

Velocity Distribution Functions of Pickup Ions with Ulysses/SWICS

Master Thesis Results

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Outline

Pickup Ions

Motivation

Ulysses SWICS

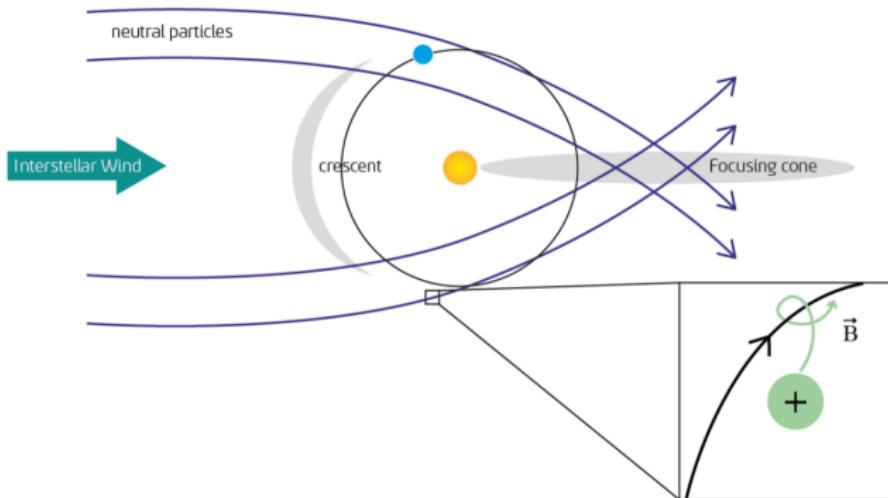
Principle of Measurement

Outlook & Conclusion

The Pickup Process

Pickup Ions:

Former neutrals that get ionised within the heliosphere



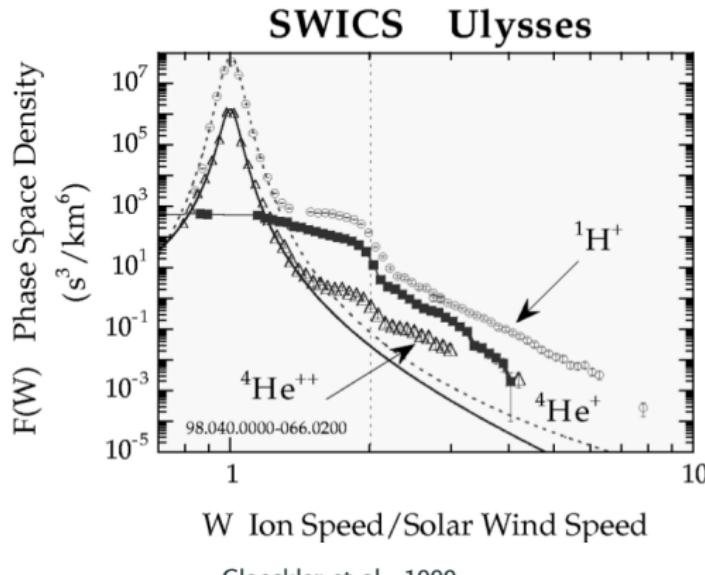
PUI – Measurement

Observed PUIs:

H^{1+} , ${}^3\text{He}^{1+}$, He^{1+} ,
 He^{2+} , C^{1+} , N^{1+} , O^{1+} ,
 Ne^{1+} , Mg^{1+} , Si^{1+} , Fe^{1+}

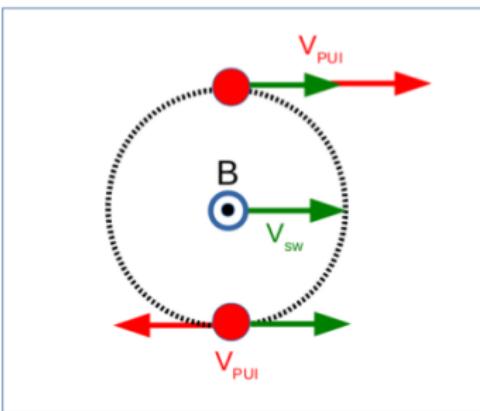
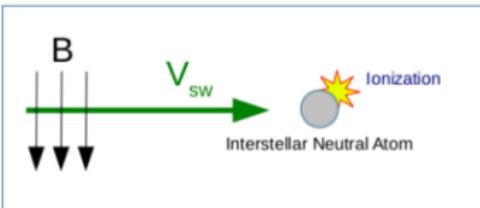
PUI or Solar Wind?

- Charge state
- Velocity distribution function (VDF)

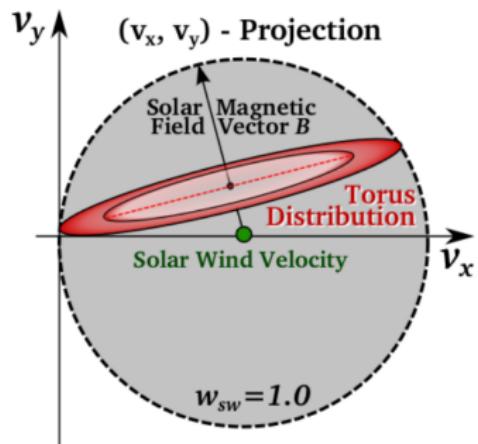


Gloeckler et al., 1999

The Pickup Process



Velocity Space:

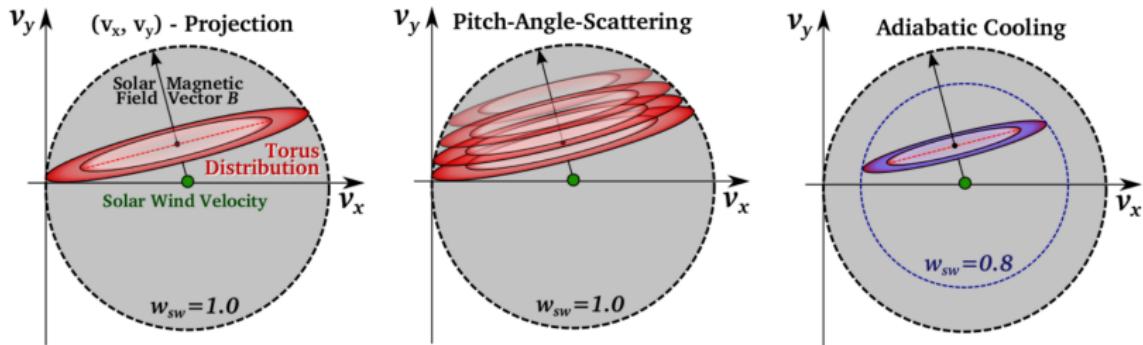


Drews et al., 2016

Taut, Drews et al., AGU Fall Meeting 2014

→ Anisotropic torus VDF

Evolution of the VDF

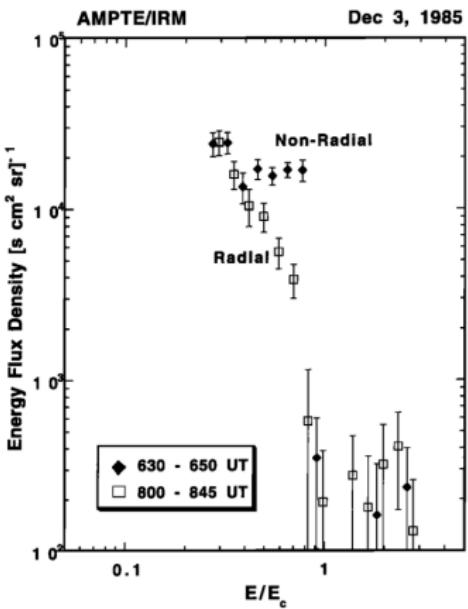


Drews, Berger et al., 2016

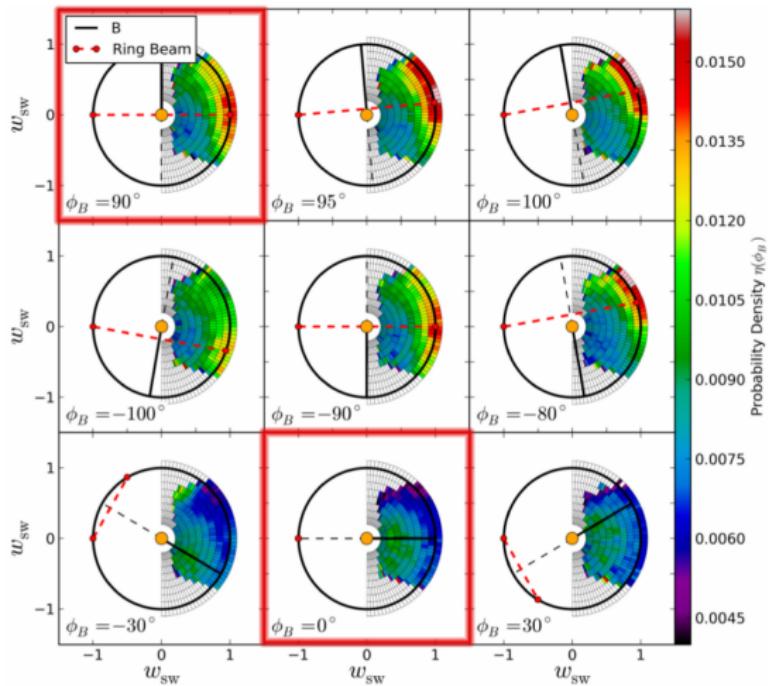
Modification of the initial torus-shaped VDF by:

- Pitch-angle scattering
→ isotropisation
- acceleration & deceleration

Anisotropic features of the VDF



Moebius et al., 1998



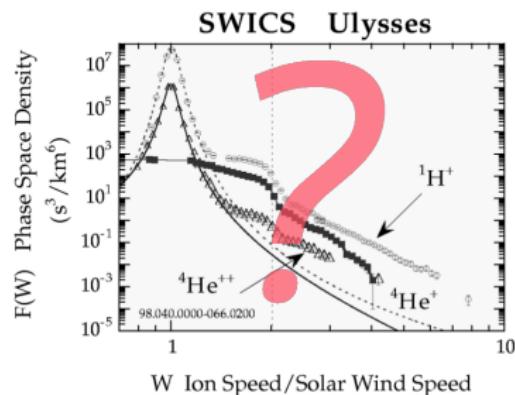
Drews, Berger et al., 2015

Motivation

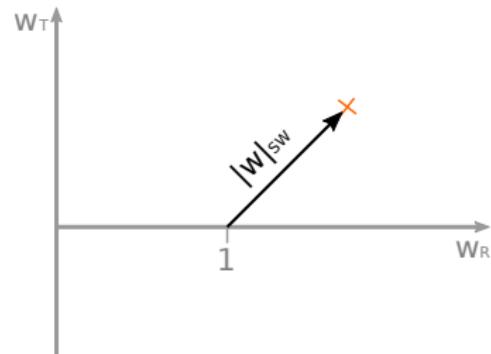
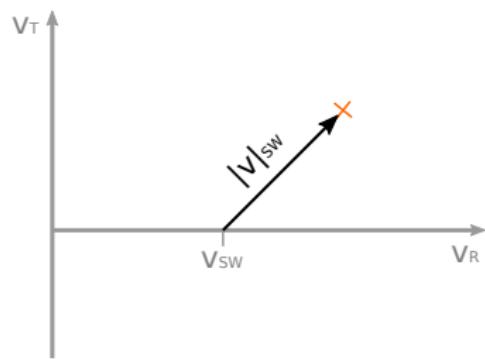
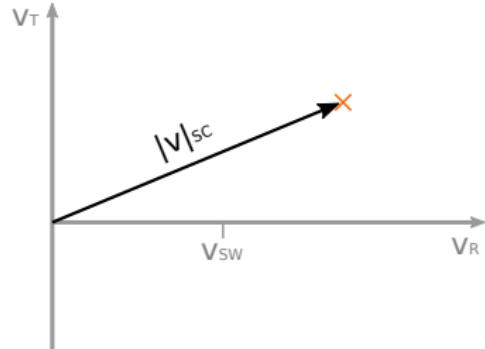
Problem:

Ambiguity of 1D reduced data

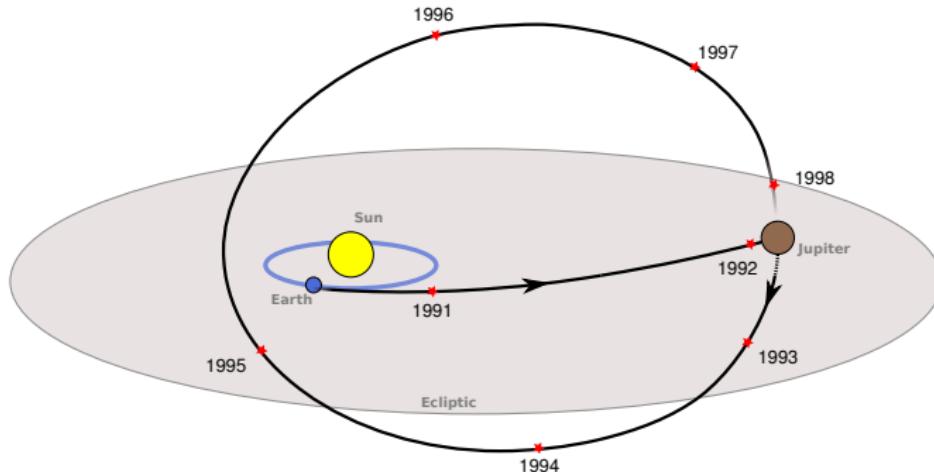
For fully understanding the
PUI transport in phase space
we need to analyse the **3D**
velocity distribution function



Frame of Reference

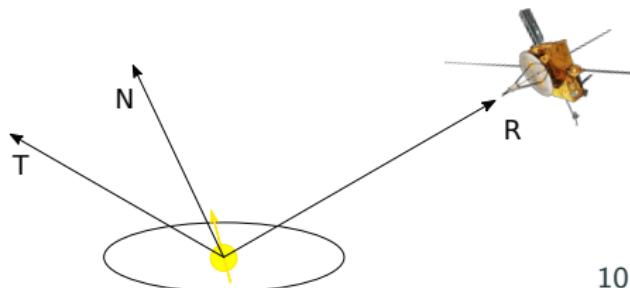


Ulysses Spacecraft (1990 – 2009)



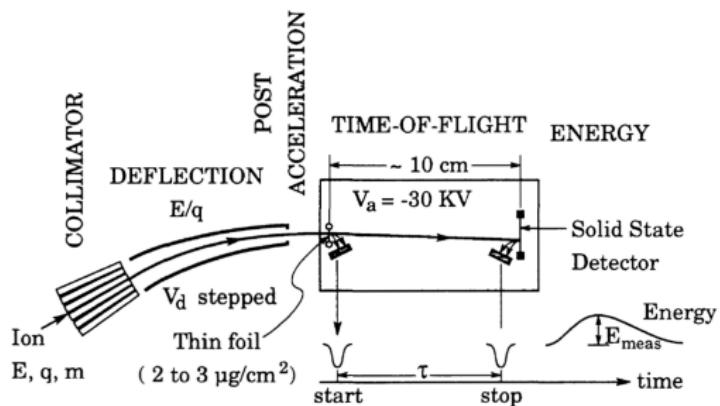
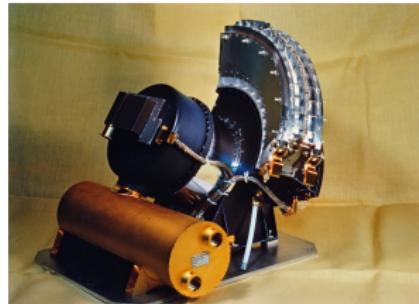
adapted from www.cosmos.esa.int, 2019

- Highly inclined orbit;
orbital period: 6.2 years
- spin-stabilized



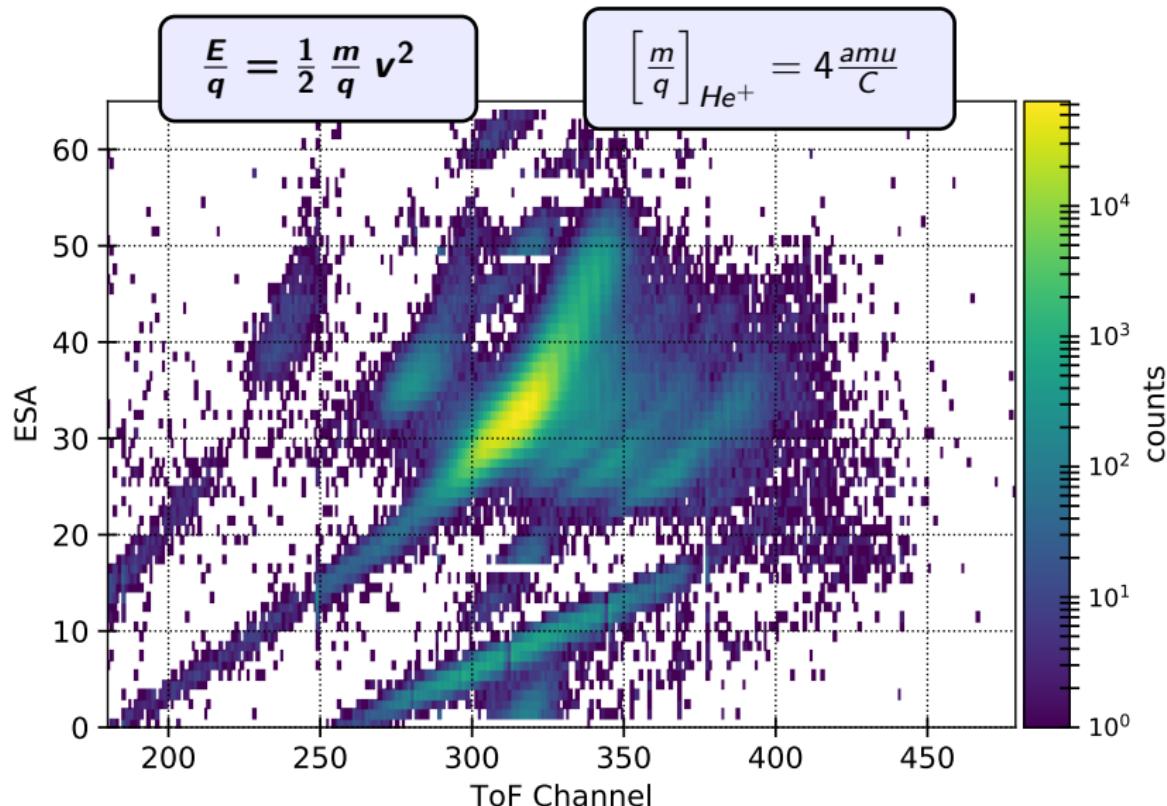
The Solar Wind Ion Composition Spectrometer

- Time-of-flight mass spectrometer
- $\left\{ \frac{E}{q}, T_{OF}, E_{SSD} \right\}$
 $\Rightarrow \left\{ \frac{M}{q}, M, |v| \right\}$
- identification & energy of the ion

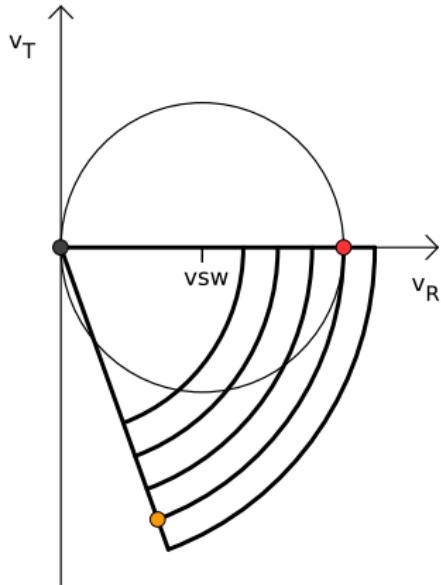
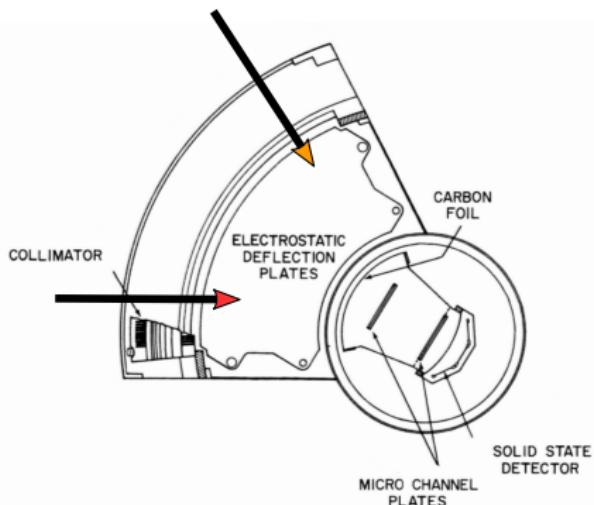


Gloeckler, Geiss et al., 1992

PHA data

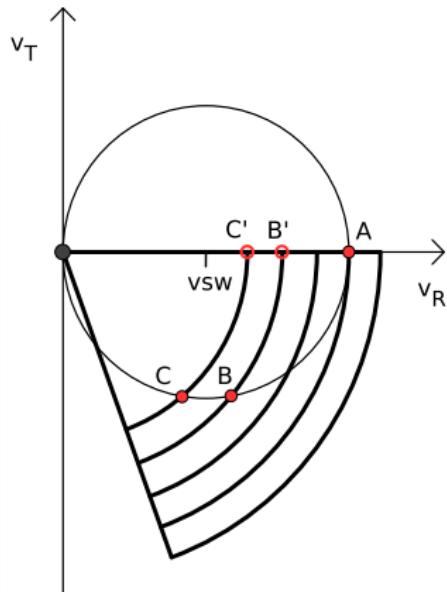
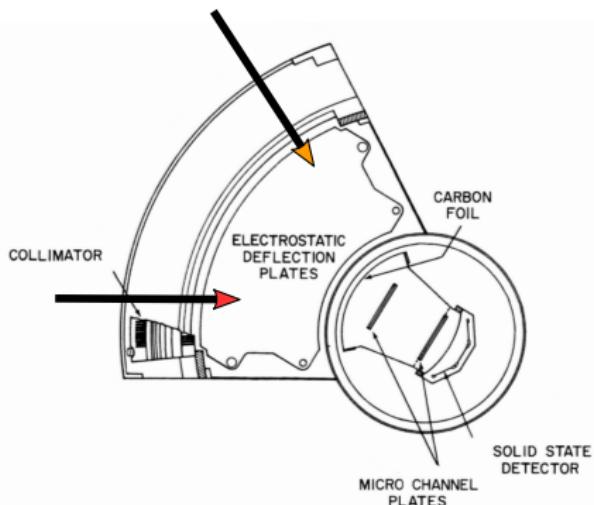


EpQ measurement



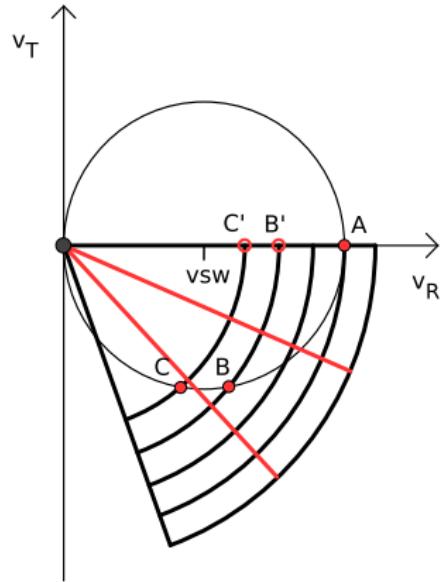
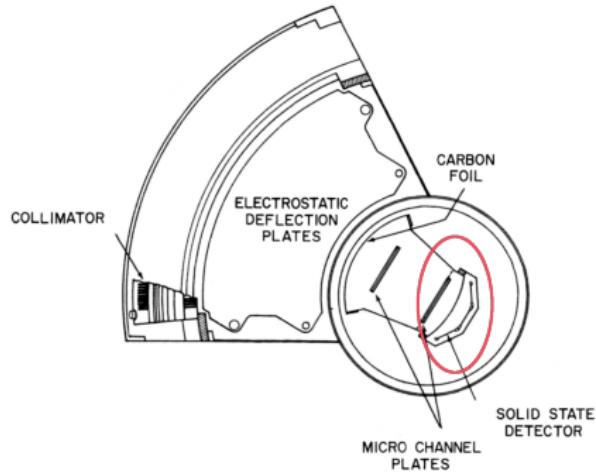
- For constant $\frac{m}{q}$: $\frac{E}{q}$ -step $\hat{=}$ absolute value of velocity
- Integration over EpQ shells \rightarrow loss of information!

EpQ measurement



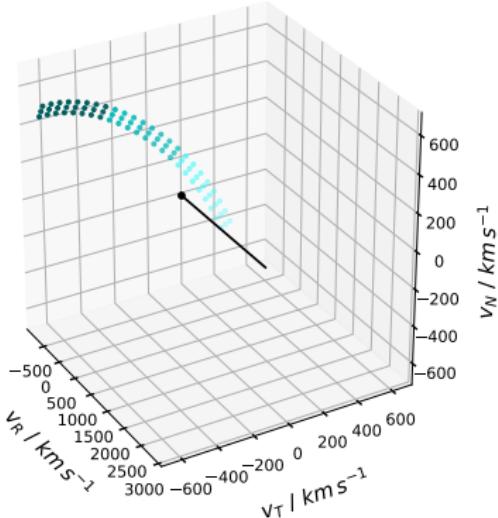
- For constant $\frac{m}{q}$: $\frac{E}{q}$ -step $\hat{=}$ absolute value of velocity
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Angular resolution

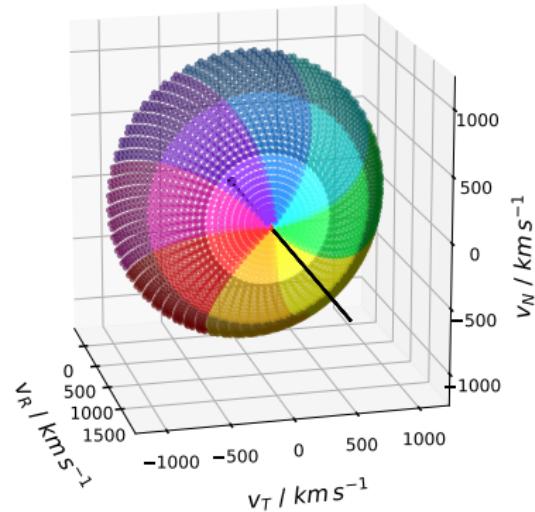


- SWICS: **3 detectors**
Rough distinction between angles of incidence
- 3rd dimension: spin of the SC
Divided into **8 sectors**

The Virtual Detector

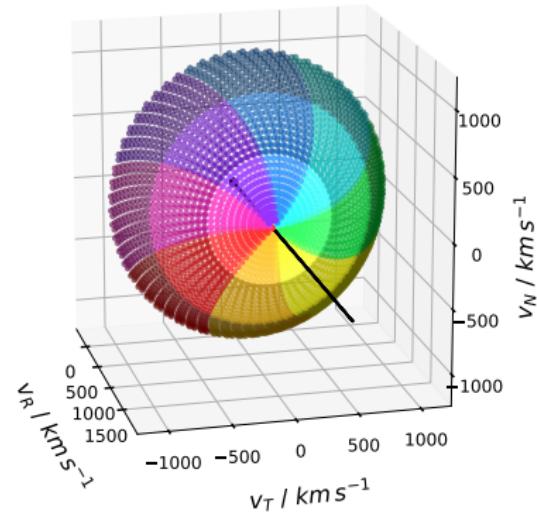
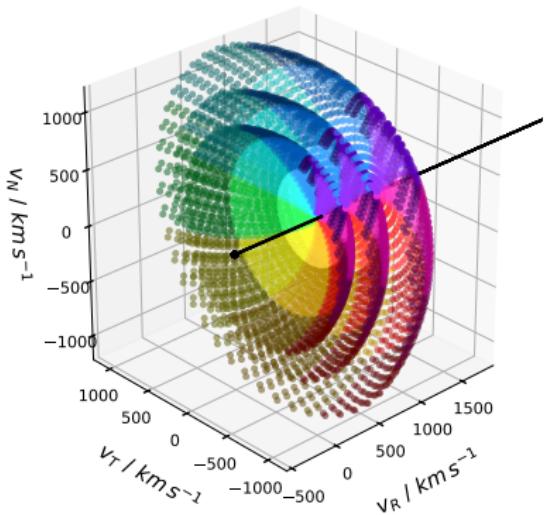


Unrotated collimator acceptance
for one EpQ step

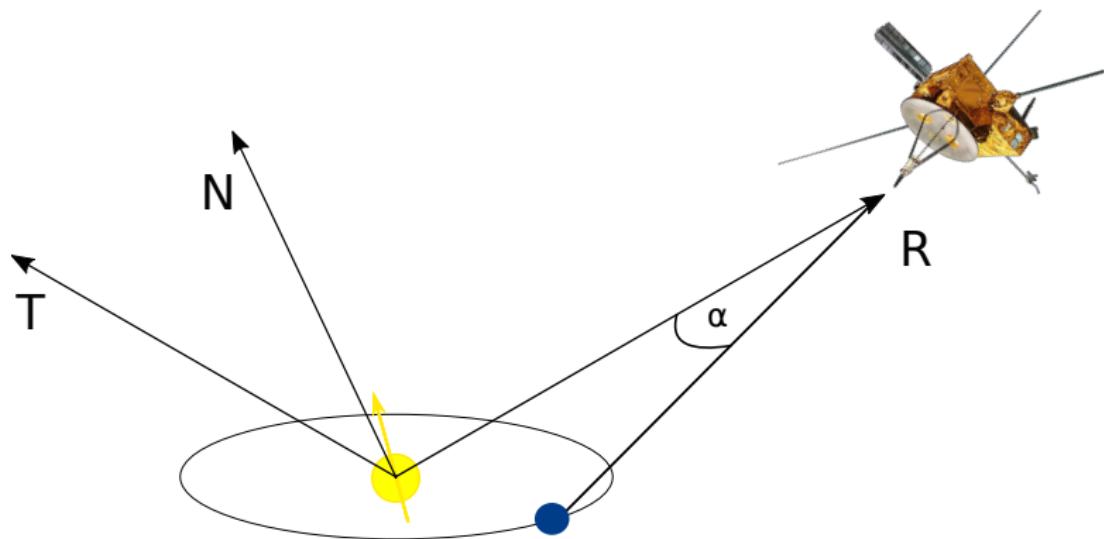


Collimator Acceptance for
one spacecraft spin and one EpQ step

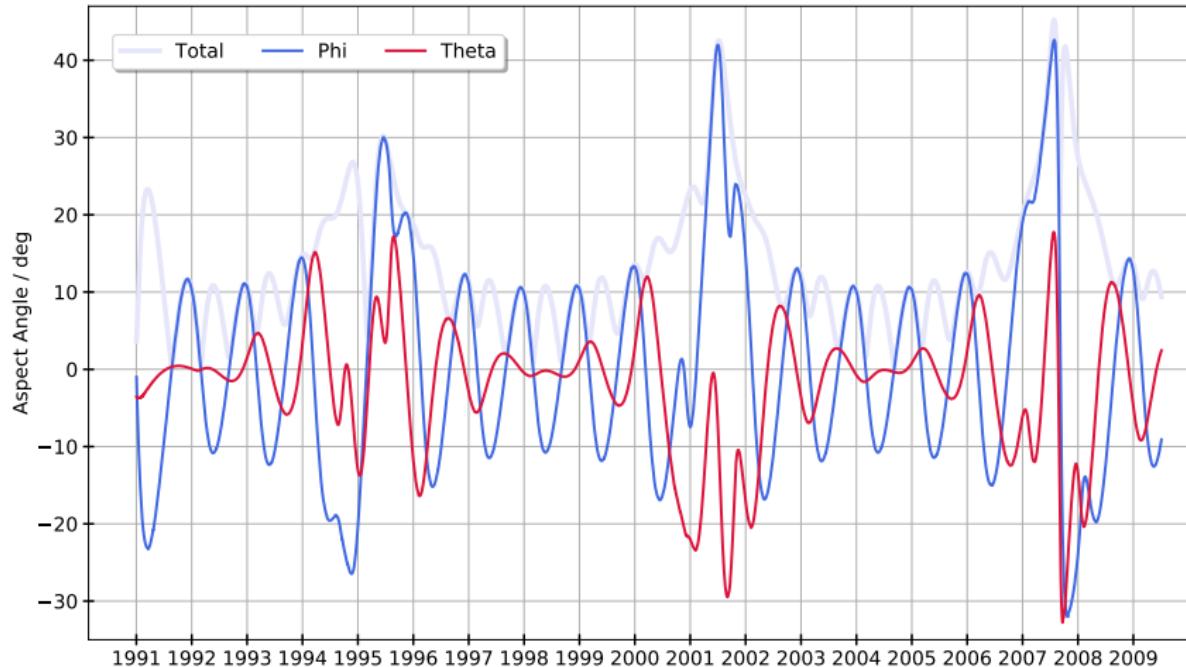
The Virtual Detector



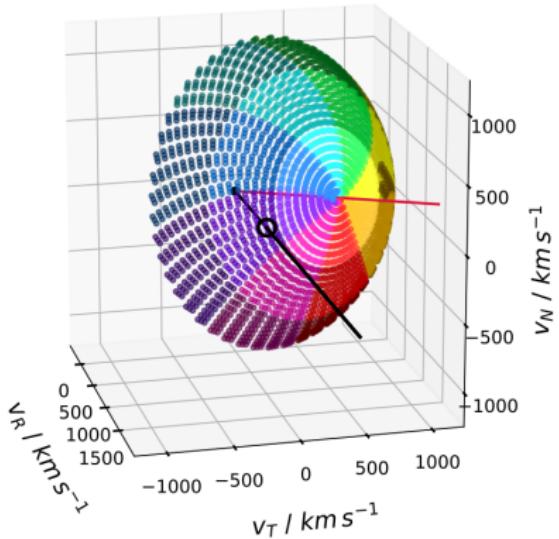
Aspect Angle



Aspect Angle

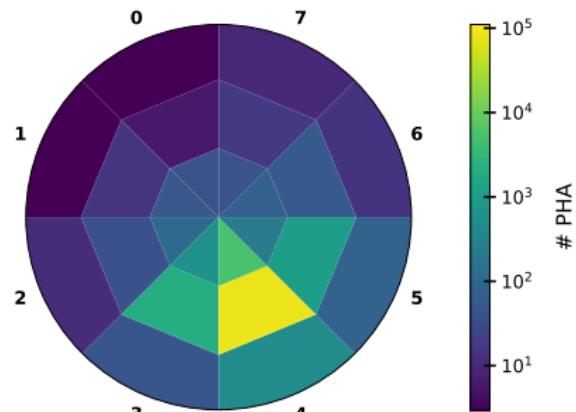


Aspect Angle



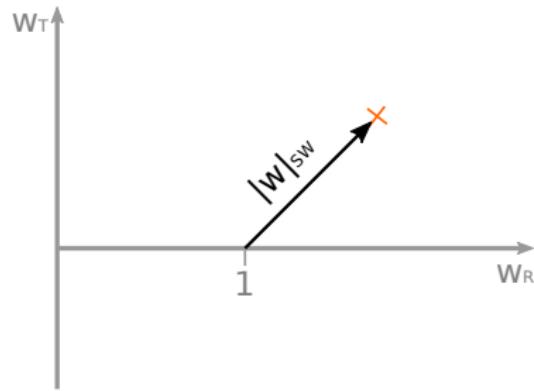
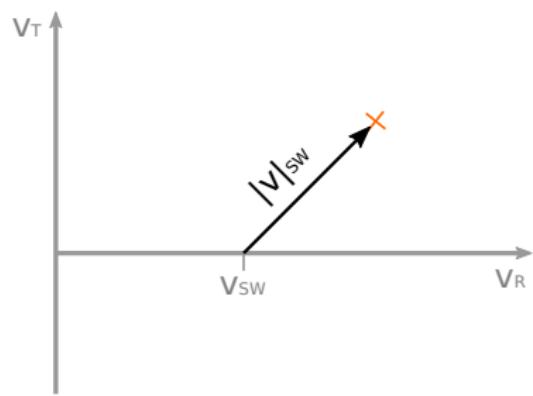
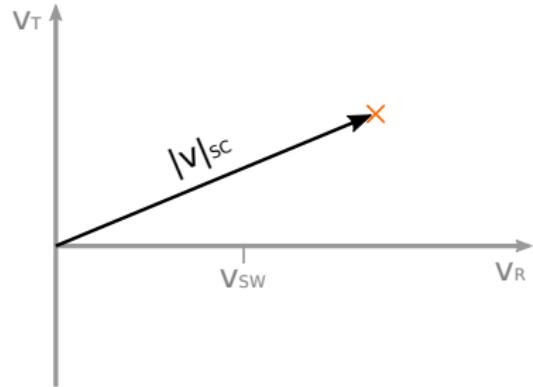
Aspect angle:

$$\varphi = 25^\circ, \vartheta = -10^\circ$$

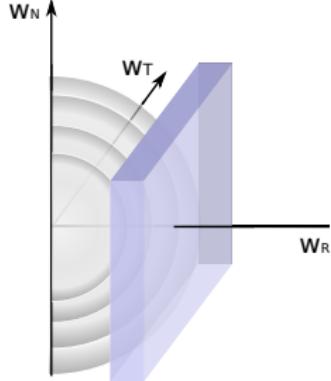


⇒ Plausibility check with
solar wind He²⁺ PHA data
(120 days in 2001)

Frame of Reference



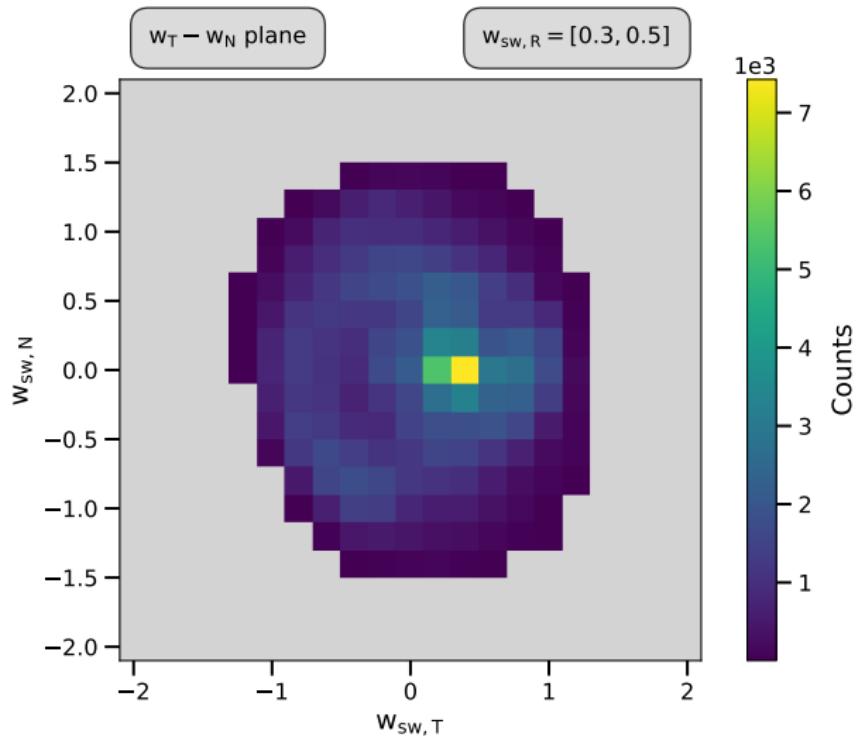
Spherical Cut through VDF



50 days in 1994

vsw :

$760 - 780 \text{ km s}^{-1}$

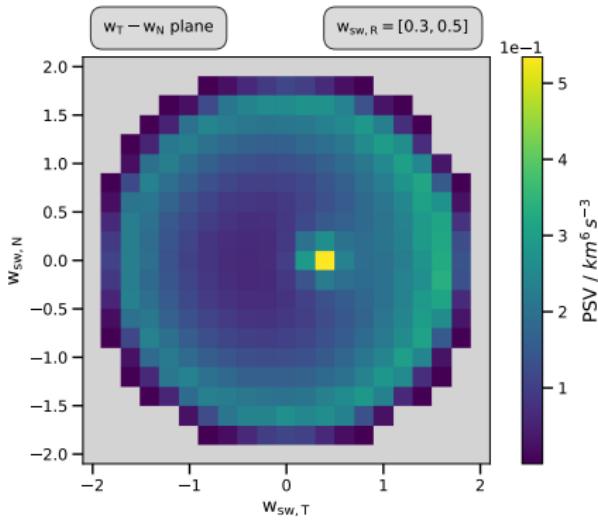
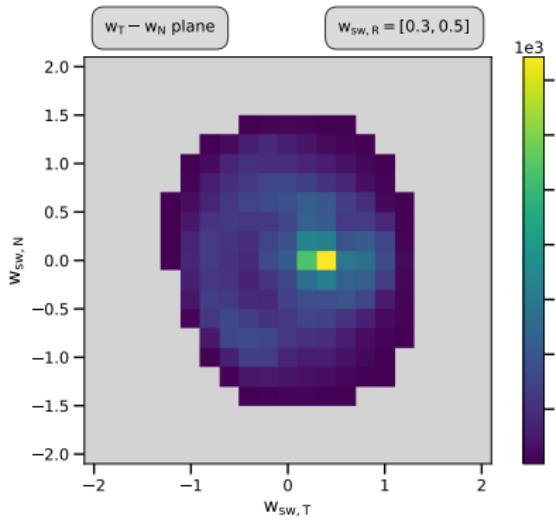


From Counts to Phase Space Density

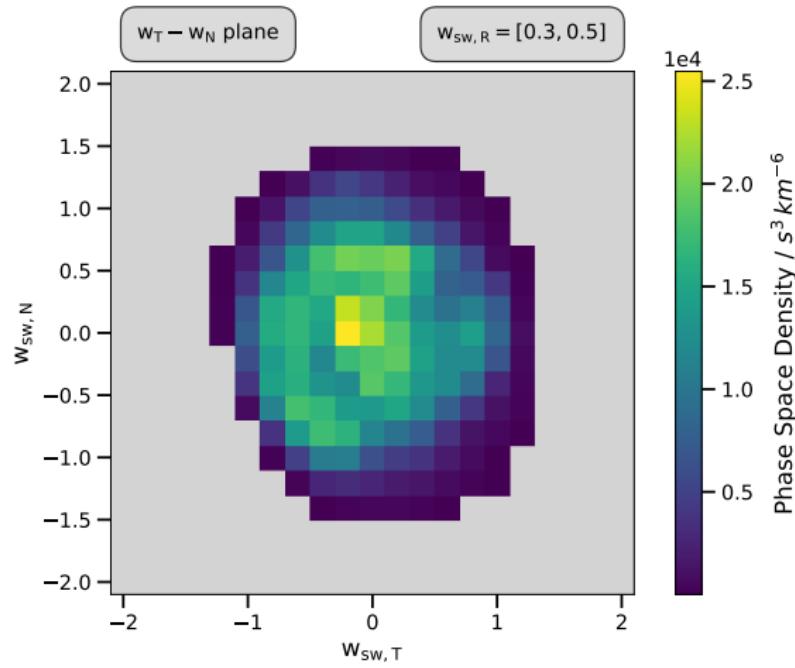
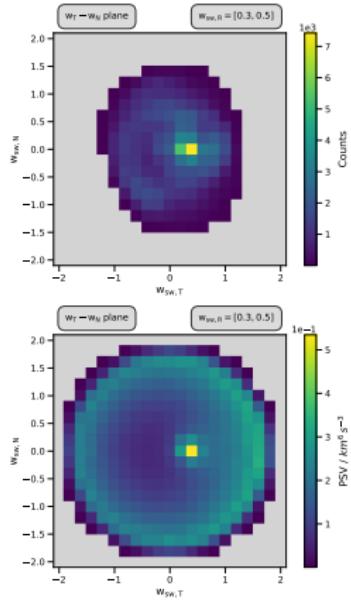
$$PSD = \frac{N}{PSV}$$



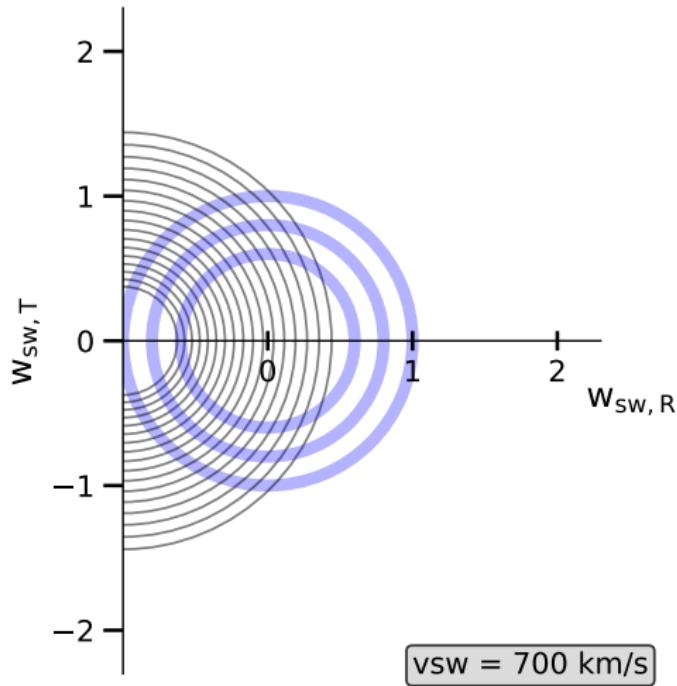
T



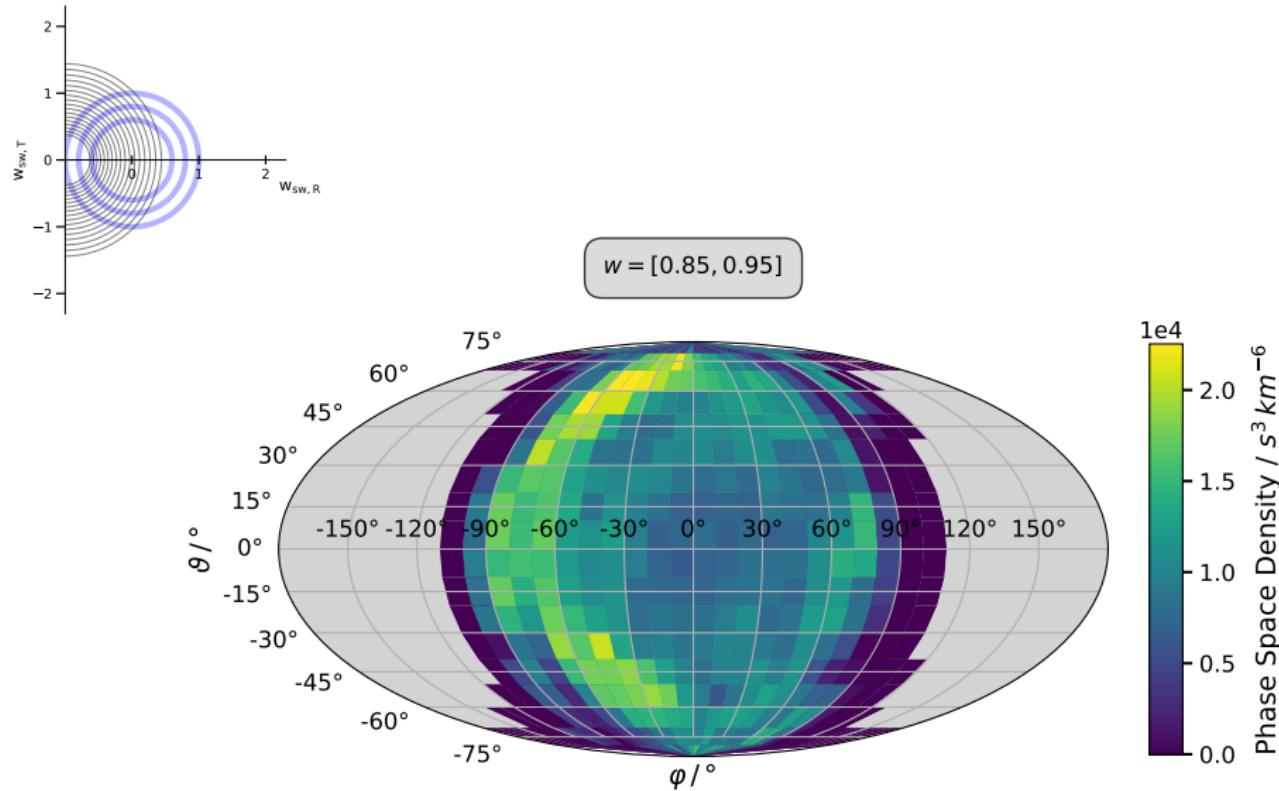
Spherical Cut through VDF



Spherical Surface



Spherical Surface



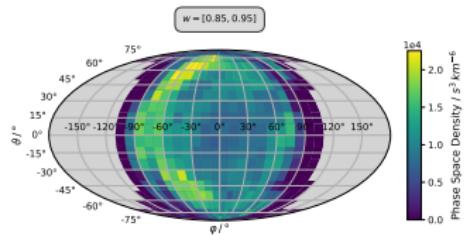
vsw : $760 - 780 \text{ km s}^{-1}$ – 50 days in 1993

1D

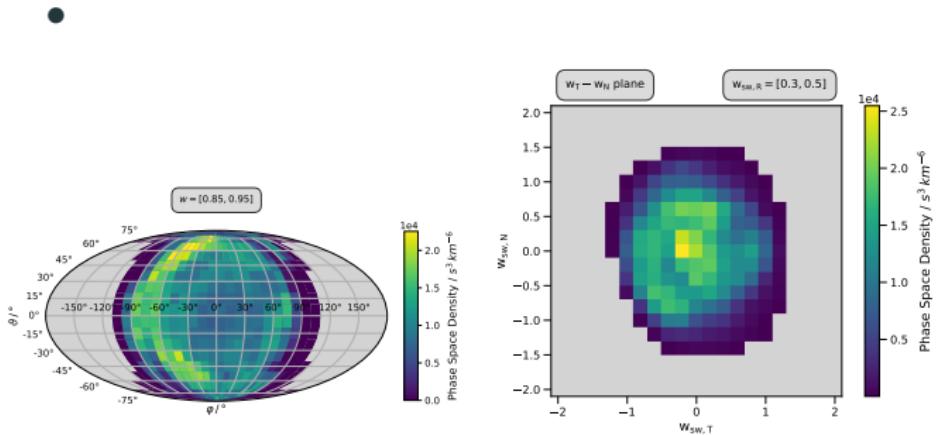
TODO

Outlook

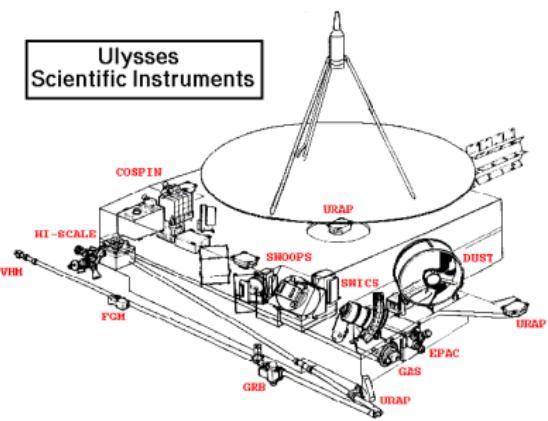
- 1D radiale Abhangigkeit untersuchen
- B-Feld abhangige Torusverteilung



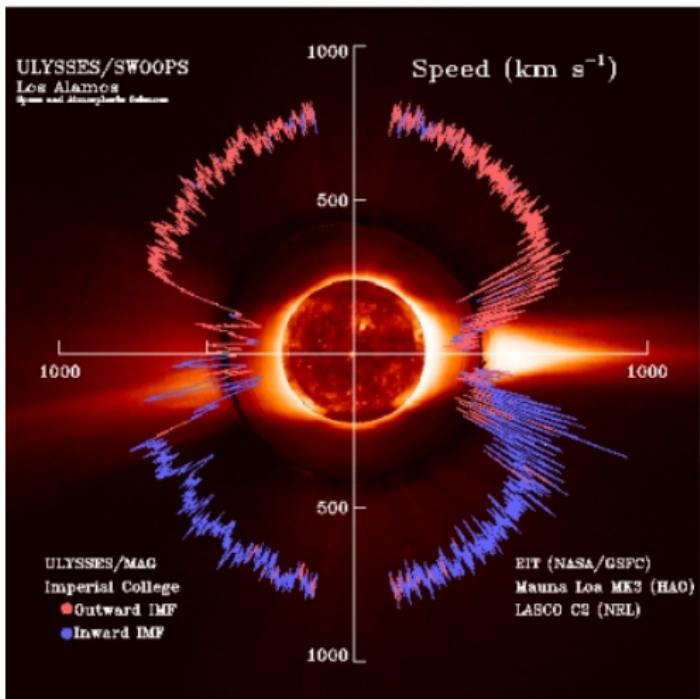
Conclusion



BACKUP

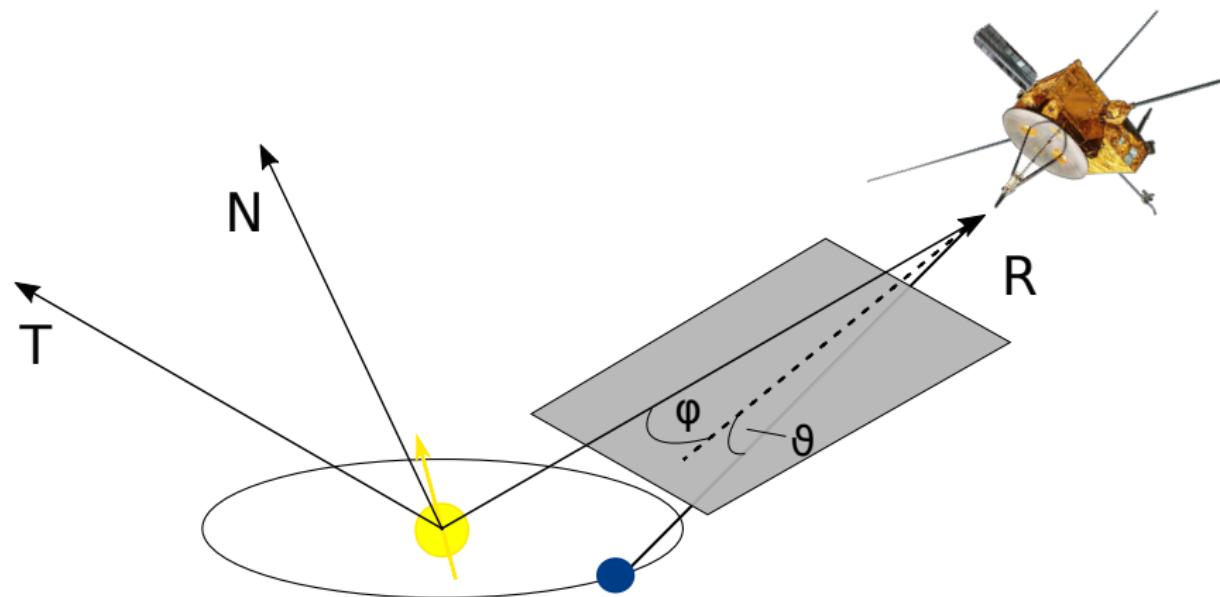


www.cosmos.esa.int, 2019

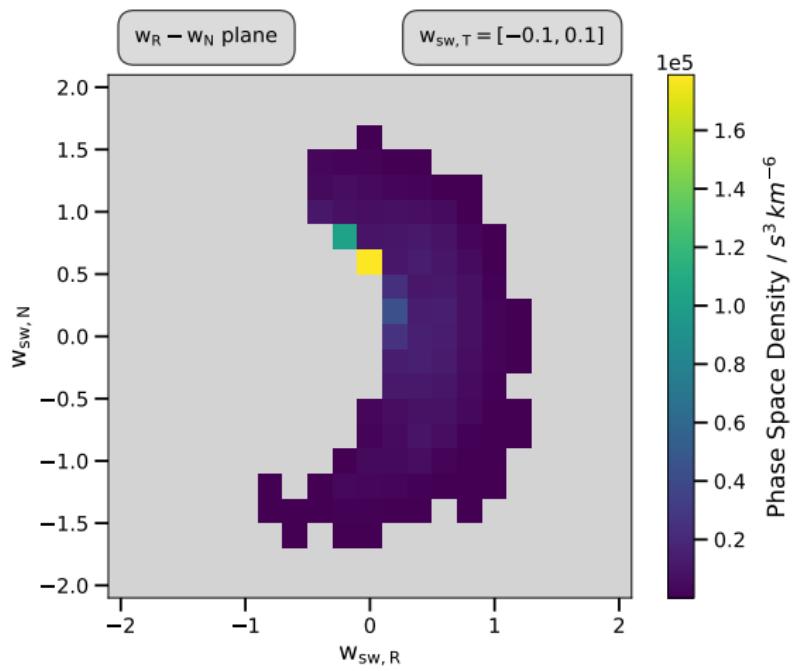
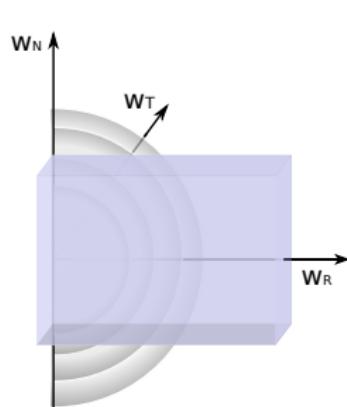


McComas et al., 2000

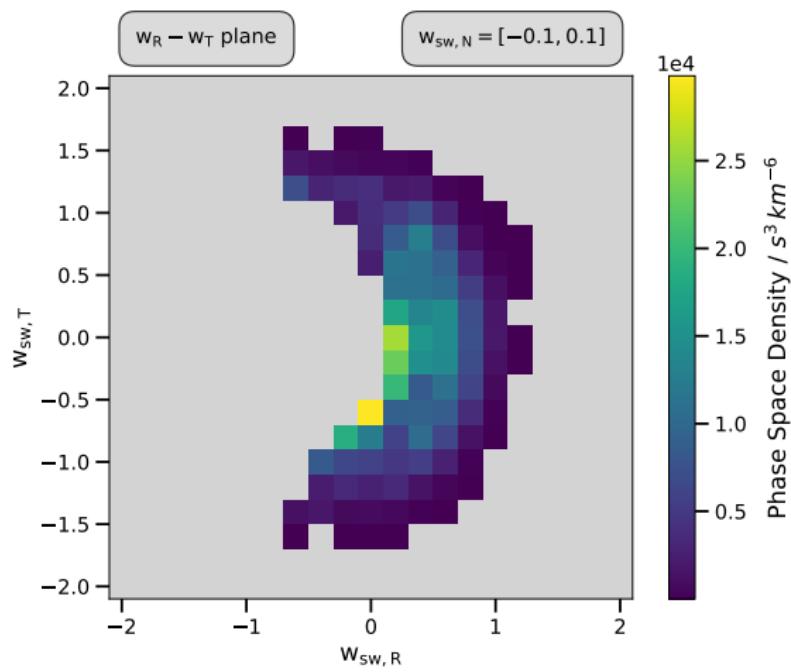
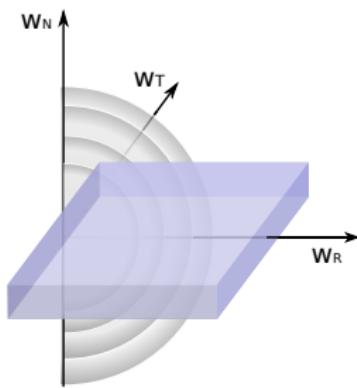
Aspect Angle



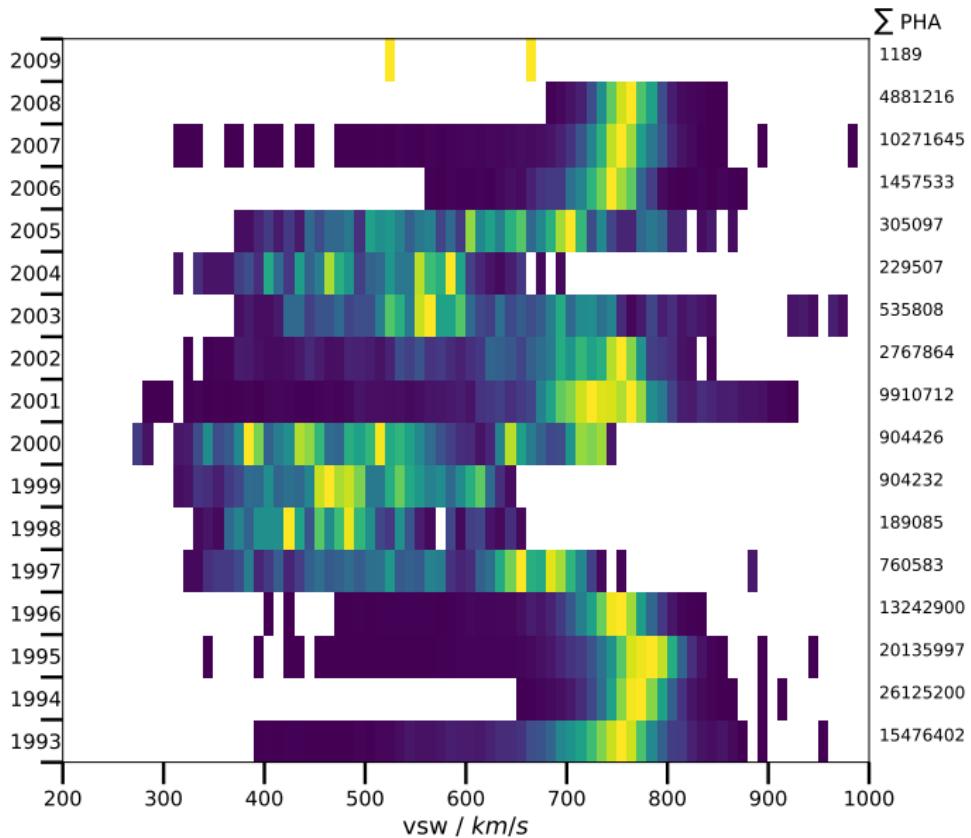
Cartesian Cut: T

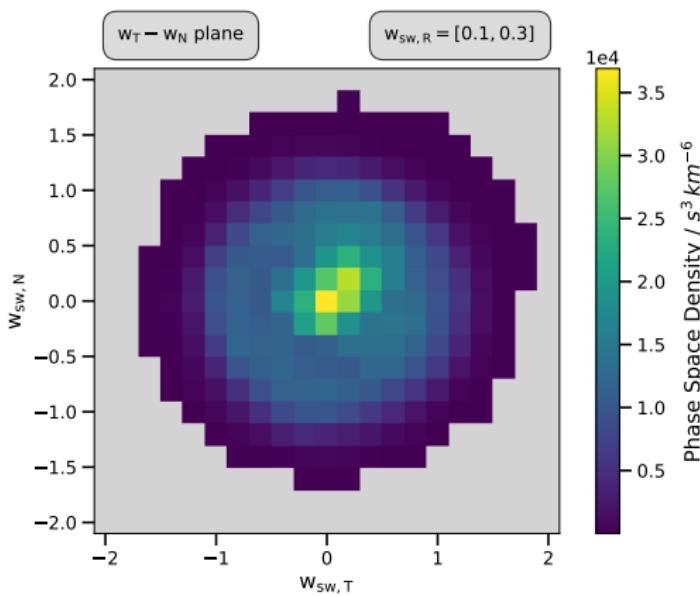
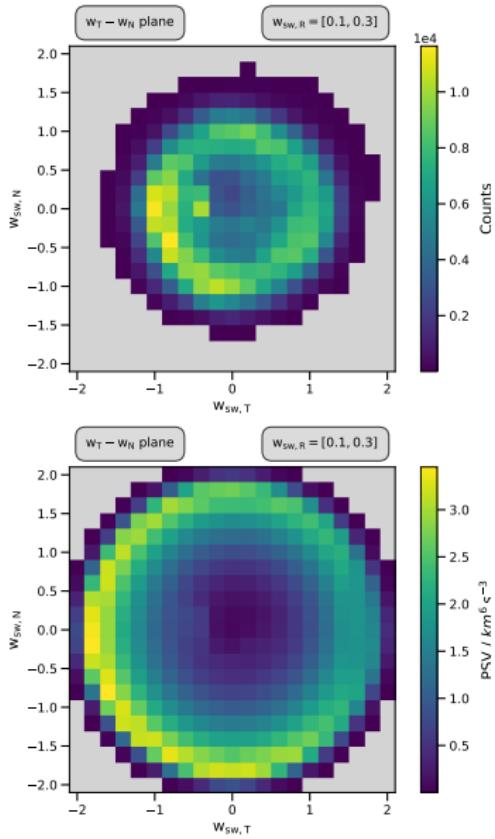


Cartesian Cut: N



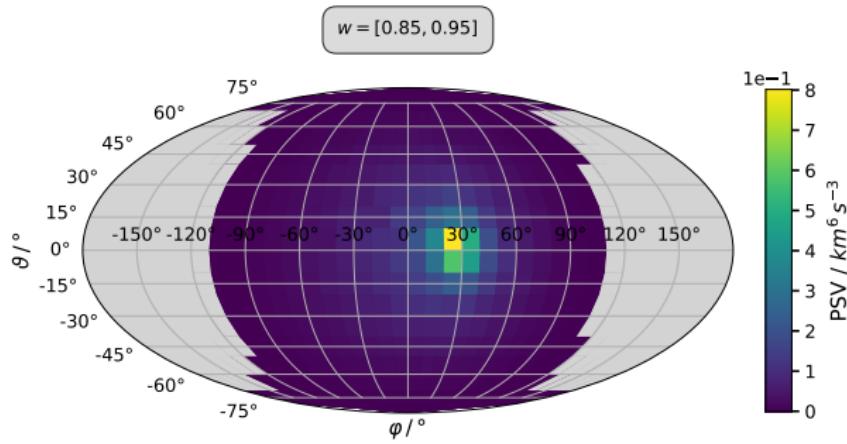
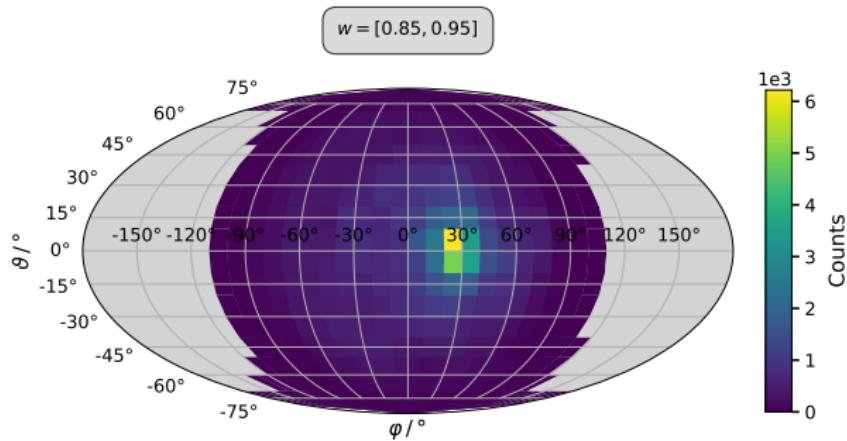
He^+ Data: vsw





v_{SW} : 740 – 780 km s⁻¹

250 days in 1994



vsw : $760 - 780 \text{ km s}^{-1}$

50 days in 1993