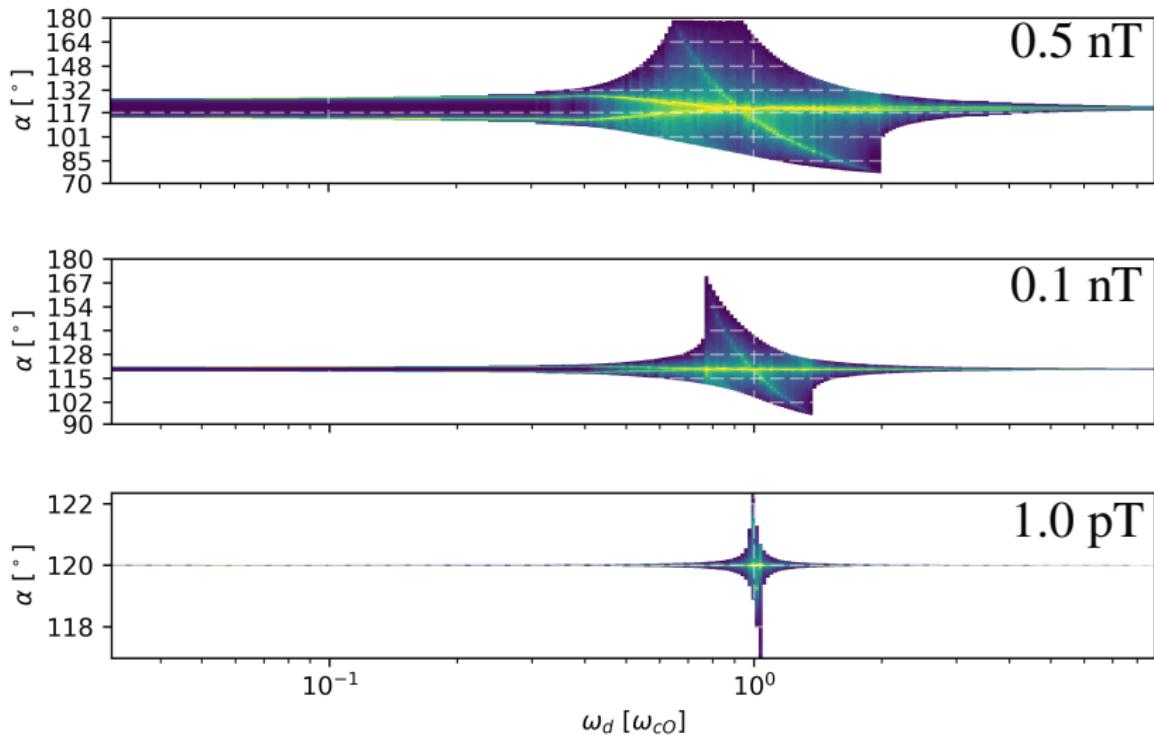
# Boundary Conditions

-2 test particles, initial pitch angle: 120°



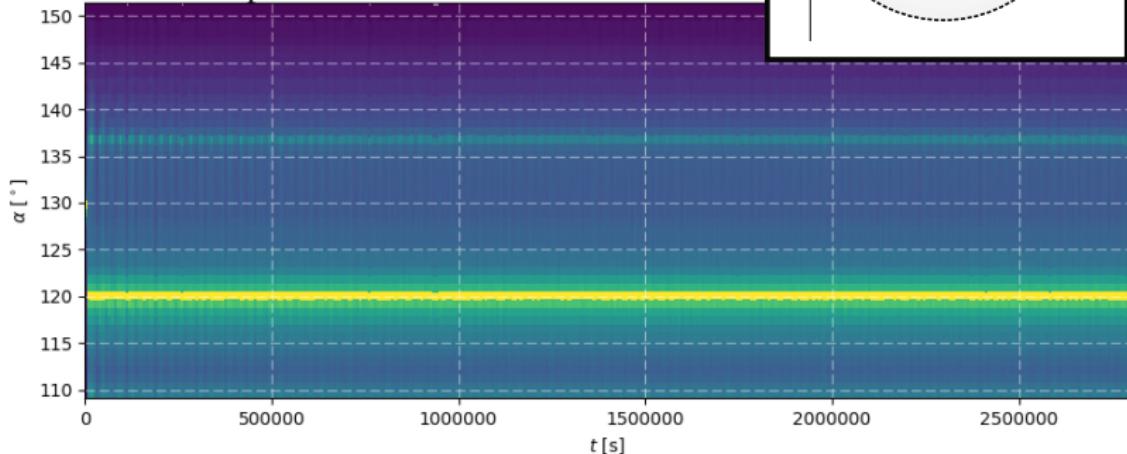
## Results

### Interaction with continuous, monofrequent waves



# Boundary Conditions

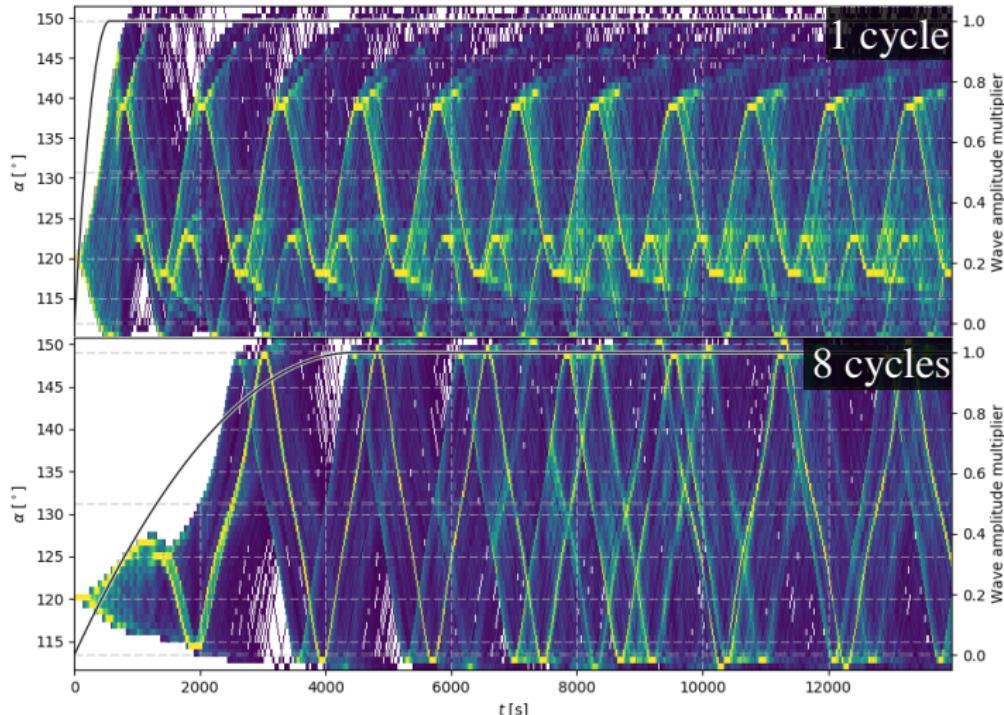
- Background field: 5 nT
- Wave amplitude: 0.1nT
- Wave frequency: 0.28 of oxygen gyro frequency
- Simulated time: 1000 wave cycles
- 500 test particles, initial pitch angle: 120°
- solar wind speed: 450 km/s



## Results

### Interaction with intermittent waves

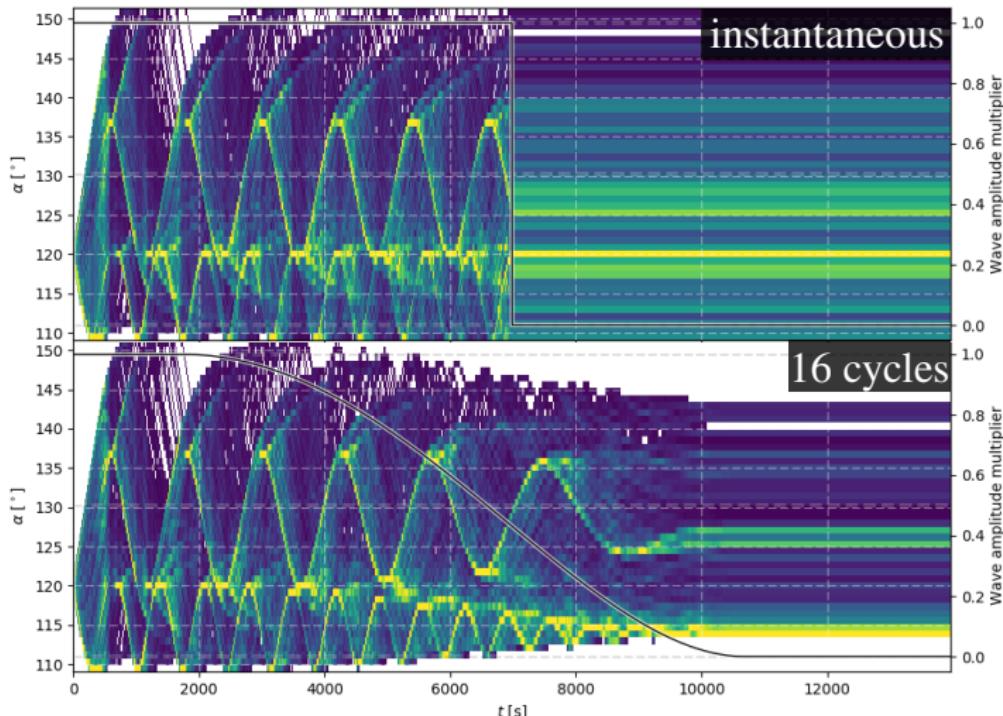
# Turning waves on



## Results

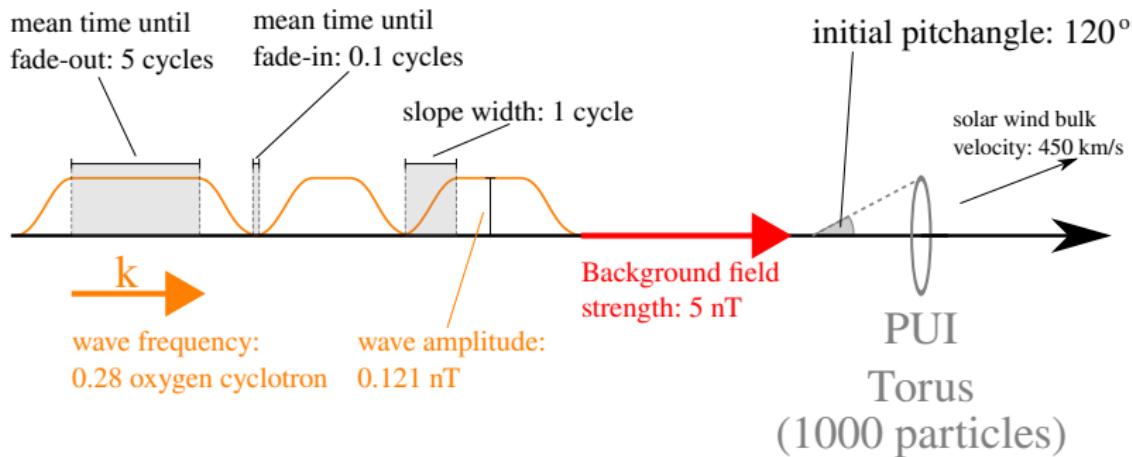
### Interaction with intermittent waves

# Turning waves off



# Turning waves on and off randomly

## Boundary conditions:

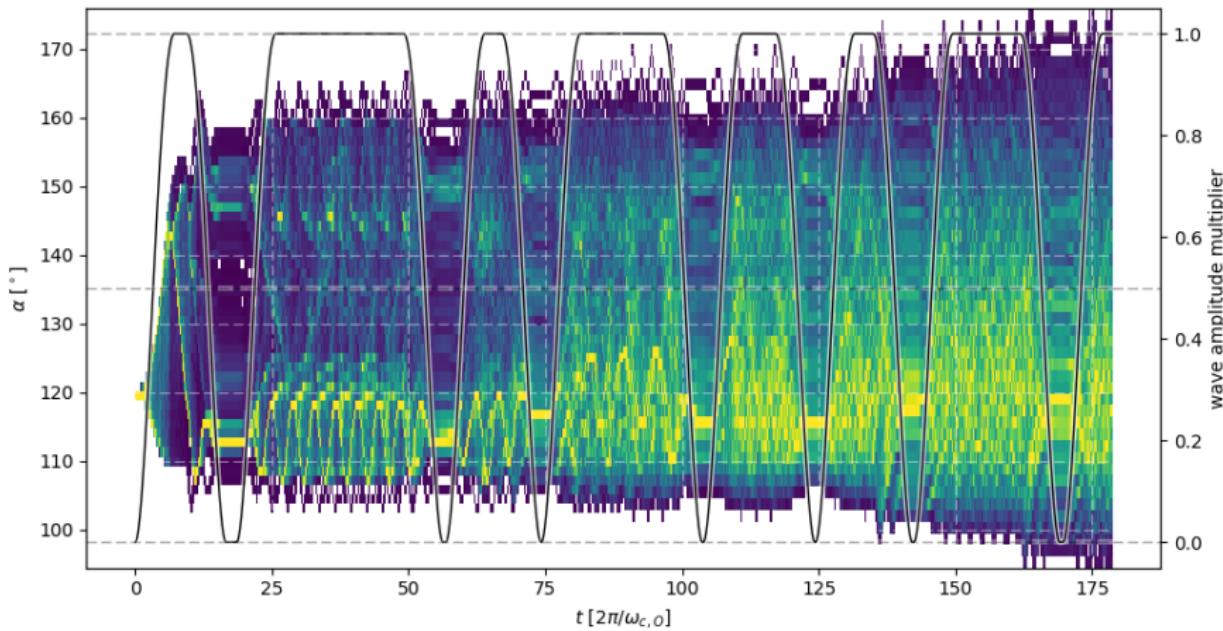


200 consecutive runs of 25 wave cycles each.

## Results

### Interaction with intermittent waves

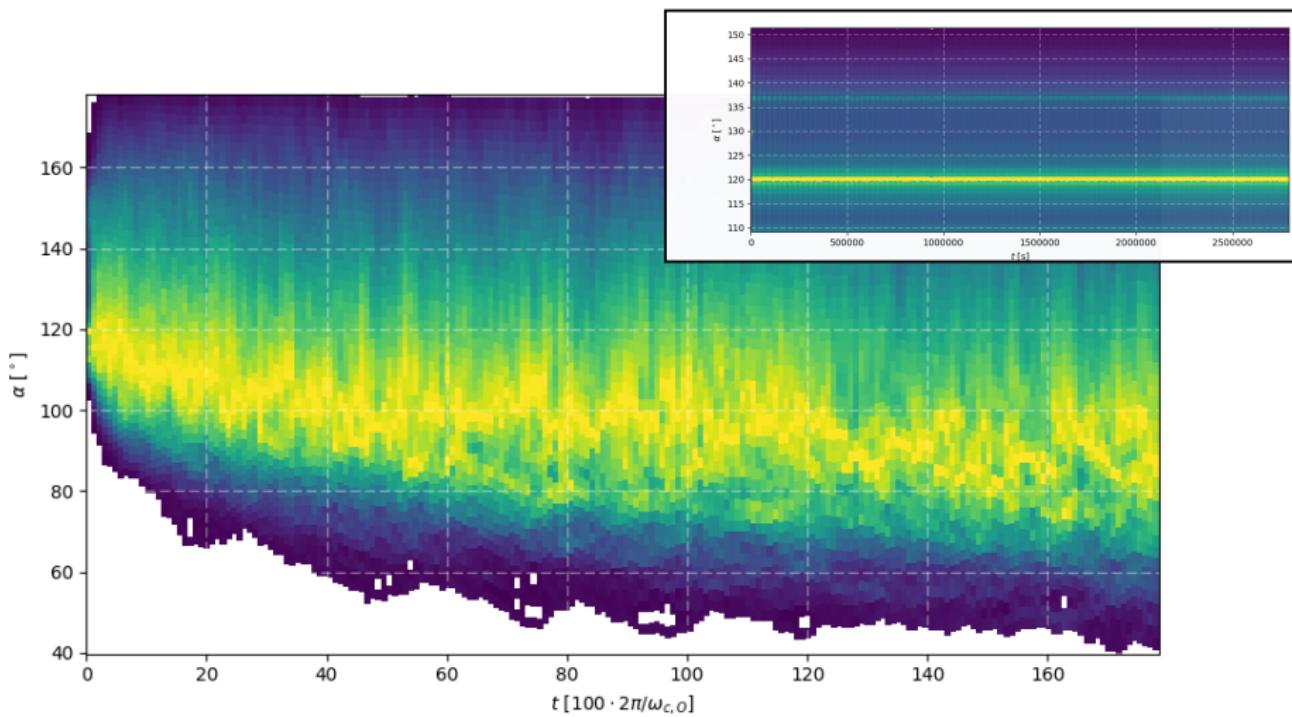
# Turning waves on and off randomly - first 50 cycles



## Results

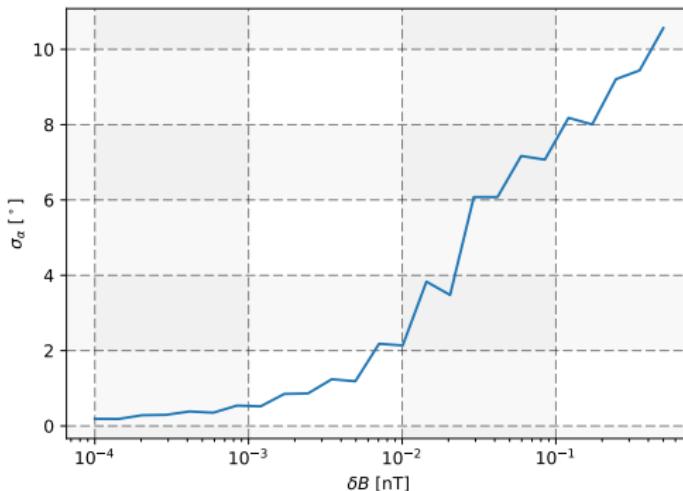
### Interaction with intermittent waves

# Turning waves on and off randomly - full simulation



# Dependency on the wave amplitude

Standard deviation of the pitch angle distribution after 5000 wave cycles

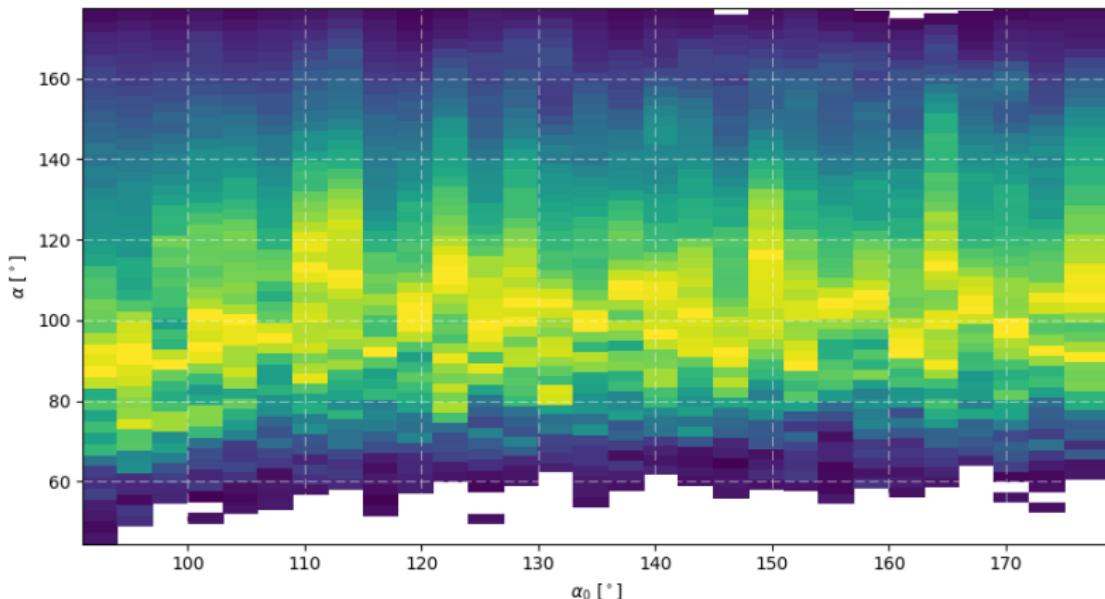


## Results

### Interaction with intermittent waves

# Dependency on the initial pitch angle

Pitch angle distribution after 2500 wave cycles

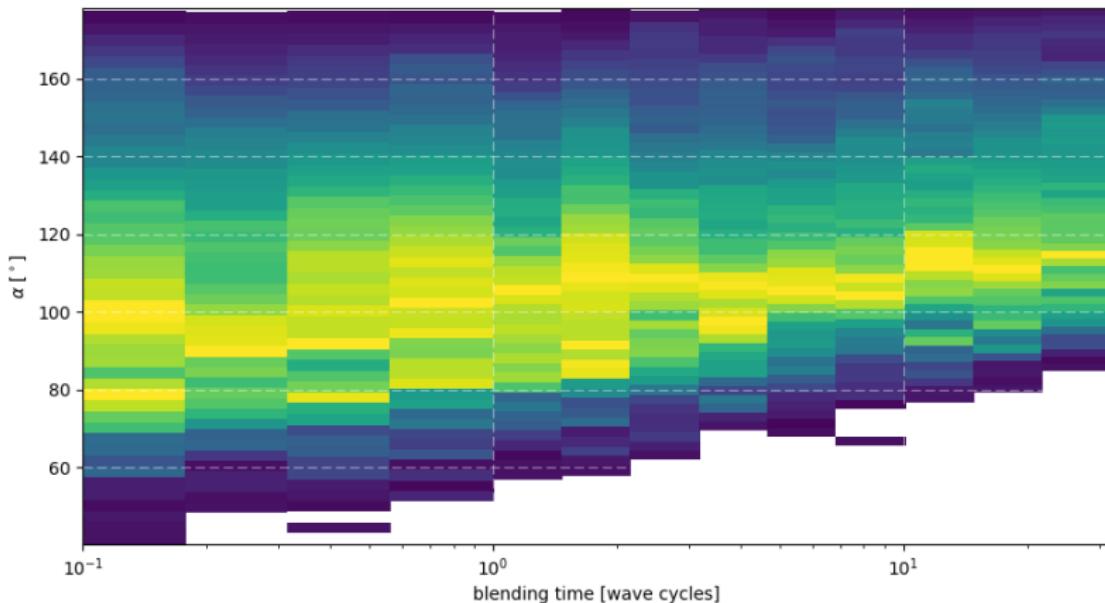


## Results

### Interaction with intermittent waves

# Dependency on the fading - time

Pitch angle distribution after 2500 wave cycles

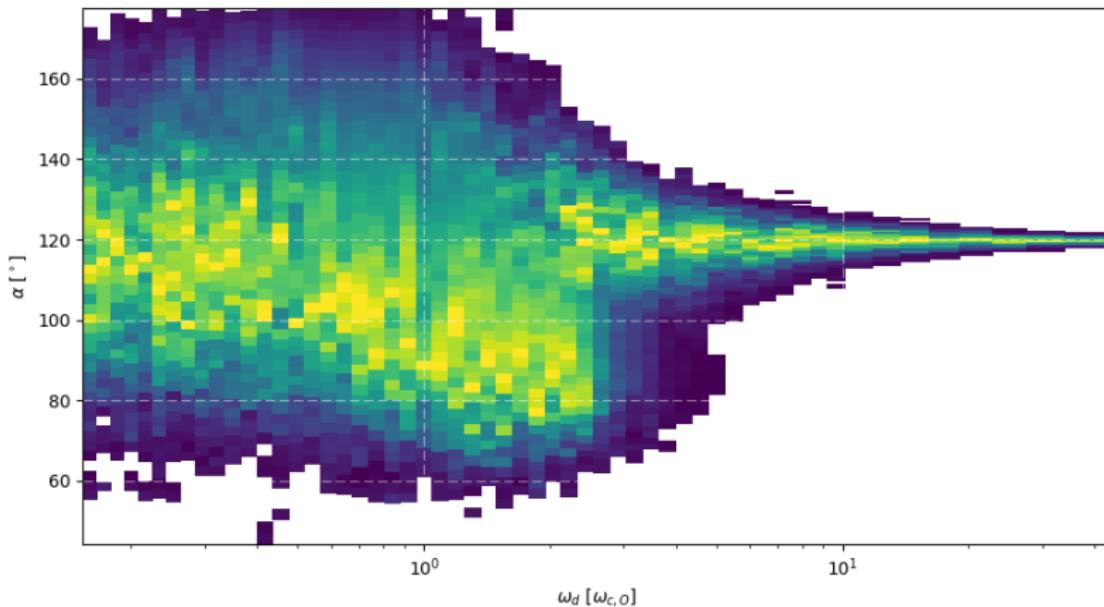


## Results

### Interaction with intermittent waves

# Dependency on the wave frequency

Pitch angle distribution after 2500 wave cycles



# Conclusions and Outlook

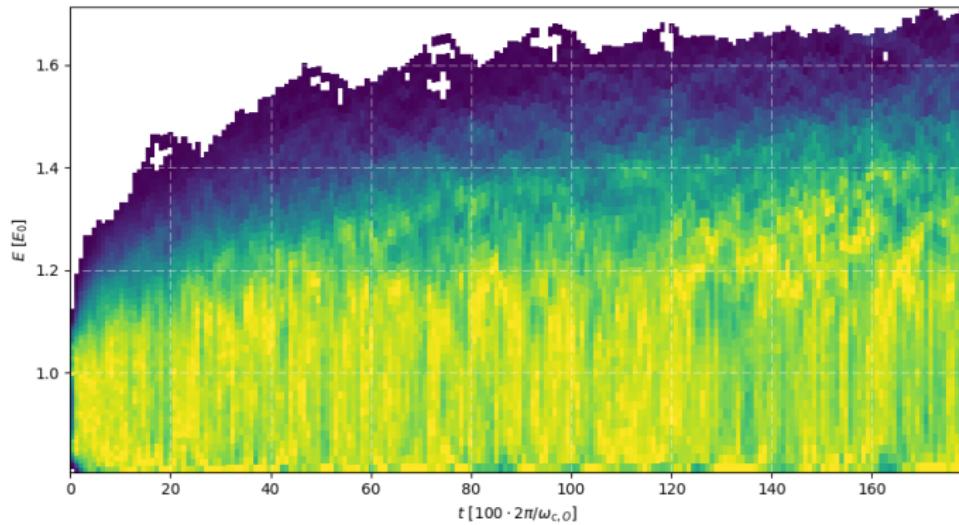
A numerical study to understand wave particle interactions qualitatively on microscopic scales has been performed.

- **Continuous, monofrequent wave fields...**
  - cause stationary oscillations in pitch angle.
  - do not cause the distribution to become broader over time.
- **Intermittent, monofrequent wave fields...**
  - freeze and restore the oscillations at random times
  - cause a systematic broadening of the pitch angle distribution over time

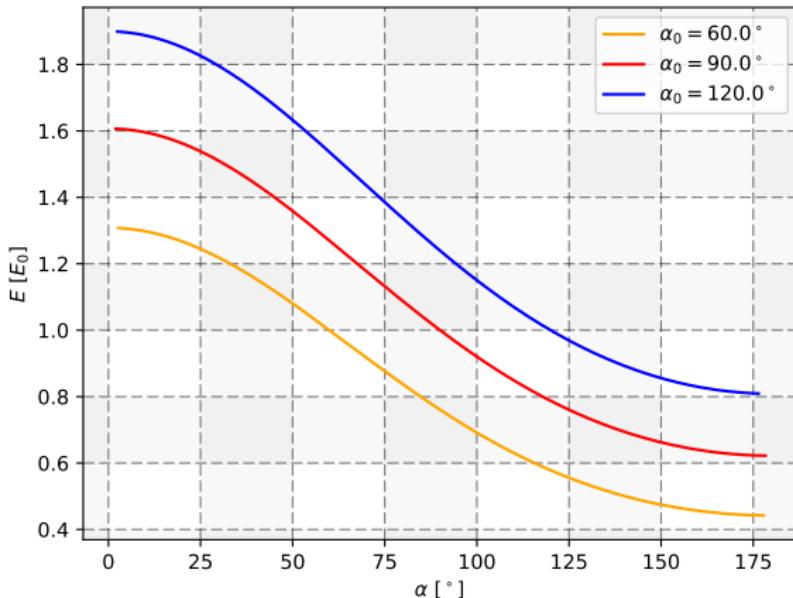
## Outlook

- Gradient and curvature in the background field
- Interactions with polyfrequent waves
- Modification by collisions with solar wind particles
- Amplitude modulation

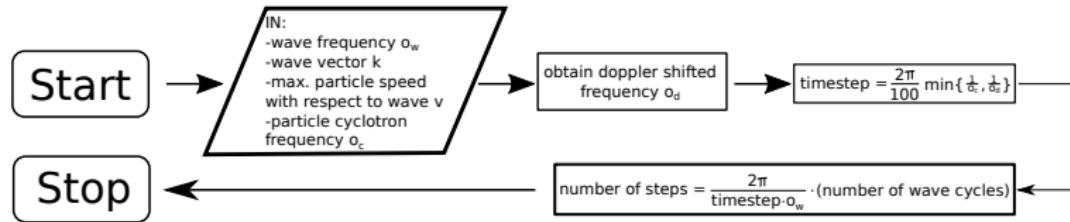
# Particle energy distribution



# Energy as a function of the pitch angle



# Choice of time step



## Basic idea:

At least 100 points per gyro period or particle-sampled wave period.

The shortest of these time scales is considered for the time step.

# Time derivatives of $v_{||}$ and $v_{\perp}$

From Lorentz-Force:

$$\dot{\vec{v}} = \frac{q}{m} \left( \delta \vec{E} + \vec{v} \times (\vec{B}_0 + \delta \vec{B}) \right) \quad (1)$$

$$\dot{\vec{v}} = \frac{q}{m} \delta \vec{E} + \frac{q}{m} \vec{v} \times \vec{B}_0 + \frac{q}{m} \vec{v} \times \delta \vec{B}. \quad (2)$$

Obtain the z-Component:

$$\dot{v}_{||} = \frac{q}{m} v_{\perp} \delta B \sin(\Delta\varphi). \quad (3)$$

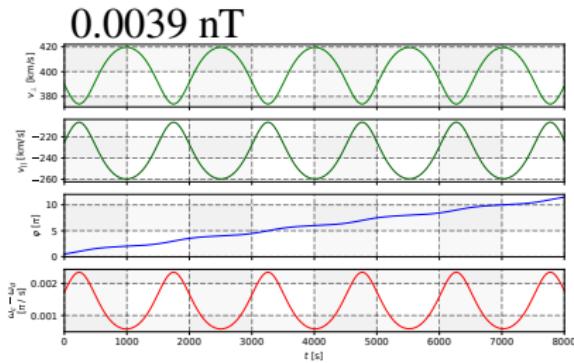
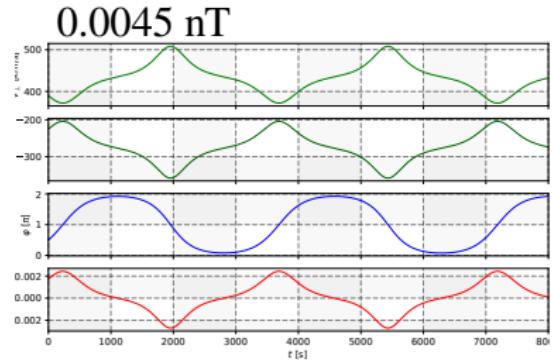
Conservation of energy in wave frame:

$$\dot{v}_{\perp} = \dot{v}_{||} \frac{v_{||} - v_{\varphi}}{v_{\perp}} \quad (4)$$

Derivative of  $v_{\perp}$ :

$$\dot{v}_{\perp} = \frac{q \delta B}{m} \sin(\Delta\varphi) (v_{||} - v_{\varphi}) \quad (5)$$

# Euler-integration



# Overtaking-criterion

Criterion:

$k|v_{||}| > |\omega_c - \omega|$ , before  $\Delta\varphi$  reaches  $2\pi$ .

Linearization:  $v_{||} = v_{||,0} + \vartheta \cdot t$  with

$$\vartheta := \dot{v}_{||}(t=0) = \frac{q\delta B}{m} \sqrt{\varepsilon - (v_{||} - v_\varphi)^2} \quad (6)$$

Change in phase difference:

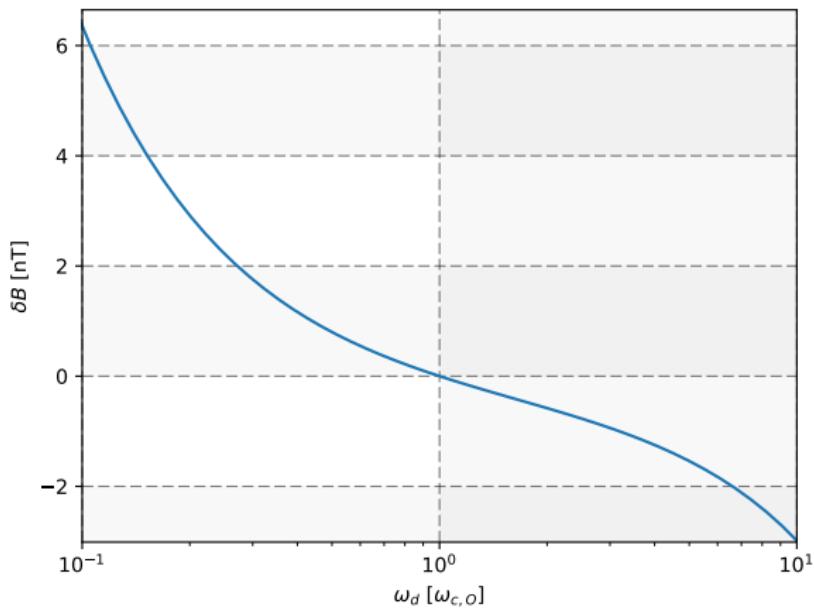
$$\frac{d}{dt}\varphi = \omega_c - \omega = kv_{||,0} + k\vartheta t. \quad (7)$$

Result: Required amplitude for overtaking:

$$\delta B = \frac{m}{q} \cdot \frac{\left(\frac{\omega_c - \omega}{k} - v_{||,0}\right) \cdot \frac{\omega_c - \omega + kv_{||,0}}{2\pi}}{\sqrt{\varepsilon - (v_{||,0} - v_\varphi)^2}} \quad (8)$$

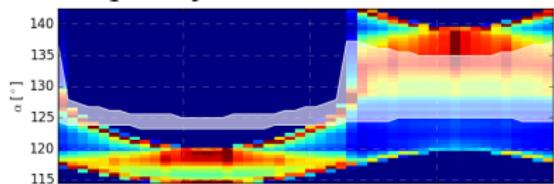
# Overtaking-criterion

Required amplitude for overtaking

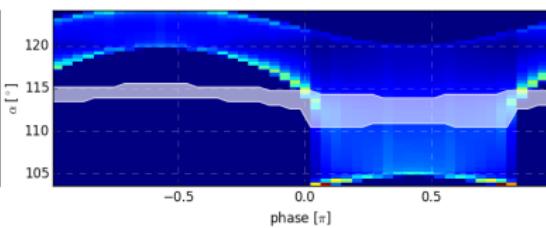
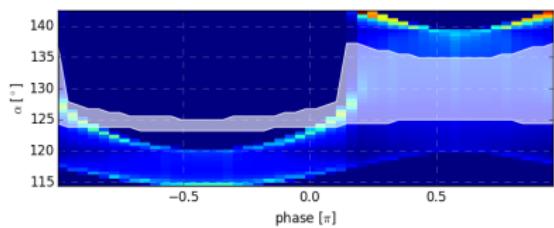
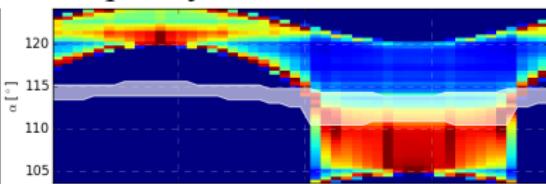


# Time-averaged spectra of the pitch angle as a function of $\Delta\varphi$

frequency below first resonance

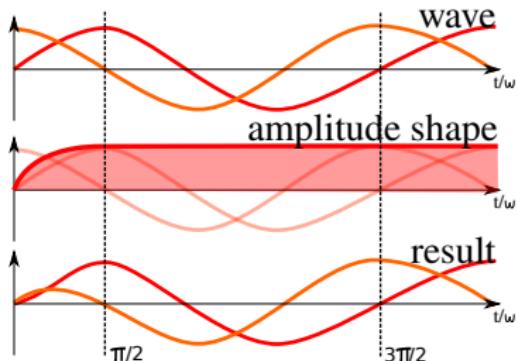


frequency above first resonance



# Implementation of the fields - Remarks

- **Coupling with electrical field:**  $\vec{E} = -\vec{v} \times \vec{B}$ ,  $\delta v = \pm \frac{\delta B}{\sqrt{\mu_0 \rho}}$
- **Fade-in:** Wave faded in with a parabolic shape over the first quarter period (for continuous waves).
- Parabolic or fourth-order shape functions have been used to model intermittent waves.



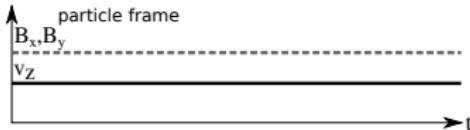
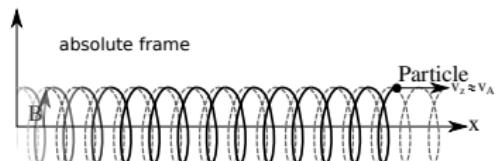
**shape functions:**

$$f_1(x) = 1 - \left( \left( \frac{x}{a} \right) \right)^2 \quad (9)$$

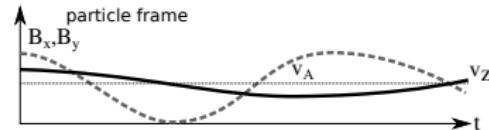
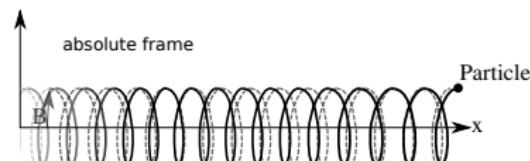
$$f_2(x) = \left( 1 - \left( \frac{x}{a} \right)^2 \right)^2 \quad (10)$$

# $0^{\text{th}}$ order resonance

without interaction



with interaction



Condition:  $v_z \approx v_A$   
 Particle is „surfing“on the wave.  
 $v_z$ : Synchronization with  $v_A$ .

# General theoretical considerations

- **initial parallel velocity**

from resonance condition

$$v_{||} = v_{ph} - \frac{n\omega_c}{k_{||}} \quad (11)$$

- **The energy in the waveframe is conserved**

(absolute) energy gain:  $\Delta E = h\omega/2\pi$

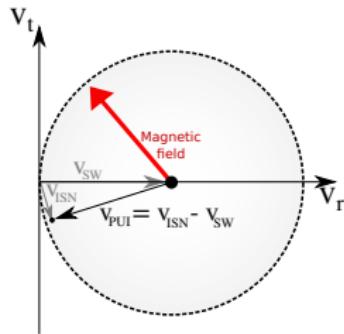
$$\text{const.} = \frac{m}{2} \left( (v_{||} - v_{ph})^2 + v_{\perp}^2 \right) \quad (12)$$

Implications:

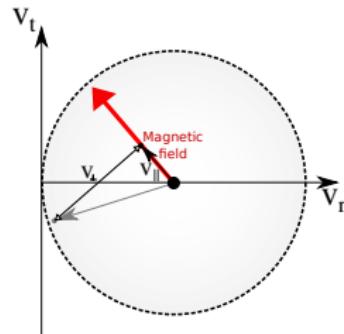
- Parameter for "goodness" of simulation:  
 $\varepsilon := (v_{||} - v_{ph})^2 + v_{\perp}^2$
- Change in parallel velocity difference causes perpendicular acceleration/deceleration.

# Sampling of torus distributions

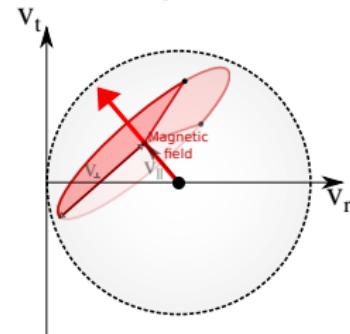
Translation into solar wind system



Computation of parallel and perpendicular speed

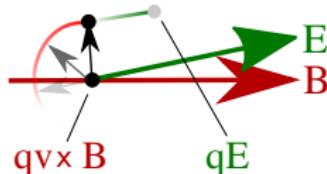


Rotation about the local magnetic field



## Optimization (Boris-Scheme)

- $q\vec{E}$ : Accelleration along  $\vec{E}$
- $q\vec{v} \times \vec{B}$ : Rotation around  $\vec{B}$



**Rotation around  $\vec{B}$**

**Accelleration along  $\vec{E}$**

- ① Add half electric impulse to  $\vec{v}_{old}$ :  $\vec{v}^- := \vec{v}_{old} + q\Delta t \vec{E}/2m$ .
- ② Rotate  $\vec{v}^-$  around  $\vec{B}$  for  $\vec{v}^+$ :  $\vec{v}^+ := \vec{v}^- + \vec{v}' \times \vec{s}$ , where  $\vec{s} = 2\vec{t}/(1 + \vec{t}^2)$  and  $\vec{t} = q\vec{B}\Delta t/2m$ .
- ③ Add the other half electric impulse to get  $\vec{v}_{new}$ :  $\vec{v}_{new} := \vec{v}^+ + q\Delta t \vec{E}/2m$

# Correllation between $\delta \vec{u}$ and $\delta \vec{B}$

Equation of motion and equation of induction:

$$\rho_m \frac{\partial \vec{v}_m}{\partial t} = \vec{j} \times \vec{B} \quad (13)$$

$$\frac{\partial \vec{B}}{\partial t} = \nabla \times (\nabla \times \vec{B}). \quad (14)$$

Assumptions:

$$\vec{B} = \vec{B}_0 + \vec{B}_1, \vec{v}_m = \vec{v}_0 + \vec{v}_1 \quad (15)$$

$$\vec{B}_0 = (0, 0, B_0), \vec{B}_1 = (B_{1x}, 0, 0), \vec{v}_0 = 0, \vec{v}_1 = (v_{1x}, 0, 0) \quad (16)$$

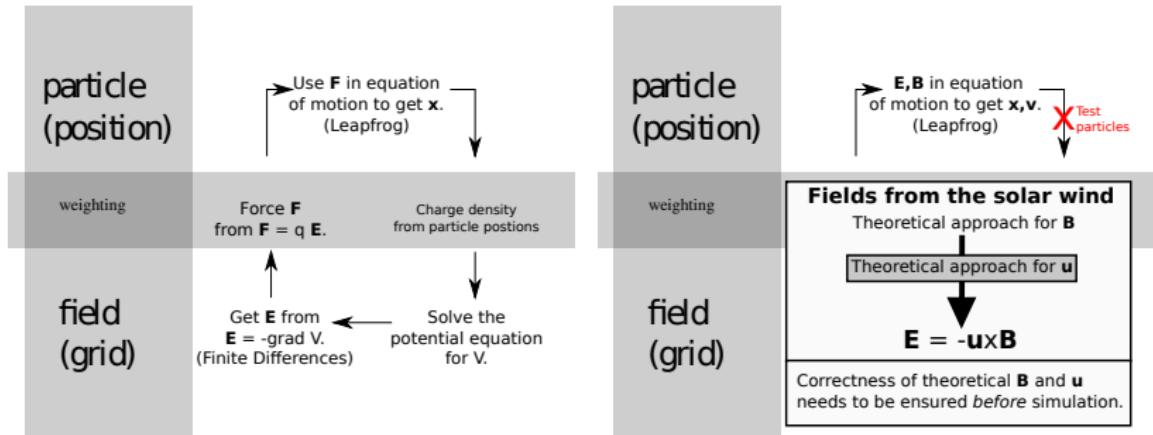
Equations simplify to:  $\rho_m \frac{\partial B}{\partial t} = \mu_0 \frac{\partial u}{\partial z}$

Wave equations for  $\vec{u}_1$  and  $\vec{B}_1$  predict a sinusoidal disturbance, which is correllated.

This yields  $\vec{u}_1, \vec{B}_1 \propto \sin(\omega t \pm \vec{k} \cdot \vec{x})$ . Amplitudes:  $\delta \vec{u}, \delta \vec{B}$ .

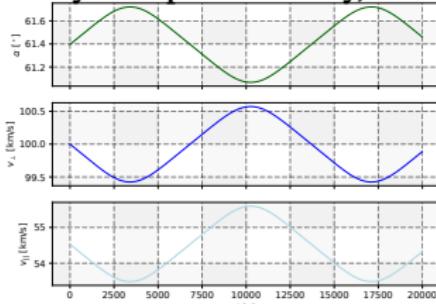
$$\Rightarrow \delta \vec{u} = \pm \delta \vec{B} / \sqrt{\mu_0 \rho_m}$$

# Comparison: PIC vs. test particles

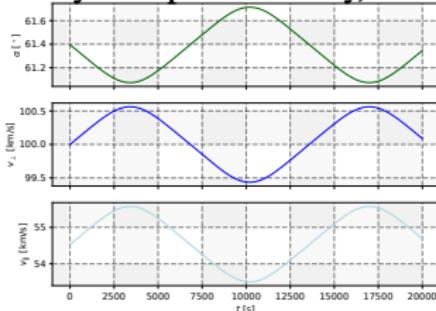


**Boundary conditions:**  $\omega = 0.5 \omega_{c,O}$ ,  $\delta B = 5 \cdot 10^{-4}$  nT

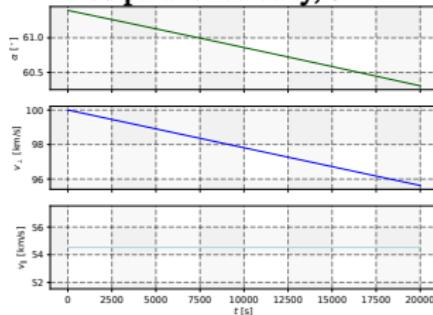
**Dynamic parallel velocity,  $0^\circ$**



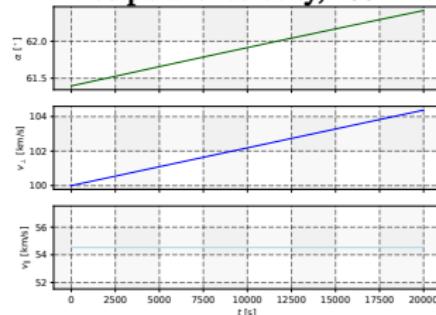
**Dynamic parallel velocity,  $180^\circ$**



**Fixed parallel velocity,  $0^\circ$**



**Fixed parallel velocity,  $180^\circ$**



## Resonant frequencies and wavevectors

