

Mark Scheme (Results)

Summer 2021

Pearson Edexcel International Advanced Level in Physics (WPH15) Paper 05 Thermodynamics, Radiation, Oscillations and Cosmology

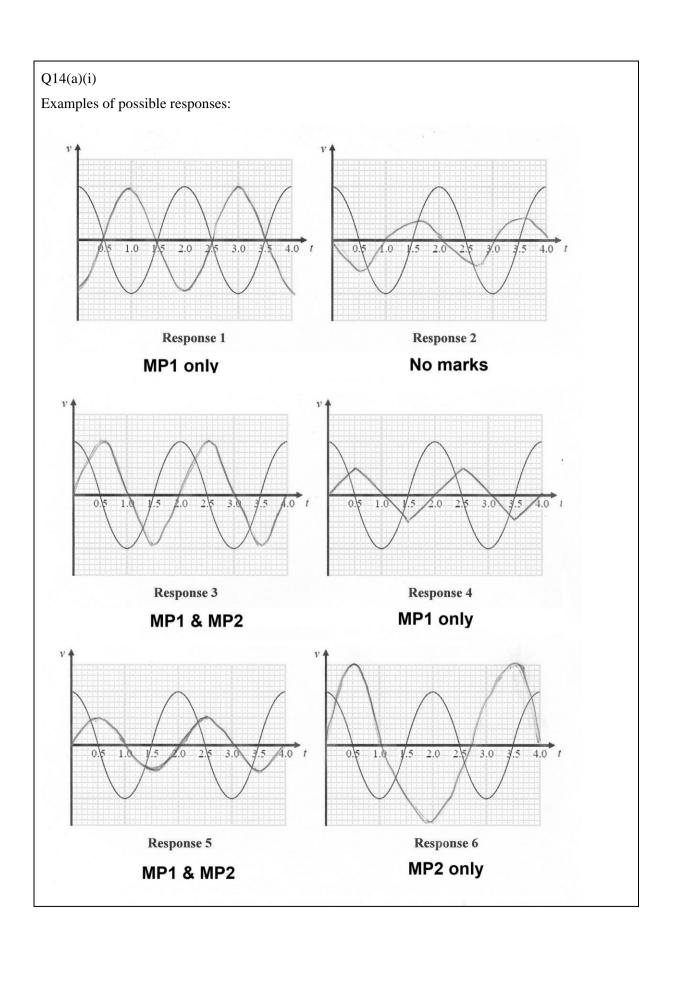
Question Number	Answer	Mark
1	D is the correct answer	(1)
	A is not the correct answer as the background is already included in the count	
	B is not the correct answer as the background will still add a systematic error	
	C is not the correct answer as the background will still add a systematic error	
2	B is the correct answer	(1)
	A is not the correct answer as H_0 does not give the size of the universe	
	C is not the correct answer as $1/H_0$ gives the age of the universe	
	D is not the correct answer as H_0 does not give the size of the universe	
3	D is the correct answer	(1)
	A is not the convert an arrange of all for many in	
	A is not the correct answer as damping occurs at all frequencies	
	B is not the correct answer as energy is transferred at all frequencies	
4	C is not the correct answer as energy is dissipated at all frequencies D is the correct answer	(1)
4	D is the correct answer	(1)
	A is not the correct answer as helium is not being fused in the Sun	
	B is not the correct answer as fusion doesn't require a large number of H nuclei	
	C is not the correct answer as fusion does not require a large mass of H	
5	B is the correct answer as $g = \frac{GM}{r^2}$ and $M \propto \rho$ (as both have the same volume)	(1)
6	A is the correct answer	(1)
	B is not the correct answer as this would have a much lower temperature than the Sun	
	C is not the correct answer as this would have a much higher luminosity than the Sun	
	D is not the correct answer as this would have a much lower luminosity than the Sun	
7	C is the correct answer	(1)
	A is not the correct encurer as mean square valuative increases as the gas is heated	
	A is not the correct answer as mean square velocity increases as the gas is heated B is not the correct answer as $p \propto T$, so T quadruples when p quadruples	
	D is not the correct answer as $p \propto T$, so T quadruples when p quadruples	
8	D is the correct answer as $p \propto T$, so T quadruples when p quadruples D is the correct answer as $L = \sigma A T^4$, so $L \propto T^4$ (as both have the same radius)	(1)
O	D is the correct answer as $L = 0AT$, so $L \propto T$ (as both have the same radius)	(1)
9	B is the correct answer as $v_{\text{max}} = \omega A$ and $\omega = \frac{2\pi}{T}$, so $v_{\text{max}} = \left(\frac{2\pi}{T}\right) \times A$	(1)
10	A is the correct answer	(1)
	B is not the correct answer as λ_{max} is less for X, so surface temperature is higher	
	C is not the correct answer as the max intensity of X (hence luminosity) is higher	
	D is not the correct answer as the max intensity of X (hence luminosity) is higher	
	and λ_{max} for X is less, so surface temperature must be higher	

Question Number	Answer	Mark
11(a)	The atoms/molecules make more frequent collisions with the glass tube Or The atoms/molecules have a higher rate of collision with the glass tube Or The atoms/molecules make more collisions per second with the glass tube (Do not accept collisions between molecules) The rate of change of momentum of the atoms/molecules increases (1) The force exerted on the glass tube increases (1) (Pressure exerted by the gas increases) as pressure is force per unit area)
11(b)	Use of $pV = NkT$ (1 $N = 6.3 \times 10^{22}$ (1 Example of calculation $N = \frac{1.05 \times 10^5 \text{ Pa} \times 2.43 \times 10^{-3} \text{ m}^3}{1.38 \times 10^{-23} \text{ J K}^{-1} \times 293 \text{ K}} = 6.31 \times 10^{22}$	
	Total for question 11	6

Question Number	Answer		Mark
12(a)	A standard candle is a (astronomical) object of known <u>luminosity</u>	(1)	1
12(b)(i)	Use of $P = \frac{\Delta E}{\Delta t}$	(1)	
	Use of $I = \frac{P}{A}$	(1)	
	Use of $I = \frac{L}{4\pi d^2}$	(1)	
	$L = 2.2 \times 10^{35} \text{ (W)}$	(1)	4
	Example of calculation		
	$P = \frac{9.40 \times 10^{-23} \text{ J}}{1.15 \times 10^{-3} \text{ s}} = 8.17 \times 10^{-20} \text{ W}$		
	$I = \frac{8.17 \times 10^{-20} \text{ W}}{1.00 \times 10^{-4} \text{ m}^2} = 8.17 \times 10^{-16} \text{ W m}^{-2}$		
	$L = 4\pi d^2 I = 4\pi \times (4.60 \times 10^{24} \text{ m})^2 \times 8.17 \times 10^{-16} = 2.17 \times 10^{35} \text{ W}$		
12(b)(ii)	Source luminosity is much larger than the luminosity of the Sun		
	Or source is equivalent to the combined output of many Suns		
	Or $L_{\text{FRB}}/L_{\text{Sun}} \sim 5 \times 10^8$	(1)	
	So such a large power output is unlikely to be artificially produced. Or the temperature would be much greater than that of the Sun (so not likely to be artificially produced) [dependent on MP1]	(1)	2
	Response consistent with their calculated value in (b)(i) Total for question 12		7

Question Number	Answer		Mark
13	Use of $\rho = \frac{m}{V}$	(1)	
	·	(1)	
	Use of $\Delta E = mc\Delta\theta$	(1)	
	Use of $\Delta E = L\Delta m$		
	Use of $P = \frac{\Delta E}{\Delta t}$ [to calculate time to melt completely]		
	Or use of $P = \frac{\Delta E}{\Delta t}$ to calculate energy received from the Sun in 1 day	(1)	
	$t = 1.21 \times 10^5 \text{ s or}$ $\mathbf{Or} \ \Delta E = 6.48 \times 10^{10} \text{ J}$	(1)	
	t = 33.7 hours, so palace would not melt completely in a day	(1)	
	Or energy required is $9.09 \times 10^{10} J$, so more energy required than would be transferred in 1 day, so palace would not melt completely.	(1)	6
	(Allow full credit for responses in which 1 day is 12 hours)		
	Example of calculation		
	$m = \rho V = 1325 \text{ kg m}^{-3} \times 1250 \text{ m}^3 = 1.66 \times 10^6 \text{ kg}$		
	$\Delta E = 1.66 \times 10^6 \times 1.30 \times 10^3 \text{J kg}^{-1} \text{K}^{-1} \times (36.0 - 28.5) \text{ K} = 1.62 \times 10^{10} \text{ J}$		
	$\Delta E = 4.5 \times 10^4 \text{ J kg}^{-1} \times 1.66 \times 10^6 \text{ kg} = 7.47 \times 10^{10} \text{ J}$		
	Energy required = $1.62 \times 10^{10} \text{ J} + 7.47 \times 10^{10} \text{ J} = 9.09 \times 10^{10} \text{ J}$		
	$t = \frac{(1.62 + 7.47) \times 10^{10} \text{J}}{7.5 \times 10^5 \text{ W}} = 1.21 \times 10^5 \text{ s}$		
	$t = \frac{1.21 \times 10^5 \text{ s}}{3600 \text{ s hour}^{-1}} = 33.7 \text{ hour}$		
	In 1 day, $\Delta E = 7.5 \times 10^5 \text{ W} \times 24 \times 3600 \text{ s} = 6.48 \times 10^{10} J$		
	Total for question 13		6
	Total for Ancount 19		U

Question Number	Answer		Mark
14(a)(i)	Same time period as velocity and constant amplitude	(1)	
	Wave shifted a quarter cycle to the right [i.e. a positive sine wave, displacement is zero at time zero.]	(1)	2
14(a)(ii)	T = 2.0 s from graph	(1)	
	Use of $T = 2\pi \sqrt{\frac{\ell}{g}}$ (accept any value of T that could be read from the graph)	(1)	
	$\ell = 0.99 \text{ m}$	(1)	3
	Example of calculation		
	$2.0 \text{ s} = 2\pi \sqrt{\frac{\ell}{9.81 \text{ m s}^{-2}}}$		
	$\ell = \frac{(2.0 \text{ s})^2 \times 9.81 \text{ m s}^{-2}}{4\pi^2} = 0.994 \text{ m}$		
14(b)	EITHER		
	Suitable data logger application identified	(1)	
	Reason why data logger is an advantage in this situation	(1)	
	OR		
	Max 2 from		
	When data has to be collected over a very short time interval	(1)	
	When multiple data sets have to be collected simultaneously	(1)	
	When data has to be collected over a very long time interval	(1)	2
	Total for question 14		7



Question Number	Answer		Mark	
15(a) 15(b)	$\lambda_{\text{max}} = 0.37 \rightarrow 0.40 \text{ (}\mu\text{m)} \tag{1}$ Use of $\lambda_{\text{max}}T = 2.898 \times 10^{-3} \text{ m K} \tag{1}$ $T = 7600 \text{ K (accept answer consistent with their stated value of } \lambda_{\text{max}}) \tag{1}$ $\frac{\text{Example of calculation}}{T = \frac{2.898 \times 10^{-3} \text{ m K}}{0.38 \times 10^{-6} \text{ m}}} = 7626 \text{ K}$ Corresponding pair of wavelengths recorded (one from each spectrum) Wavelength shift calculated [wavelengths may be out of range, but must]		3	
	Use of $\frac{\Delta\lambda}{\lambda} \approx \frac{v}{c}$ [value of λ must be taken from lab spectrum] $v = 1.5 \times 10^7 \text{ m s}^{-1}$ [v will depend upon in-range values used] Star is receding $\frac{\text{Example of calculation}}{\lambda_{\text{star}} = 654 \text{ nm} \rightarrow 658 \text{ nm}} \qquad \lambda_{\text{lab}} = 622 \text{ nm} \rightarrow 626 \text{ nm}$ Or $\lambda_{\text{star}} = 479 \text{ nm or } 480 \text{ nm} \qquad \lambda_{\text{lab}} = 452 \text{ nm} \rightarrow 456 \text{ nm}$ $v = 3.00 \times 10^8 \text{ m s}^{-1} \times \frac{(656 \text{ nm} - 624 \text{ nm})}{624 \text{ nm}} = 1.54 \times 10^7 \text{ m s}^{-1}$	(1)(1)(1)	5	
	Total for question 15		8	

Question Number	Answer		Mark
16(a)	Either	1)	
	II CE GMm P2.	1)	
		1)	
	T = 5800 s		
	Or		
	Use of $F = \frac{GMm}{m}$ with $F = \frac{mv^2}{m}$	1)	
		1)	
	Use of $v = \frac{2\pi r}{T}$	1)	3
	T = 5800 s		
	Example of calculation		
	$\frac{GMm}{r^2} = m\omega^2 r$		
	$\therefore \omega = \sqrt{\frac{GM}{r^3}}$		
	$ \omega = \sqrt{\frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2} \times 6.0 \times 10^{24} \text{ kg}}{(6.4 \times 10^6 \text{ m} + 5.5 \times 10^5 \text{ m})^3}} = 1.09 \times 10^{-3} \text{ rad s}^{-1} $		
	$T = \frac{2\pi \text{ rad}}{1.09 \times 10^{-3} \text{ rad s}^{-1}} = 5755 \text{ s}$		
16(b)	Either		
	$(F = \frac{GMm}{r^2})$, so) the (gravitational) force is greater for a low Earth orbit	1)	
	$F = m \left(\frac{2\pi}{T}\right)^2 r$ [accept angular velocity is greater]	1\	
	So if Finerages when r decreases then T must decrease	1)	
	(MP3 dependent upon MP1 AND MP2)	1)	
	Or		
	$\left(\frac{2\pi}{T} = \sqrt{\frac{GM}{r^3}}, \text{ so}\right) T^2 = \frac{4\pi^2 r^3}{GM}$	1)	
	G and M are constant, so $T \propto \sqrt{r^3}$	1)	
	So when r is smaller T is smaller	1)	
	(MP3 dependent upon MP1 OR MP2)	1)	3
	[Accept converse argument]		

16(c)	Use of $V_{\text{grav}} = (-)\frac{GM}{r}$	(1)	
	Use of $\Delta E_k = GMm\left(\frac{1}{r_1} - \frac{1}{r_2}\right)$	(1)	
	$\Delta E_{\rm k}$ =1.1×10 ⁹ J	(1)	3
	[Do not credit use of $\Delta E_{\text{grav}} = mg\Delta h$, as g is not constant]		
	Example of calculation		
	$\Delta E_{\rm k} = 6.67 \times 10^{-11} \rm N m^2 kg^{-2} \times 6.0 \times 10^{24} kg \times 227 kg \left(\frac{1}{6.4 \times 10^6 \rm m} - \frac{1}{(6.4 \times 10^6 + 5.5 \times 10^5) \rm m} \right)$		
	$\therefore \Delta E_{\mathbf{k}} = 1.12 \times 10^9 \mathrm{J}$		
	Total for question 16		9

Question Number	Answer		Mark
17(a)	(The mass meets the conditions for simple harmonic motion as)		
	There is a (resultant) <u>force</u> acting on the mass which is proportional to its displacement from its equilibrium position.	(1)	
	The <u>force</u> is always directed towards the equilibrium position	(1)	2
	(An equation with symbols defined, and the negative sign justified, may be a valid response for both marks		
	For equilibrium position accept: undisplaced point/position or fixed point/position or central point/position)		
17(b)(i)	Use of $\Delta F = k \Delta x$	(1)	
	$k = 26.2 \text{ (N m}^{-1})$	(1)	2
	Example of calculation		
	$k = \frac{0.2 \text{ kg} \times 9.81 \text{ N kg}^{-1}}{7.5 \times 10^{-2} \text{ m}} = 26.16 \text{ N m}^{-1}$		
17(b)(ii)	Combine $T = 2\pi \sqrt{\frac{m}{k}}$ with $f = \frac{1}{T}$ to obtain $f^2 = \frac{k}{4\pi^2}m^{-1}$	(1)	
	Compare with $y = mx + c$ to identify gradient as $\frac{k}{4\pi^2}$	(1)	
	Gradient of graph calculated	(1)	
	Large triangle used for gradient calculation	(1)	
	$k = 26.7 \text{ N m}^{-1}$	(1)	
	A conclusion consistent with the value calculated in (i) (accept comparison with "show that" value from (i))	(1)	6
	Example of calculation		
	$T^2 = \frac{4\pi^2 m}{k} :: f^2 = \frac{k}{4\pi^2} m$		
	So gradient = $\frac{k}{4\pi^2}$		
	Gradient= $\frac{(3.25 - 0.00) \text{ s}^{-2}}{(5.00 - 0.20) \text{ kg}^{-1}} = 0.677 \text{ kg s}^{-2}$		
	$k=4\pi^2 \times 0.677 \text{ kg s}^{-2}=26.7 \text{ N m}^{-1}$		
	Total for question 17		10

Question Number	Answer		Mark
18(a)	A massive/large nucleus splits into smaller fragments	(1)	1
18(b) (i)	Steeply rising curve near to origin	(1)	
	Slowly decreasing curve after peak	(1)	2
18(b) (ii)	Iron-56 marked at peak of curve	(1)	1
	Example of graph for (i) and (ii)		
	Binding energy per nucleon Iron-56		
18(c)	Top line correct	(1)	
	Bottom line correct	(1)	2
	$^{236}_{92}U \rightarrow ^{93}_{38}Sr + ^{141}_{54}Xe + 2 \times ^{1}_{0}n$		
18(d)	Calculation of mass defect	(1)	
	Binding energy per nucleon = 7.38 (MeV)	(1)	2
	Example of calculation		
	Mass defect = $(92 \times 0.93827 + 144 \times 0.93956 - 219.8750) \text{ GeV}/c^2$		
	Mass defect = $1.74248 \text{ GeV}/c^2$		
	Binding energy/nucleon = 1.74248 GeV/236 = 7.383 MeV		

*18(e)

This question assesses a student's ability to show a coherent and logically structured answer with linkages and fully-sustained reasoning.

Marks are awarded for indicative content and for how the answer is structured and shows lines of reasoning.

	Number of marks awarded for structure of answer and sustained line of reasoning
Answer shows a coherent and logical structure with linkages and fully sustained lines of reasoning demonstrated throughout	2
Answer is partially structured with some linkages and lines of reasoning	1
Answer has no linkages between points and is unstructured	0

The following table shows how the marks should be awarded for structure and lines of reasoning.

Total marks awarded is the sum of marks for indicative content and the marks for structure and lines of reasoning

IC points	IC mark	Max linkage mark	Max final mark
6	4	2	6
5	3	2	5
4	3	1	4
3	2	1	3
2	2	0	2
1	1	0	1
0	0	0	0

Indicative content

- IC1 Energy from the α particles is transferred to atoms/molecules in the
- IC2 An electron in the atom/molecule is promoted to a higher energy state **Or** the atom/molecule/electron is excited
- IC3 When the electron return to a lower energy state a photon (of uvradiation) is emitted

Or when the atom/molecule/electron de-excites, a photon (of uvradiation) is emitted

- IC4 α radiation is strongly ionising and so has a short range (in air)
- IC5 Ultraviolet radiation is weakly ionising (and has long range in air)
- IC6 UV-radiation can be detected much further from the source so is safer

6

Question Number	Answer		Mark
19(a)	Calculation of mass difference	(1)	
	Conversion from u to kg, using a conversion factor of 1.66 ×10 ⁻²⁷ kg u ⁻¹	(1)	
	Use of $\Delta E = c^2 \Delta m$	(1)	
	Conversion of energy to eV	(1)	
	$\Delta E = 5.61 \text{ (MeV)}$	(1)	5
	Example of calculation		
	Mass difference = $237.999089 \text{ u} - 233.991578 \text{ u} - 4.001506 \text{ u} = 6.005 \times 10^{-3} \text{ u}$		
	Mass difference = $6.005 \times 10^{-3} \text{ u} \times 1.66 \times 10^{-27} \text{ kg} = 9.9683 \times 10^{-30} \text{ kg u}^{-1}$		
	$\Delta E = c^2 \Delta m = (3.00 \times 10^8 \text{ m s}^{-1})^2 \times 9.9683 \times 10^{-30} \text{ kg} = 8.9715 \times 10^{-13} \text{ J}$		
	$\Delta E = \frac{8.9715 \times 10^{-13} \text{ J}}{1.60 \times 10^{-13} \text{ J MeV}^{-1}} = 5.607 \text{ MeV}$		
19(b)	Convert α-particle energy from MeV to J	(1)	
	Use of $\lambda = \frac{\ln 2}{t_{1/2}}$	(1)	
	Use of $A = A_0 e^{-\lambda t}$	(1)	
	Use of $P = \frac{\Delta E}{\Delta t}$	(1) (1)	
	P = 0.083 (W)	(1)	5
		(1)	
	Example of calculation		
	$5.6 \text{ MeV} = 5.6 \times 1.60 \times 10^{-19} \text{ J MeV}^{-1} = 8.96 \times 10^{-13} \text{ J}$		
	$\lambda = \frac{\ln 2}{t_{1/2}} = \frac{0.693}{87.7 \text{ year}} = 7.90 \times 10^{-3} \text{ year}^{-1}$		
	$6.75 \times 10^{10} \text{ Bq} = A_0 e^{-7.90 \times 10^{-3} \text{ year}^{-1} \times 40 \text{ year}}$		
	$A_0 = 9.26 \times 10^{10} \text{ Bq}$		
	So $P = 9.26 \times 10^{10} \text{s}^{-1} \times 8.96 \times 10^{-13} \text{ J} = 0.0830 \text{ W}$		

19(c)	Maximum energy of beta particles read from graph 1 (in range 210 keV \rightarrow 225 keV)	(1)	
	Beta particle range read from graph 2 (in range $0.05 \text{ cm} \rightarrow 0.08 \text{ cm}$)		
	Or max. energy for 0.5 cm polyethylene read from graph.2 (in range $1000 \text{ keV} \rightarrow 1200 \text{ keV}$)	(1)	
	Conclusion that 0.5 cm polyethylene would be sufficient	(1)	3
	MP3 dependent on MP1 and MP2		
	Total for question 19		13