

Mark Scheme (Results)

January 2024

Pearson Edexcel International Advanced Level In Physics (WPH15)

Paper 01: Thermodynamics, Radiation,

Oscillations and Cosmology

Question Number	Answer	Mark
1	B is the only correct answer	(1)
	A is not the correct answer, as molecular potential energy increases C is not the correct answer, as molecular kinetic energy doesn't change and molecular potential energy increases D is not the correct answer, as molecular kinetic energy doesn't change	
2	A is the only correct answer	(1)
	B is not the correct answer, as parallax can only be used for nearby stars C is not the correct answer, as parallax measurements do not involve luminosity D is not the correct answer, as the parallax effect is greatest for nearby stars	
3	C is the only correct answer, as the mass-energy equivalent is calculated by converting the mass to kg and then using $\Delta E = c^2 \Delta m$	(1)
4	D is the only correct answer	(1)
	A is not the correct answer, as acceleration can be in the same or the opposite direction to velocity  B is not the correct answer, as acceleration can be in the same or the opposite direction to velocity  C is not the correct answer, as acceleration is always towards the equilibrium point	
5	C is the only correct answer	(1)
	A is not the correct answer, as this does not make the process spontaneous B is not the correct answer, as this does not make the process spontaneous D is not the correct answer, as this does not make the process spontaneous	
6	B is the only correct answer	(1)
	A is not the correct answer, as speed increases C is not the correct answer, as gravitational potential energy decreases and speed increases D is not the correct answer, as gravitational potential energy decreases	
7	<b>B</b> is the only correct answer, as $v_{\text{max}} = \omega A$	(1)
8	D is the only correct answer	(1)
	A is not the correct answer, as the mean kinetic energy is the same for both gases B is not the correct answer, as the mean kinetic energy is the same for both gases C is not the correct answer, as He molecules are more massive than H molecules	
9	D is the only correct answer	(1)
	A is not the correct answer, as the peak for star Y is at the higher frequency B is not the correct answer, as the peak for star Y is at the lower frequency C is not the correct answer, as the peak for star X is at the higher frequency	
10	A is the only correct answer, as $g = \frac{GM}{r^2}$	(1)

Question Number	Answer		
11	Stars in area P are (massive) <u>main sequence</u> stars (1)		
	Stars in area P evolve into area S and stars in area S evolve into area Q [Allow $P \to S \to Q$ , or arrows on diagram] (1)		
	Stars in area S are <u>red</u> giant stars <b>and</b> stars in area Q are <u>white</u> dwarf stars [accept red supergiant for red giant]	3	
	Total for question 11	3	

Question Number	Answer		Mark
12	EITHER		
	Use of $g = \frac{GM}{r^2}$	(1)	
	Use of $T = 2\pi \sqrt{\frac{\ell}{g}}$ with $g = 9.81$ N kg <sup>-1</sup> to calculate length [ $\ell = 0.994$ m]	(1)	
	Use of $T = 2\pi \sqrt{\frac{\ell}{g}}$ with $g_{\text{venus}}$ to calculate period	(1)	
	$T_{\text{Venus}} = 2.1 \text{ s [Do not accept } 2.0 \text{ s]}$	(1)	
	OR		
	Use of $g = \frac{GM}{r^2}$	(1)	
	Ratio of periods with $T = 2\pi \sqrt{\frac{\ell}{g}}$	(1)	
	Re-arrangement to obtain $\frac{T_{\text{Venus}}}{T_{\text{Earth}}} = \sqrt{\frac{g_{\text{Earth}}}{g_{\text{Mars}}}}$	(1)	
	$T_{\text{Venus}} = 2.1 \text{ s} \text{ [Do not accept 2.0 s]}$	(1)	4
	Example of calculation		
	$g = \frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2} \times 4.87 \times 10^{24} \text{ kg}}{(6.05 \times 10^6 \text{ m})^2} = 8.87 \text{ N kg}^{-1}$		
	$\frac{T_{\text{Venus}}}{T_{\text{Earth}}} = \sqrt{\frac{g_{\text{Earth}}}{g_{Mars}}} = \sqrt{\frac{9.81 \text{ N kg}^{-1}}{8.87 \text{ N kg}^{-1}}} = 1.05$		
	$T_{\text{Venus}} = 1.05 \times 2.00 \text{s} = 2.10 \text{s}$		
	Total for question 12		4

<b>Question Number</b>	Answer		Mark		
13	EITHER				
	Use of $\frac{\Delta f}{f} \approx \frac{v}{c}$ [must have $f_{\text{lab}}$ in denominator]	1)			
	$v = (-)1.5 \times 10^6 \mathrm{m  s^{-1}}$	1)			
	Galaxy is receding (from the Earth) (1)				
	OR				
	Use of $z = \frac{\Delta f}{f}$ [must have $f_{lab}$ in denominator]	1)			
	$z = (-) 5.0 \times 10^{-3} $	1)			
	Galaxy is receding (from the Earth)	1)	3		
	Example of calculation $v = 3.0 \times 10^8 \text{ m s}^{-1} \times \left(\frac{6.142 \times 10^{14} \text{Hz} - 6.173 \times 10^{14} \text{ Hz}}{6.173 \times 10^{14} \text{ Hz}}\right) = -1.51 \times 10^6 \text{ m s}^{-1}$				
	Total for question 13		3		

Question Number	Answer		Mark
14	Use of mass of molecule to calculate $N$	(1)	
	Use of $pV = NkT$	(1)	
	Conversion of temperature to kelvin	(1)	
	Difference in pressures calculated	(1)	
	$\Delta p = 3.6 \times 10^4 \text{ Pa}$	(1)	5
	Example of calculation $N = \frac{2.00 \times 10^{-3} \text{kg}}{4.67 \times 10^{-26} \text{kg}} = 4.28 \times 10^{22}$		
	$p = \frac{4.28 \times 10^{22} \times 1.38 \times 10^{-23} \text{ J K}^{-1} \times (22 + 273) \text{ K}}{2.85 \times 10^{-4} \text{ m}^3} = 6.11 \times 10^5 \text{ Pa}$		
	$\Delta p = 6.11 \times 10^5 - 5.75 \times 10^5 \text{ Pa} = 3.64 \times 10^4 \text{Pa}$		
	Total for question 14		5

Question Number	Answer					
15(a)	Energy is transferred from banana to liquid nitrogen	(1)				
	The molecular potential energy of nitrogen molecules increases (as the nitrogen boils)					
	Or This provides the latent heat of vaporisation (of the nitrogen)		2			
	Or This provides the latent heat to change state (of the nitrogen)	(1)	-			
15(b)	Use of $\Delta E = mc\Delta\theta$	(1)				
	Use of $\Delta E = mL$	(1)				
	$m = 0.23$ kg, which is less that 0.5 kg so the teacher's estimate was inaccurate Or $9.9 \times 10^4$ J > $4.5 \times 10^4$ J so the teacher's estimate was inaccurate	(1)	3			
	Example of calculation $\Delta E = 0.118 \text{ kg} \times 1.76 \times 10^3 \text{ J kg}^{-1} \text{ K}^{-1} \times (292 - 77.4) \text{ K} = 4.46 \times 10^4 \text{ J}$					
	$m = \frac{4.46 \times 10^4 \text{ J}}{1.98 \times 10^5 \text{ J kg}^{-1}} = 0.225 \text{ kg}$					
	Total for question 15		5			

Question Number	Answer		Mark
16(a)	Calculates surface area of Sun [using $A = 4\pi r^2$ ]	(1)	
	Use of $L = \sigma A T^4$	(1)	
	Use of $I = \frac{L}{4\pi d^2}$	(1)	
	$I = 0.89 \text{ W m}^{-2}$	(1)	4
	Example of calculation		
	$A = 4\pi \times (6.96 \times 10^8 \text{ m})^2 = 6.09 \times 10^{18} \text{ m}^2$		
	$L = 5.67 \times 10^{-8} \text{W m}^{-2} \text{K}^{-4} \times 6.09 \times 10^{18} \text{ m}^{2} \times (5800 \text{ K})^{4} = 3.91 \times 10^{26} \text{ W}$		
	$I = \frac{3.91 \times 10^{26} \text{ W}}{4\pi \times (5.91 \times 10^{12} \text{ m})^2} = 0.891 \text{ W m}^{-2}$		
16(b)	EITHER Equates $F = \frac{GMm}{r^2}$ with $F = m\omega^2 r$	(1)	
	Use of $\omega = \frac{2\pi}{T}$	(1)	
	Conversion of T to years	(1)	
	T = 250  (years)	(1)	
	OR Use of $F = \frac{GMm}{r^2}$ with $F = \frac{mv^2}{r}$	(1)	
	Use of $v = \frac{2\pi r}{T}$	(1)	
	Conversion of T to years	(1)	
	T = 250  (years)	(1)	4
	Example of calculation	(1)	•
	$\frac{GMm}{r^2} = m\omega^2 r$ $\omega = \sqrt{\frac{GM}{r^3}} = \sqrt{\frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-1} \times 1.99 \times 10^{30} \text{ kg}}{(5.91 \times 10^{12} \text{ m})^3}}$ $\therefore \omega = 8.02 \times 10^{-10} \text{ rad } s^{-1}$ $T = \frac{2\pi}{\omega} = \frac{2\pi \text{ rad}}{8.02 \times 10^{-10} \text{ rad s}^{-1}} = 7.84 \times 10^9 \text{ s} = \frac{7.84 \times 10^9 \text{ s}}{3.15 \times 10^7 \text{ s year}^{-1}} = 249 \text{ year}$		
	Total for question 16		8

Question Number	Answer	Mark
17(a)	The astronauts must experience a centripetal force as they orbit the Earth	
	Or The weight of the astronaut acts as a centripetal force Or The astronauts are in free fall as they orbit the Earth  (1)	)
	They do not experience (normal) contact forces (1)	2
17(b)	Use of $\Delta F = -k\Delta x$ (1)	)
	Use of $T = 2\pi \sqrt{\frac{m}{k}}$ to calculate mass of chair	
	Use of $T = 2\pi \sqrt{\frac{m}{k}}$ to calculate mass of chair plus astronaut (1)	
	Mass difference calculated (1)	)
	Mass of astronaut = $67 \text{ kg}$ (1)	5
	Example of calculation $k = \frac{175 \text{ N}}{0.289 \text{ m}} = 606 \text{ N m}^{-1}$ $m = \frac{kT^2}{4\pi^2} = \frac{606 \text{ N m}^{-1} \times (0.858 \text{ s})^2}{4\pi^2} = 11.3 \text{ kg}$ $m = \frac{606 \text{ N m}^{-1} \times (2.26 \text{ s})^2}{4\pi^2} = 78.3 \text{ kg}$ mass of astronaut = $(78.3 - 11.3) \text{ kg} = 67.0 \text{ kg}$	
	Total for question 17	7

Question Number	Answer					
18(a)	EITHER	- (4)				
	Use of $\lambda_{\text{max}}T = 2.898 \times 10^{-3} \text{ m K}$	(1)				
	$\left  \frac{T_{\rm A}}{T_{\rm B}} = 1.76 \right $	(1)				
	Comparison of their ratio with 2 and conclusion	(1)				
	OR	(1)				
	Use of $\lambda_{\text{max}}T = 2.898 \times 10^{-3} \text{ m K}$					
	$T_{\rm A} = 3410 \; { m K} \; { m and} \; T_{ m B} = 1932 \; { m K}$	(1)	3			
	Comparison of their value of $T_A$ with $2T_B$ and conclusion	(1)				
	$\frac{\text{Example of calculation}}{\frac{T_{\text{A}}}{T_{\text{B}}}} = \frac{\lambda_{\text{max,B}}}{\lambda_{\text{max,A}}} = \frac{1500 \times 10^{-9} \text{ m}}{850 \times 10^{-9} \text{ m}} = 1.76$					
18(b)	Standard candles are stellar objects with a known luminosity	(1)				
	Locate a standard candle in the galaxy	(1)				
	Measure the intensity of the standard candle (on Earth)	(1)				
	Use the inverse square law to calculate the distance to the standard candle <b>Or</b> Use $I = \frac{L}{4\pi d^2} [I \text{ and } L \text{ must be defined}]$	(1)	4			
18(c)	Use of $v = H_0 d$	(1)				
	Use of $t = \frac{1}{H_0}$	(1)				
	Conversion from s to year	(1)				
	$t = 1.37 \times 10^{10} \text{ year}$	(1)	4			
	Example of calculation $H_0 = \frac{2.02 \times 10^6 \text{ m s}^{-1}}{8.70 \times 10^{23} \text{ m}} = 2.32 \times 10^{-18} \text{ s}^{-1}$ $t = \frac{1}{2.32 \times 10^{-18} \text{ s}^{-1}} = 4.31 \times 10^{17} \text{ s}$ $4.31 \times 10^{17} \text{ s}$					
	$t = \frac{4.31 \times 10^{17} \text{ s}}{3.15 \times 10^{7} \text{ s year}^{-1}} = 1.37 \times 10^{10} \text{ year}$					
	Total for question 18		11			

Question Number	Answer		Mark
19(a)	A large nucleus splits (into two nuclei plus neutrons)	(1)	
	The binding energy increases, (so energy is released) [Accept binding energy per nucleon increases]  Or There is a decrease in (total) mass, (so energy is released)	(1)	2
19(b)	Binding energy per nucleon read from graph for before and after fission	(1)	
	Binding energy before and after fission calculated  Or Binding energy increase calculated	(1)	
	Energy released $\approx 240 \text{ (MeV)}$	(1)	3
	Example of calculation Increase in binding energy per nucleon = $(8.5 - 7.5)$ MeV = $1.0$ MeV		
	Increase in binding energy $\approx 238 \times 1.0 \text{ MeV}$		
	Energy released $\approx 240 \text{ MeV}$		
19(c)(i)	Top line correct Bottom line correct	(1) (1)	2
	$\frac{\text{Example of completed equation}}{{}_{38}^{90}\text{Sr}} \rightarrow {}_{39}^{90}\text{Y} + {}_{-1}^{0}\beta^{-} + {}_{0}^{0}\overline{\nu}_{e}$		
19(c)(ii)	Calculates total energy released	(1)	
	Conversion between eV and J	(1)	
	P = 110  (W)	(1)	3
	Example of calculation $P = 1295 \times 10^{12} \text{ s}^{-1} \times 0.546 \times 10^{6} \text{ eV} \times 1.6 \times 10^{-19} \text{ J eV}^{-1} = 113 \text{ W}$	, ,	
19(c)(iii)	Use of $\lambda = \frac{\ln 2}{t_{1/2}}$	(1)	
	Use of $A = A_0 e^{-\lambda t}$	(1)	
	$t = 44 \text{ (years)} \neq 50 \text{ (years)}$ , so claim inaccurate <b>Or</b> (After 50 years), $A = 1.1 \times 10^{15} \text{ (Bq)} \neq 1.295 \times 10^{15} \text{ (Bq)}$ so claim inaccurate <b>Or</b> (If $A = 1.295 \times 10^{15} \text{ Bq}$ ), $A_0 = 4.3 \times 10^{15} \text{ (Bq)} \neq 3.7 \times 10^{15} \text{ (Bq)}$ so claim	(1)	
	inaccurate	(1)	3
	Example of calculation		
	Example of calculation $\lambda = \frac{\ln 2}{28.8} = 0.0241 \text{ years}^{-1}$		
	$1295 \text{ TBq} = 3700 \text{ TBq} \times e^{-0.0241 \text{ years}^{-1} \times t}$		
	$t = \frac{\ln\left(\frac{1295 \text{ TBq}}{3700 \text{ TBq}}\right)}{-0.0241 \text{ years}^{-1}} = 43.6 \text{ years}$		
	Total for question 19		13

Question Number	Answer					
20(a)	Alpha radiation would not pass through the sheet of paper	(1)				
	No change in count rate when paper placed (between source and GM tube), so there can't be any alpha radiation	(1)				
	There can't be any gamma radiation, as gamma radiation would pass through the aluminium sheet  Or There can't be any gamma radiation, as count rate decreases to background with aluminium sheet.	(1)				
	It must be beta radiation as beta radiation would not pass through the aluminium sheet.	(1)	4			
20(b)	EITHER	(4)				
	Gradient of 1st graph calculated	(1)				
	Gradient = $(-) \mu$	(1)				
	$\mu = 60 \text{ (cm}^{-1})$	(1)				
	Photon energy read from 2 <sup>nd</sup> graph	(1)				
	Photon energy = 70 (keV) which is closest to 80 (keV) so source is <sup>133</sup> Xe	(1)				
	OR					
	Gradient of 1st graph calculated	(1)				
	Gradient = $(-) \mu$	(1)				
	$\mu = 60 \text{ (cm}^{-1})$	(1)				
	$\mu$ read from 2 <sup>nd</sup> graph for at least one source in table	(1)				
	$\mu$ for <sup>133</sup> Xe = 50 (cm <sup>-1</sup> ), which is closest to 60 (cm <sup>-1</sup> ) so source is <sup>133</sup> Xe	(1)	5			
	Example of calculation					
	Gradient = $\frac{(60 - 540)}{8 \text{ cm}} = -60 \text{ cm}^{-1}$					
	Photon energy = 70 keV					
	Total for question 20		9			

Question Number			Answer			Mark		
*21(a)	structured ans	wer with link arded for indi	dent's ability to show ages and fully-sustair cative content and for g.	ned reasoni	ing.			
	The following table shows how the marks should be awarded for structure and lines of reasoning.							
	Number of marks awarded for structure of answer and sustained line of reasoning							
	linkages and throughout	fully sustaine	and logical structure ed lines of reasoning of	demonstrat	2			
	of reasoning	•	red with some linkag	•				
	Answer has	no linkages b	etween points and is u	unstructure	0			
	Total marks a for structure a		sum of marks for ind asoning	icative cor	ntent and the marks			
	IC points	IC mark	Max linkage mark	Max fina	ll mark			
	6	4	2	6				
	5 4	3 3	2	5 4				
	3	2	1	3				
	2	2	0	2				
	1	1	0	1				
	0	0	0	0				
	Indicative co	ontent						
	IC1 There must be a (very) high temperature (in the core) IC2 To give nuclei/protons large/enough kinetic energy [accept K.E. or E <sub>k</sub> ]							
	IC3 So that nuclei/protons get close enough to fuse							
			n electrostatic repul	sion betwe	een nuclei/protons			
			very) high density					
	_	_	llision rate to mainta					
	Or T	o give a high	n collision rate to ma	ıntaın hıg	h temperature	6		

21(b)(i)	Mass difference calculated	(1)	
	Energy in MeV	(1)	
	Conversion from (M)eV to J	(1)	
	B.E./nucleon = $1.12 \times 10^{-12}$ (J)	(1)	4
	Example of calculation Mass difference = $3727.6 \text{ MeV/c}^2 - 2 \times 938.28 \text{ MeV/c}^2 - 2 \times 939.57 \text{ MeV/c}^2 = 28.1 \text{ MeV/c}^2$		
	Binding energy = 28.1 MeV		
	Binding energy = $(28.1 \times 10^6) \text{ eV} \times 1.60 \times 10^{-19} \text{ J eV}^{-1} = 4.496 \times 10^{-12} \text{ J}$		
	B.E./nucleon = $4.496 \times 10^{-12} \text{ J} / 4 = 1.124 \times 10^{-12} \text{ J}$		
21(b)(ii)	MAX 2		
	The B.E./nucleon is the energy required to remove a nucleon from the nucleus	(1)	
	Large energy is required to break the helium nucleus apart	(1)	
	Decay products (of helium) have a lower binding energy per nucleon	(1)	2
	Total for question 21		12